

General aspects of QCD at future colliders

Gavin Salam (CERN)

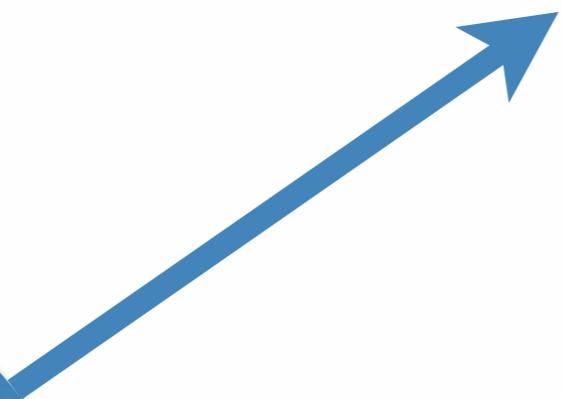
SLAC Workshop on Physics at a 100 TeV Collider
23-25 April 2014

We have seen / will see talks on many of the
key topics of QCD:

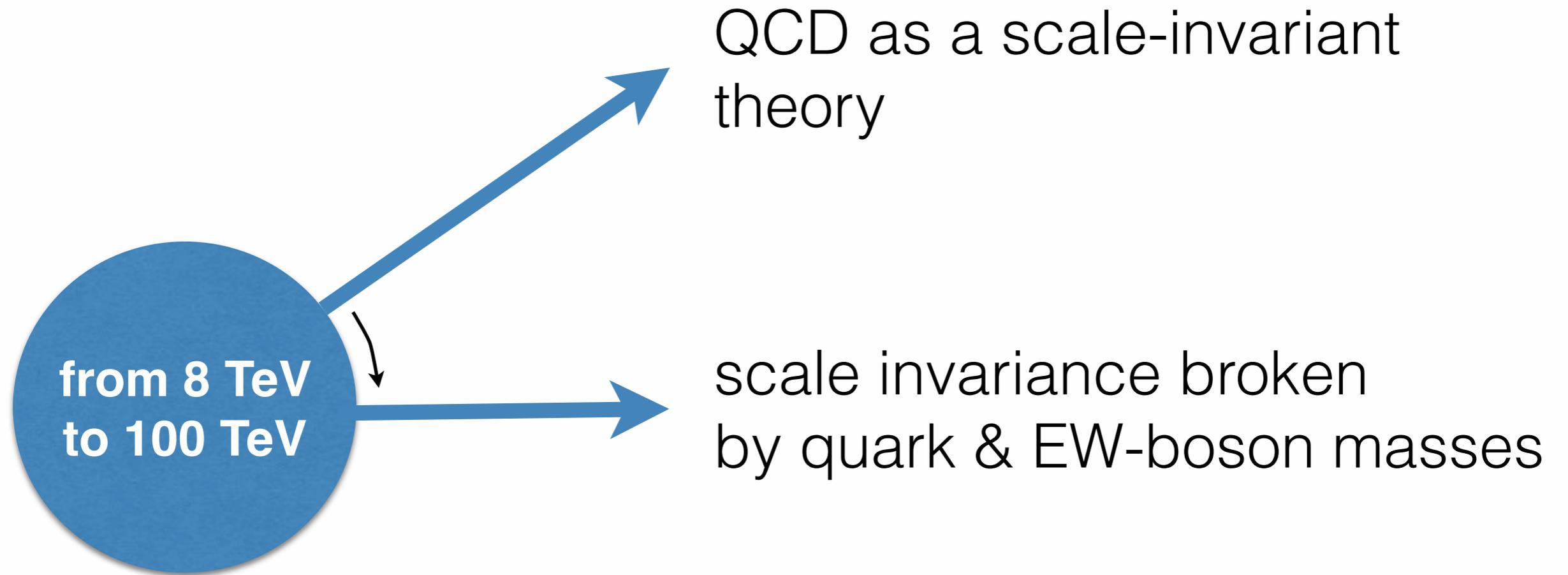
PDFs, MC matching, Jets

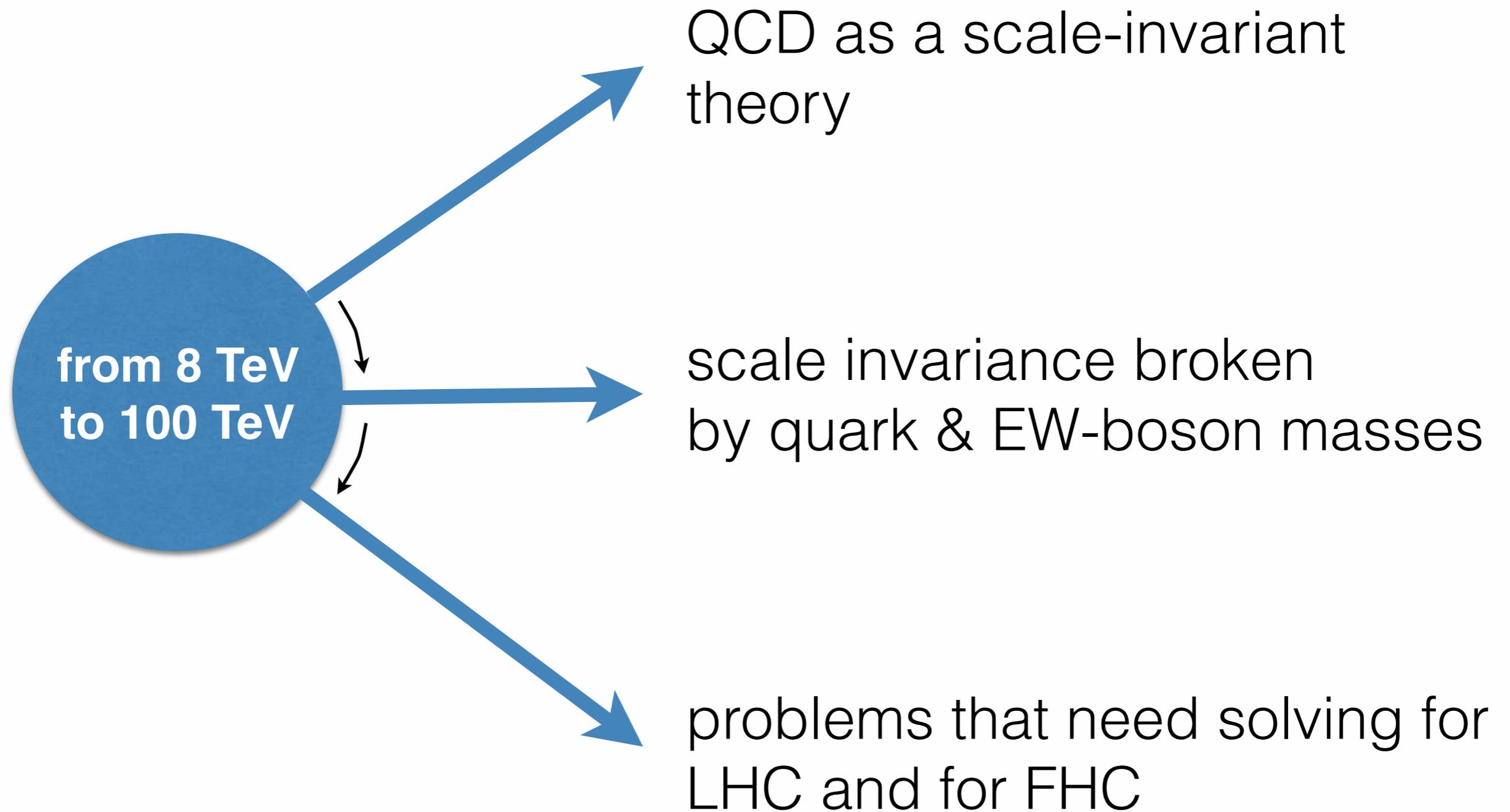
What is there left for me to tell you?

**from 8 TeV
to 100 TeV**

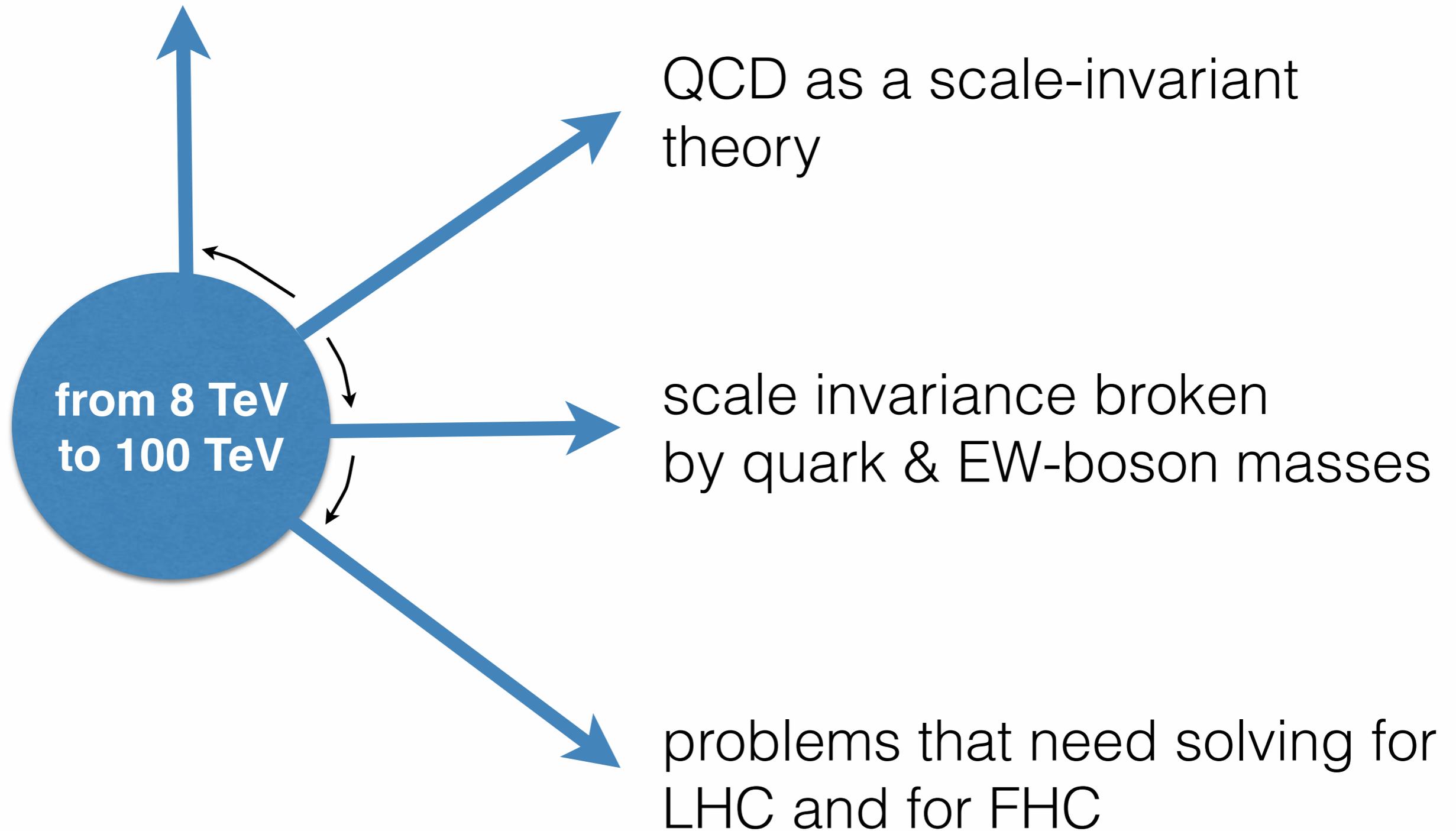


QCD as a scale-invariant
theory

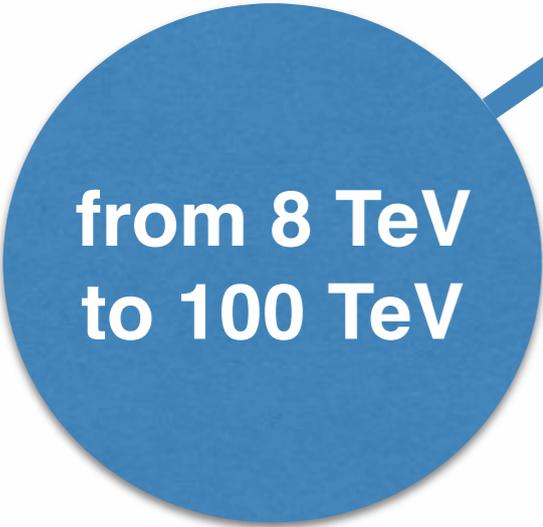




Collider Reach



QCD as a scale-invariant theory



from 8 TeV
to 100 TeV

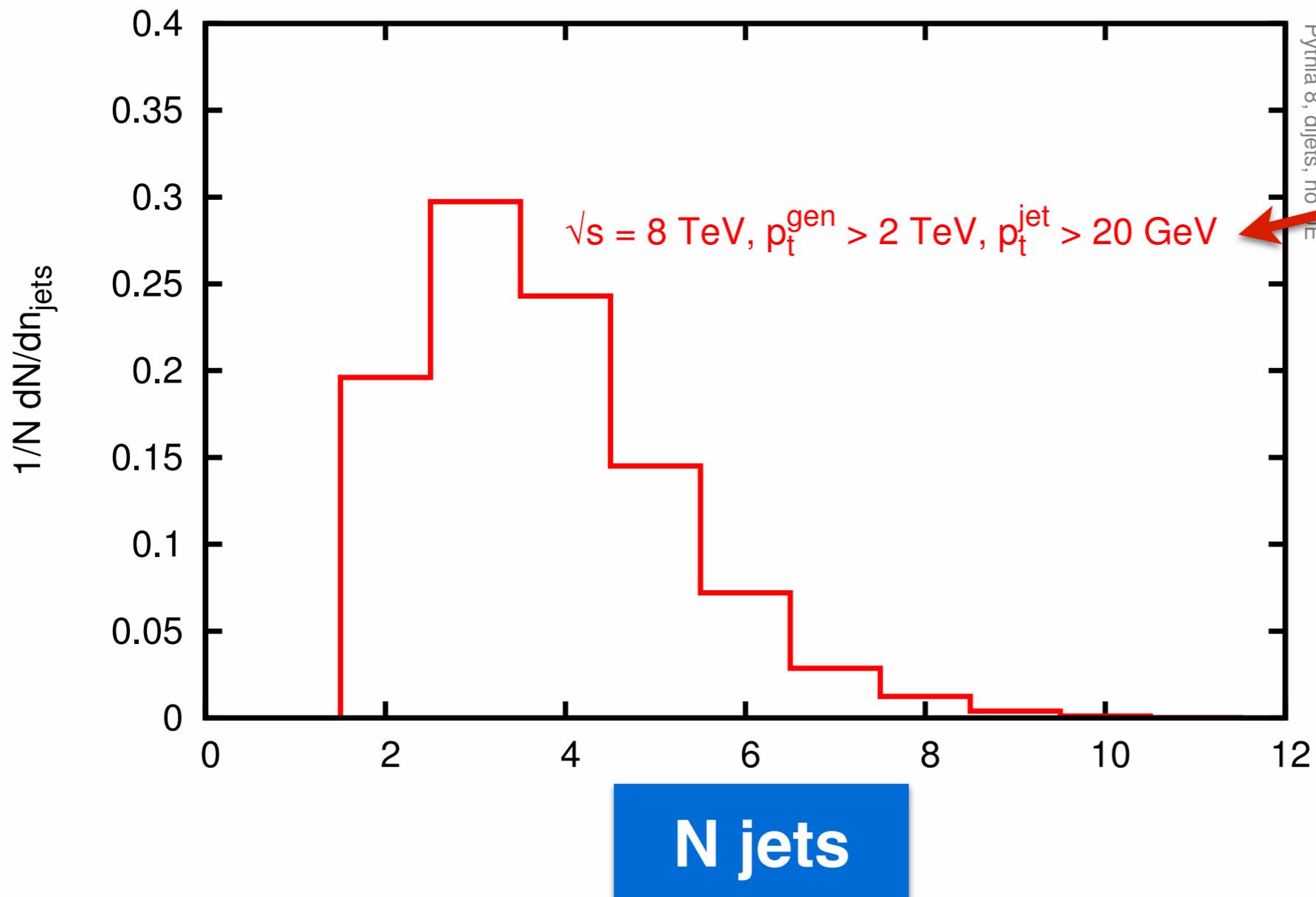
At high scales, α_s runs slowly, as do PDFs

Little difference between
2 TeV physics at an 8 TeV collider
and 25 TeV physics at a 100 TeV collider

$$\alpha_s (2 \text{ TeV}) = 0.083$$

$$\alpha_s (25 \text{ TeV}) = 0.067$$

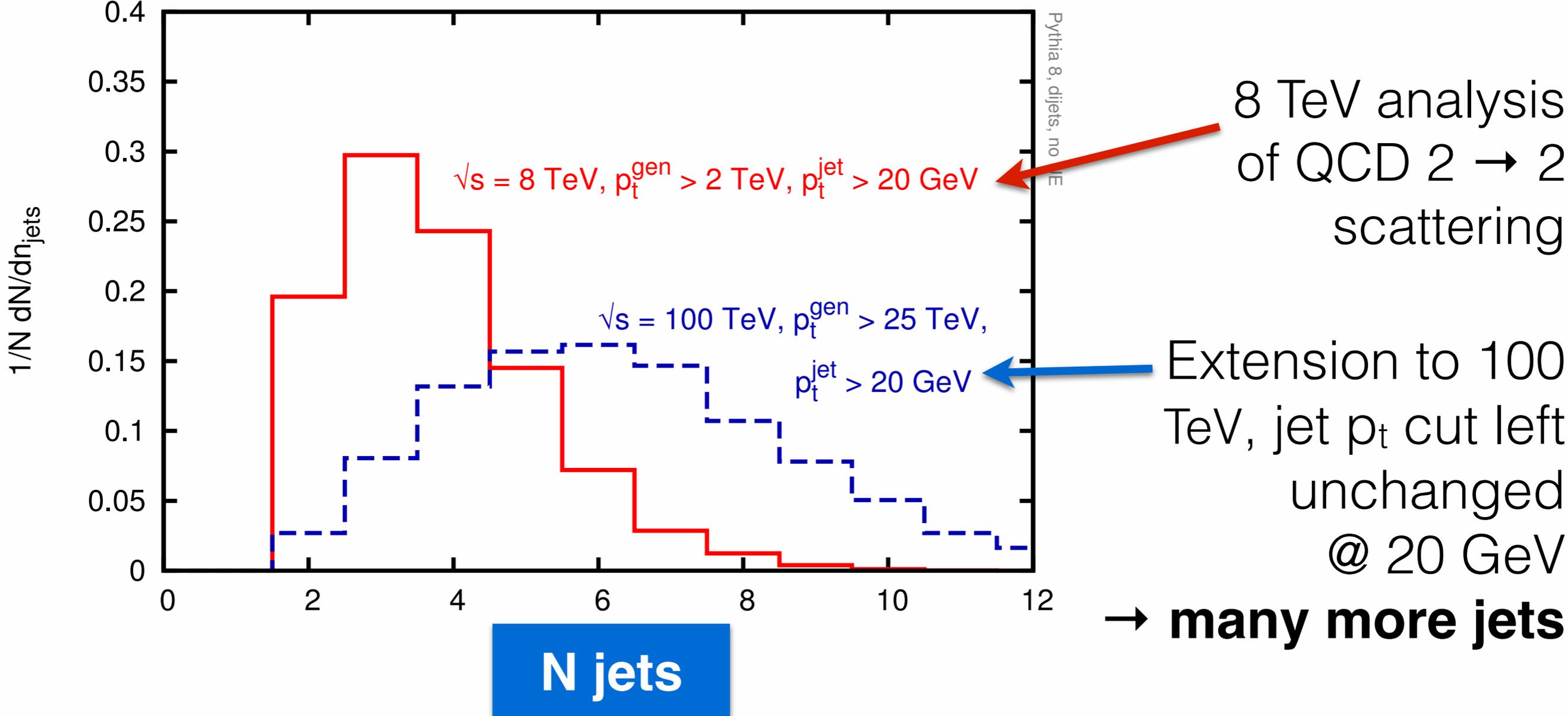
Scale invariance holds if **all ratios of scales kept fixed**



8 TeV analysis
of QCD $2 \rightarrow 2$
scattering

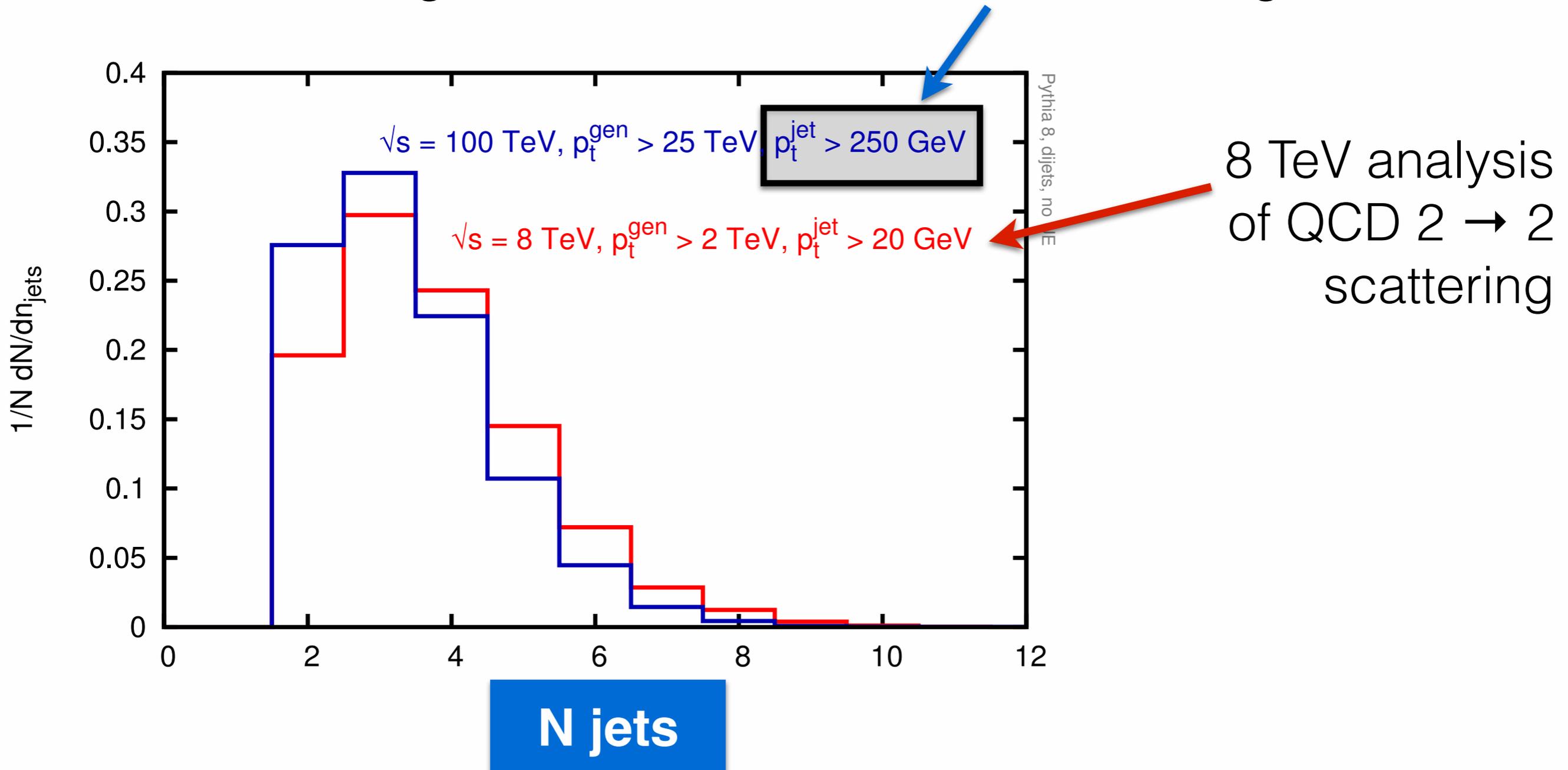
Count the jets

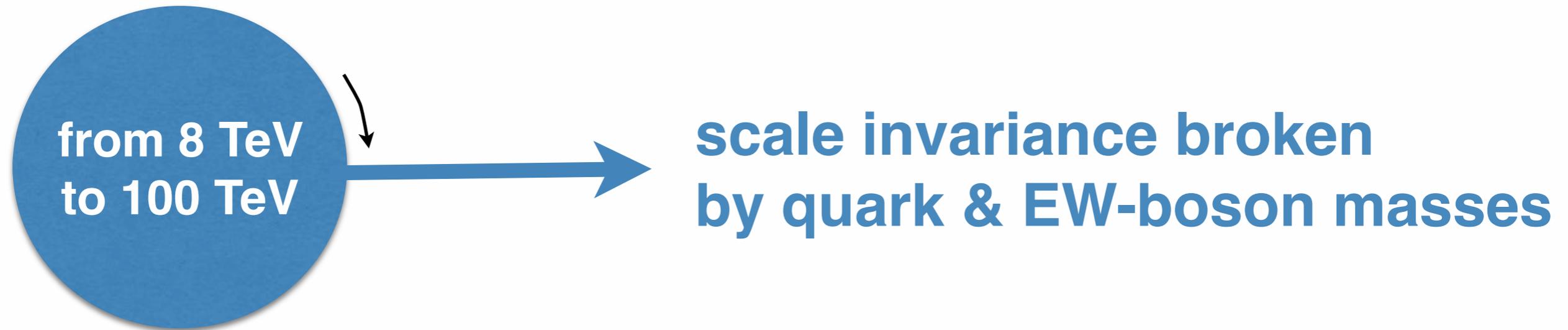
Scale invariance holds if **all ratios of scales kept fixed**



Scale invariance holds if **all ratios of scales kept fixed**

100 TeV, scaling also the **jet cut** \rightarrow **250 GeV**
gives a distribution much like the original





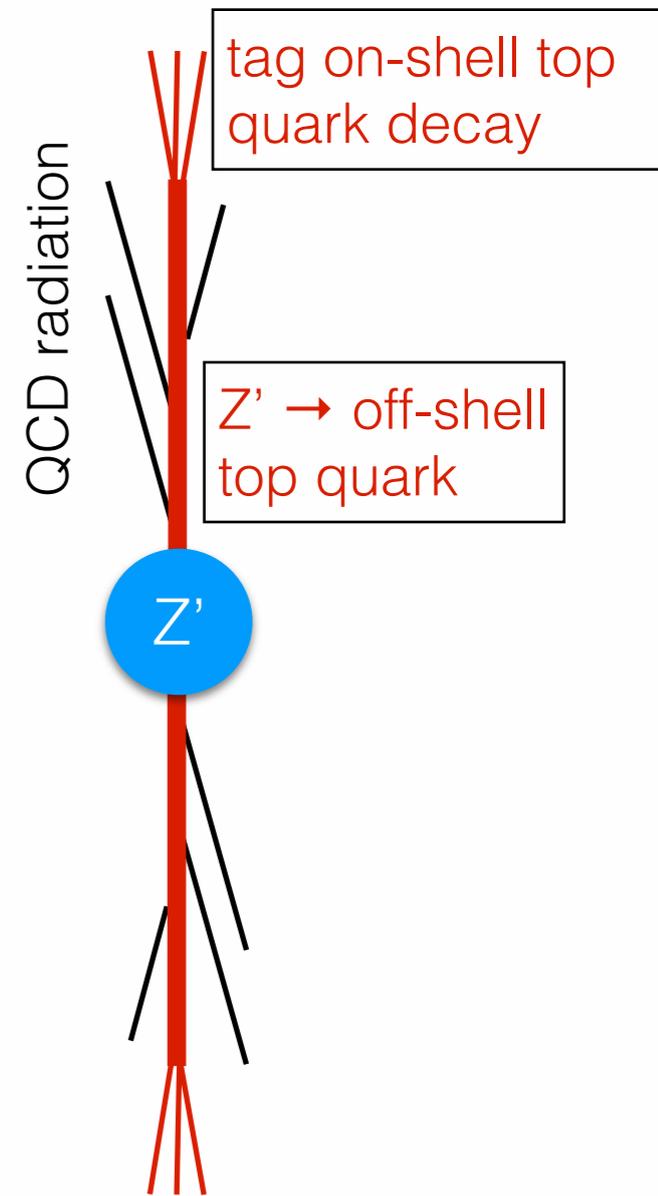
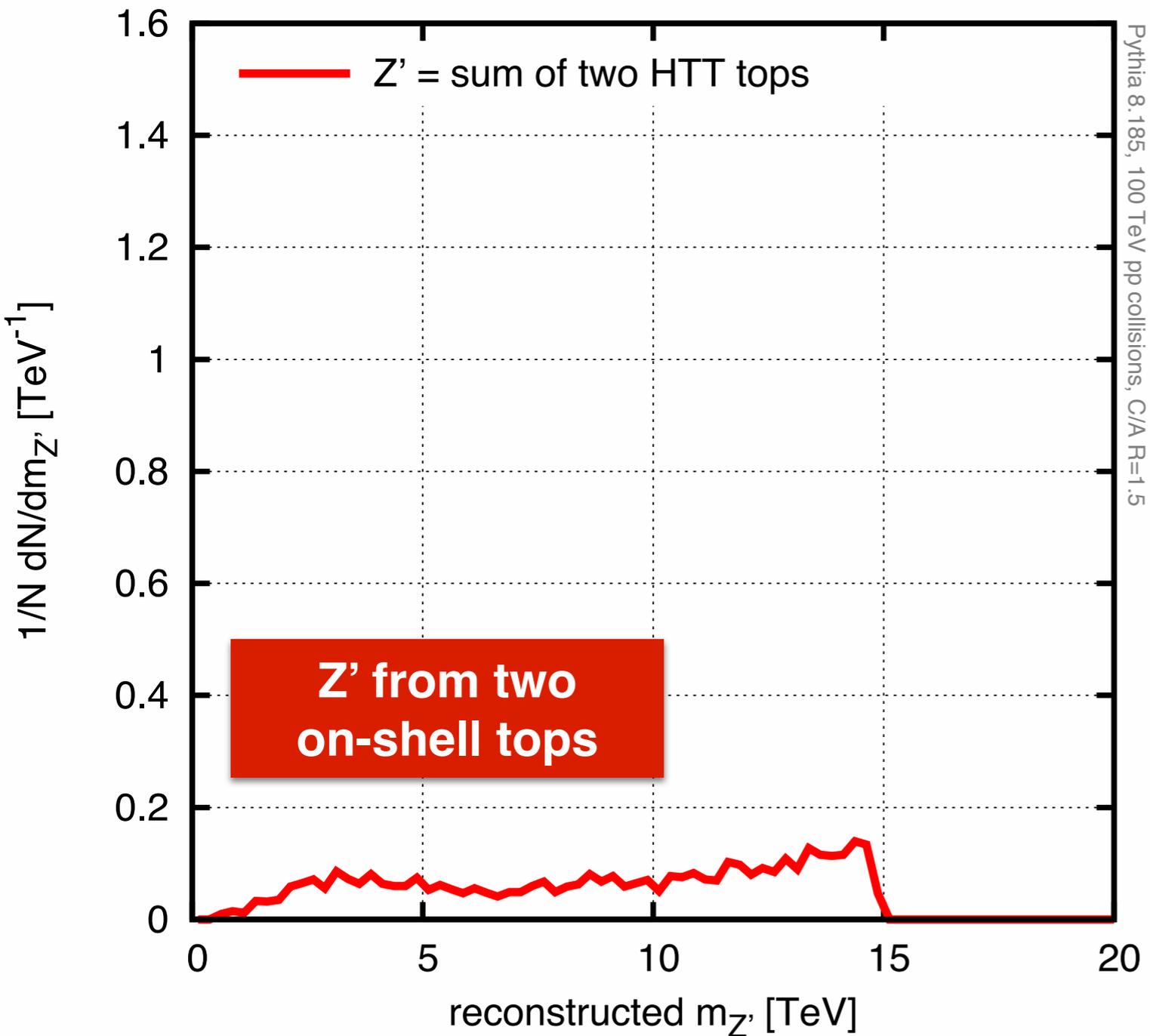
top quarks v. top jets

