

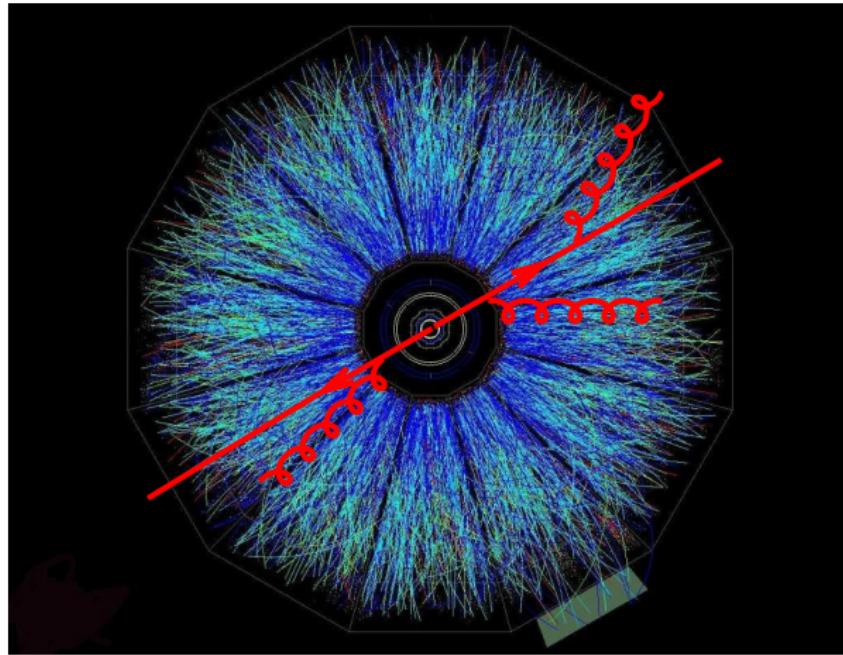
Jet reconstruction with FastJet

Gavin P. Salam

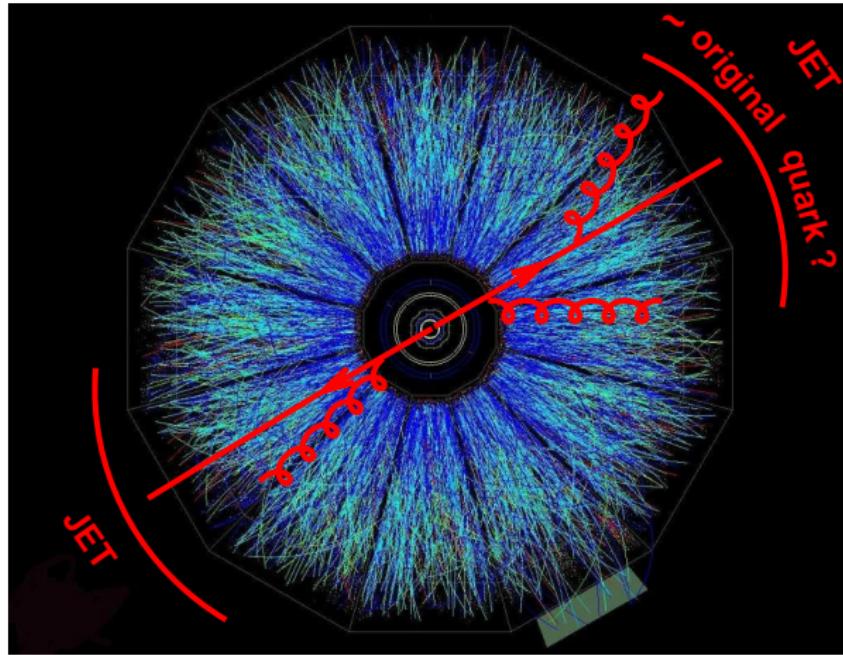
LPTHE, UPMC Paris 6 & CNRS

Jets in Proton-Proton and Heavy-Ion Collisions
Prague, 12 August 2010

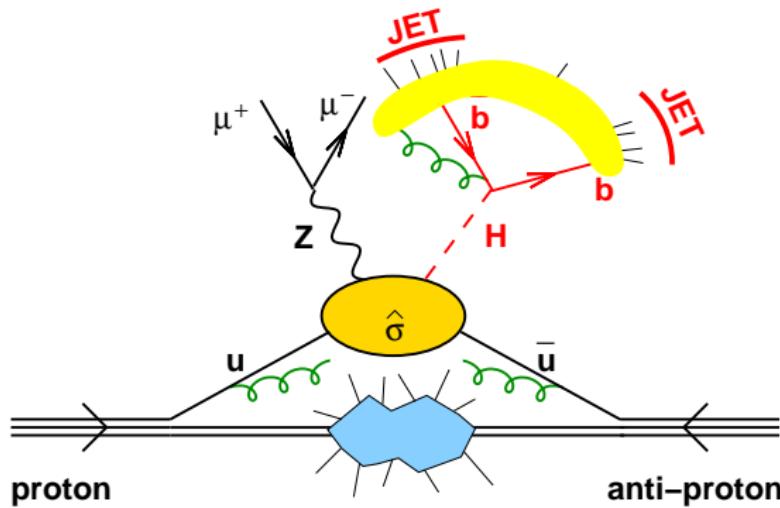
Based on work (some preliminary) with
Matteo Cacciari, Juan Rojo, Sebastian Sapeta, Gregory Soyez



Radiation from high-momentum quarks & gluons traversing hot medium can tell us about the medium



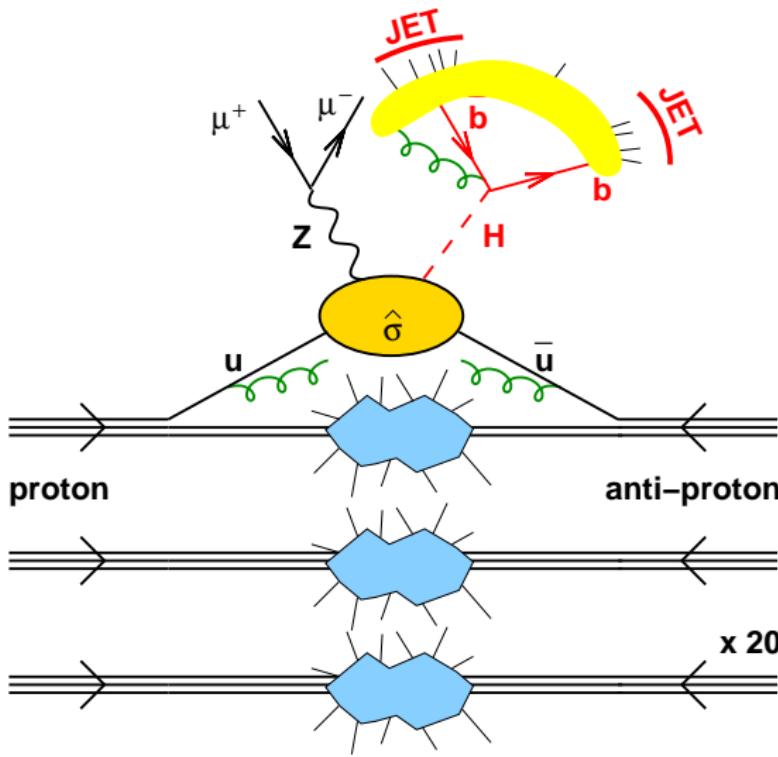
Radiation from high-momentum quarks & gluons traversing hot medium can tell us about the medium



Use jets to reconstruct
quarks from decay of some
new heavy object

e.g. a Higgs boson

At high luminosity, many simultaneous pp collisions –
not unlike AuAu/PbPb collision