EW radiation in jets

The top quark as a light parton

[informal studies with Tilman Plehn]

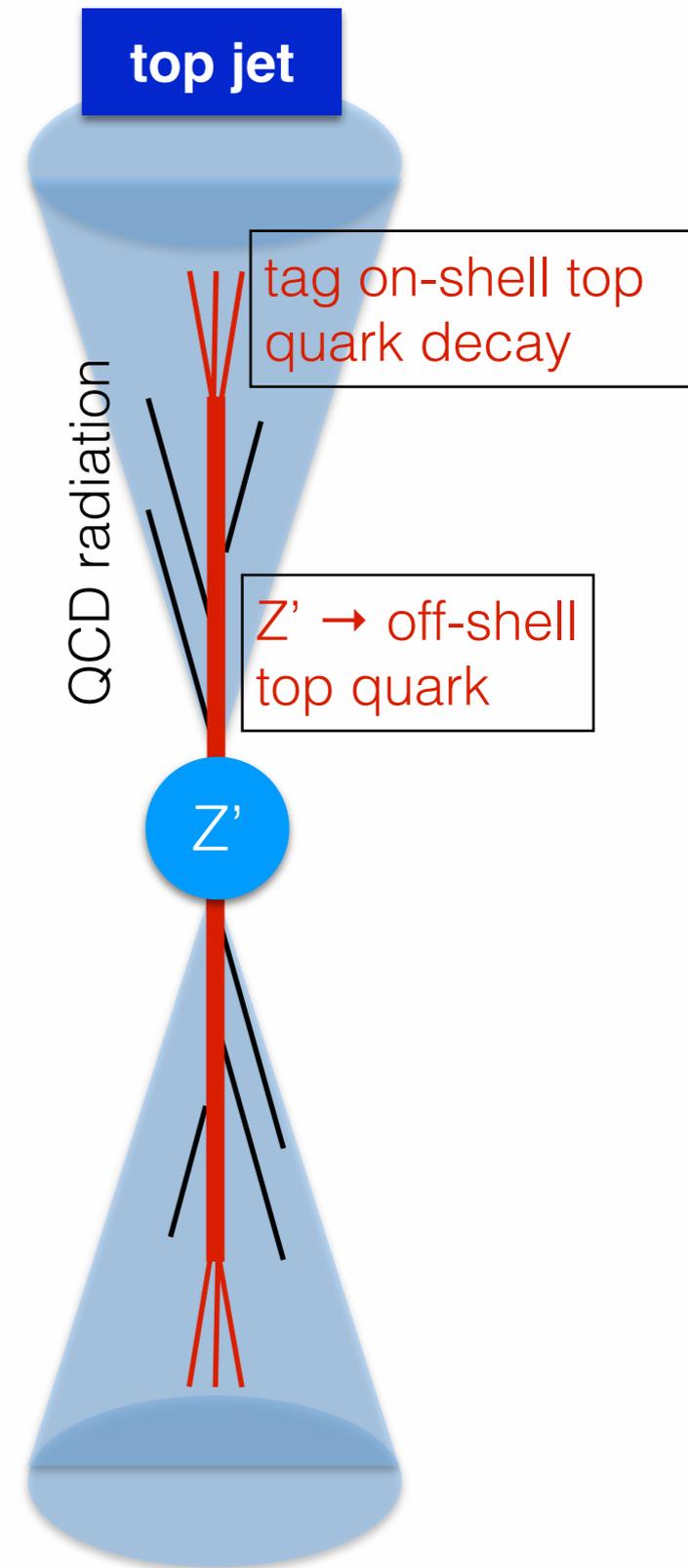
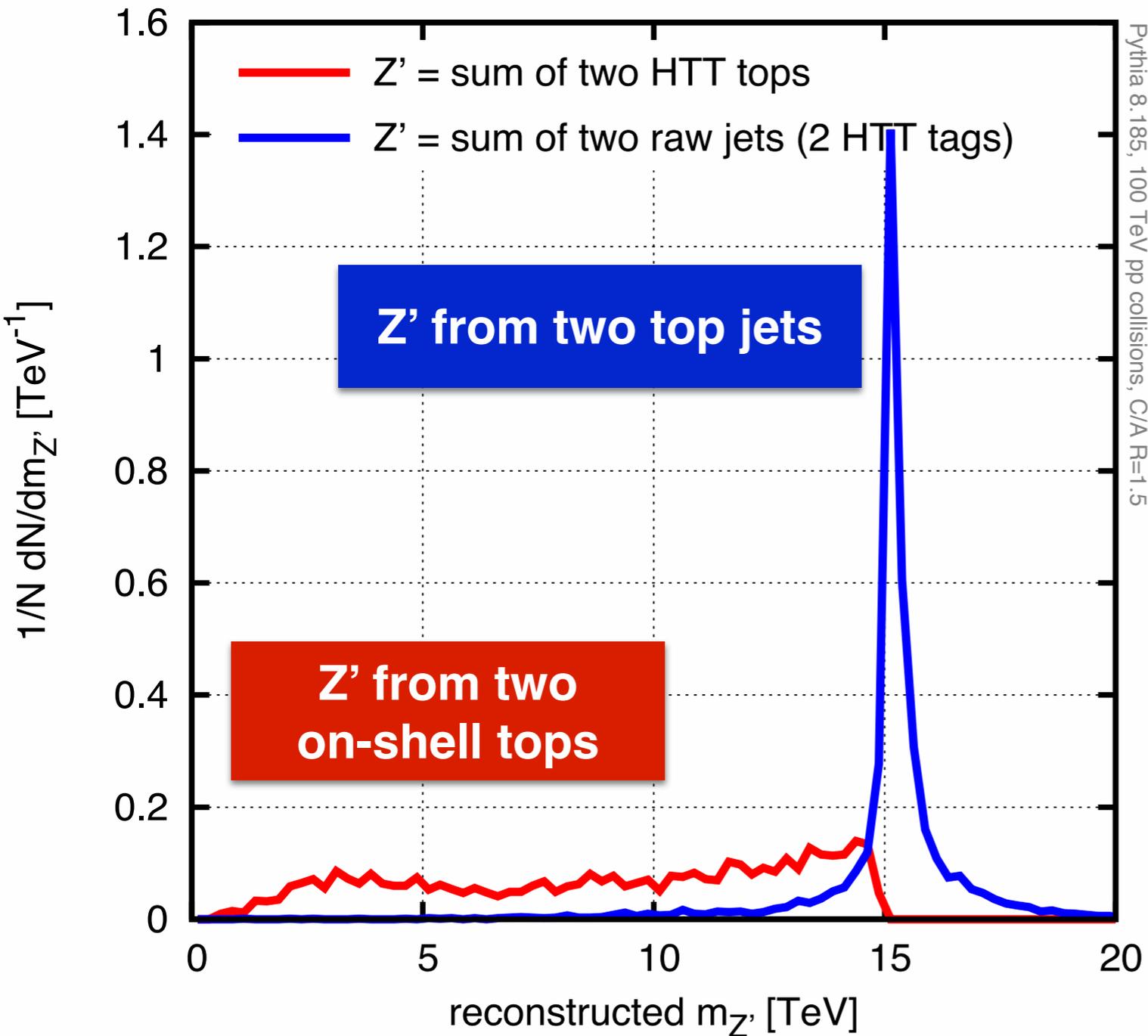
tag two hadr. tops (HEPTopTagger), reconstruct "tt" mass



The top quark as a light parton

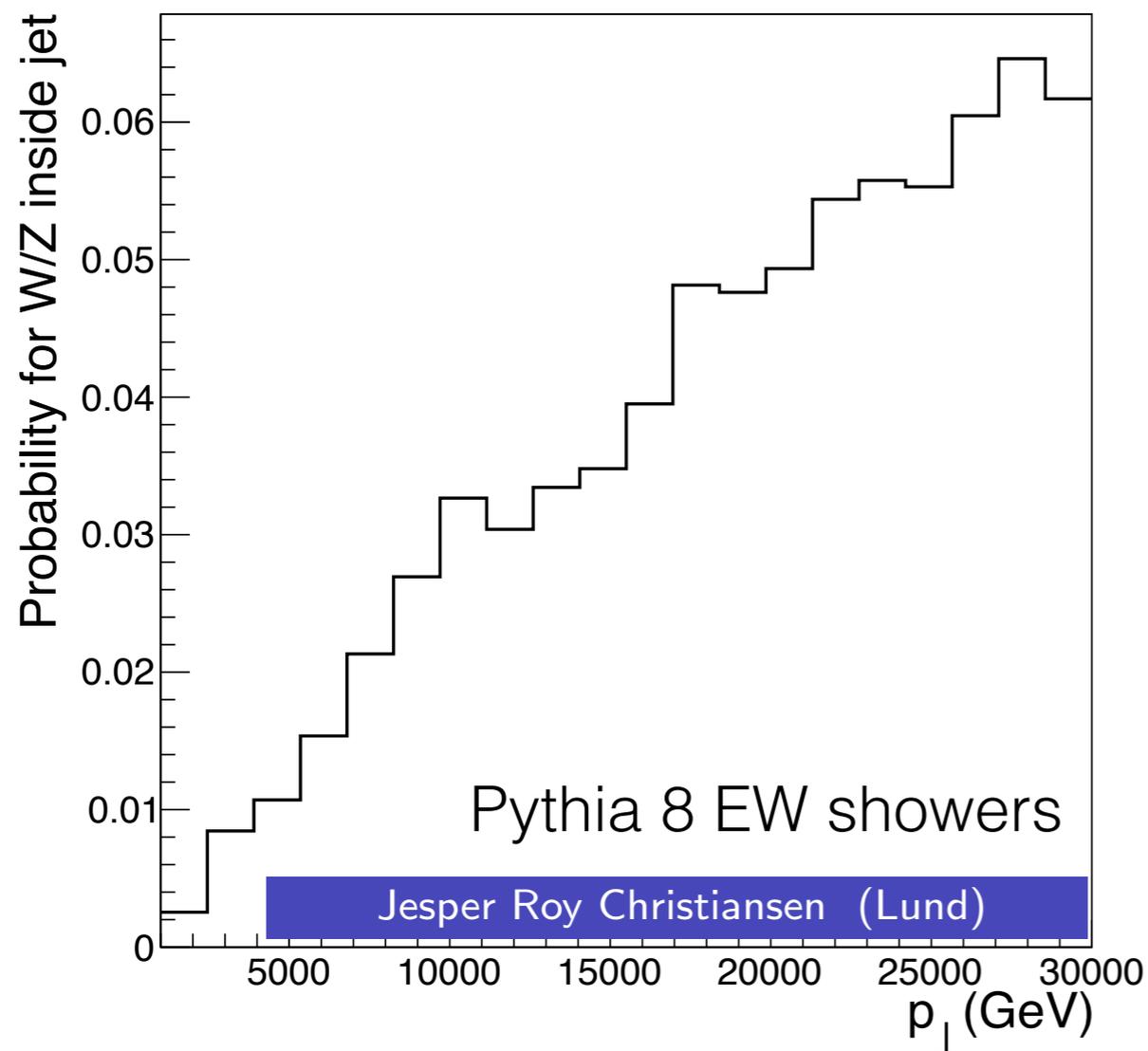
[informal studies with Tilman Plehn]

tag two hadr. tops (HEPTopTagger), reconstruct "tt" mass



EW radiation in jets

Prob. of Z/W in jet v. jet p_t



+ analogous plots at fixed order from MLM

Significant enhancement of W's and Z's in jets:

$$\sim \alpha_{EW} \ln^2 \frac{p_t}{M_W}$$

New, fun theory!

Experimentally, how different is this from bottom and charm in jet?

cf. 20% BR for $b \rightarrow c \nu l^{\pm}$ and $O(10\%)$ of 1 TeV gluon jets having b's inside

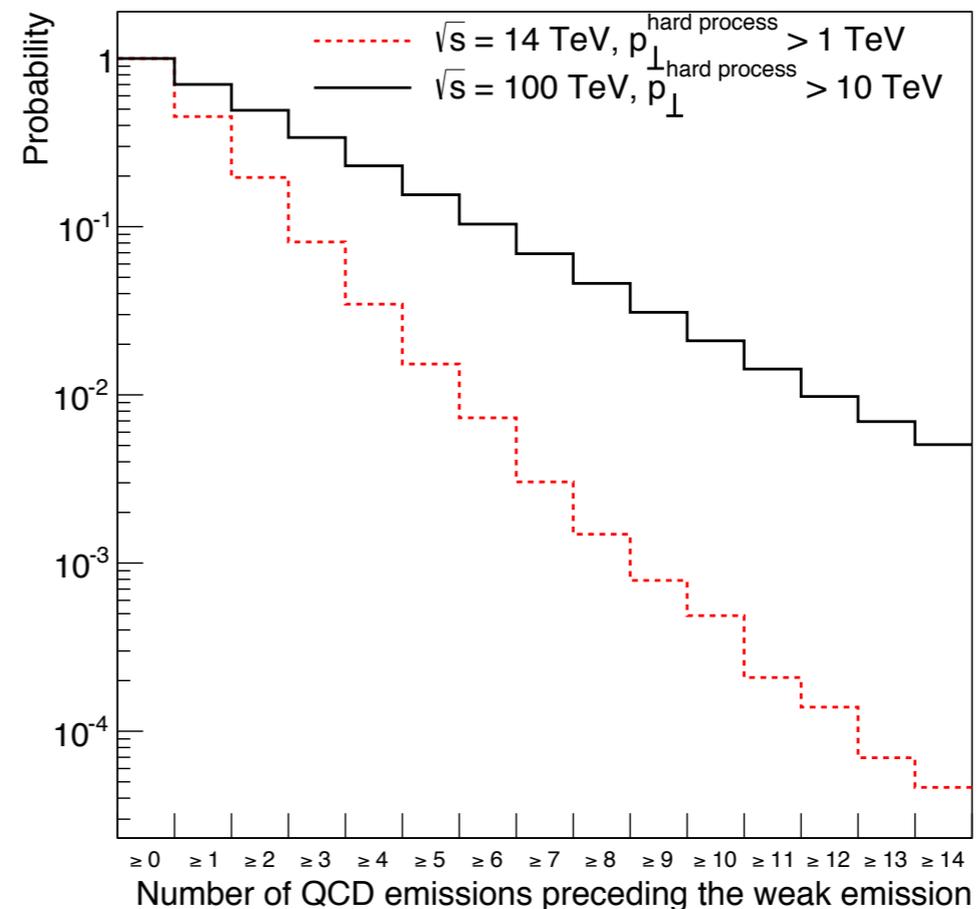
EW radiation in initial state

It polarizes the incoming partons
Can we make use of this?

How to simulate EW radiation correctly?

Competition between QCD and weak emissions

- Need to include up to 11 emissions, to only miss 1 %
⇒ Does this become problematic for merging techniques?

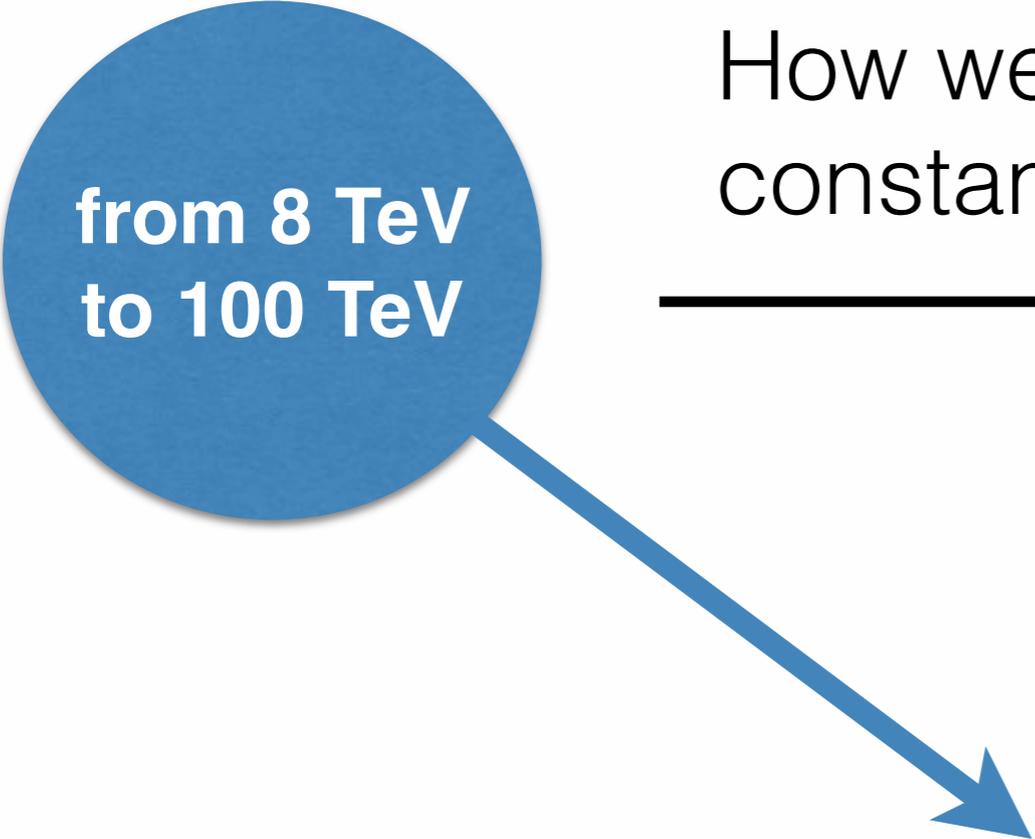


Some long-standing problems:

Why is QCD so poorly convergent at hadron colliders?

How well do we know our fundamental constants?

**from 8 TeV
to 100 TeV**



**problems that need solving for
LHC and for FHC**

Radically worse perturbative series at hadron colliders:

e^+e^- collisions: $R_{\text{hadrons}} \propto 1 + 0.32\alpha_s + 0.14\alpha_s^2 + \dots$

pp collisions: $\sigma_{gg \rightarrow H} \propto 1 + 9.8\alpha_s + 33\alpha_s^2 + ?$

C_A / C_F explains twice worse convergence

But convergence is 10–30 times worse

WHY?

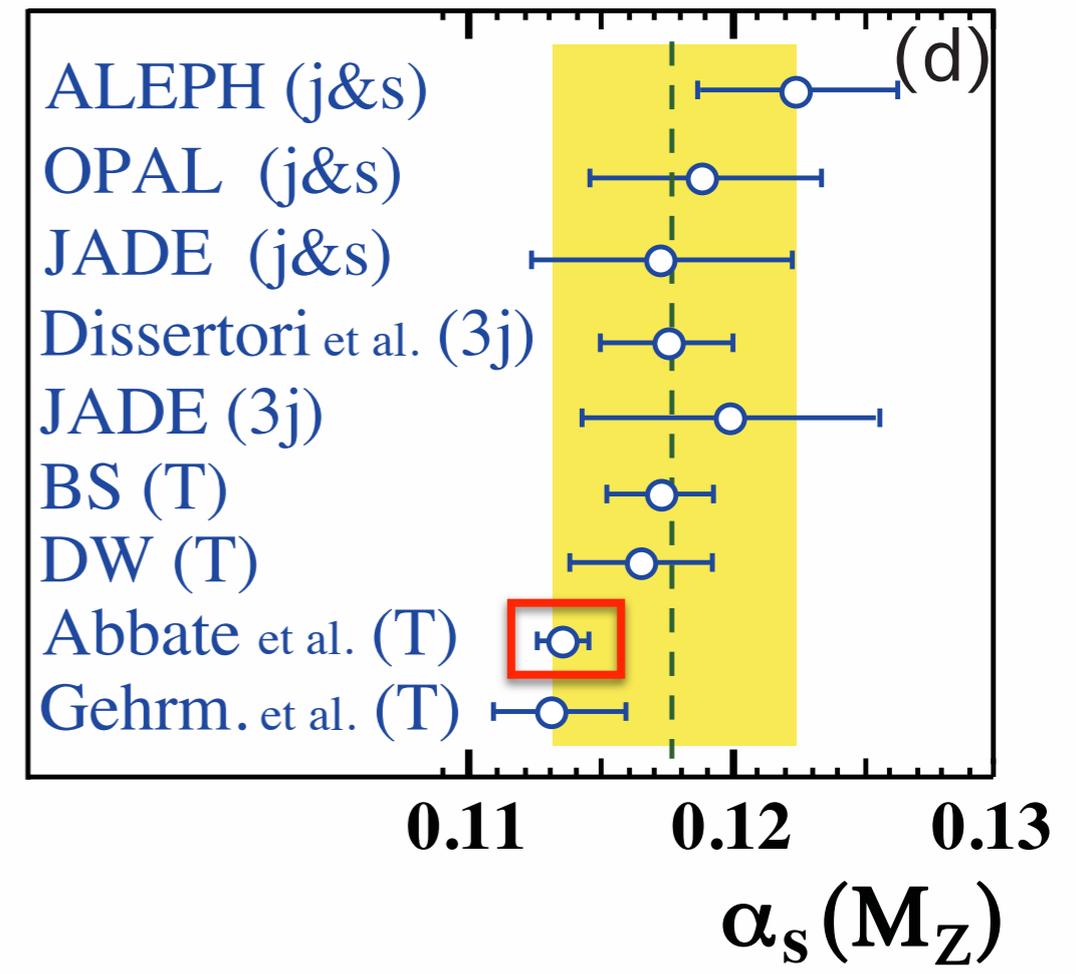
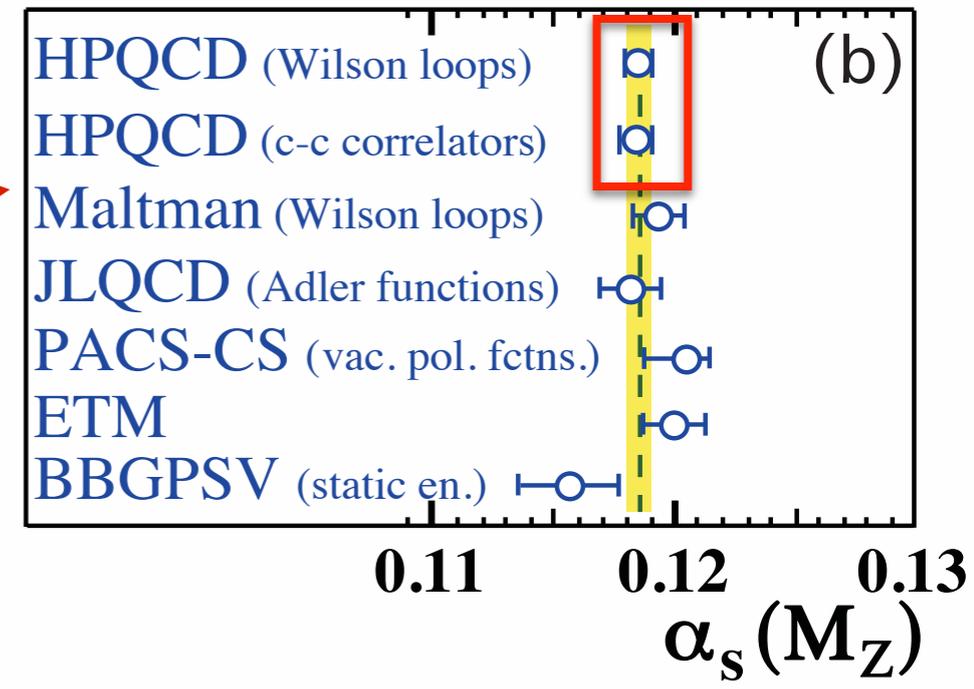
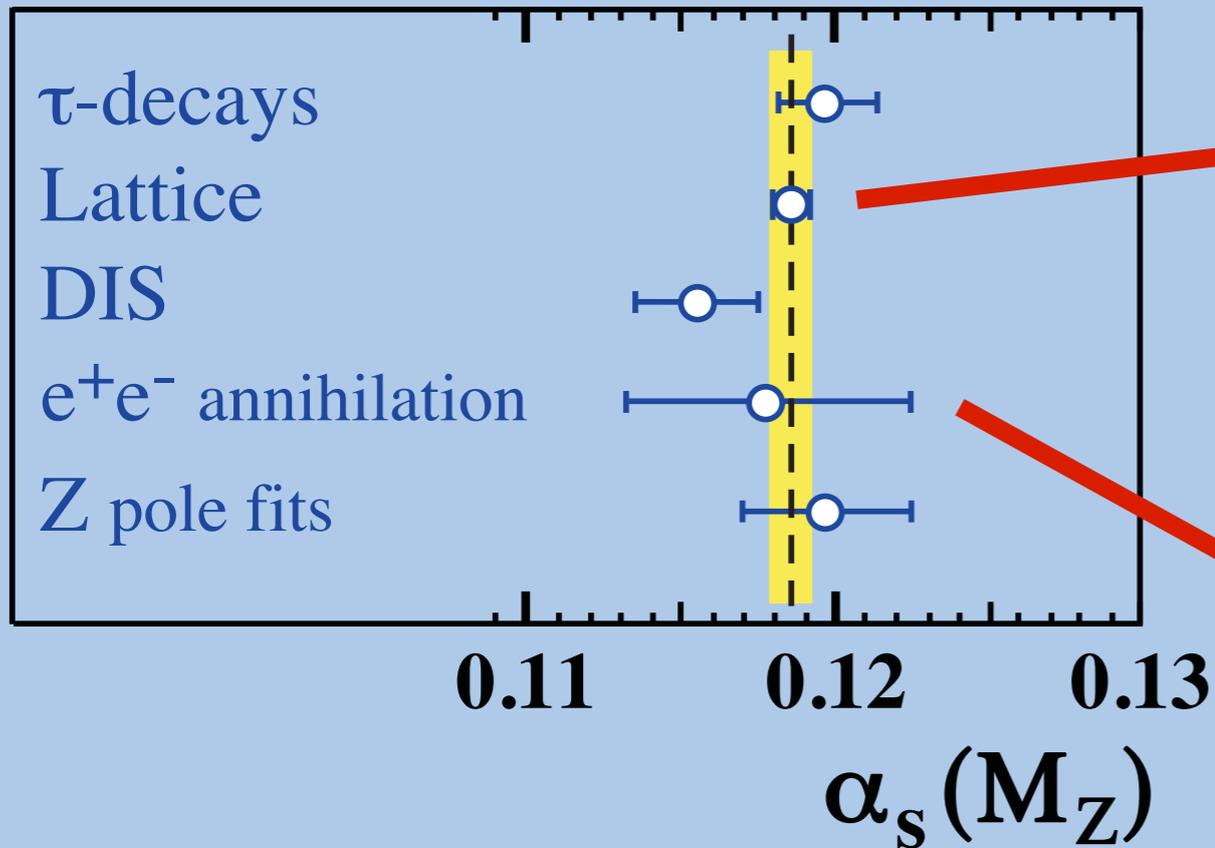
Strong Coupling Constant

PDG world average: $\alpha_s(M_Z) = 0.1184 \pm 0.0006$

w/o lattice inputs:
(~ choice by PDF4LHC) $\alpha_s(M_Z) = 0.1183 \pm 0.0012$

Uncertainty gets amplified in cross sections, e.g. $gg \rightarrow H$

$$\frac{\delta\sigma_{ggH}}{\sigma_{ggH}} \sim 3 \frac{\delta\alpha_s}{\alpha_s}$$



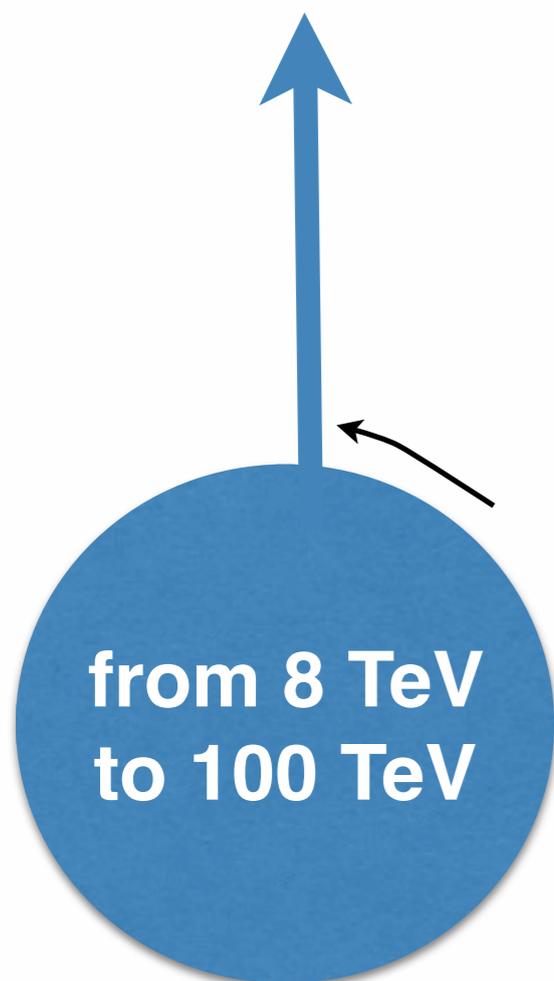
Different extractions quoting small uncertainties are not consistent

Thrust + SCET: 0.1135 ± 0.0010

Wilson loops: 0.1184 ± 0.0006

Differ at 4σ — how do we resolve this?

Collider Reach



Quick and dirty estimates of the reach of future colliders based on existing limits

with Andi Weiler

How soon will LHC@13TeV beat 8TeV searches?

What can high-lumi LHC (3000fb^{-1}) do compared to original LHC plan (300fb^{-1})?

What is the gain from a future 33/50/100/150 TeV collider?

There are already many well-designed searches

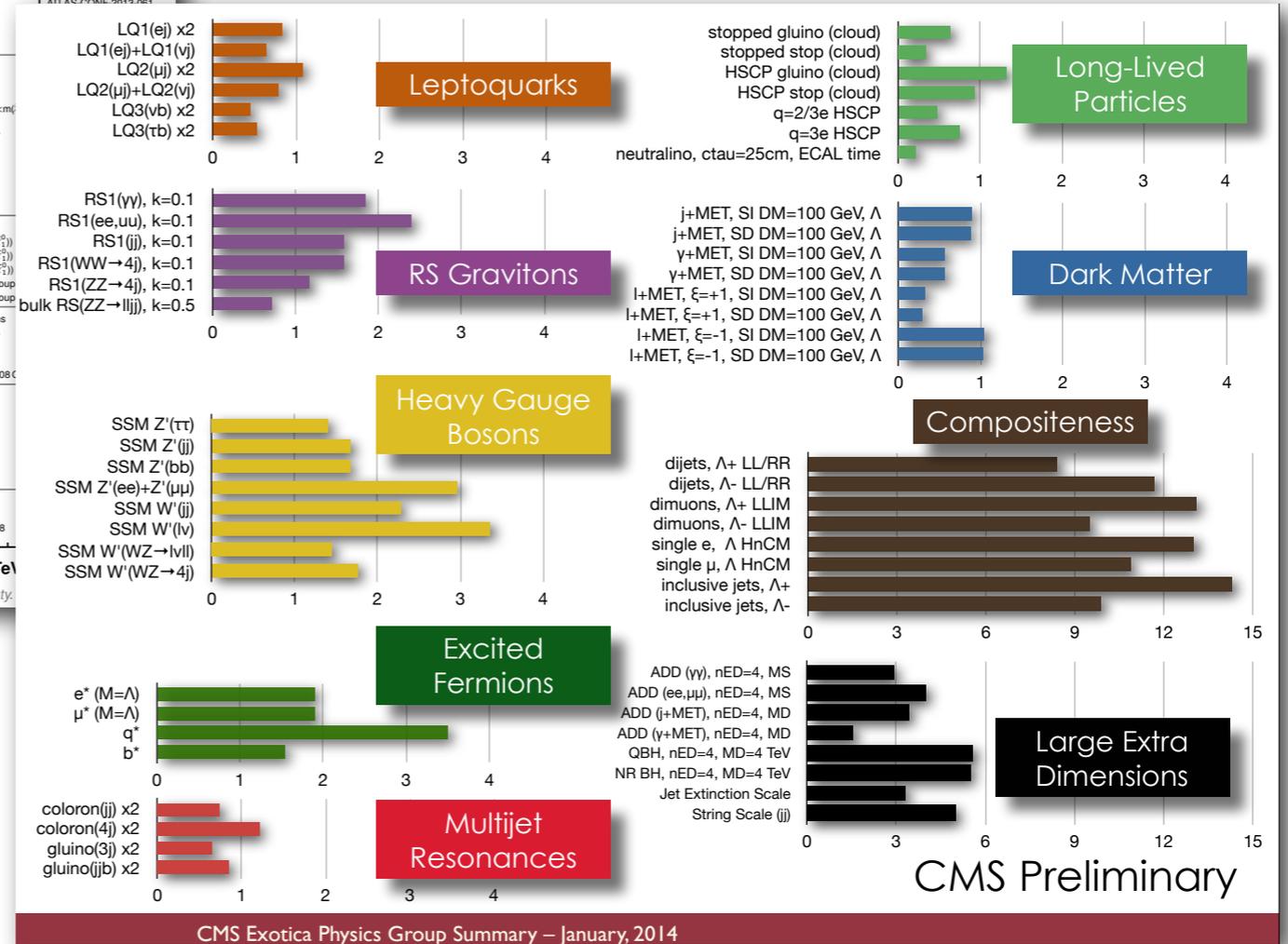
ATLAS SUSY Searches* - 95% CL Lower Limits
 Status: SUSY 2013 ATLAS Preliminary
 $\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1}$ $\sqrt{s} = 7, 8 \text{ TeV}$

| Model | e, μ, τ, γ | Jets | E_T^{miss} | $\int \mathcal{L} dt [\text{fb}^{-1}]$ | Mass limit | Reference |
|---|------------------------|-----------|---------------------|--|--------------------------------|--|
| Inclusive Searches | | | | | | |
| MSUGRA/CMSSM | 0 | 2-6 jets | Yes | 20.3 | \tilde{g}, \tilde{g} 1.7 TeV | ATLAS-CONF-2013-047 |
| MSUGRA/CMSSM | 1 e, μ | 3-6 jets | Yes | 20.3 | \tilde{g} 1.2 TeV | ATLAS-CONF-2013-062 |
| MSUGRA/CMSSM | 0 | 7-10 jets | Yes | 20.3 | \tilde{g} 1.1 TeV | 1308.1841 |
| $\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{q}\tilde{q}_1^0$ | 0 | 2-6 jets | Yes | 20.3 | \tilde{q} 740 GeV | ATLAS-CONF-2013-047 |
| $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{q}_1^0$ | 0 | 2-6 jets | Yes | 20.3 | \tilde{g} 1.3 TeV | ATLAS-CONF-2013-047 |
| $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{q}_1^0 + qqW^{\pm}\tilde{q}_1^0$ | 1 e, μ | 3-6 jets | Yes | 20.3 | \tilde{g} 1.18 TeV | ATLAS-CONF-2013-062 |
| GMSB (\tilde{L} NLSP) | 2 e, μ | 0-3 jets | - | 20.3 | \tilde{g} 1.12 TeV | ATLAS-CONF-2013-089 |
| GMSB (\tilde{L} NLSP) | 2 e, μ | 2-4 jets | Yes | 4.7 | \tilde{g} 1.24 TeV | 1208.4688 |
| GMSB (\tilde{L} NLSP) | 1-2 τ | 0-2 jets | Yes | 20.7 | \tilde{g} 1.4 TeV | ATLAS-CONF-2013-026 |
| GGM (bino NLSP) | 2 γ | - | Yes | 4.8 | \tilde{g} 1.07 TeV | 1209.0753 |
| GGM (wino NLSP) | 1 $e, \mu + \gamma$ | - | Yes | 4.8 | \tilde{g} 619 GeV | ATLAS-CONF-2012-144 |
| GGM (higgsino-bino NLSP) | γ | 1 b | Yes | 4.8 | \tilde{g} 900 GeV | 1211.1167 |
| GGM (higgsino NLSP) | 2 $e, \mu (Z)$ | 0-3 jets | Yes | 5.8 | \tilde{g} 690 GeV | ATLAS-CONF-2012-152 |
| Gravitino LSP | 0 | mono-jet | Yes | 10.5 | \tilde{g} 645 GeV | ATLAS-CONF-2012-147 |
| 3rd gen. \tilde{g} med. | | | | | | |
| $\tilde{g} \rightarrow b\tilde{b}\tilde{q}_1^0$ | 0 | 3 b | Yes | 20.1 | \tilde{g} 1.2 TeV | $m(\tilde{q}_1^0) < 600 \text{ GeV}$ |
| $\tilde{g} \rightarrow t\tilde{t}\tilde{q}_1^0$ | 0 | 7-10 jets | Yes | 20.1 | \tilde{g} 1.1 TeV | $m(\tilde{q}_1^0) < 350 \text{ GeV}$ |
| $\tilde{g} \rightarrow t\tilde{t}\tilde{q}_1^0$ | 0-1 e, μ | 3 b | Yes | 20.1 | \tilde{g} 1.34 TeV | $m(\tilde{q}_1^0) < 400 \text{ GeV}$ |
| $\tilde{g} \rightarrow b\tilde{t}\tilde{q}_1^0$ | 0-1 e, μ | 3 b | Yes | 20.1 | \tilde{g} 1.3 TeV | $m(\tilde{q}_1^0) < 300 \text{ GeV}$ |
| 3rd gen. squarks direct production | | | | | | |
| $\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{q}_1^0$ | 2 $e, \mu (SS)$ | 2 b | Yes | 20.1 | \tilde{b}_1 100-620 GeV | $m(\tilde{q}_1^0) < 90 \text{ GeV}$ |
| $\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{q}_1^0$ | 2 $e, \mu (SS)$ | 0-3 b | Yes | 20.7 | \tilde{b}_1 275-430 GeV | $m(\tilde{q}_1^0) = 2 m(\tilde{q}_1^0)$ |
| $\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow b\tilde{q}_1^0$ | 1-2 e, μ | 1-2 b | Yes | 4.7 | \tilde{t}_1 110-167 GeV | $m(\tilde{q}_1^0) = 55 \text{ GeV}$ |
| $\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow W\tilde{b}_1^0$ | 2 e, μ | 0-2 jets | Yes | 20.3 | \tilde{t}_1 130-220 GeV | $m(\tilde{q}_1^0) = m(\tilde{t}_1) - m(W) - 50 \text{ GeV}, m(\tilde{q}_1^0) < m(\tilde{t}_1)$ |
| $\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow t\tilde{q}_1^0$ | 2 e, μ | 2 jets | Yes | 20.3 | \tilde{t}_1 225-525 GeV | $m(\tilde{q}_1^0) = 0 \text{ GeV}$ |
| $\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow b\tilde{q}_1^0$ | 1 e, μ | 2 b | Yes | 20.1 | \tilde{t}_1 150-580 GeV | $m(\tilde{q}_1^0) < 200 \text{ GeV}, m(\tilde{q}_1^0) - m(\tilde{q}_1^0) = 5 \text{ GeV}$ |
| $\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{q}_1^0$ | 0 | 2 b | Yes | 20.7 | \tilde{t}_1 200-610 GeV | $m(\tilde{q}_1^0) = 0 \text{ GeV}$ |
| $\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{q}_1^0$ | 0 | 2 b | Yes | 20.5 | \tilde{t}_1 320-660 GeV | $m(\tilde{q}_1^0) = 0 \text{ GeV}$ |
| $\tilde{t}_1\tilde{t}_1$ (natural GMSB) | 2 $e, \mu (Z)$ | 1 b | Yes | 20.7 | \tilde{t}_1 500 GeV | $m(\tilde{q}_1^0) - m(\tilde{q}_1^0) < 85 \text{ GeV}$ |
| $\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$ | 3 $e, \mu (Z)$ | 1 b | Yes | 20.7 | \tilde{t}_2 271-520 GeV | $m(\tilde{q}_1^0) > 150 \text{ GeV}$ |
| EW direct | | | | | | |
| $\tilde{L}_L\tilde{L}_L, \tilde{L} \rightarrow \tilde{L}\tilde{q}_1^0$ | 2 e, μ | 0 | Yes | 20.3 | \tilde{L} 85-315 GeV | $m(\tilde{q}_1^0) = 0 \text{ GeV}$ |
| $\tilde{L}_L\tilde{L}_L, \tilde{L} \rightarrow \tilde{L}\nu(\tilde{\nu})$ | 2 e, μ | 0 | Yes | 20.3 | \tilde{L} 125-450 GeV | $m(\tilde{q}_1^0) = 0 \text{ GeV}, m(\tilde{L}) = 0.5(m(\tilde{L}) + m(\tilde{q}_1^0))$ |
| $\tilde{L}_L\tilde{L}_L, \tilde{L} \rightarrow \tilde{L}\nu(\tilde{\nu})$ | 2 τ | - | Yes | 20.7 | \tilde{L} 180-330 GeV | $m(\tilde{q}_1^0) = 0 \text{ GeV}, m(\tilde{L}) = 0.5(m(\tilde{L}) + m(\tilde{q}_1^0))$ |
| $\tilde{L}_L\tilde{L}_L, \tilde{L} \rightarrow \tilde{L}\nu(\tilde{\nu})$ | 3 e, μ | 0 | Yes | 20.7 | \tilde{L} 315 GeV | $m(\tilde{q}_1^0) = m(\tilde{q}_1^0), m(\tilde{q}_1^0) = 0, \text{ sleptons decoupled}$ |
| $\tilde{L}_L\tilde{L}_L, \tilde{L} \rightarrow W\tilde{Z}$ | 3 e, μ | 0 | Yes | 20.7 | \tilde{L} 285 GeV | $m(\tilde{q}_1^0) = m(\tilde{q}_1^0), m(\tilde{q}_1^0) = 0, \text{ sleptons decoupled}$ |
| $\tilde{L}_L\tilde{L}_L, \tilde{L} \rightarrow W\tilde{Z}$ | 1 e, μ | 2 b | Yes | 20.3 | \tilde{L} 270 GeV | $m(\tilde{q}_1^0) - m(\tilde{q}_1^0) = 160 \text{ MeV}, \tau(\tilde{q}_1^0) = 0.2 \text{ ns}$ |
| Long-lived particles | | | | | | |
| Direct $\tilde{L}_L\tilde{L}_L$ prod., long-lived \tilde{L}_L | Disapp. trk | 1 jet | Yes | 20.3 | \tilde{L} 230 GeV | $m(\tilde{q}_1^0) = 100 \text{ GeV}, 10 \mu\text{s} < \tau(\tilde{L}) < 1000 \text{ s}$ |
| Stable, stopped \tilde{g} R-hadron | 0 | 1-5 jets | Yes | 22.9 | \tilde{g} 475 GeV | $10^{-4} < \tau(\tilde{g}) < 50$ |
| GMSB, stable $\tilde{L}, \tilde{L} \rightarrow \tilde{L}(e, \mu) + \tau$ | 1-2 μ | - | - | 15.9 | \tilde{L} 832 GeV | $0.4 < \tau(\tilde{L}) < 2 \text{ ns}$ |
| GMSB, $\tilde{L} \rightarrow \gamma G$, long-lived \tilde{L} | 2 γ | - | - | 4.7 | \tilde{L} 230 GeV | $1.5 < \tau < 156 \text{ mm}, \text{BR}(\mu) = 1, m(\tilde{q}_1^0) = 108 \text{ GeV}$ |
| $\tilde{q}\tilde{q}, \tilde{L} \rightarrow q\tilde{q}\mu$ (RPV) | 1 μ , displ. vtx | - | - | 20.3 | \tilde{L} 1.0 TeV | |
| RPV | | | | | | |
| LFV $pp \rightarrow \tilde{\nu}_e + X, \tilde{\nu}_e \rightarrow e + \mu$ | 2 e, μ | - | - | 4.6 | $\tilde{\nu}_e$ 1.61 TeV | $A_{131} = 0.10, A_{132} = 0.05$ |
| LFV $pp \rightarrow \tilde{\nu}_e + X, \tilde{\nu}_e \rightarrow e(\mu) + \tau$ | 1 $e, \mu + \tau$ | - | - | 4.6 | $\tilde{\nu}_e$ 1.1 TeV | $A_{131} = 0.10, A_{132} = 0.05$ |
| Bilinear RPV CMSSM | 1 e, μ | 7 jets | Yes | 4.7 | \tilde{g}, \tilde{g} 1.2 TeV | $m(\tilde{q}) = m(\tilde{g}), c\tau_{\text{LSP}} < 1 \text{ mm}$ |
| $\tilde{L}_L\tilde{L}_L, \tilde{L} \rightarrow W\tilde{L}_L, \tilde{L} \rightarrow ee\tilde{\nu}_e, e\mu\tilde{\nu}_e$ | 4 e, μ | - | Yes | 20.7 | \tilde{L} 760 GeV | $m(\tilde{q}_1^0) > 300 \text{ GeV}, A_{131} > 0$ |
| $\tilde{L}_L\tilde{L}_L, \tilde{L} \rightarrow W\tilde{L}_L, \tilde{L} \rightarrow \tau\tilde{\nu}_e, e\tau\tilde{\nu}_e$ | 3 $e, \mu + \tau$ | - | Yes | 20.7 | \tilde{L} 350 GeV | $m(\tilde{q}_1^0) > 80 \text{ GeV}, A_{131} > 0$ |
| $\tilde{g} \rightarrow q\tilde{q}$ | 0 | 6-7 jets | - | 20.3 | \tilde{g} 916 GeV | $\text{BR}(\tau) = \text{BR}(\mu) = \text{BR}(e) = 0\%$ |
| $\tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow b\tilde{s}$ | 2 $e, \mu (SS)$ | 0-3 b | Yes | 20.7 | \tilde{g} 880 GeV | |
| Other | | | | | | |
| Scalar gluon pair, sgluon $\rightarrow q\tilde{q}$ | 0 | 4 jets | - | 4.6 | sgluon 100-287 GeV | incl. limit from 1110.2693 |
| Scalar gluon pair, sgluon $\rightarrow t\tilde{t}$ | 2 $e, \mu (SS)$ | 1 b | Yes | 14.3 | sgluon 800 GeV | |
| WIMP interaction (D5, Dirac χ) | 0 | mono-jet | Yes | 10.5 | \tilde{L} scale 704 GeV | $m(\chi) < 80 \text{ GeV}$, limit of $\sim 687 \text{ GeV}$ for D8 |

Mass scale [TeV] 10^{-1} 1

$\sqrt{s} = 7 \text{ TeV}$ full data $\sqrt{s} = 8 \text{ TeV}$ partial data $\sqrt{s} = 8 \text{ TeV}$ full data

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1σ theoretical signal cross section uncertainty.



How do we leverage that experience to guesstimate future reaches?

A rough way of doing it

Suppose ATLAS/CMS are currently sensitive to gluinos of 1250 GeV (95% CL_s , 8 TeV, 20 fb⁻¹)



Work out how many signal events that corresponds to



Find out for what gluino mass you would get the same number of signal events at 14 TeV with 300 fb⁻¹ (assume # of background events scales same way)

What we're discussing is solution of the following equation for M_{high}

$$\frac{N_{\text{signal-events}}(M_{\text{high}}^2, 14 \text{ TeV}, \text{Lumi})}{N_{\text{signal-events}}(M_{\text{low}}^2, 8 \text{ TeV}, 19 \text{ fb}^{-1})} = 1$$

Many complications (e.g. coupling constants & other prefactors) mostly cancel in the ratio.

Dependence on M and on \sqrt{s} mostly comes about through parton distribution functions (PDFs) & simple dimensions.

Instead of cross section ratio, use **parton luminosity ratio**

Assume dominance of a single partonic scattering channel, ij (you have to know enough physics to figure out which is most appropriate).

Equation we solve to find M_{high} is then

$$\frac{\mathcal{L}_{ij}(M_{\text{high}}^2, s_{\text{high}})}{\mathcal{L}_{ij}(M_{\text{low}}^2, s_{\text{low}})} \times \frac{\text{lumi}_{\text{high}}}{\text{lumi}_{\text{low}}} = \frac{M_{\text{high}}^2}{M_{\text{low}}^2}$$

The tools we use for this are
LHAPDF and HOPPET
most plots with MSTW2008 NNLO PDFs

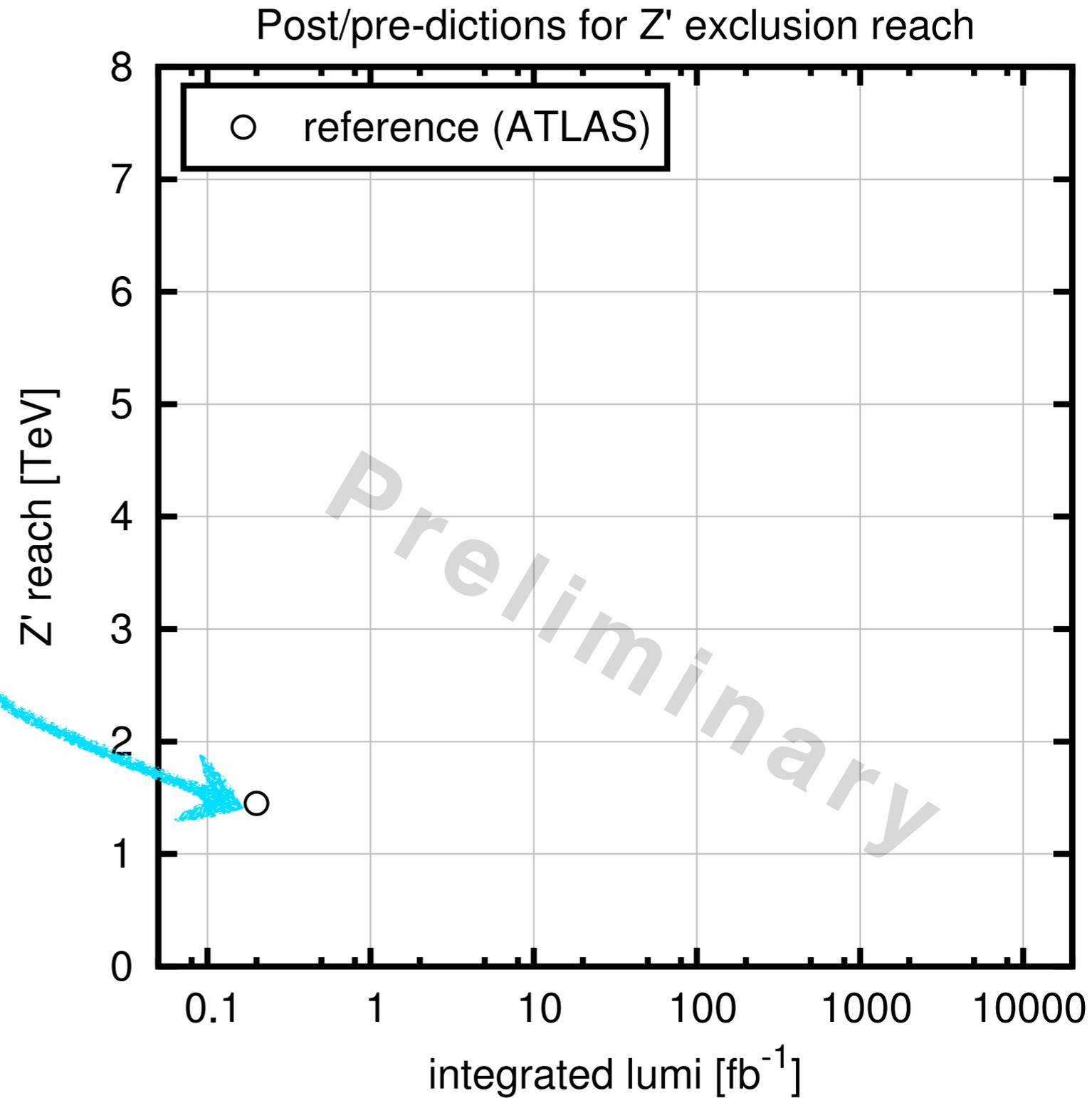
$$\mathcal{L}_{ij}(M^2, s) = \int_{\tau}^1 \frac{dx}{x} x f_i(x, M^2) \frac{\tau}{x} f_j\left(\frac{\tau}{x}, M^2\right) \quad \tau \equiv \frac{M^2}{s}$$

i & j parton

Does it work?

Try a Z' search. Take a baseline analysis:

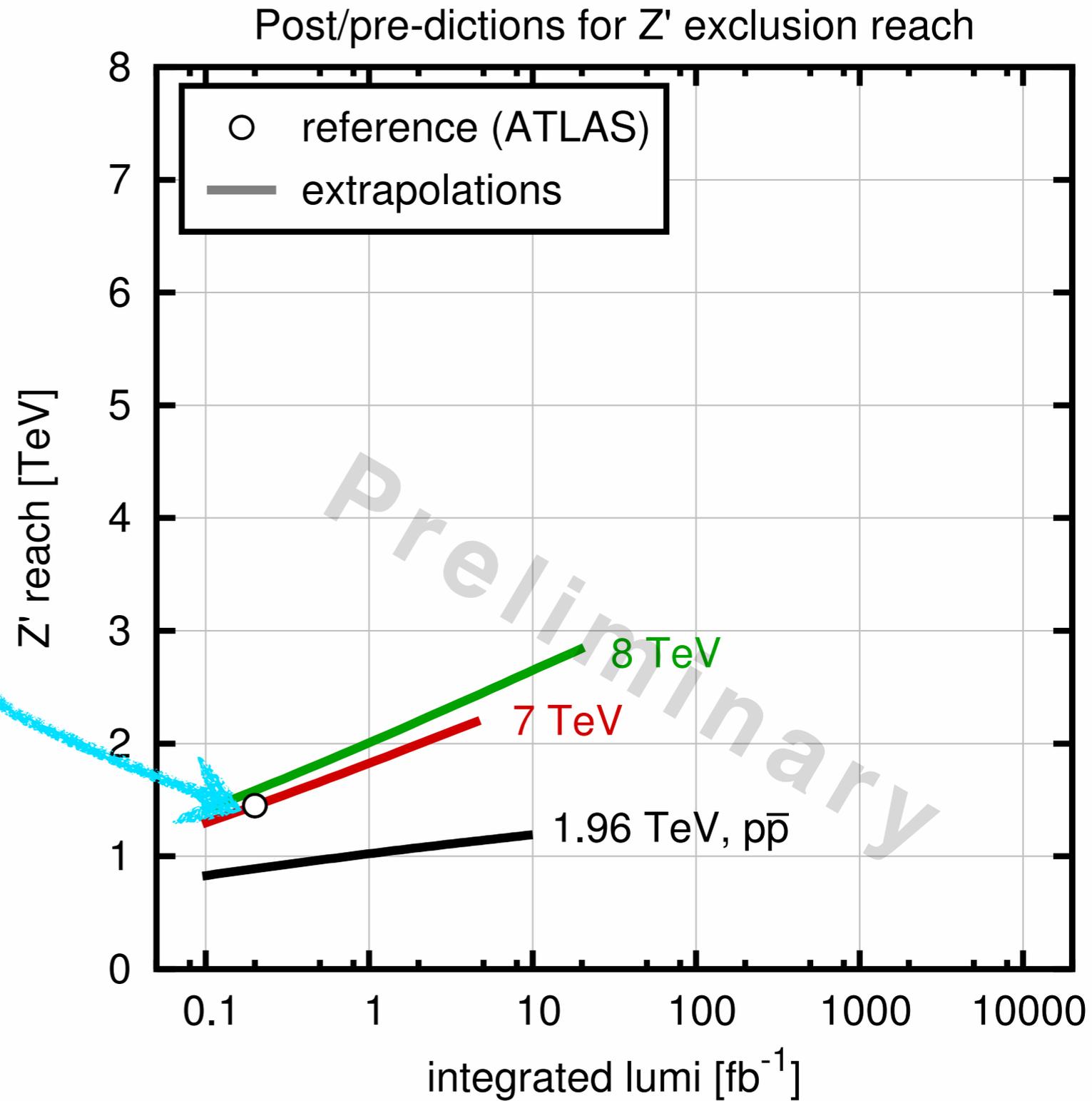
ATLAS,
0.2 fb⁻¹ @ 7 TeV
excludes M < 1450 GeV



Try a Z' search. Take a baseline analysis:

ATLAS,
 0.2 fb^{-1} @ 7 TeV
excludes $M < 1450 \text{ GeV}$

“Predict” exclusions
at other lumis &
energies (assume $q\bar{q}$)

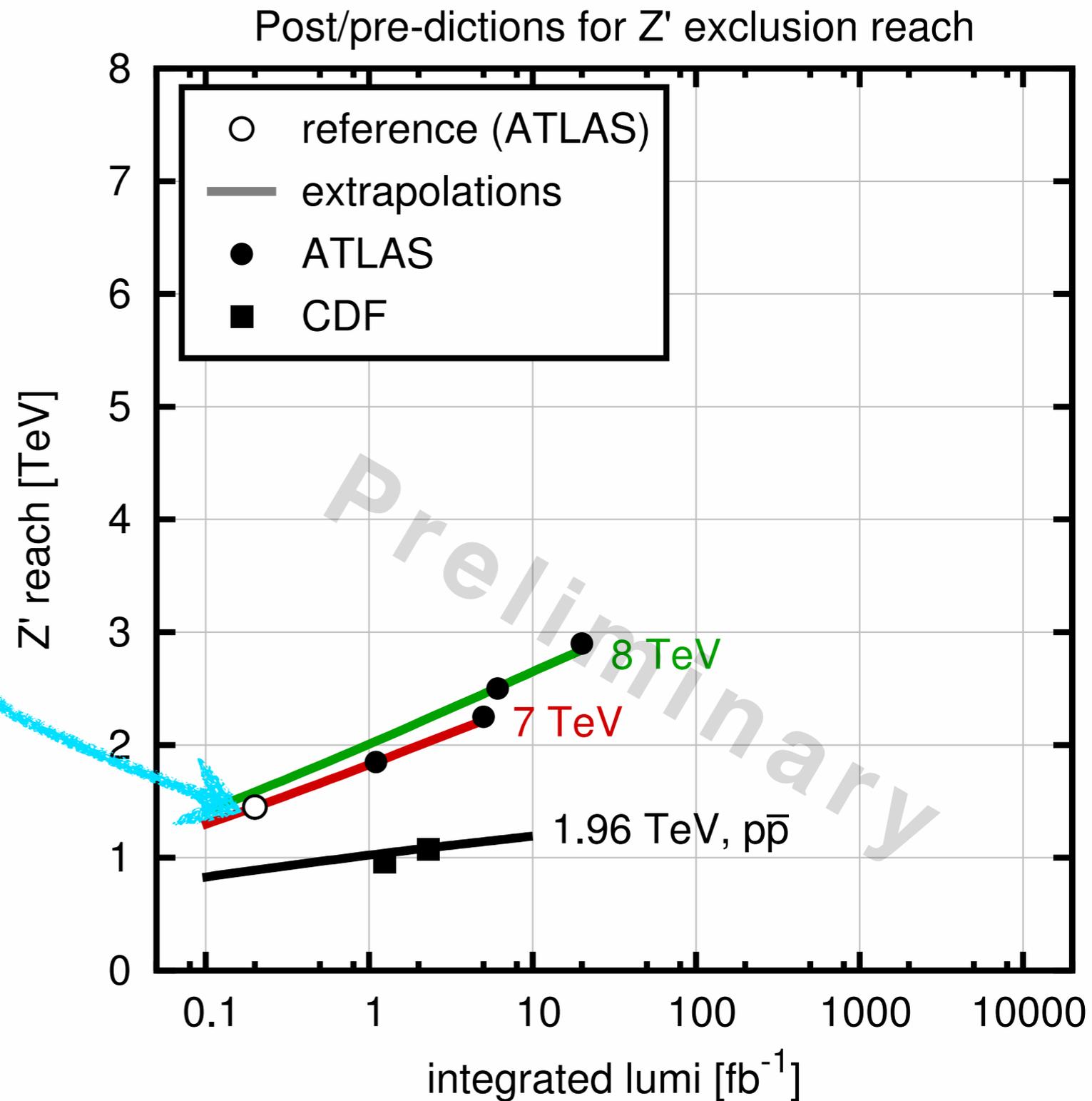


Try a Z' search. Take a baseline analysis:

ATLAS,
 0.2 fb^{-1} @ 7 TeV
excludes $M < 1450 \text{ GeV}$

“Predict” exclusions
at other lumis &
energies (assume $q\bar{q}$)

Compare to actual
exclusions

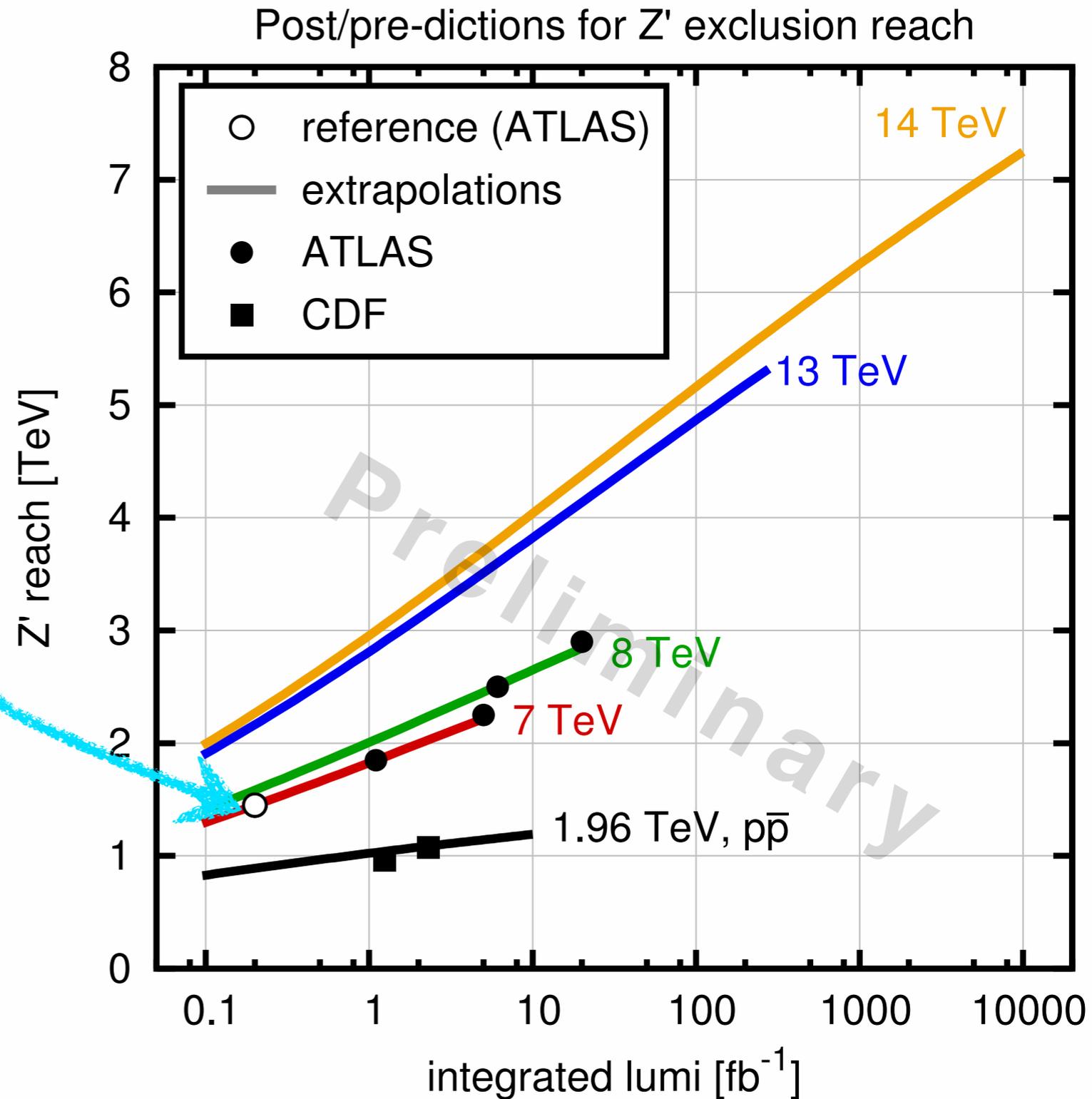


Try a Z' search. Take a baseline analysis:

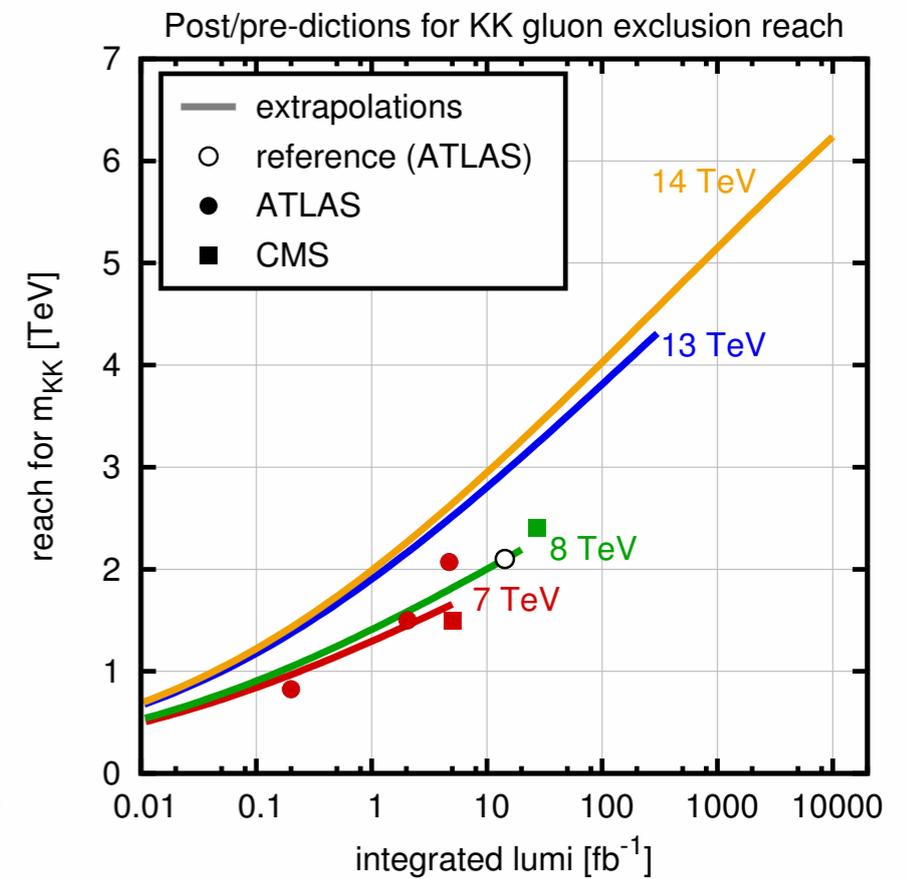
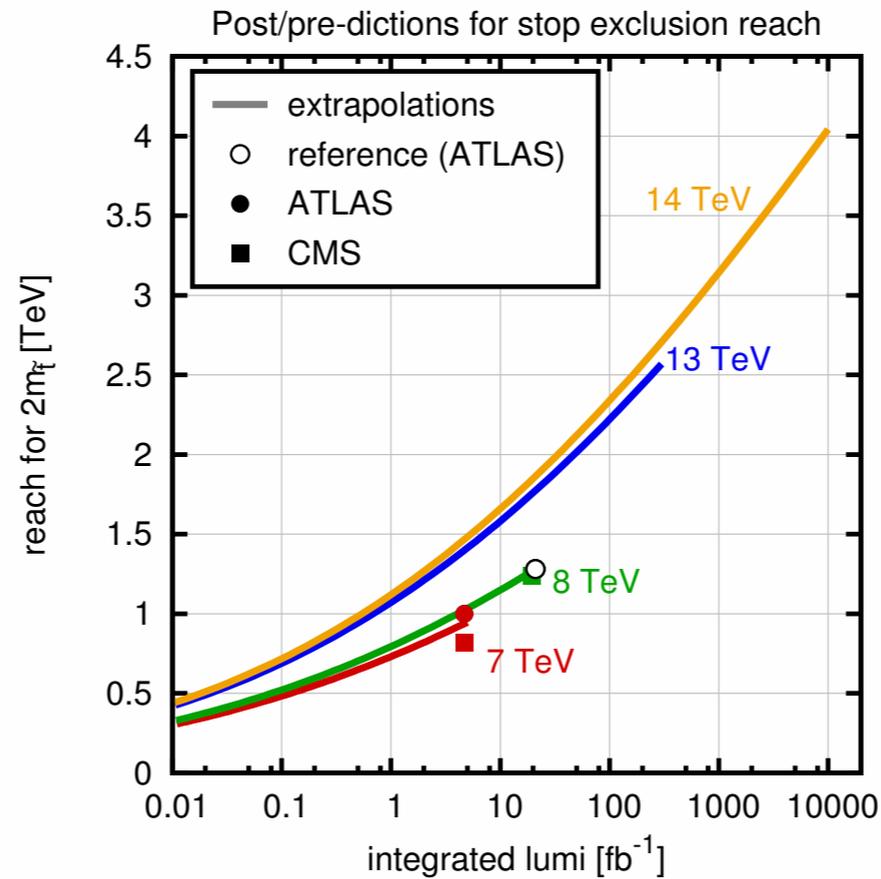
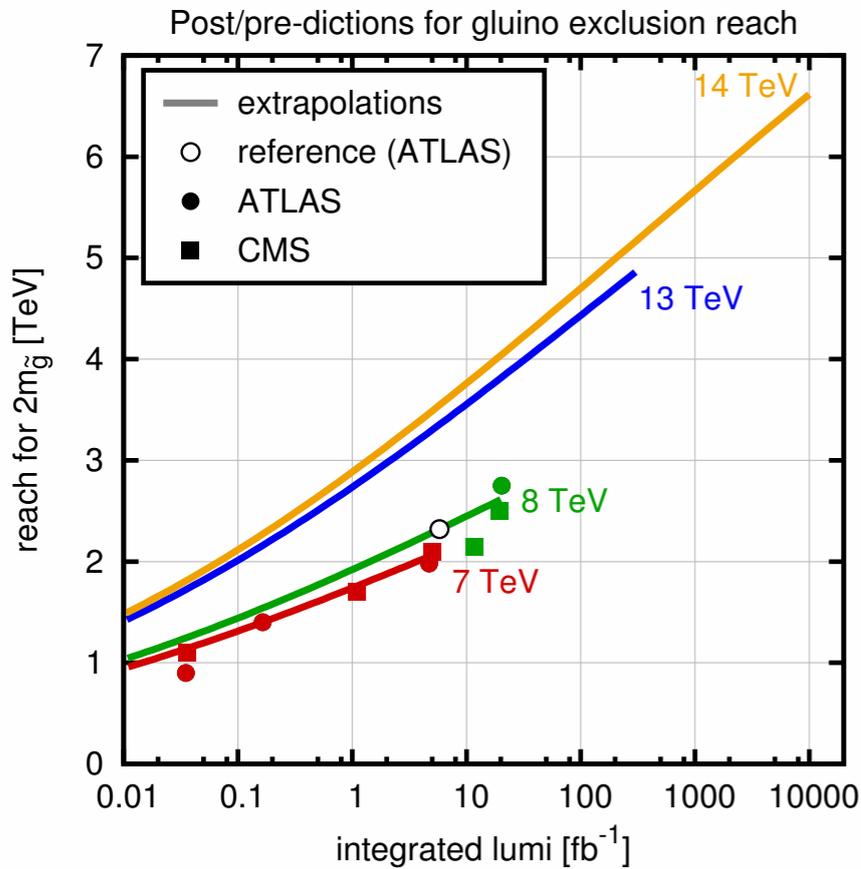
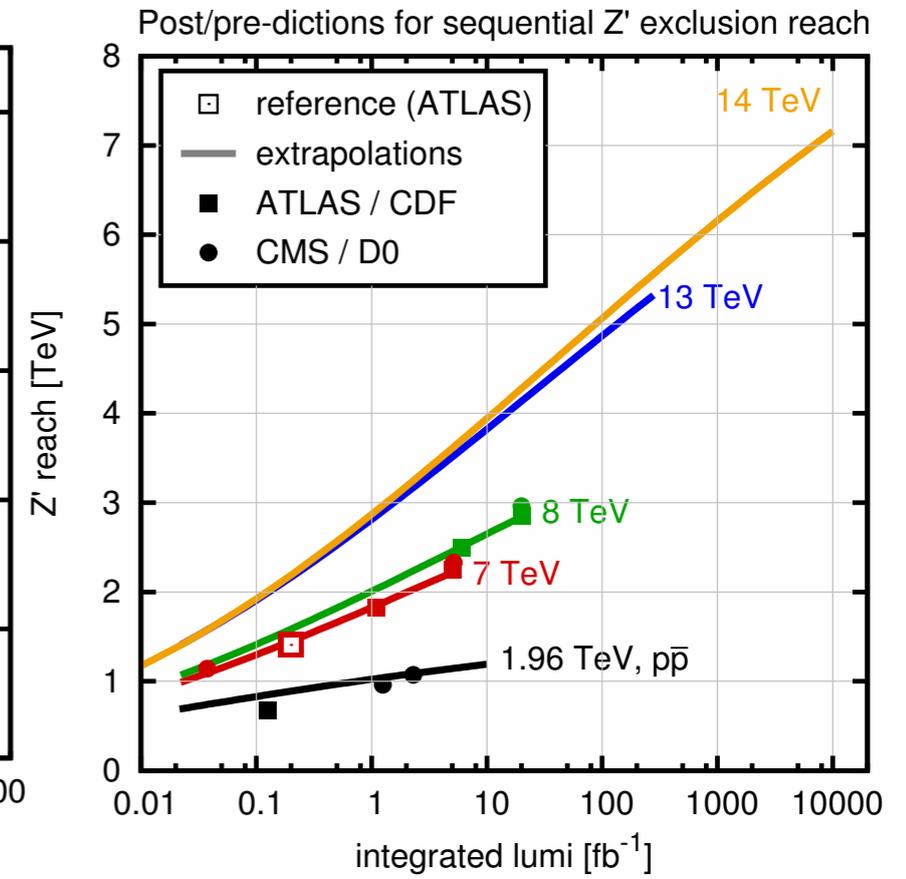
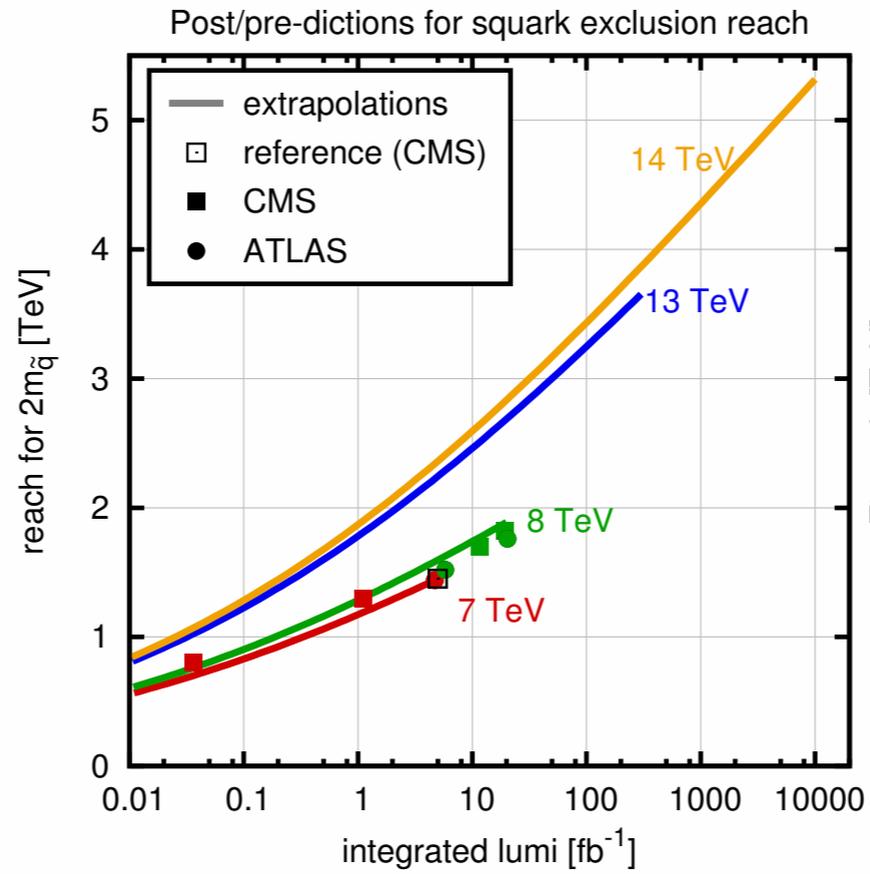
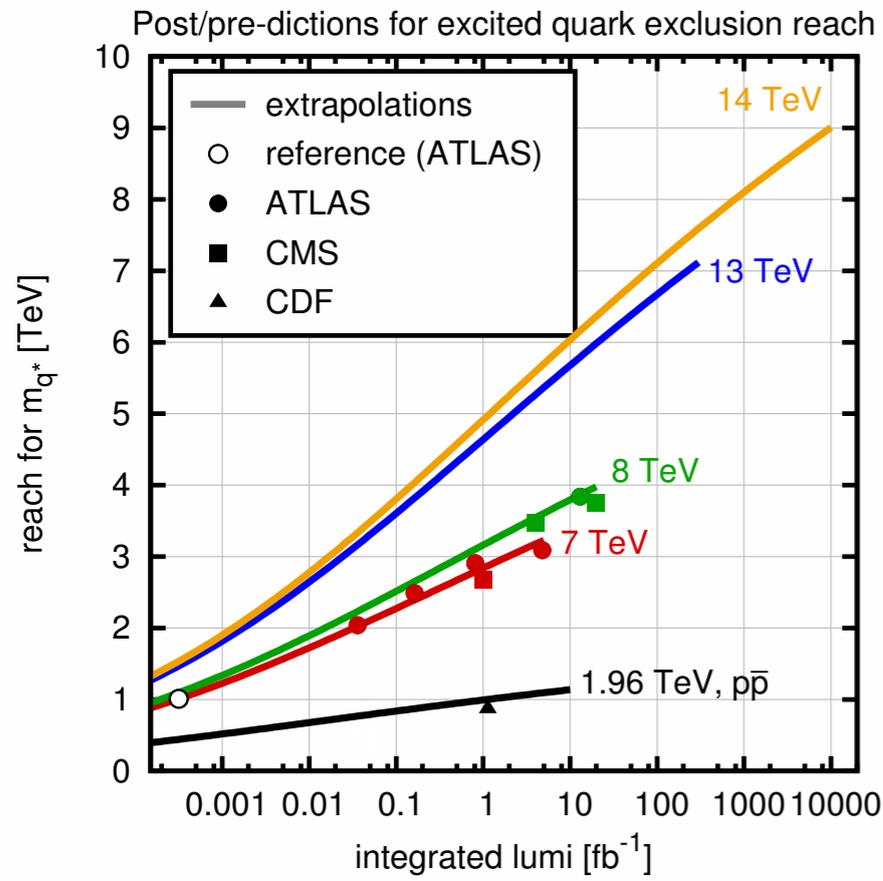
ATLAS,
 0.2 fb^{-1} @ 7 TeV
excludes $M < 1450 \text{ GeV}$

“Predict” exclusions
at other lumis &
energies (assume $q\bar{q}$)

Compare to actual
exclusions



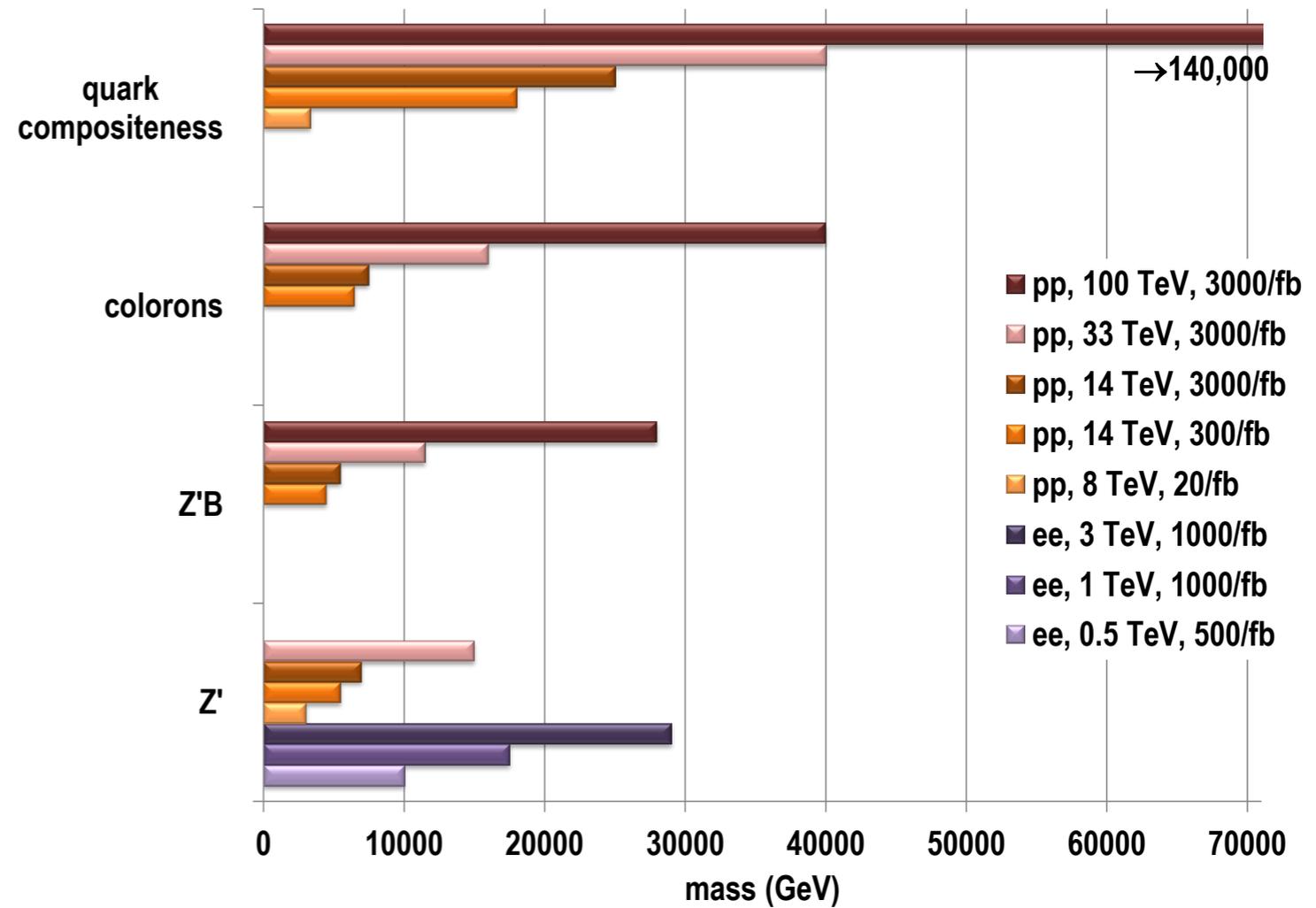
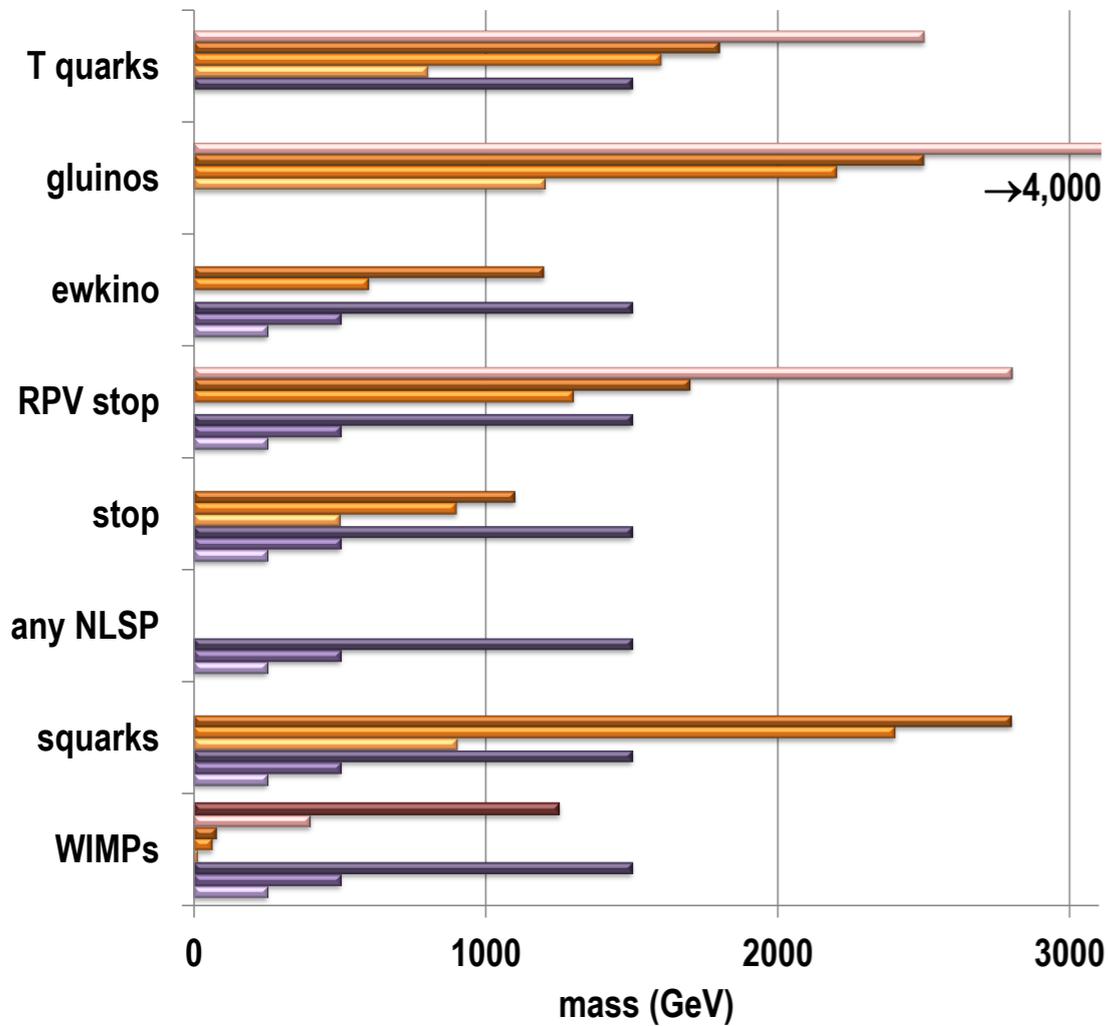
Maybe it only works so well because it's a simple search?
(Signal & Bkgd are both $q\bar{q}$ driven)



Future colliders

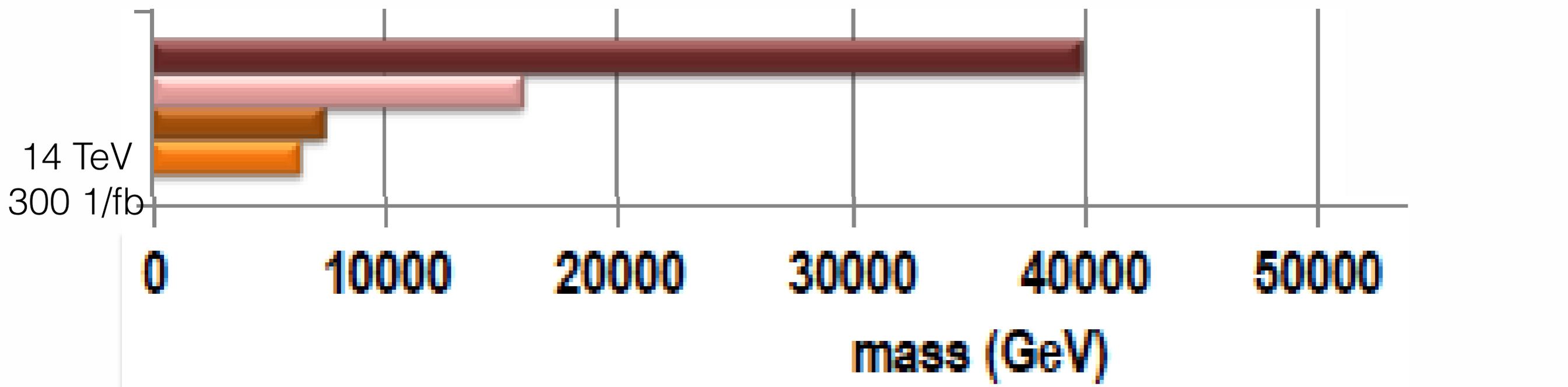
- We're ignoring all subtleties, just going for a baseline check
- If our estimate differs a lot from sophisticated simulations, something interesting has happened:
 - brick-wall (new irreducible backgrounds, granularity of assumed detectors, ...)
 - overly conservative or non-optimal estimates

Future colliders comparison



Energy Frontier Snowmass study ([1311.0299](https://arxiv.org/abs/1311.0299))