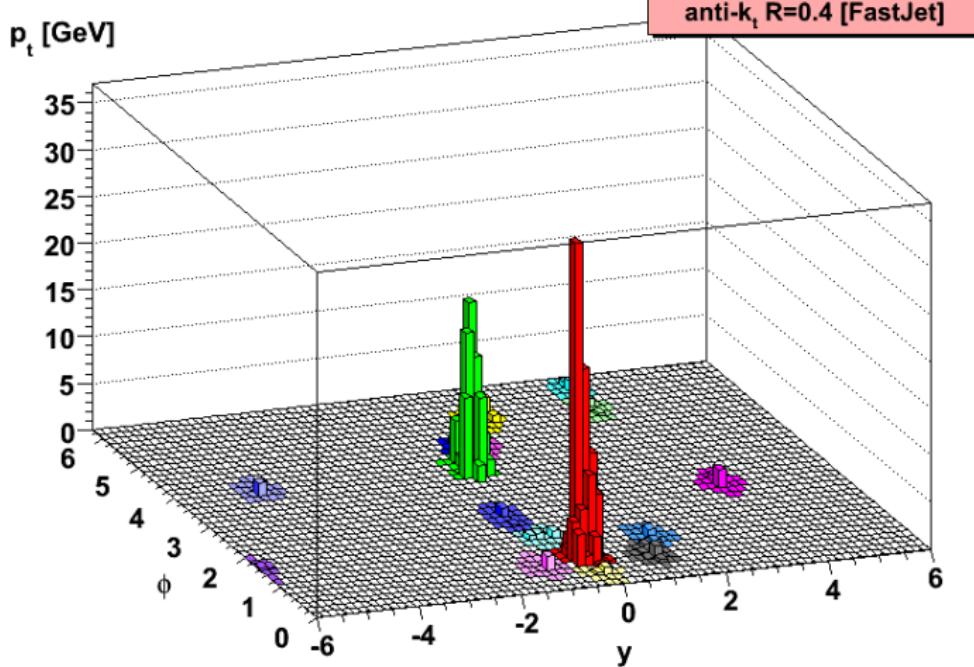


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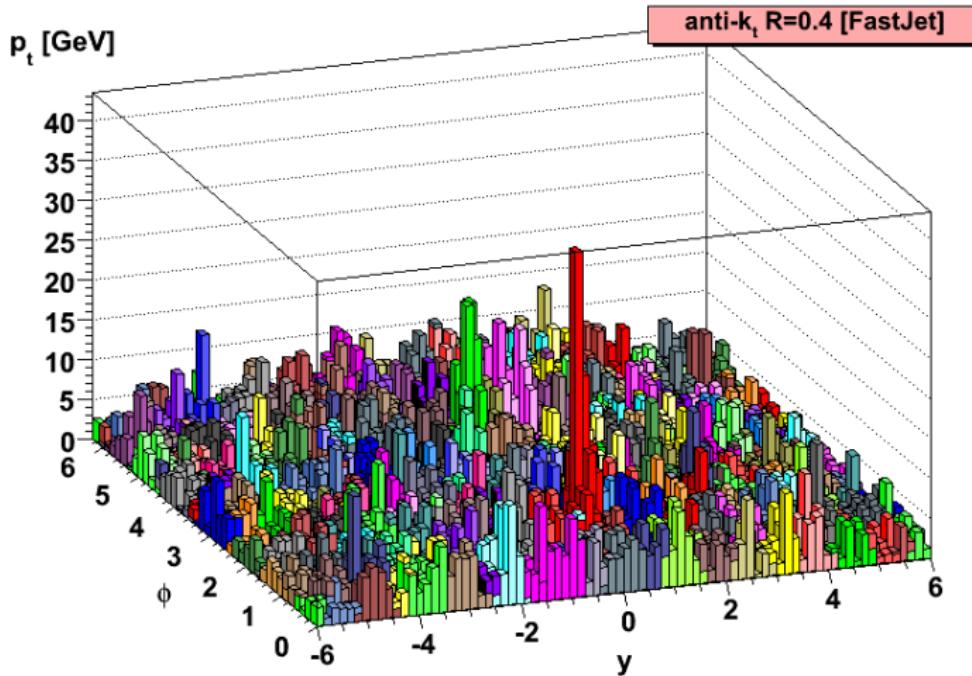
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Common challenge: large contamination



A pp event (LHC 5.5 TeV, Pythia)

Common challenge: large contamination



Contamination
in jet

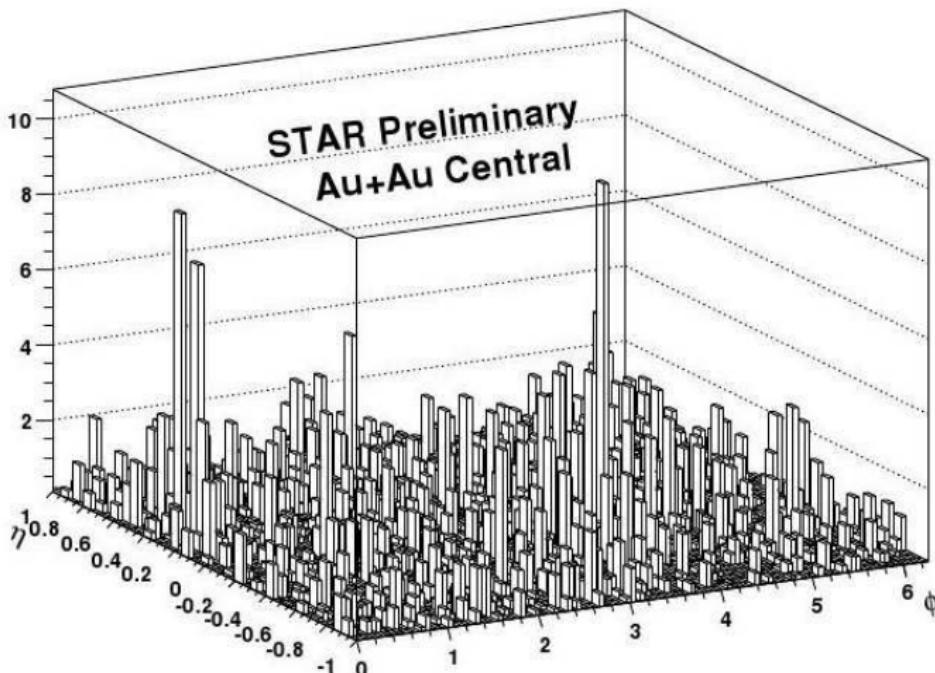
RHIC AuAu:
 $\mathcal{O}(40 \text{ GeV})$

LHC PbPb:
 $\mathcal{O}(100 \text{ GeV})$

LHC pp
(hi-lumi)
 $\mathcal{O}(5 - 40 \text{ GeV})$

A pp event (LHC 5.5 TeV, Pythia), embedded in a HI collision background (Hydjet 1.5)

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**Contamination
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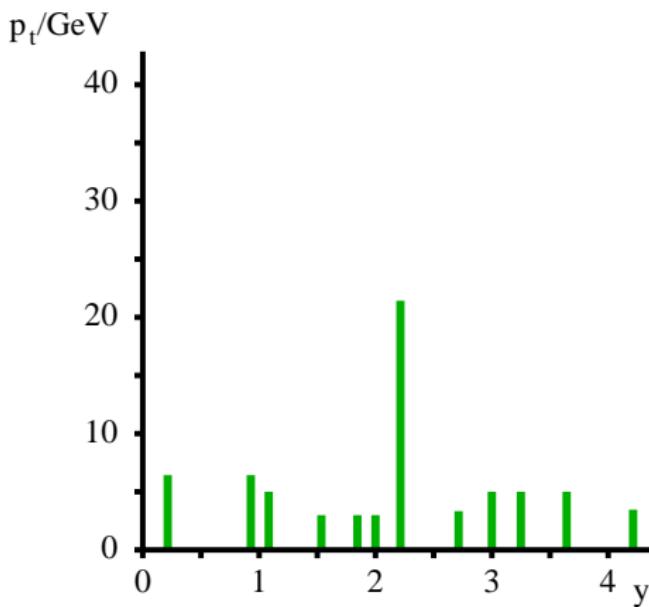
A pp event (LHC 5.5 TeV, Pythia), embedded in a HI collision background (Hydjet 1.5) and an actual STAR event

What are ingredients of jet finding in noisy environments?

1. Jets
2. Jet areas
3. Noise estimation
4. Noise subtraction
- [5. Noise suppression]

1. Jet algorithms

A jet algorithms provides a mapping:



particles \longrightarrow jets
jet.def.

Simplest pp jet algorithm is
“Cambridge/Aachen”