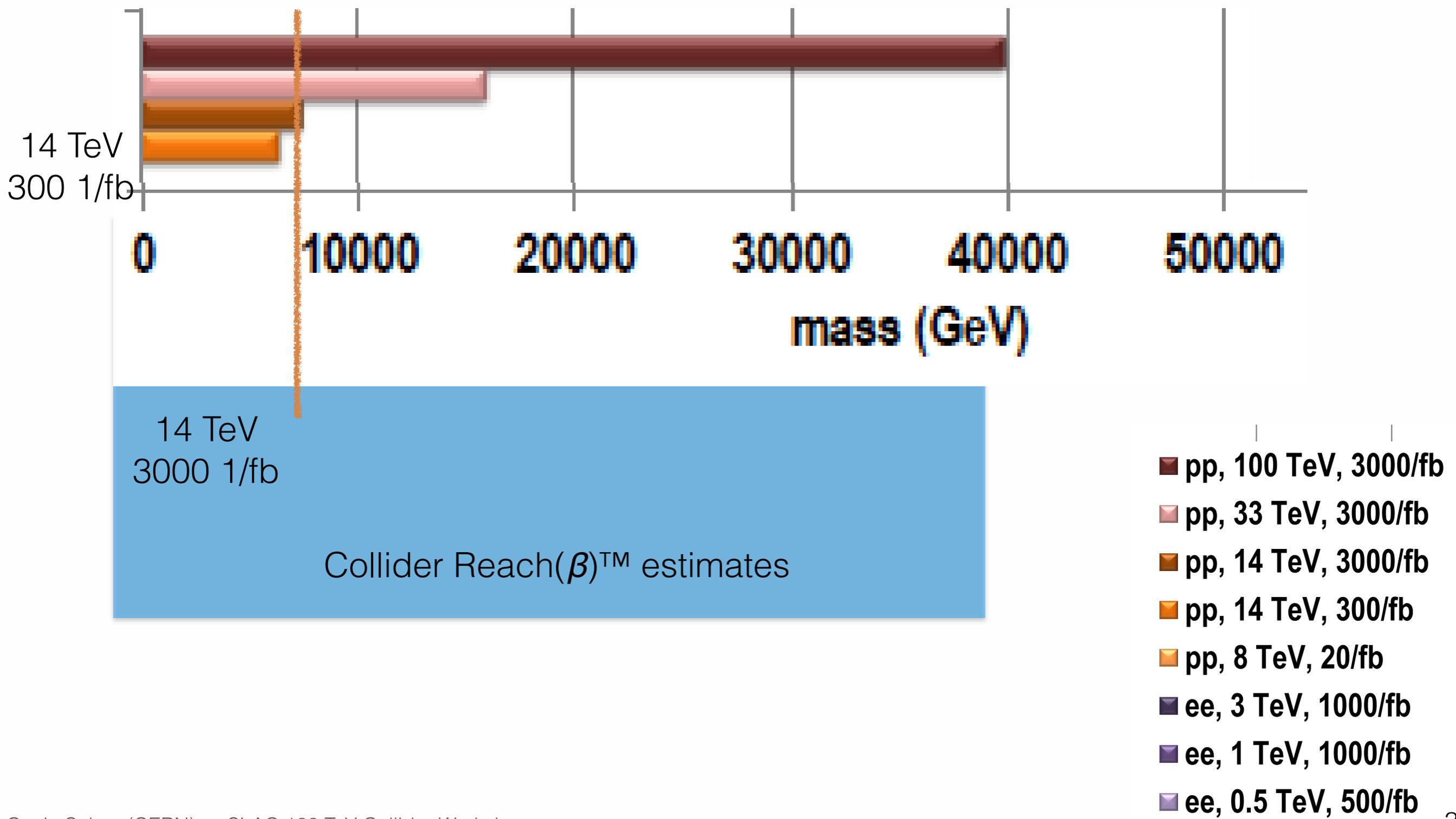
Colorons



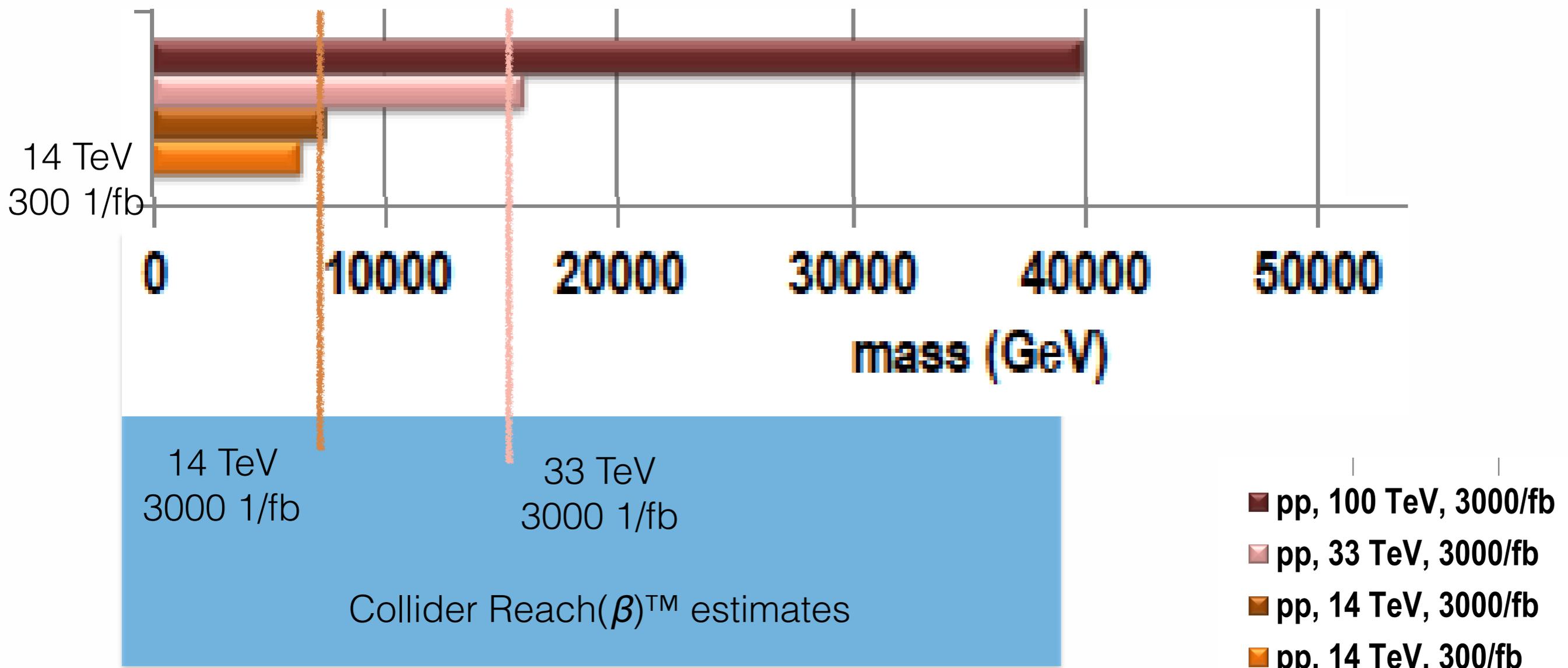
Collider Reach(β)™ estimates

- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

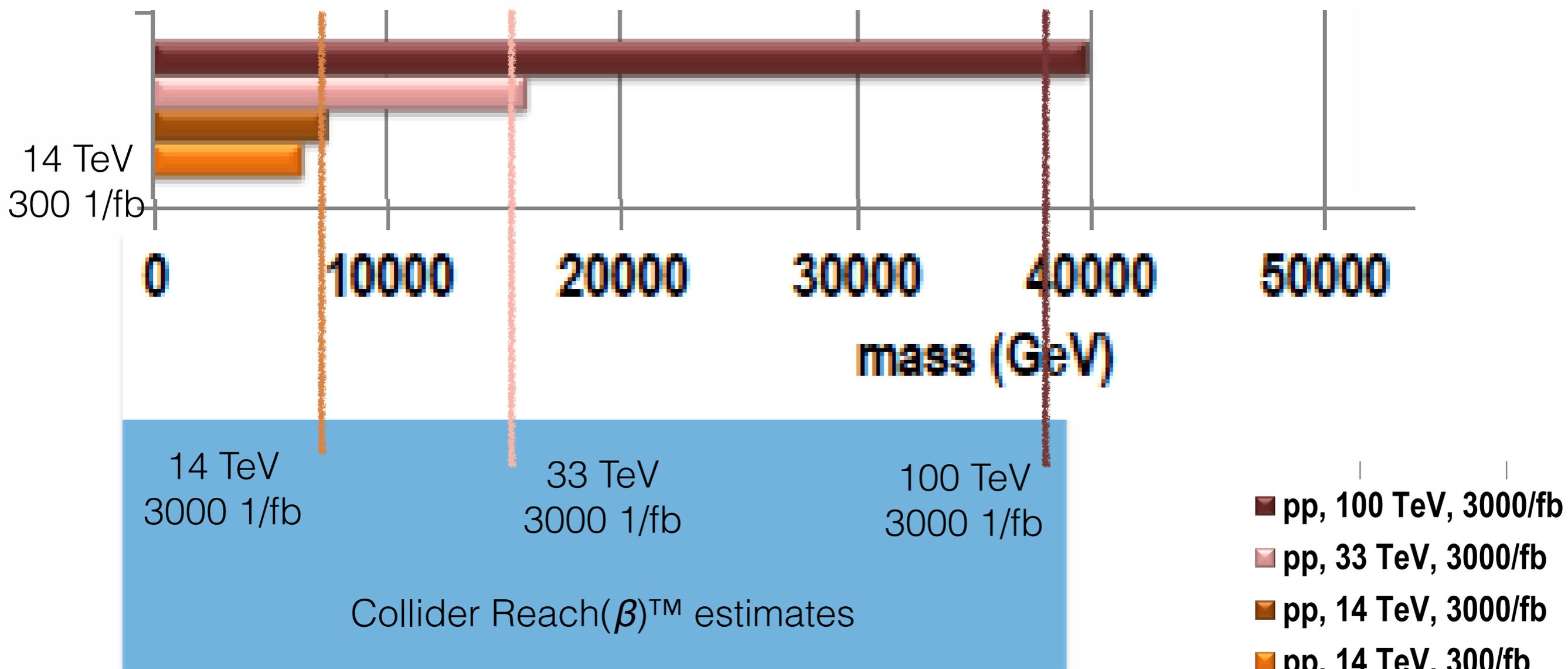
Colorons



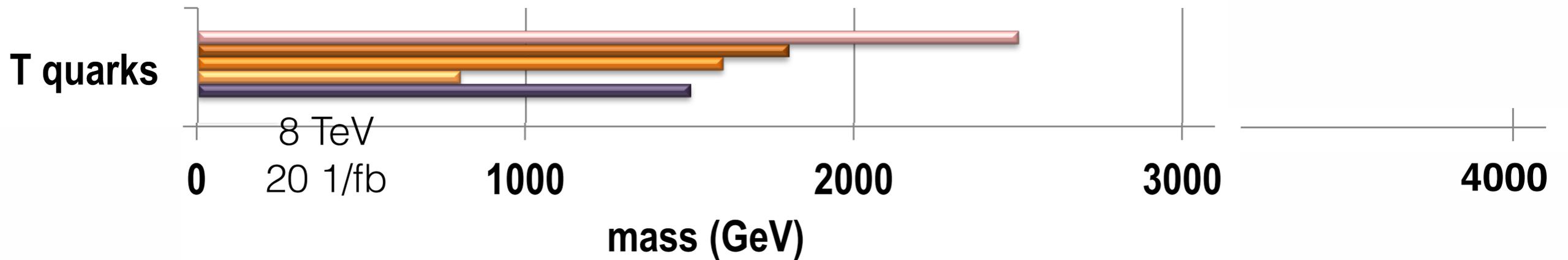
Colorons



Colorons



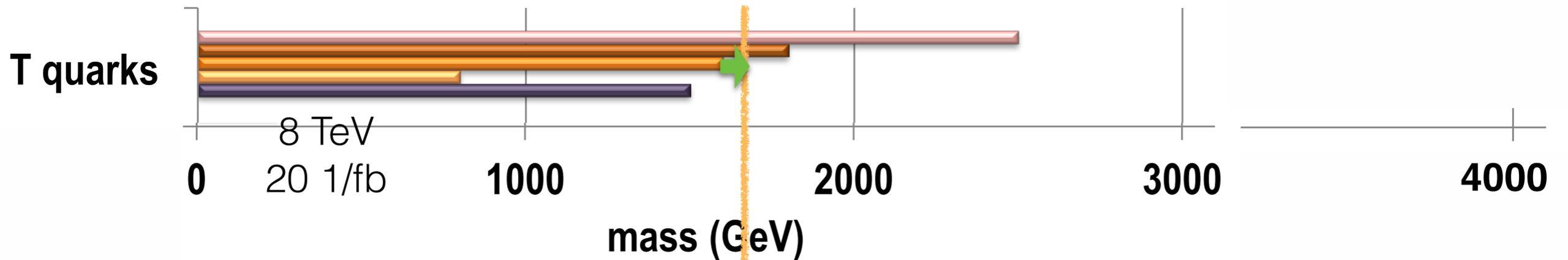
T Quarks



- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

Collider Reach(β)TM estimates

T Quarks

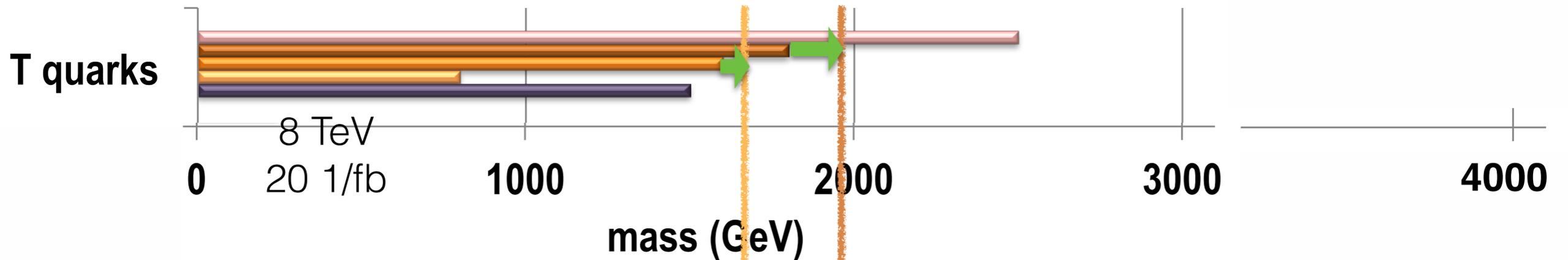


- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

14 TeV
300 1/fb

Collider Reach(β)TM estimates

T Quarks



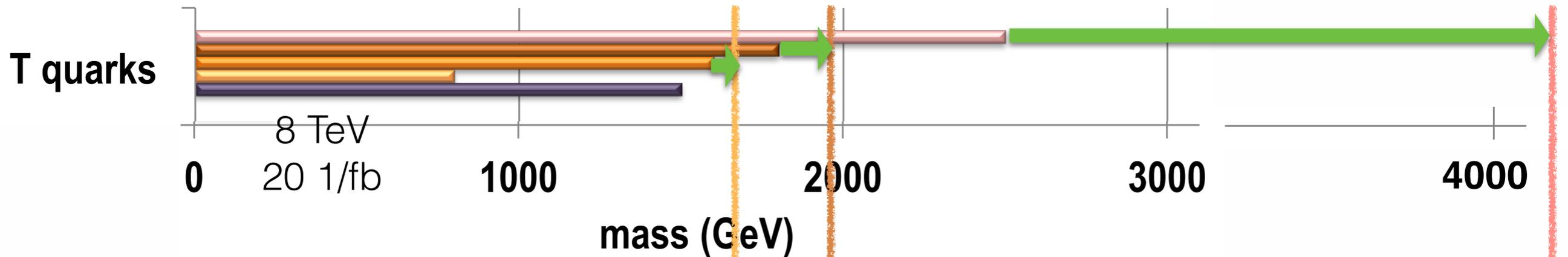
- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

14 TeV
300 1/fb

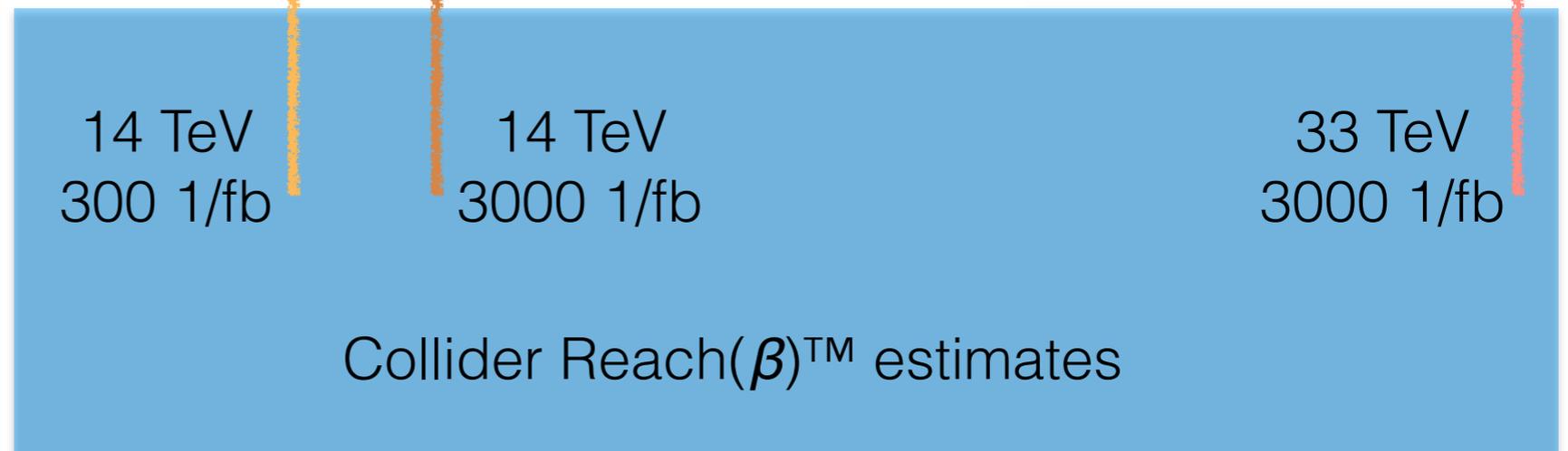
14 TeV
3000 1/fb

Collider Reach(β)TM estimates

T Quarks



- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb



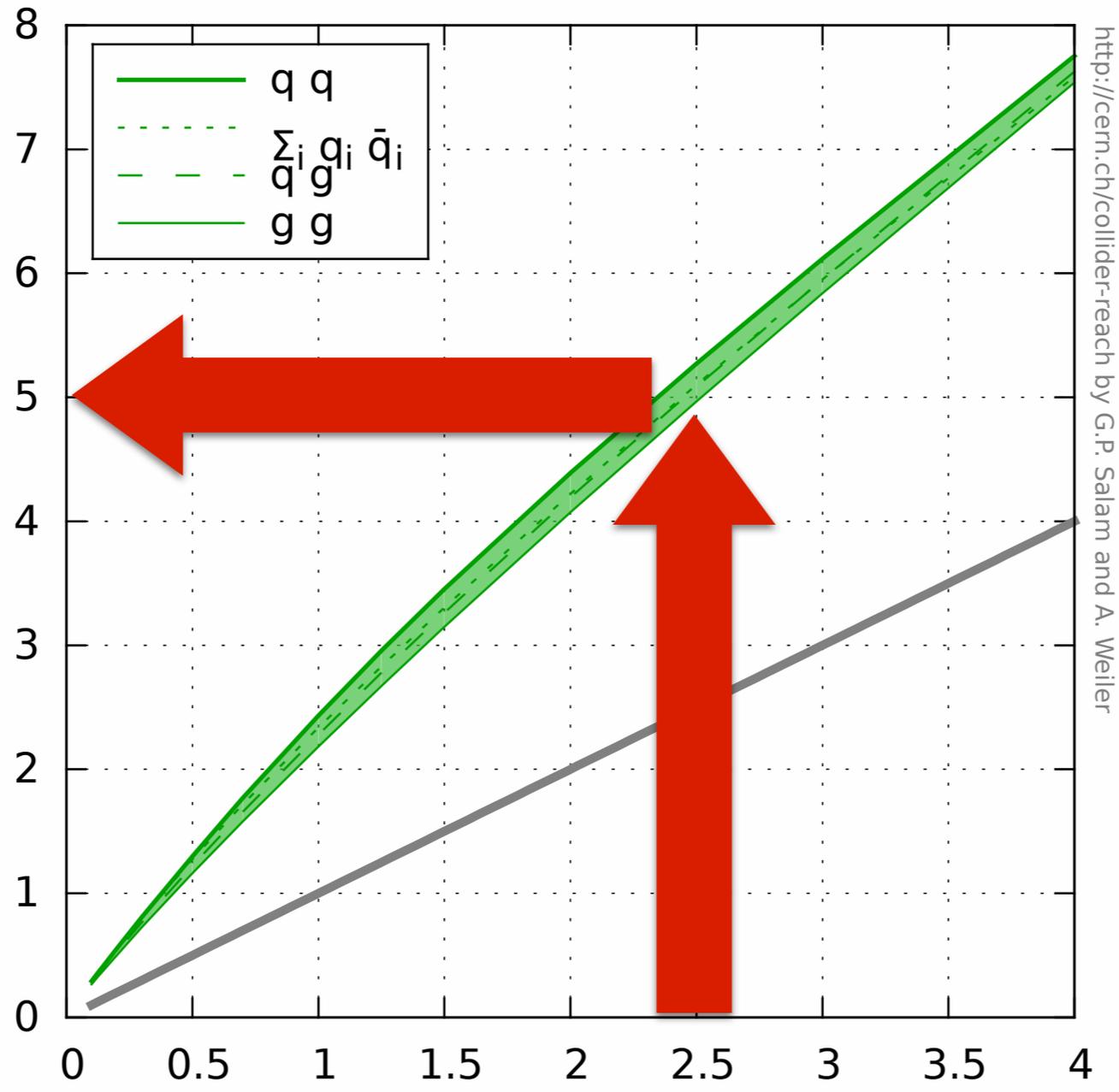
Issue seems to be detector granularity

From your iPhone
(or a generic browser)
cern.ch/collider-reach

From your Android Phone
(or a generic browser)
cern.ch/collider-reach

Collider 1: CoM energy TeV, integrated luminosity fb⁻¹
 Collider 2: CoM energy TeV, integrated luminosity fb⁻¹
 PDF:

**Mass [TeV] at
collider #2**



<http://cern.ch/collider-reach> by G.P. Salam and A. Weiler

Mass [TeV] at collider #1

Collider 1: CoM energy

8

TeV, integrated luminosity

20

fb^{-1}

Collider 2: CoM energy

14

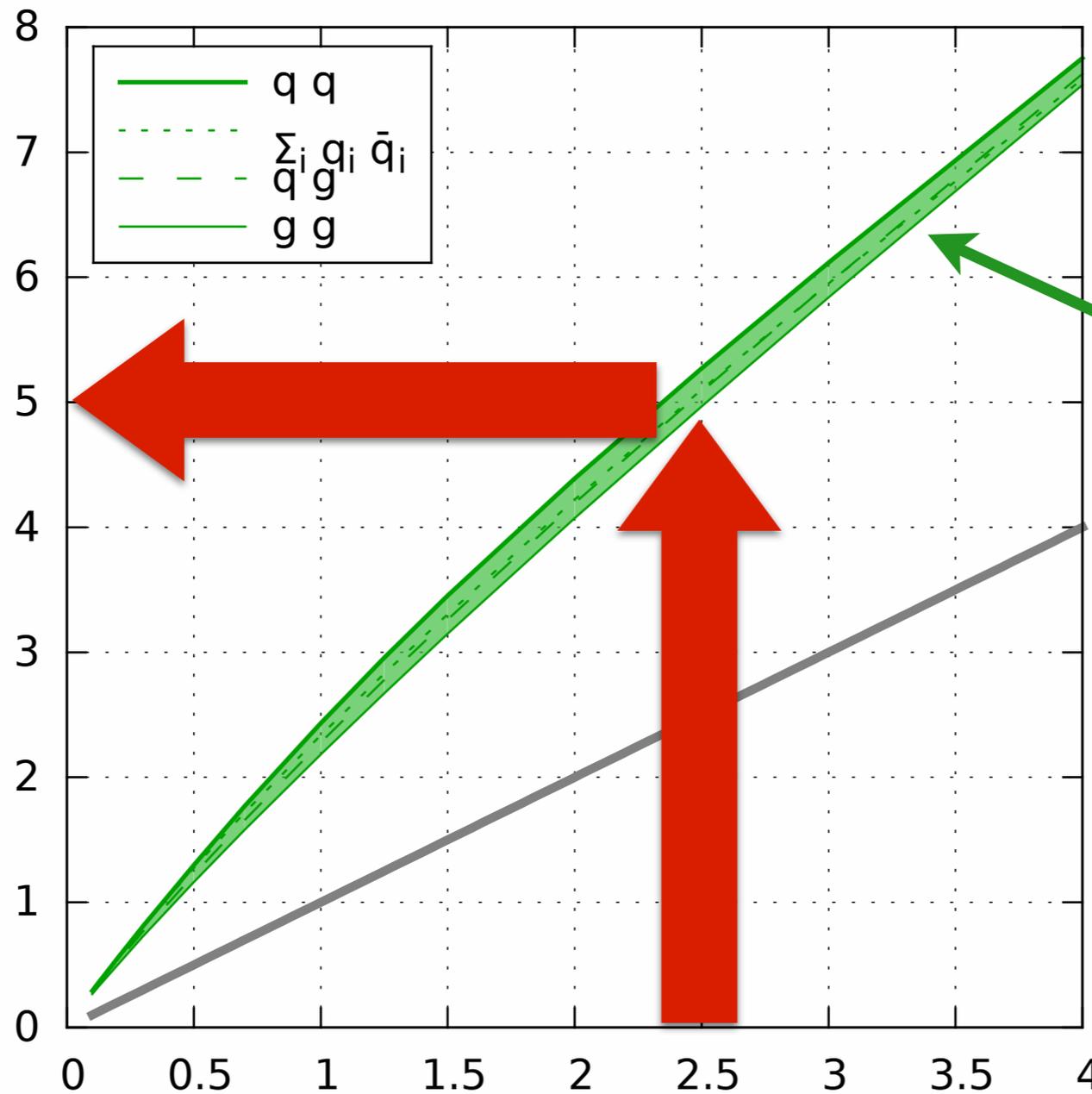
TeV, integrated luminosity

300

fb^{-1}

PDF:

MSTW2008nnlo68cl



Mass [TeV] at
collider #2

Spread of
partonic
channels
(assume same
channel for
S & B)

Mass [TeV] at collider #1

The Collider Reach tool gives you a quick (and dirty) estimate of the relation between the mass reaches of different proton-proton collider setups.

Collider 1: CoM energy TeV, integrated luminosity fb⁻¹

Collider 2: CoM energy TeV, integrated luminosity fb⁻¹

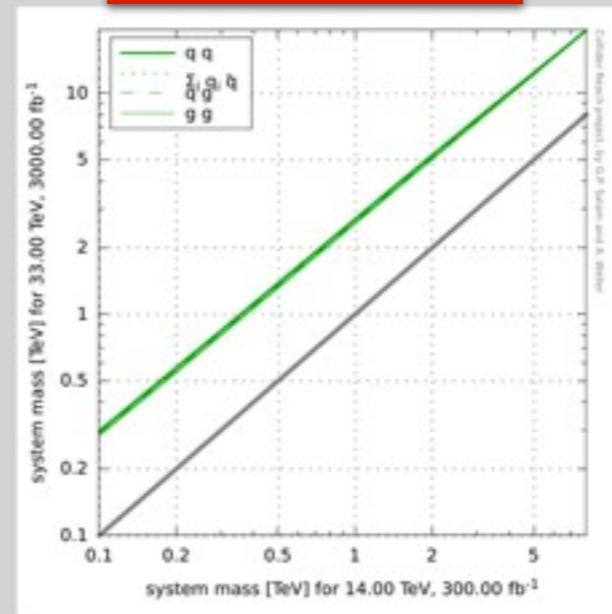
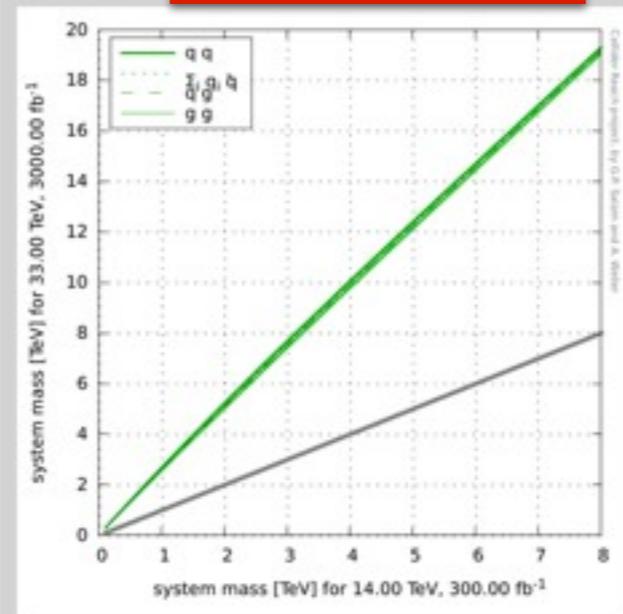
PDF:

Submit

linear plot

log-log plot

Plots

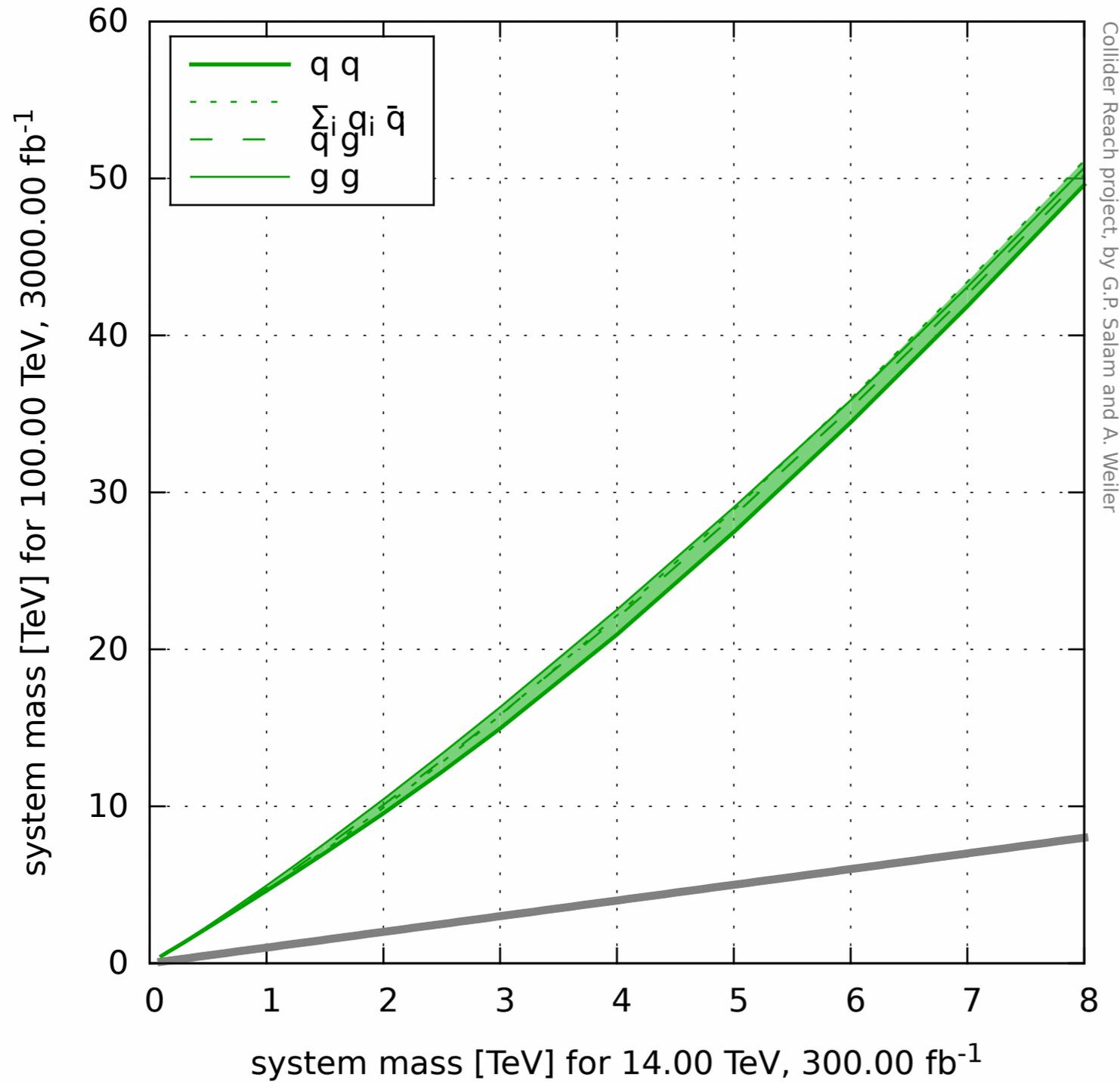


Download: [collider.pdf](#), [colliderloglog.pdf](#), plot generation [log file](#)

The PDF choice was CT10nlo.LHgrid

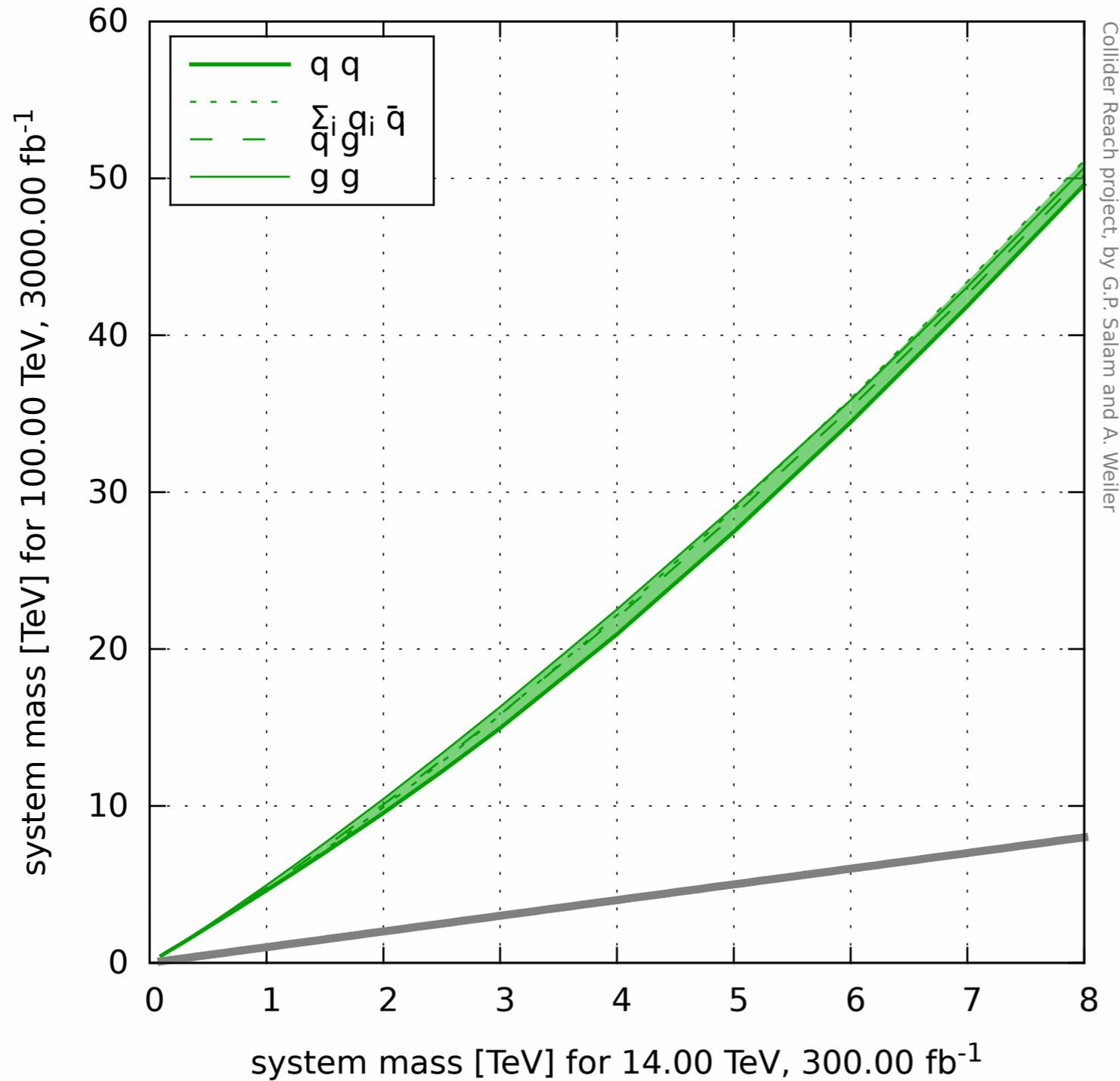
| Original mass | gg | qg | allqq | qqbar |
|---------------|--------|--------|--------|--------|
| 100. | 283. | 291. | 298. | 297. |
| 125. | 350. | 359. | 368. | 367. |
| 150. | 416. | 427. | 438. | 437. |
| 200. | 547. | 562. | 576. | 575. |
| 300. | 806. | 827. | 848. | 847. |
| 500. | 1317. | 1350. | 1386. | 1382. |
| 700. | 1822. | 1866. | 1916. | 1907. |
| 1000. | 2570. | 2628. | 2702. | 2680. |
| 1250. | 3188. | 3256. | 3349. | 3314. |
| 1500. | 3802. | 3879. | 3990. | 3939. |
| 2000. | 5018. | 5110. | 5251. | 5169. |
| 2500. | 6223. | 6327. | 6488. | 6380. |
| 3000. | 7417. | 7530. | 7703. | 7578. |
| 4000. | 9782. | 9904. | 10082. | 9945. |
| 5000. | 12120. | 12246. | 12417. | 12284. |
| 6000. | 14439. | 14565. | 14726. | 14601. |
| 7000. | 16748. | 16871. | 17021. | 16905. |
| 8000. | 19053. | 19169. | 19310. | 19206. |