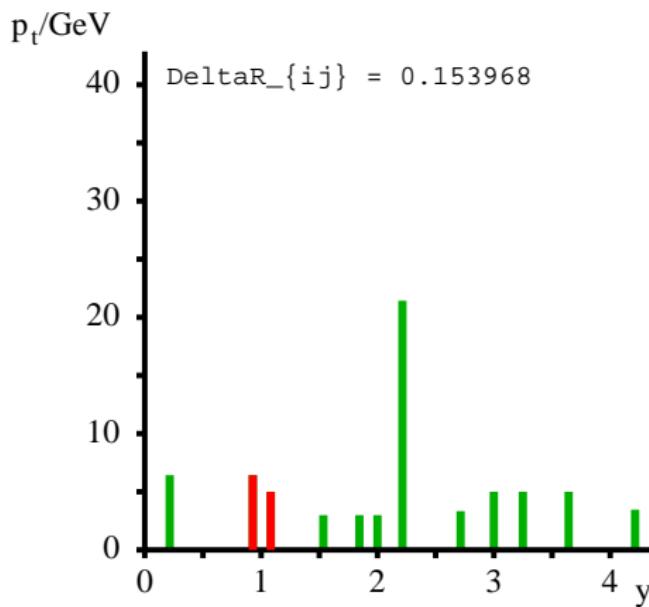
Dokshitzer et al '97

Wengler & Wobisch '98

Repeatedly recombine closest pair
of objects, until all separated by
 $\Delta R_{ij}^2 = \Delta y_{ij}^2 + \Delta\phi_{ij}^2 > R^2$.