14 TeV_{300 fb⁻¹} → 100 TeV_{3 ab⁻¹}



Collider Reach project, by G.P. Salam and A. Weiler

14 TeV_{300 fb⁻¹} → 100 TeV_{3 ab⁻¹}



The PDF choice was CT10nlo.LHgrid

| Original mass | gg | qg | allqq | qqbar |
|---------------|--------|--------|--------|--------|
| 100. | 469. | 465. | 462. | 457. |
| 125. | 585. | 579. | 575. | 568. |
| 150. | 702. | 693. | 687. | 679. |
| 200. | 937. | 923. | 912. | 902. |
| 300. | 1414. | 1386. | 1365. | 1350. |
| 500. | 2394. | 2332. | 2279. | 2261. |
| 700. | 3401. | 3300. | 3206. | 3194. |
| 1000. | 4956. | 4793. | 4619. | 4640. |
| 1250. | 6287. | 6072. | 5818. | 5892. |
| 1500. | 7647. | 7382. | 7038. | 7187. |
| 2000. | 10444. | 10090. | 9552. | 9905. |
| 2500. | 13337. | 12908. | 12185. | 12781. |
| 3000. | 16319. | 15833. | 14954. | 15795. |
| 4000. | 22531. | 21986. | 20933. | 22162. |
| 5000. | 29050. | 28508. | 27467. | 28894. |
| 6000. | 35863. | 35366. | 34451. | 35960. |
| 7000. | 43079. | 42620. | 41854. | 43411. |
| 8000. | 50671. | 50230. | 49590. | 51132. |

When you've lost your XPhone

Rule of Thumb #1

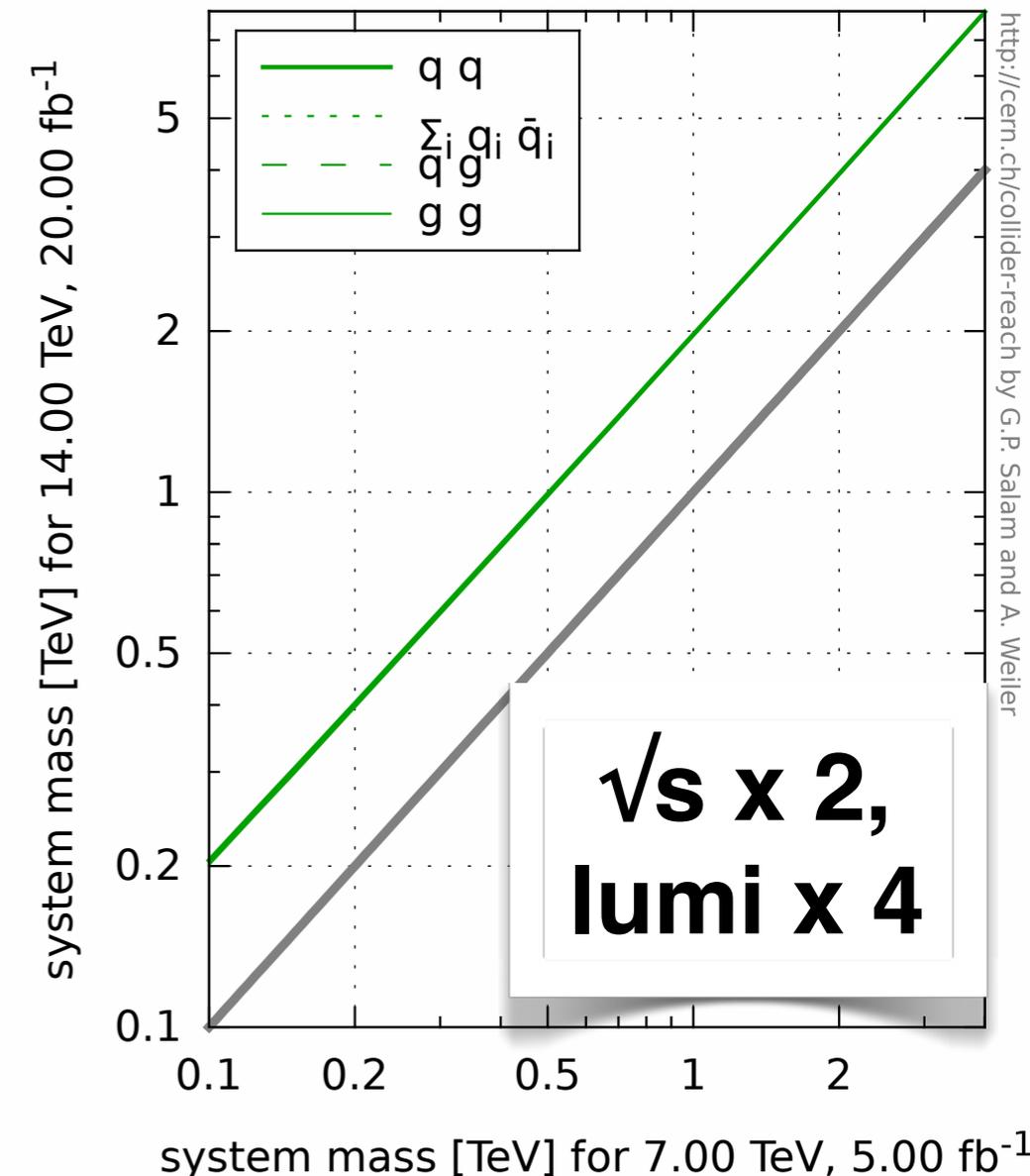
(well known among practitioners)

Increase collider energy by factor X
& increase luminosity by a factor X^2

→ **reach goes up by a factor X**

[Because you keep same Bjorken- x &
luminosity increase compensates for
 $1/\text{mass}^2$ scaling of cross sections]

PDF scaling variations are small effect



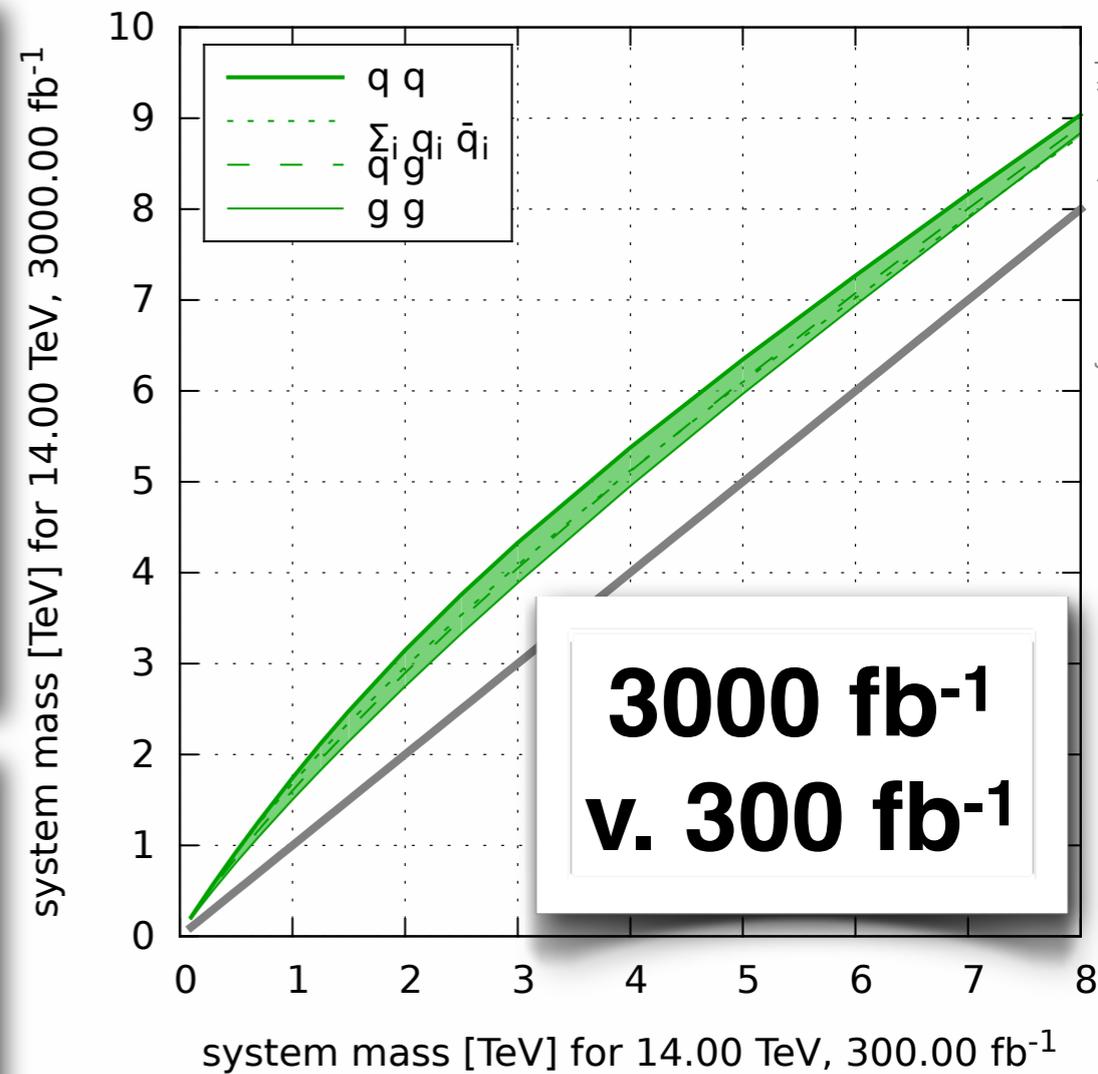
Rule of Thumb #2

(apparently not widely known previously)

Increase luminosity by factor 10
→ **reach increases by constant**
 $\Delta m \approx 0.07\sqrt{s}$

i.e. for $\sqrt{s}=14$ TeV, reach goes by up
1 TeV

No deep reason — a somewhat
random characteristic of large-x PDFs.
Only holds for $0.15 \lesssim M/\sqrt{s} \lesssim 0.6$



Consequence of rule #2

(may be a bit fragile & only for $S \approx B$)

Exclusion is $2\text{-}\sigma$

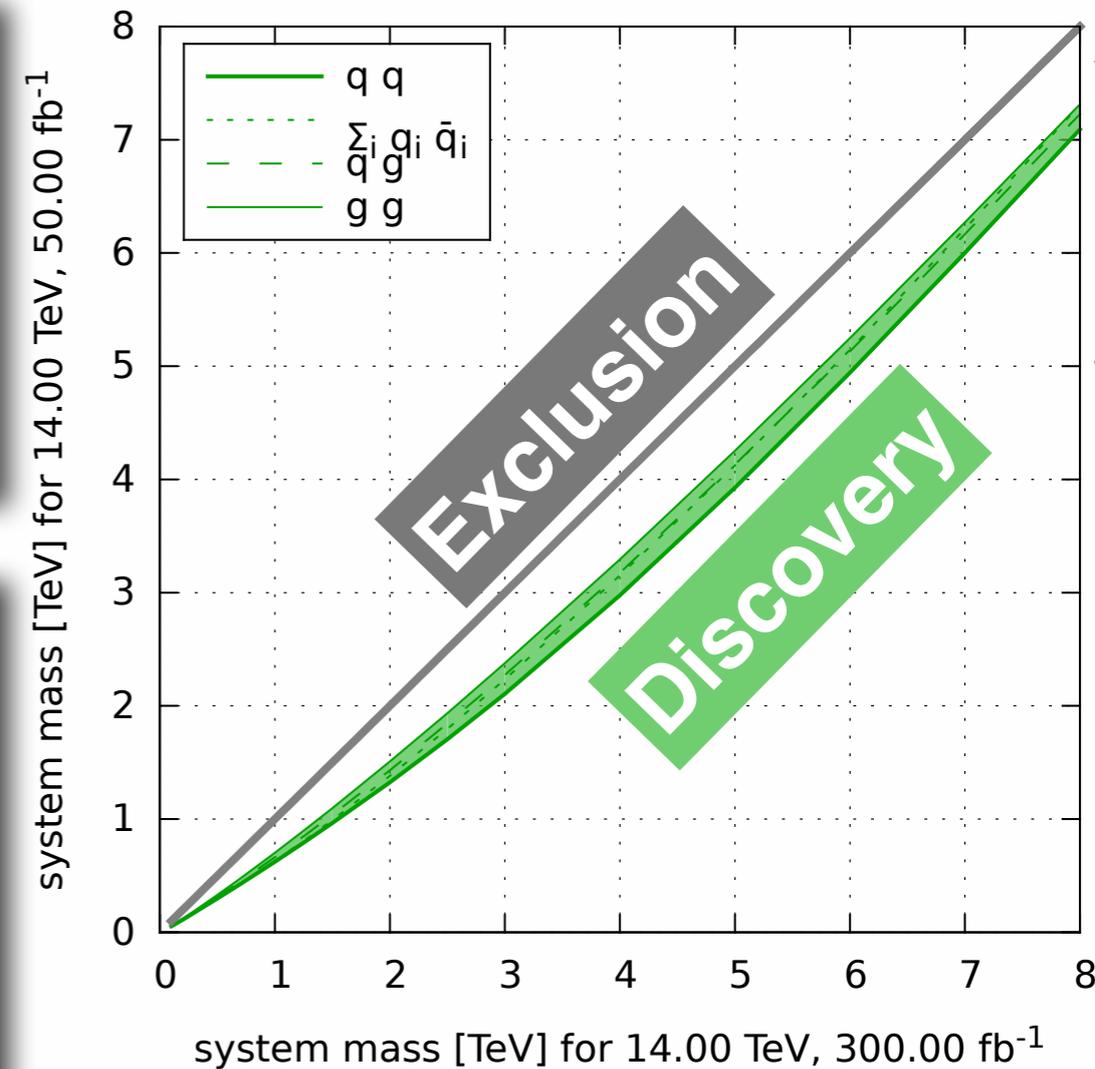
Discovery is $5\text{-}\sigma$

Need $(5/2)^2 = 6.25$ increase in lumi to go from one to the other.

Using rule #2:

discovery reach is about $0.05\sqrt{s}$
below exclusion reach

~ 0.8 TeV at 14 TeV



Conclusions

Amazing recent progress on MC merging/matching, NLO automation, high precision (N)NNLO calculations — hard to imagine how much further we will get by 100 TeV era

FHC as scaled-up LHC is probably not too bad an approx if cuts & analyses are adapted appropriately

→ part of assumption of <http://cern.ch/collider-reach>

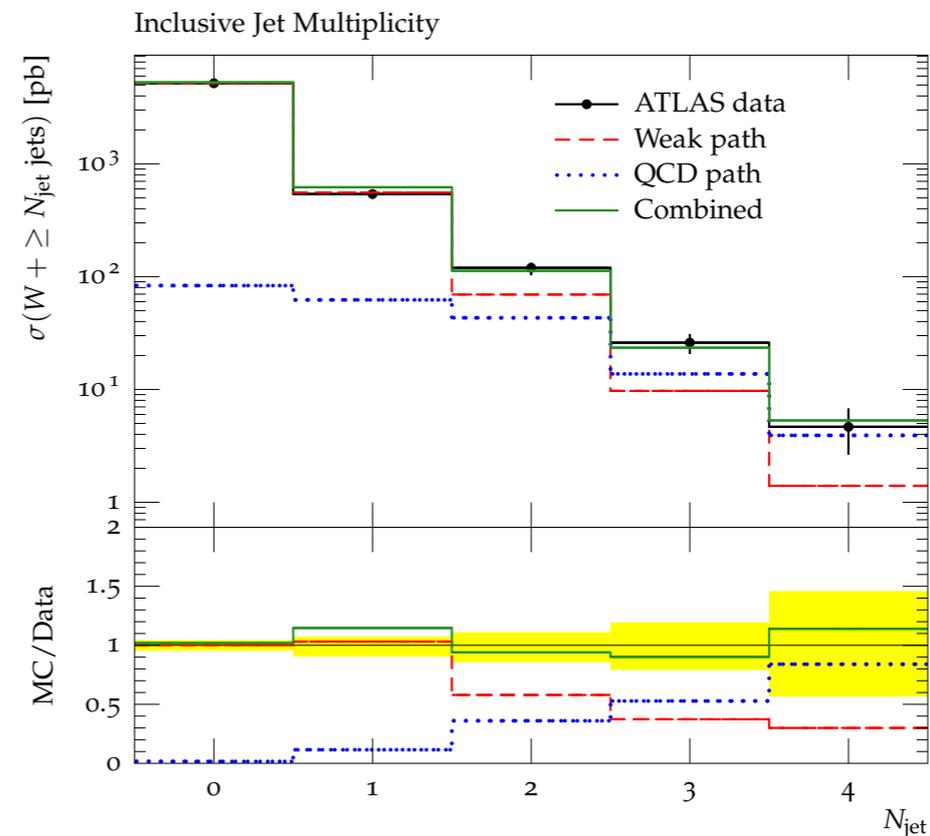
We've maybe only touched the surface on potential from $\sqrt{s} \gg m_{EW}$ — e.g. incoming parton polarization

Hard (= interesting!) problems remain in collider QCD...

BACKUP SLIDES

W + jets

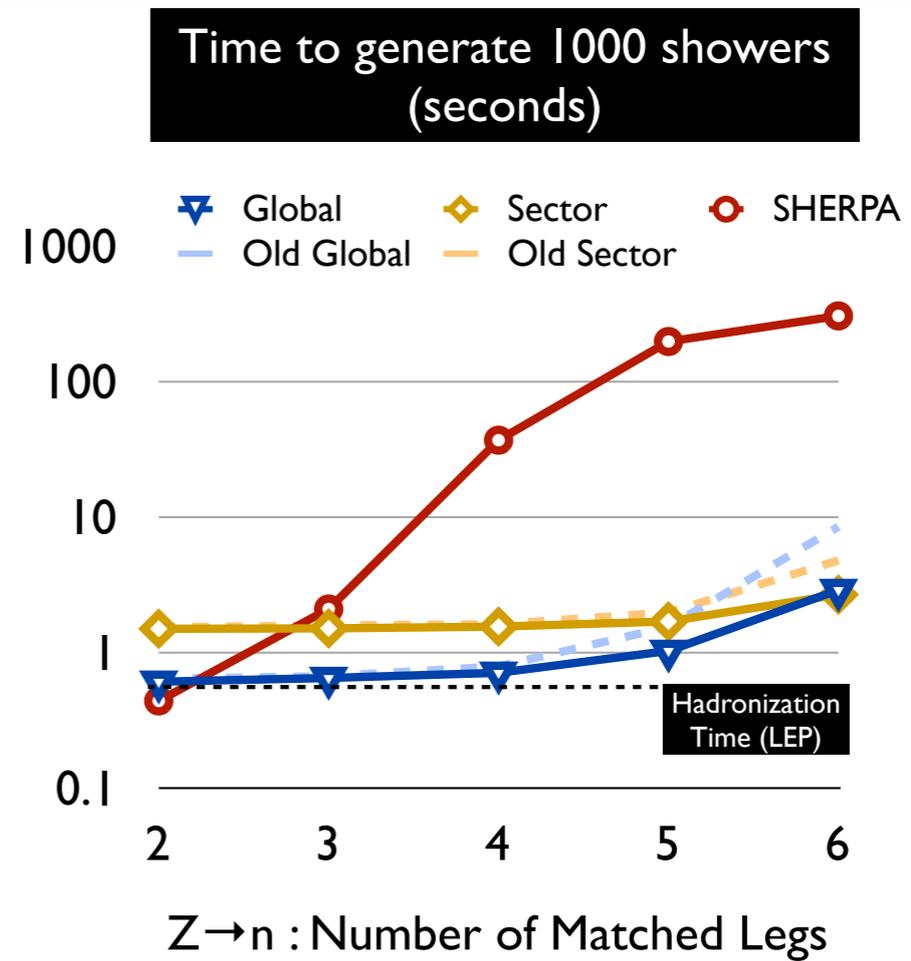
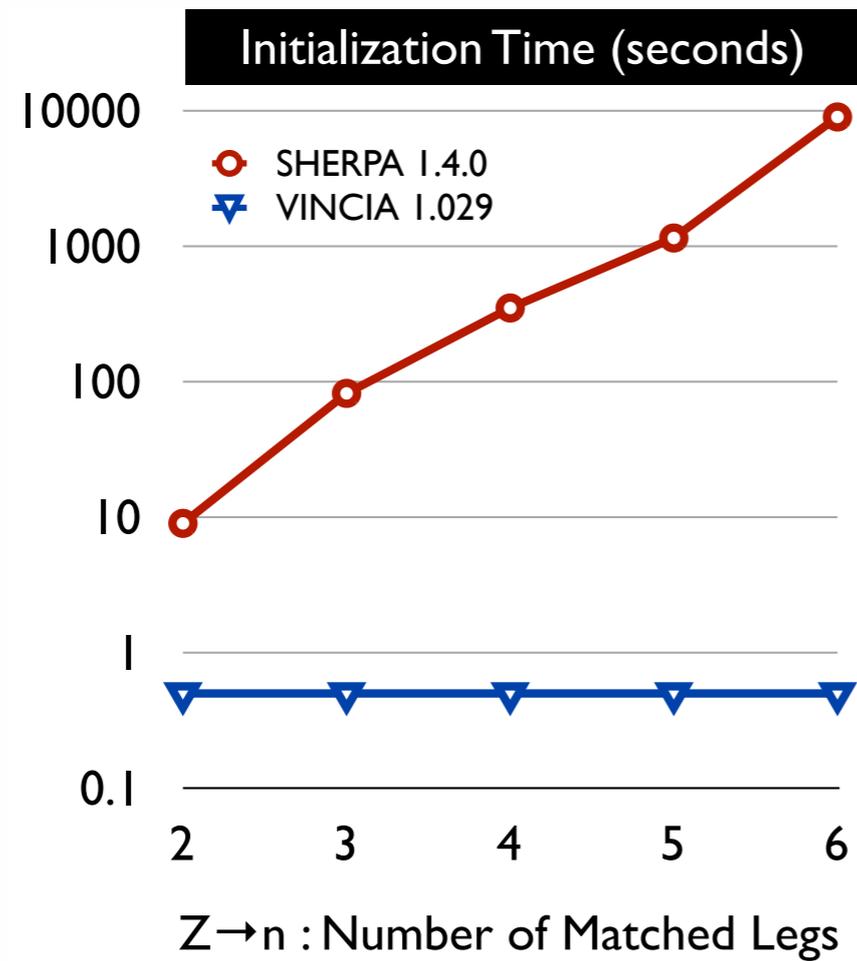
- W + jets is notoriously known for PS not describing data well
- Combine Drell-Yan W production with QCD radiation and $2 \rightarrow 2$ hard QCD processes with weak shower
- Double counting avoided by applying cuts in the spirit of the k_{\perp} jet algorithm
- k-factor applied (normalized to fit first bin)



Some Higgs reference numbers

| \sqrt{s} [TeV] | σ [pb] |
|------------------|---------------|
| 8 | 18.4 |
| 14 | 47.6 |
| 100 | 718 |

large m_{top} , NNLO, MSTW2008 ($\alpha_s = 0.117$)



1301.0933

Can this gain be replicated for pp collisions?

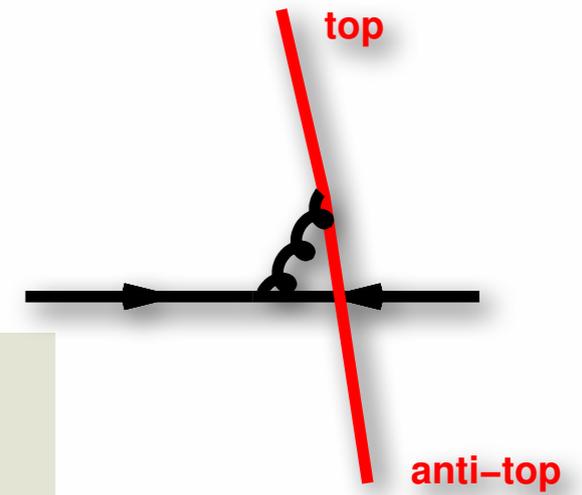
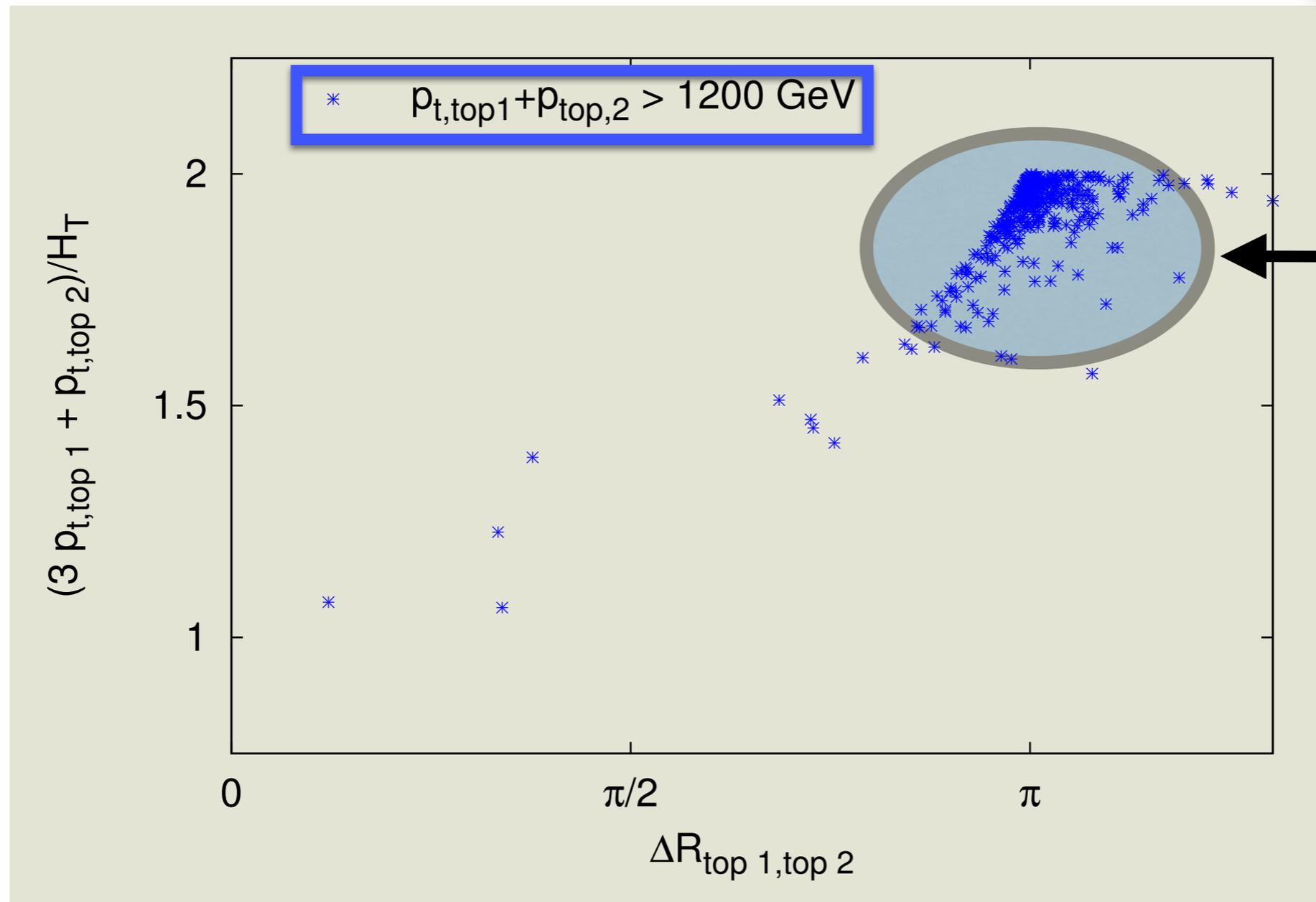
Are top pairs in high- p_t events always back-to-back?