R parameter sets angular resolution
 ϕ assumed 0 for all towers

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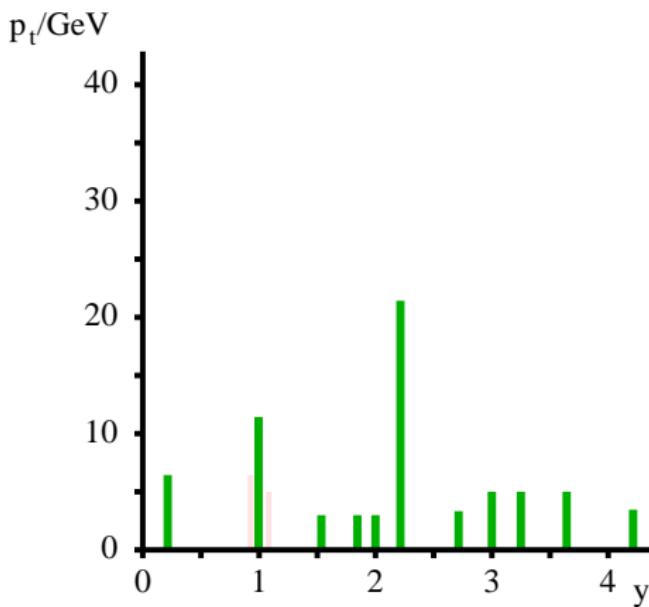
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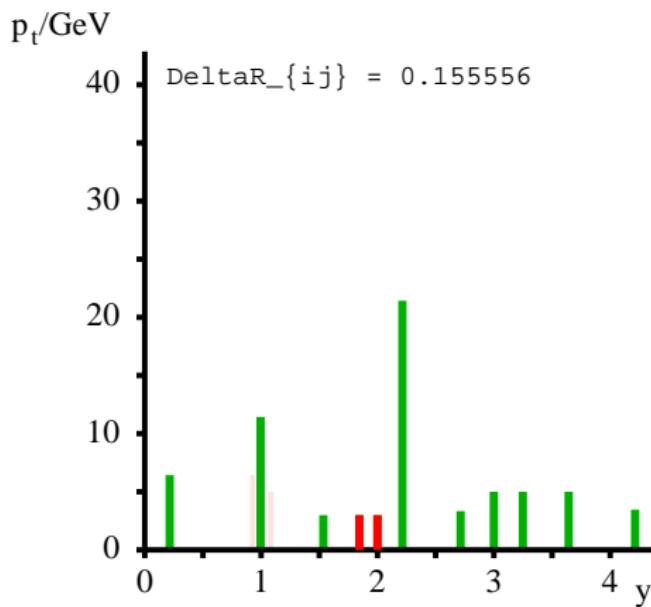
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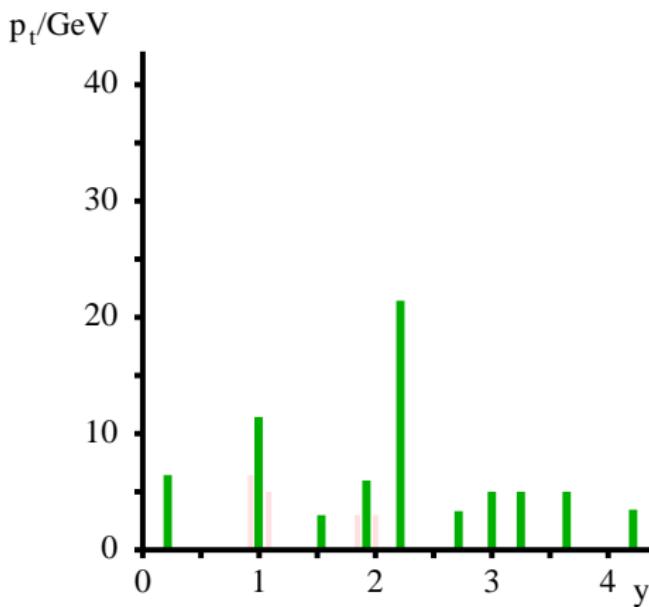