A reminder that top-quarks at LHC are almost “light”

An 8 TeV study with POWHEG, top-pair production, no decay and no parton showering (to keep things simple)

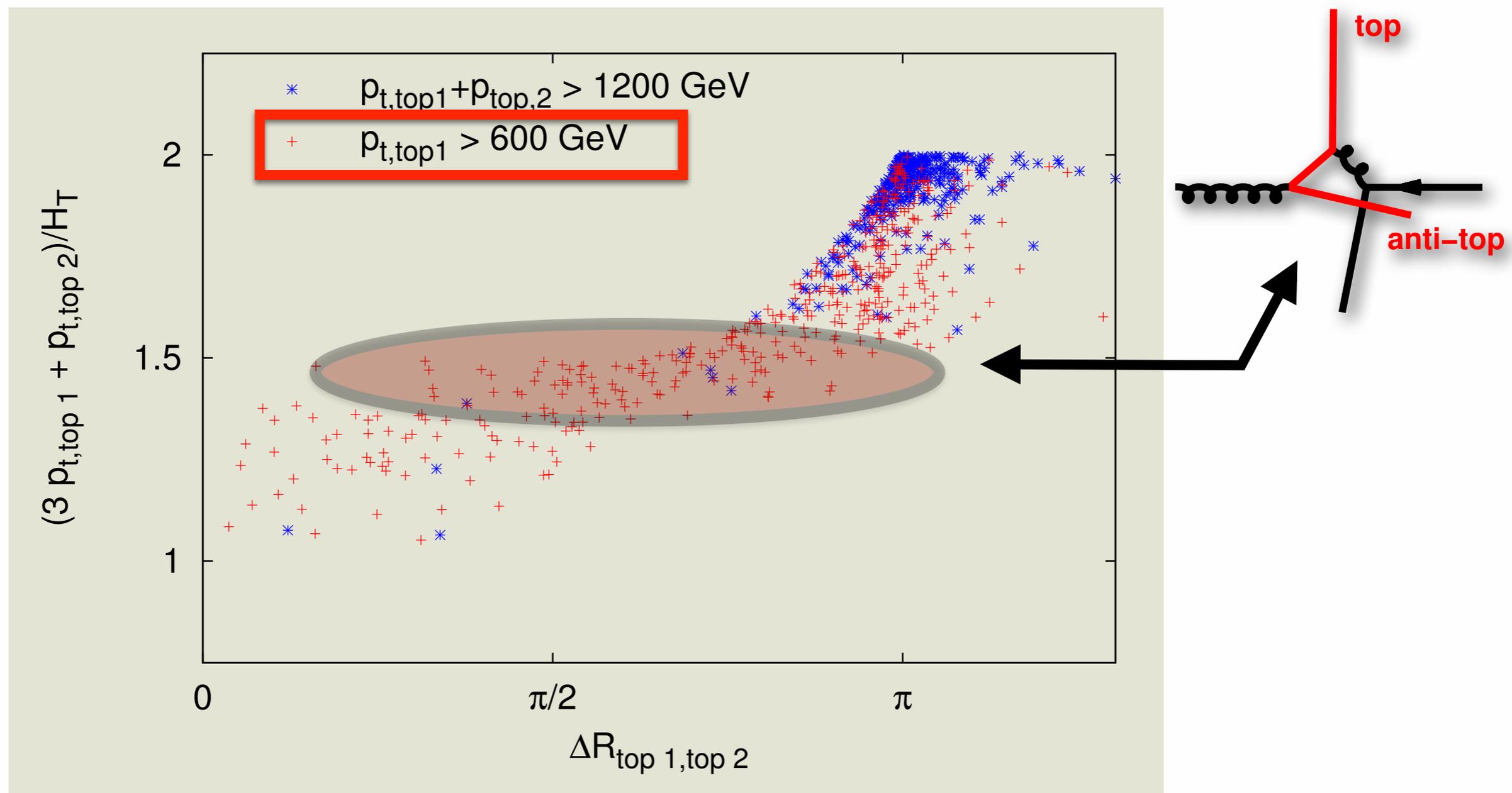
top topology v. cuts

Flavour Creation



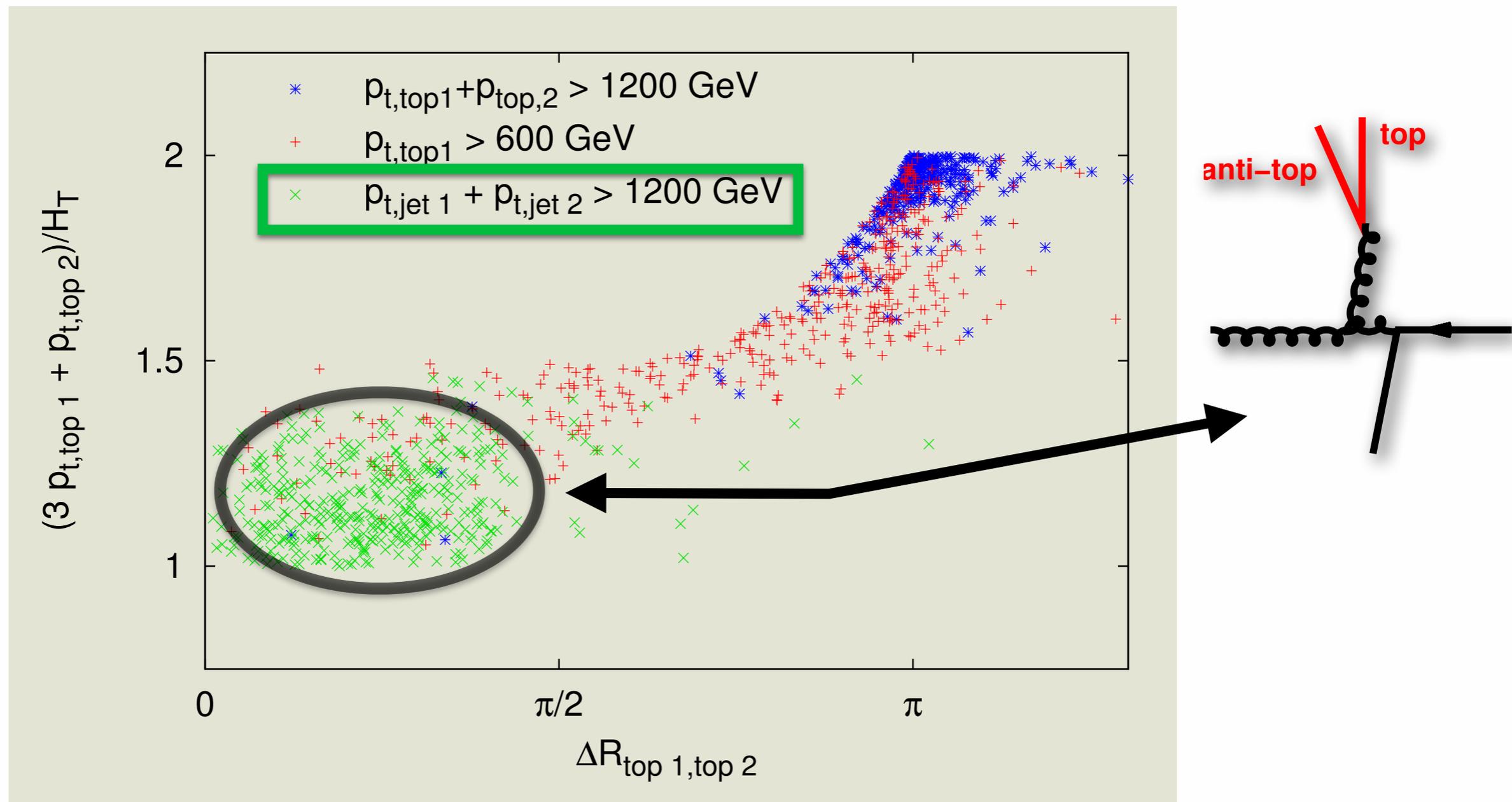
top topology v. cuts

Flavour Excitation – tops inside your PDFs



top topology v. cuts

Gluon Splitting



Assumptions

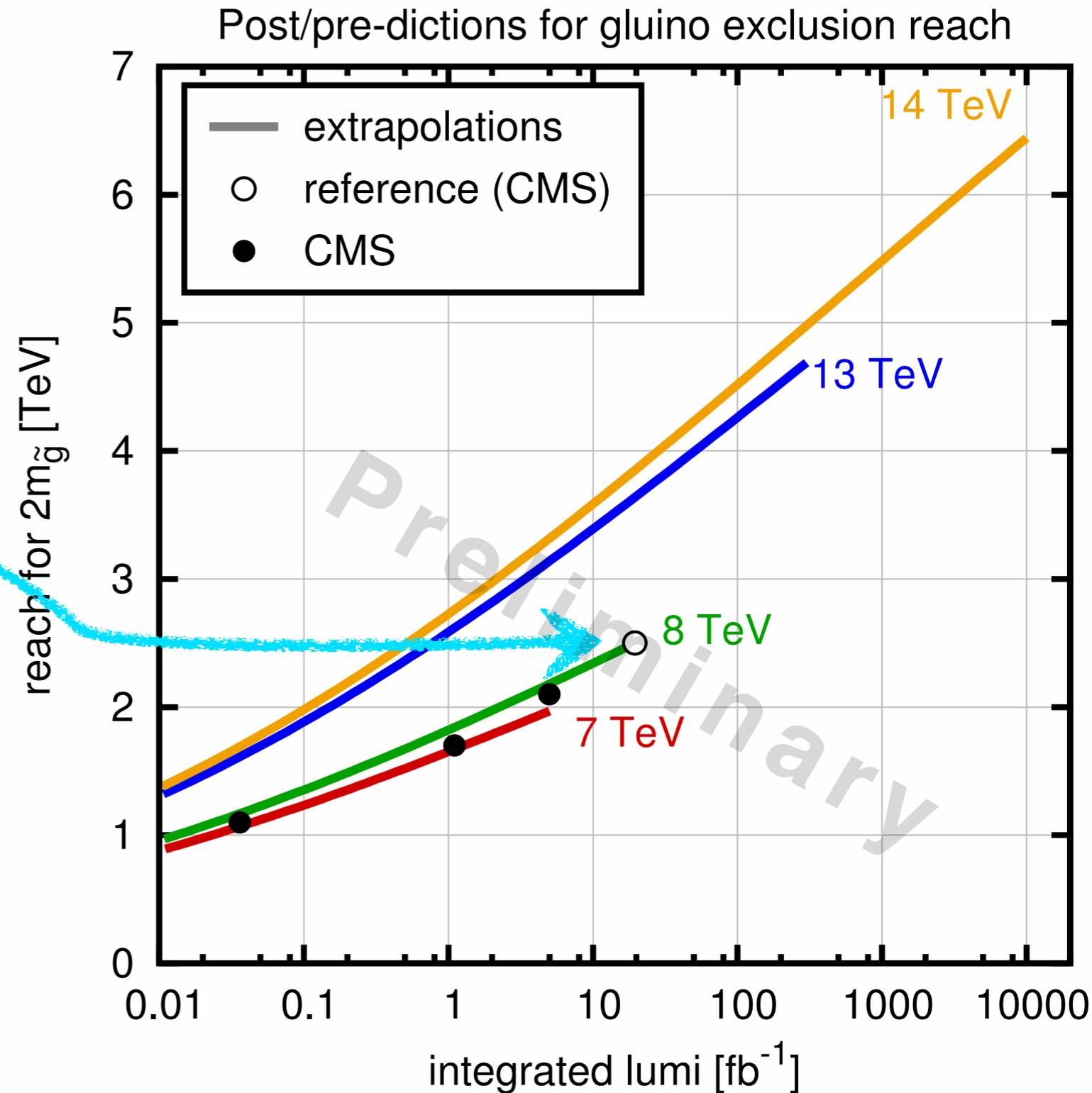
- We don't need to worry about scaling of background vs. signal
- Reconstruction efficiencies, background rejection, etc all stay reasonably constant

Try a SUSY example,
gluinos. Baseline:

CMS, 20 fb⁻¹ @ 8 TeV
excludes $M_{\tilde{g}} < 1250$ GeV
i.e. $2M_{\tilde{g}} < 2.5$ TeV

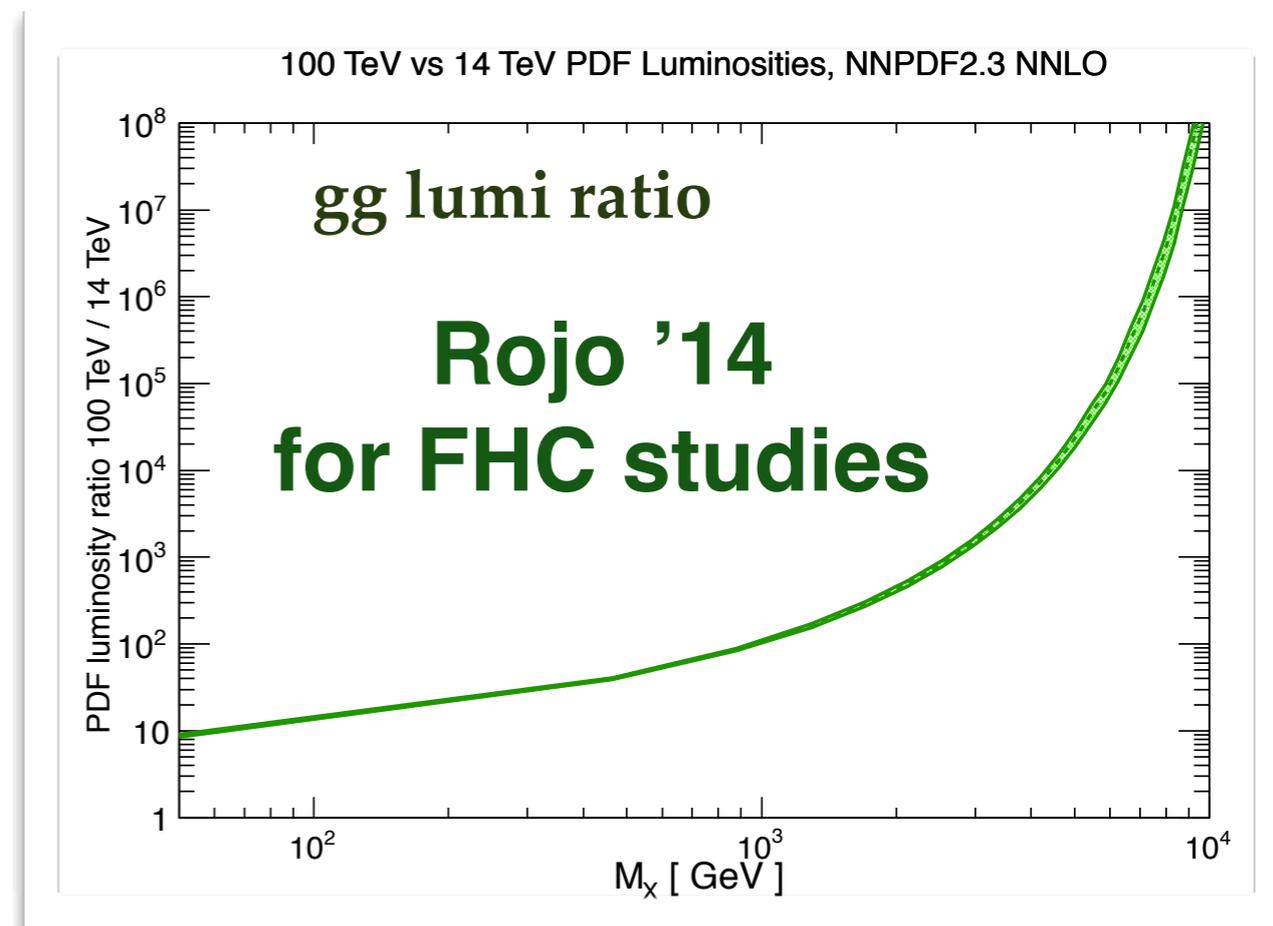
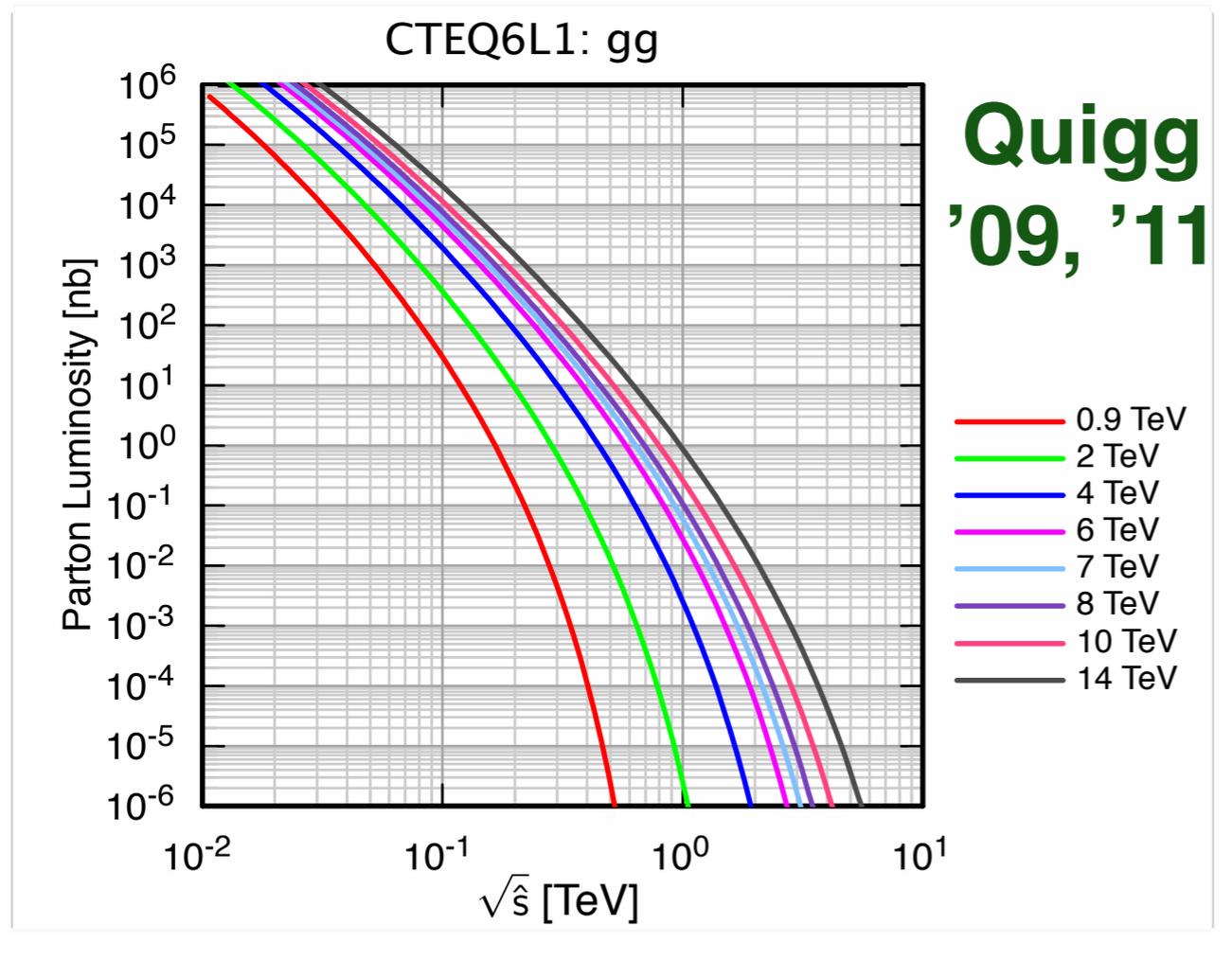
“Predict” exclusions
at other lumis &
energies (assume gg)

Compare to actual
exclusions



**Still works OK, despite (poor) assumption of same
signal and background channels [see also later]**

A side remark: Studying partonic luminosities is a standard technique

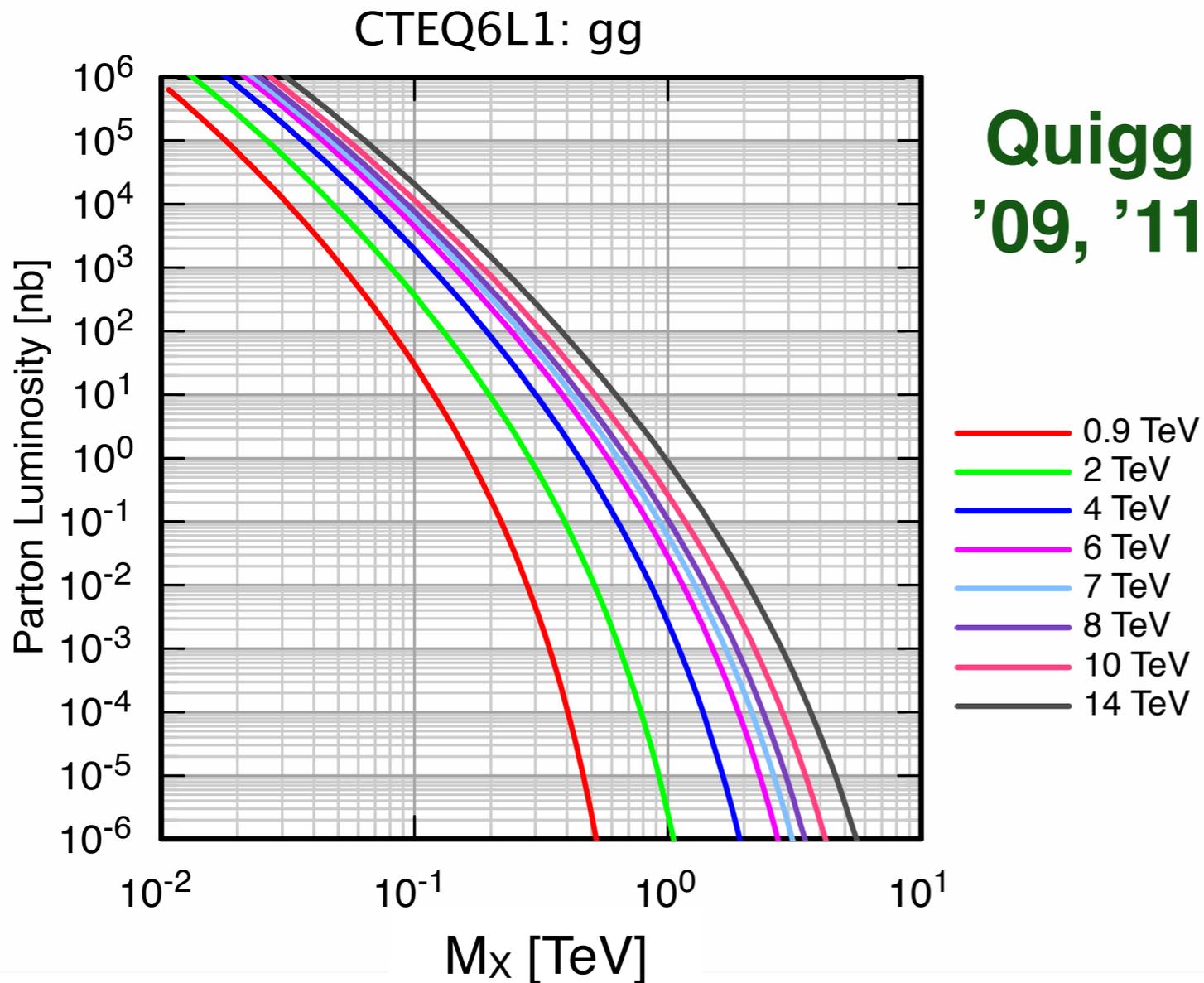


How do we differ?

Study one key question:
relate reaches [TeV] of
different colliders

Validate the approach
by postdicting LHC and
Tevatron results

Why does it work?



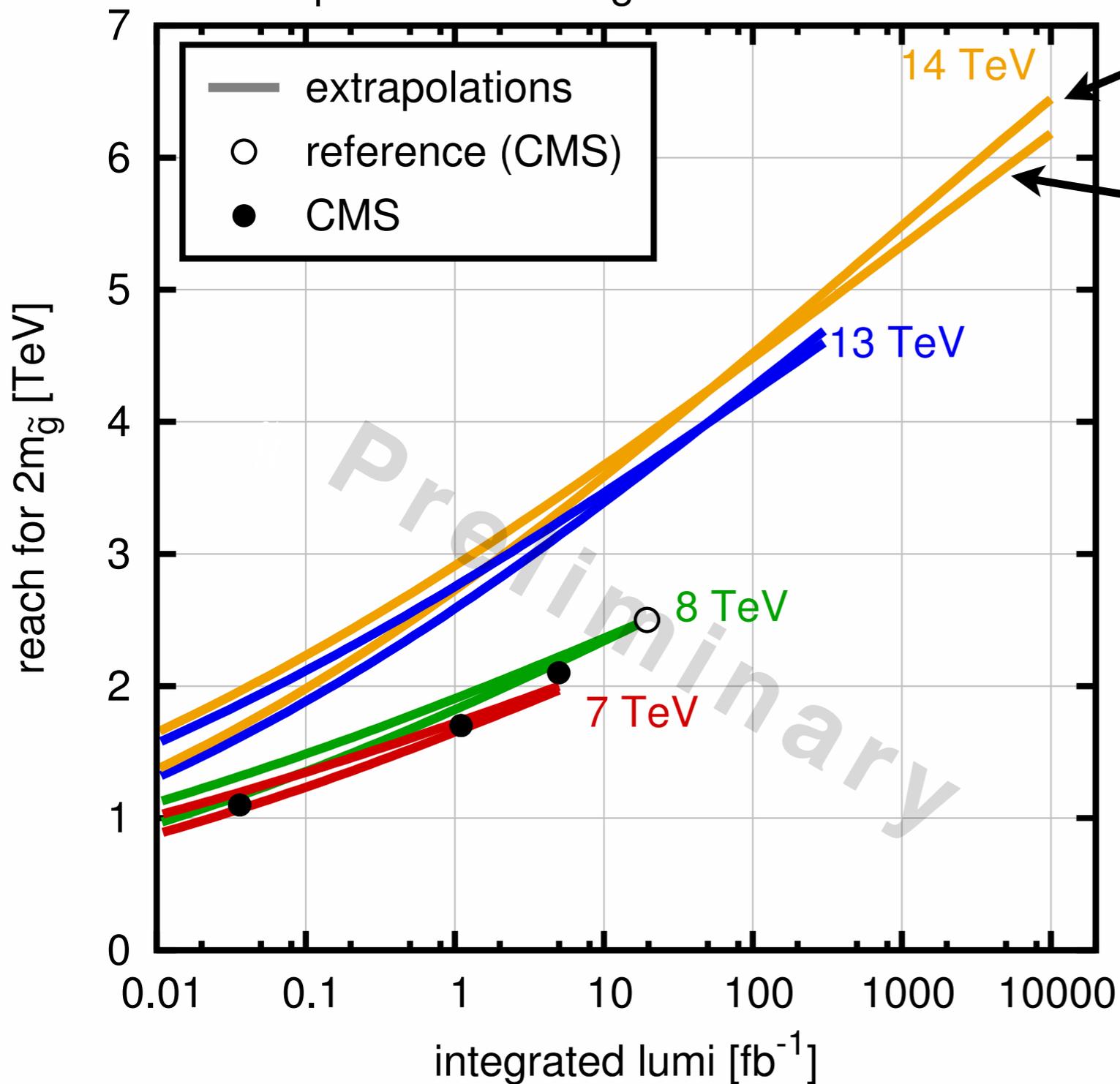
Parton luminosities fall off very fast with increasing M_X

Even when you make a mistake (e.g. wrong partonic channel) the impact on estimated M_X reach is modest

x2 in lumi \sim 10% in M_X

| ATLAS | | | | | | |
|--|----------------------|----------------------|-----------------------------|--|----------------|----------------------|
| Search | Signal | Bgd | $E_{\text{CM}}[\text{TeV}]$ | $\mathcal{L}_{\text{int}}[\text{fb}^{-1}]$ | Expected [GeV] | collider-reach [GeV] |
| Sequential Z' | $\sum \bar{q}_i q_i$ | $\sum \bar{q}_i q_i$ | 7 | 0.2 | 1450 [?] | (base-line) |
| | | | 7 | 1.1 | 1850 [?] | 1849 |
| | | | 7 | 5 | 2200 [?] | 2219 |
| | | | 8 | 6.1 | 2550 [?] | 2510 |
| | | | 8 | 20 | 2900 [?] | 2844 |
| Stop ($m_{\text{LSP}} = 0 \text{ GeV}$) | gg | gg | 7 | 4.7 | 500 [?] | (base-line) |
| | | | 8 | 20.5 | 650 [?] | 675 |
| Excited quark | gq | gg | 7 | $315 \cdot 10^{-6}$ | 1010 [?] | (base-line) |
| | | | 7 | $36 \cdot 10^{-3}$ | 2040 [?] | 2026 (gq) |
| | | | 7 | $163 \cdot 10^{-3}$ | 2490 [?] | 2395 (gq) |
| | | | 7 | 0.81 | 2910 [?] | 2790 (gq) |
| | | | 7 | 4.8 | 3090 [?] | 3220 (gq) |
| | | | 8 | 13 | 3840 [?] | 3865 (gq) |
| CMS | | | | | | |
| Search | Signal | Bgd | $E_{\text{CM}}[\text{TeV}]$ | $\mathcal{L}_{\text{int}}[\text{fb}^{-1}]$ | Expected [GeV] | collider-reach [GeV] |
| gluinos ($m_{\text{LSP}} = 100 \text{ GeV}$) | gg | $gg/gq/qq$ | 7 | 0.036 | 550 [?] | (base-line) |
| | | | 7 | 1.1 | 850 [?] | 855 |
| | | | 7 | 4.98 | 1050 [?] | 1005 |
| | | | 8 | 19.5 | 1250 [?] | 1275 |
| squarks ($m_{\text{LSP}} = 100 \text{ GeV}$) | gg | $gg/gq/qq$ | 7 | 0.036 | 400 [?] | (base-line) |
| | | | 7 | 1.1 | 650 [?] | 663 |
| | | | 7 | 4.98 | 725 [?] | 801 |
| | | | 8 | 19.5 | 910 [?] | 1033 |
| T-quarks ($\text{Br}(T \rightarrow tZ) = 1$) | gg | $gg/gq/qq$ | 7 | 1.14 | 510 [?] | (base-line) |
| | | | 7 | 5 | 550 [?] | 629 |
| | | | 8 | 19.6 | 813 [?] | 827 |

Post/pre-dictions for gluino exclusion reach

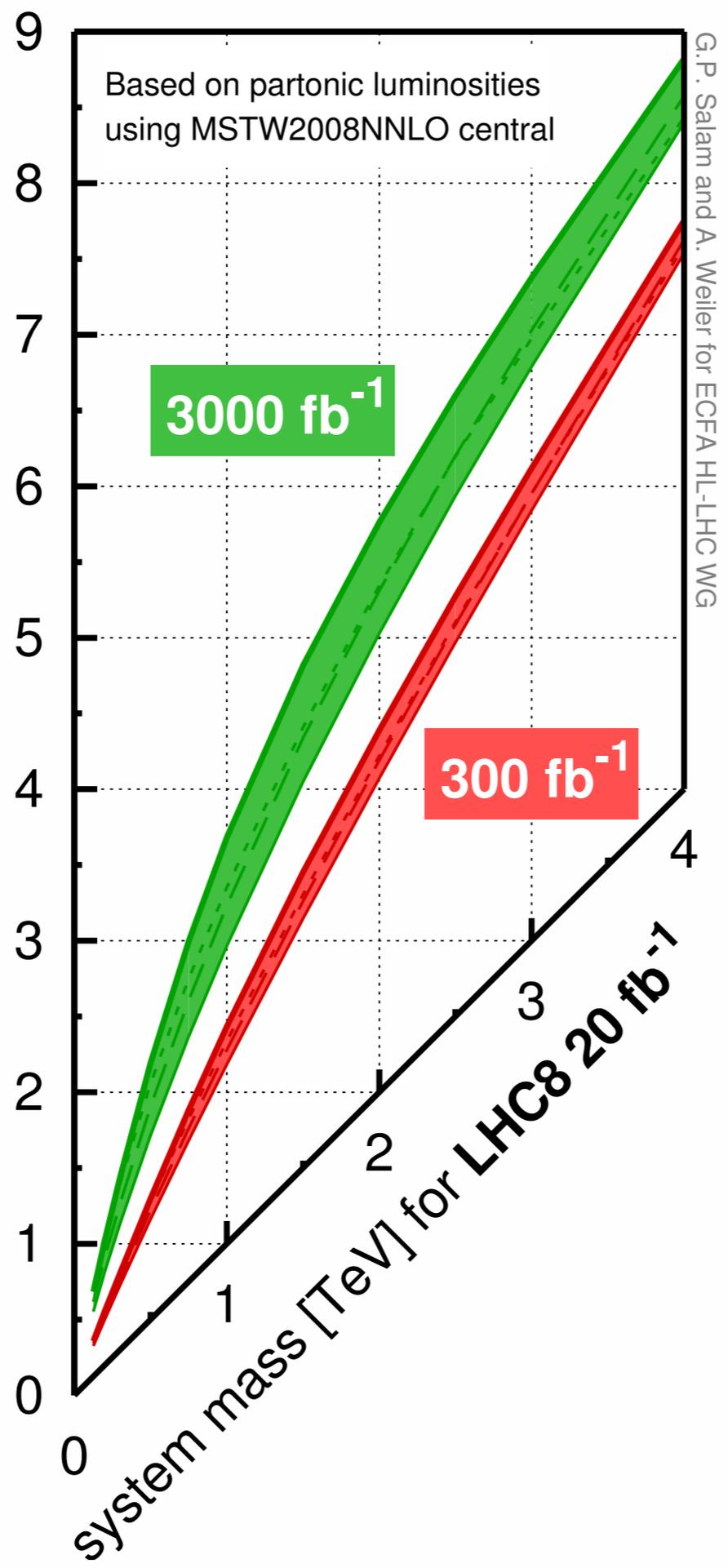


Signal gg; bkgd: gg

Signal gg; bkgd: qg

Scattering channel matters, but not too much

system mass [TeV] for LHC14



- $\Sigma\Sigma$
- - Σg
- ... $\Sigma_i q_i \bar{q}_i$
- gg

LHC comparison

1208.1447
ATLAS-CONF-2013-024

gg

| stop limits | [expected] | (lsp = 0gev) |
|----------------|------------|---------------|
| 7TeV, 4.7 ifb | 500 gev | |
| 8TeV, 20.5 ifb | 650 gev | ----> 675 GeV |

qqbar

ATLAS EXOT-2011-06
ATLAS-CONF-2012-129
ATLAS-CONF-2013-017

| sequential z-prime | [expected] | |
|--------------------|------------|----------------|
| 7TeV, 1.1 ifb | 1800 gev | |
| 8TeV, 6 ifb, | 2550 gev | ----> 2450 GeV |
| 8TeV, 20 ifb | 2800 gee | ----> 2790 GeV |

qg

EXOT-2011-07
ATLAS-CONF-2012-088
ATLAS-CONF-2012-148

| excited quark q* | [expected] | (NB, sig \neq bgd scaling) |
|------------------|------------|-------------------------------|
| 7 TeV, 1 ifb | 2900 gev | |
| 8 TeV, 5.8 ifb | 3500 gev | ----> 3700 GeV |
| 8 TeV, 13 ifb | 3700 gev | ----> 3900 GeV |

LHC comparison

1208.1447
ATLAS-CONF-2013-024

gg
stop limits [expected] (lsp = 0gev) Baseline

| | | | |
|----------------|---------|------|-----------|
| 7TeV, 4.7 ifb | 500 gev | ← | |
| 8TeV, 20.5 ifb | 650 gev | ---- | → 675 GeV |

qqbar
ATLAS EXOT-2011-06
ATLAS-CONF-2012-129
ATLAS-CONF-2013-017

sequential z-prime [expected]

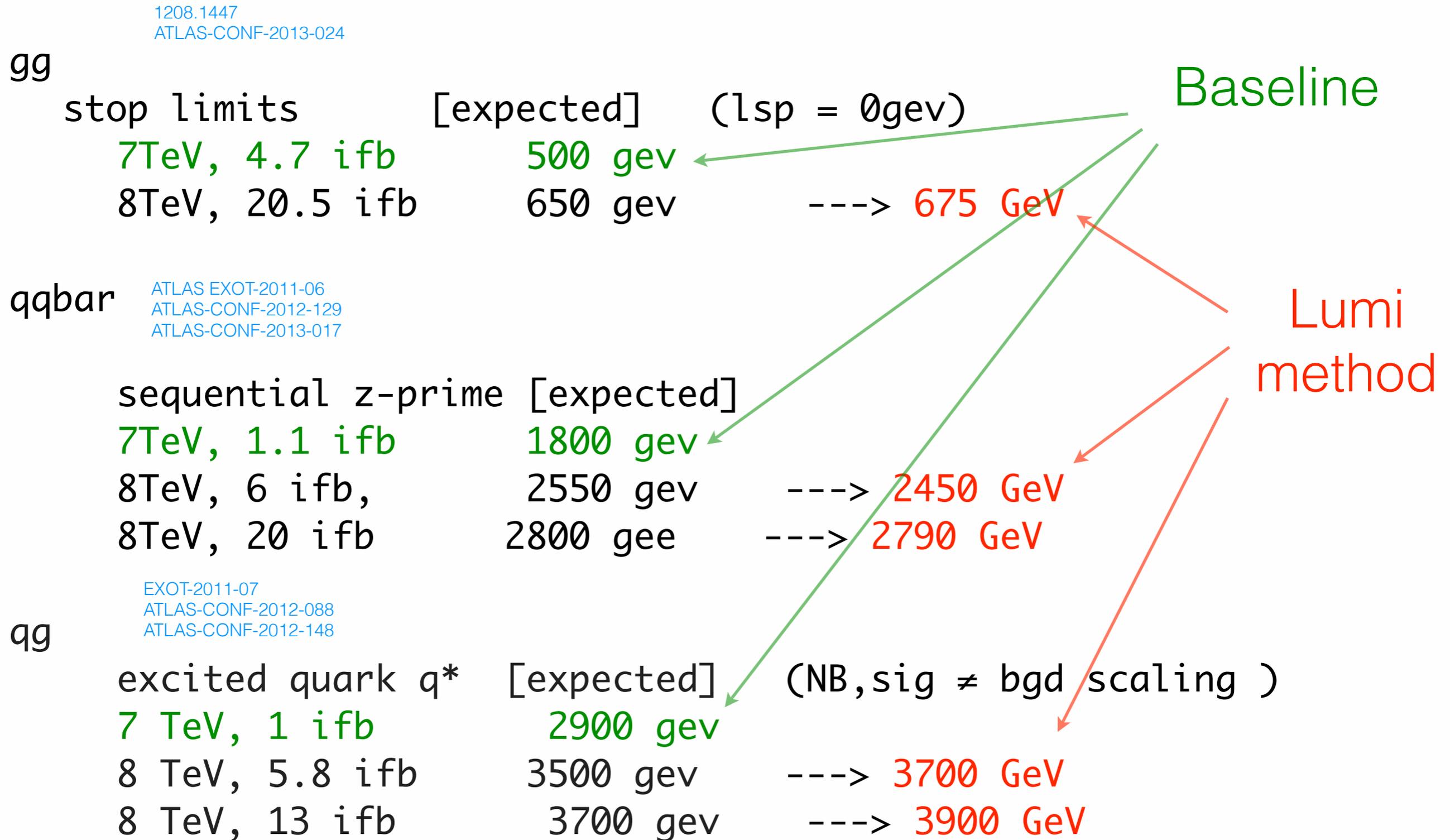
| | | | |
|---------------|----------|------|------------|
| 7TeV, 1.1 ifb | 1800 gev | ← | |
| 8TeV, 6 ifb, | 2550 gev | ---- | → 2450 GeV |
| 8TeV, 20 ifb | 2800 gee | ---- | → 2790 GeV |

qq
EXOT-2011-07
ATLAS-CONF-2012-088
ATLAS-CONF-2012-148

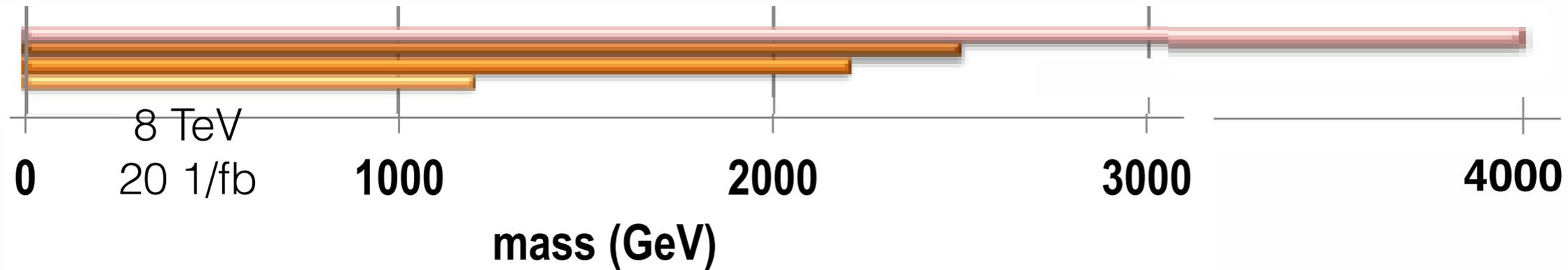
excited quark q* [expected] (NB, sig ≠ bgd scaling)

| | | | |
|----------------|----------|------|------------|
| 7 TeV, 1 ifb | 2900 gev | ← | |
| 8 TeV, 5.8 ifb | 3500 gev | ---- | → 3700 GeV |
| 8 TeV, 13 ifb | 3700 gev | ---- | → 3900 GeV |

LHC comparison



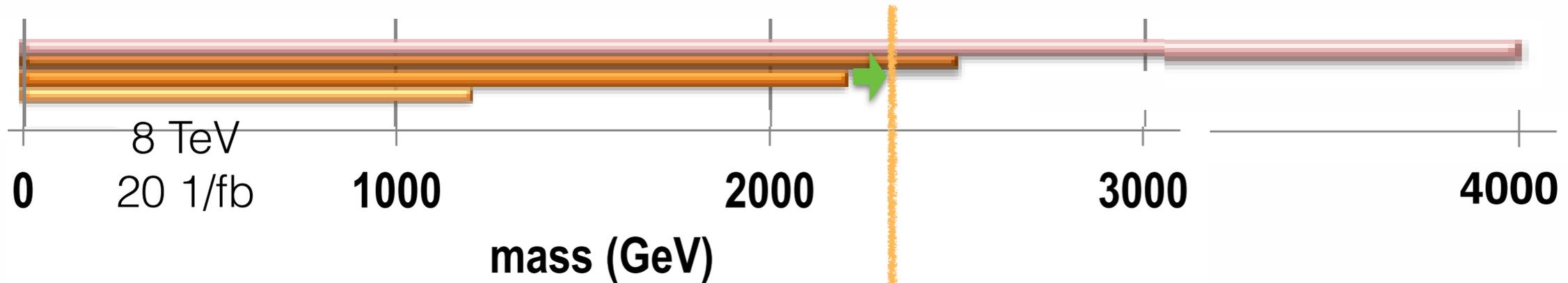
Gluinos



- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

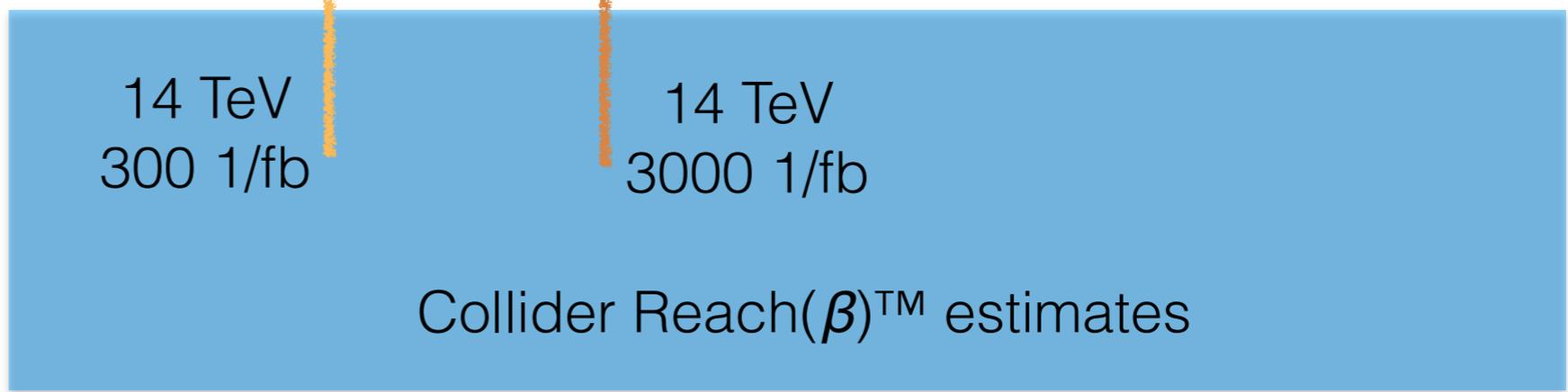
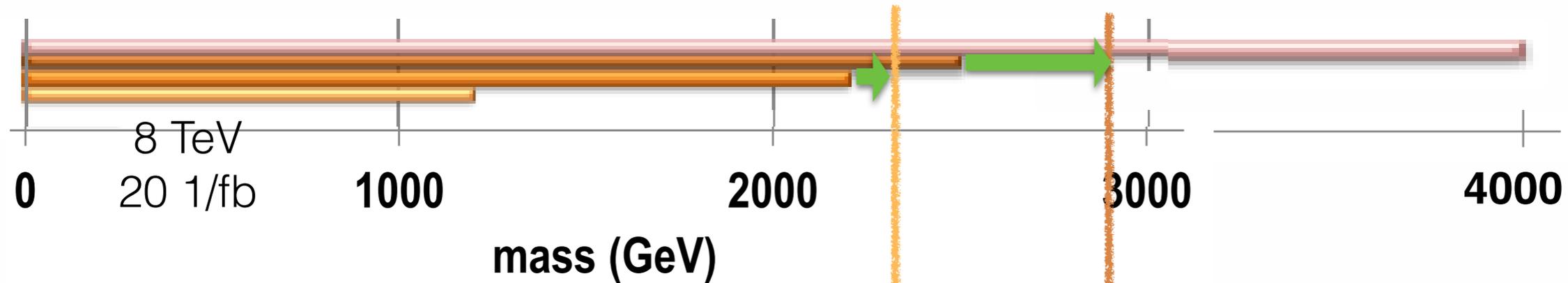
Collider Reach(β)TM estimates

Gluinos



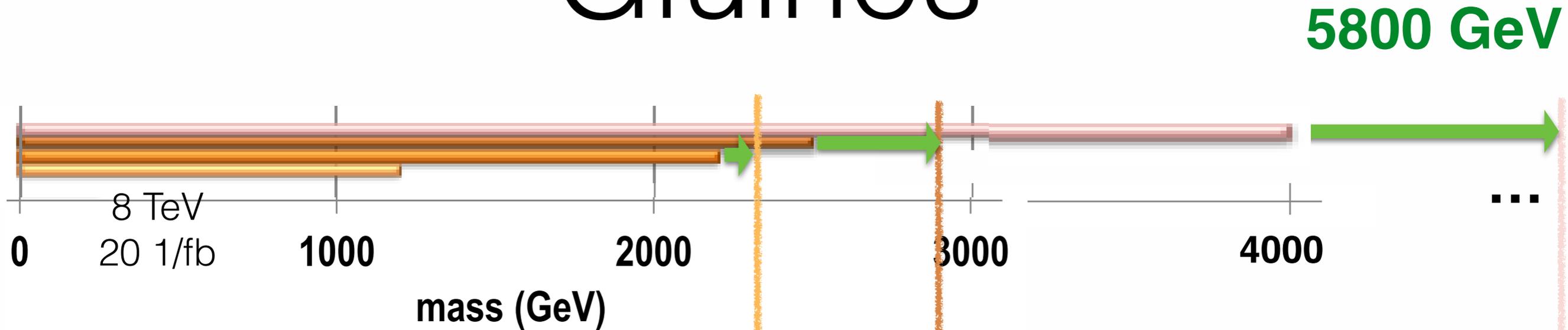
- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

Gluinos



- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

Gluinos



14 TeV
300 1/fb

14 TeV
3000 1/fb

33 TeV
3000 1/fb

■ pp, 100 TeV, 3000/fb

■ pp, 33 TeV, 3000/fb

■ pp, 14 TeV, 3000/fb

■ pp, 14 TeV, 300/fb

■ pp, 8 TeV, 20/fb

■ ee, 3 TeV, 1000/fb

■ ee, 1 TeV, 1000/fb

■ ee, 0.5 TeV, 500/fb

Collider Reach(β)TM estimates

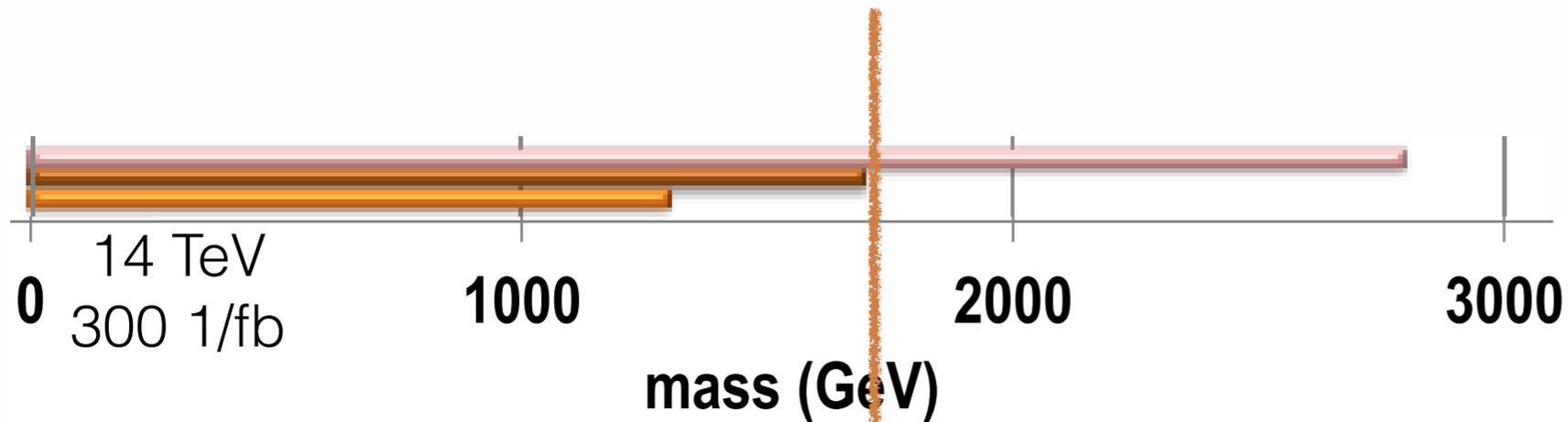
RPV stops



- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

Collider Reach(β)TM estimates

RPV stops

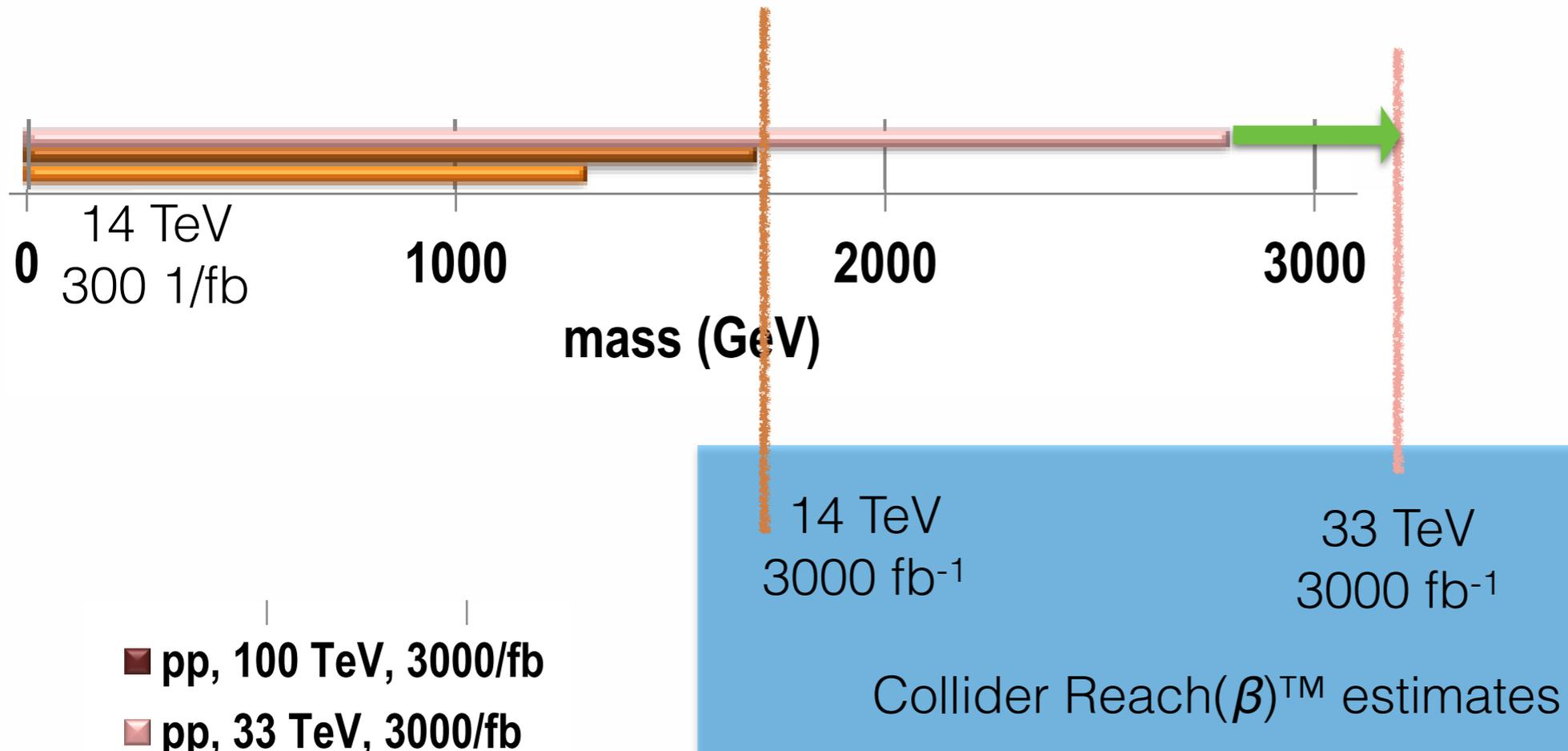


14 TeV
3000 fb⁻¹

Collider Reach(β)TM estimates

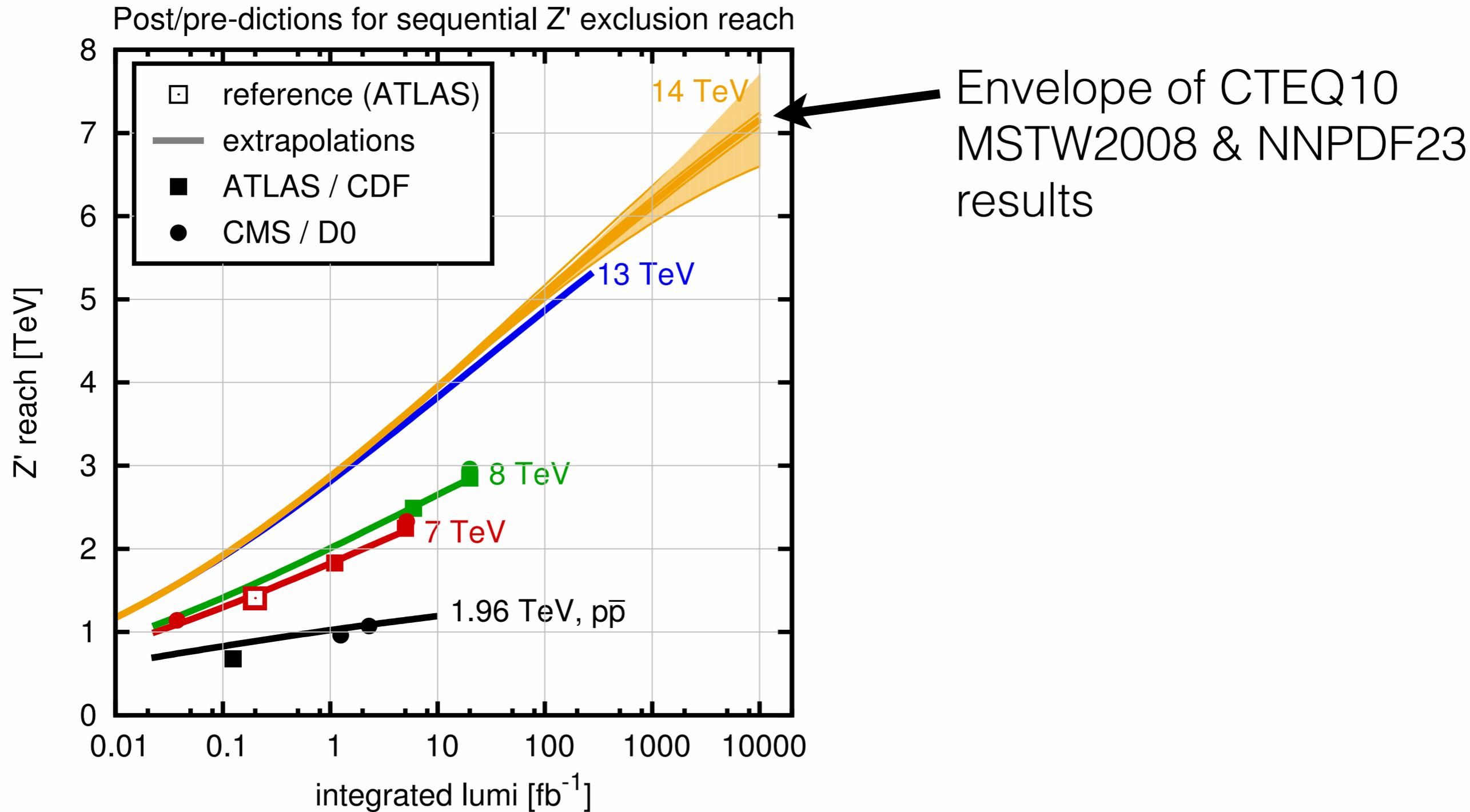
- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

RPV stops

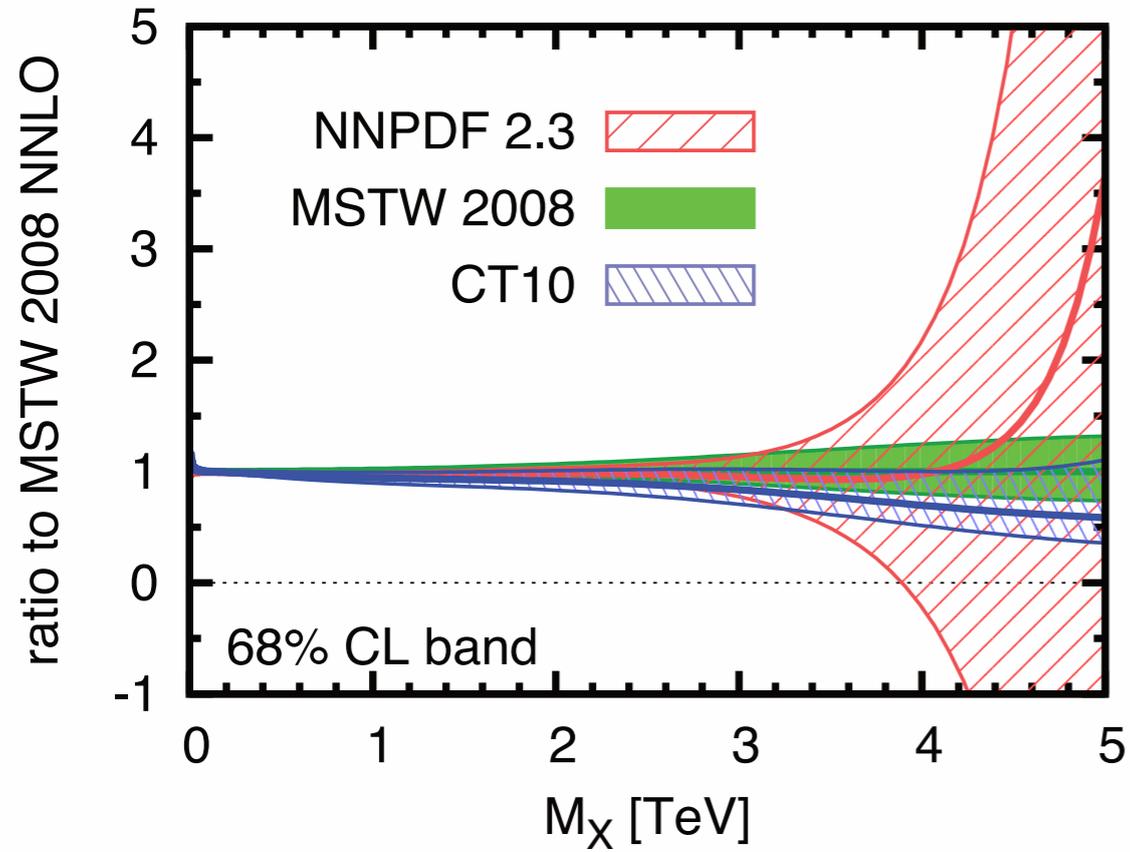


- pp, 100 TeV, 3000/fb
- pp, 33 TeV, 3000/fb
- pp, 14 TeV, 3000/fb
- pp, 14 TeV, 300/fb
- pp, 8 TeV, 20/fb
- ee, 3 TeV, 1000/fb
- ee, 1 TeV, 1000/fb
- ee, 0.5 TeV, 500/fb

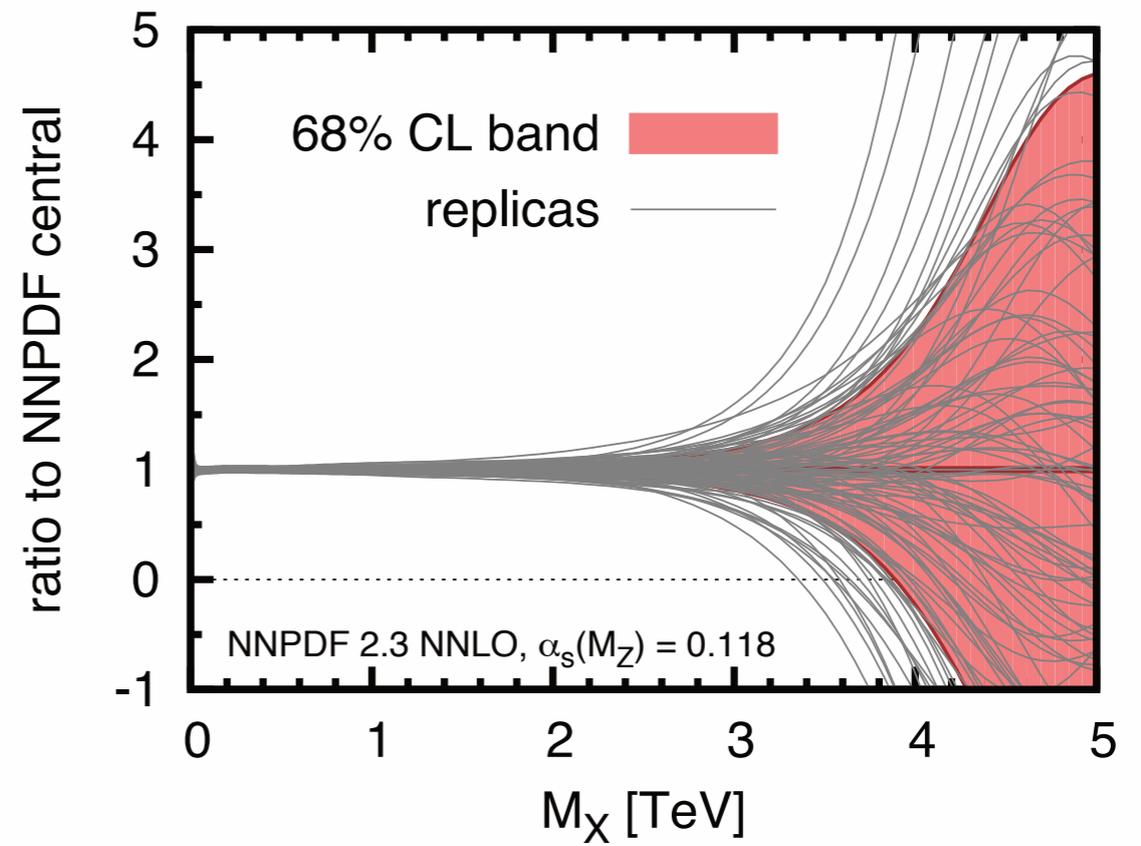
Impact of PDF uncertainties



NNLO $q\bar{q}$ luminosities (LHC 8 TeV)



NNPDF $q\bar{q}$ luminosity replicas (LHC 8 TeV)



$q\bar{q}/qq$ luminosity ratios (LHC 8 TeV)