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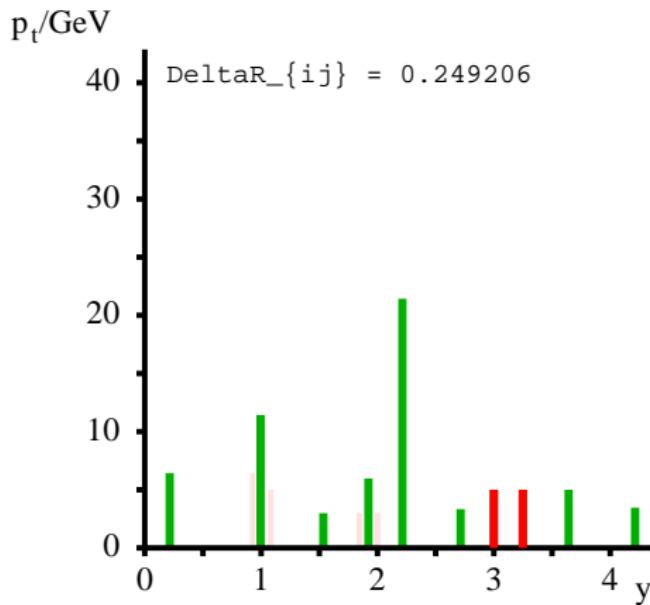
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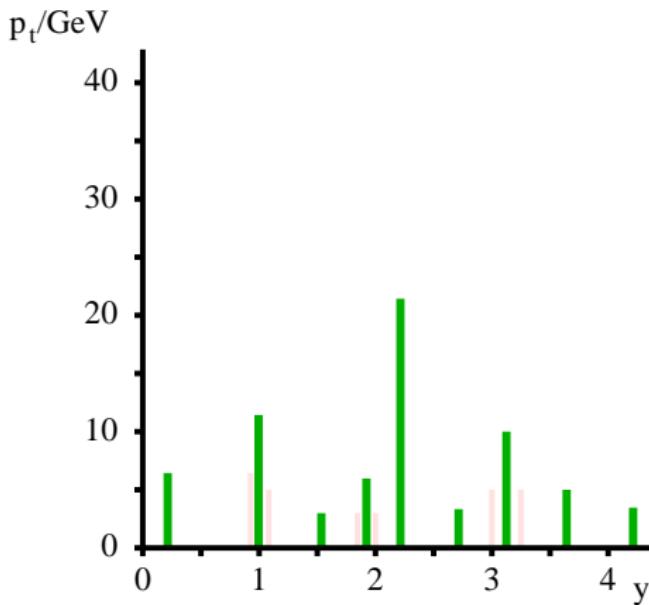
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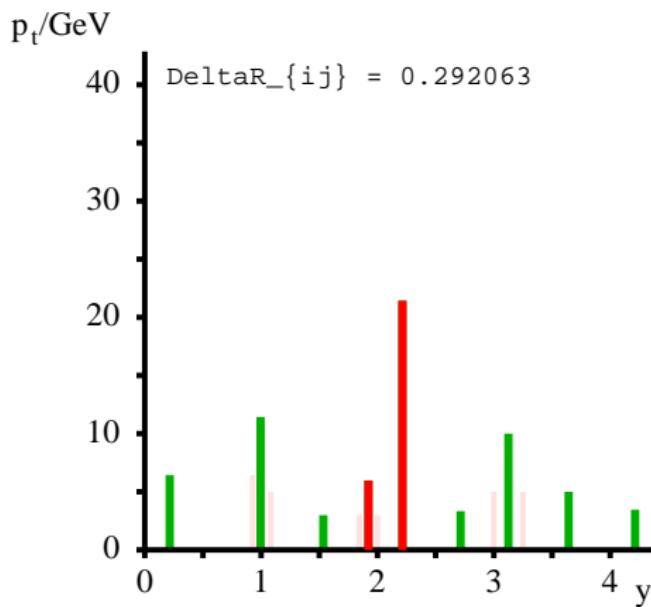
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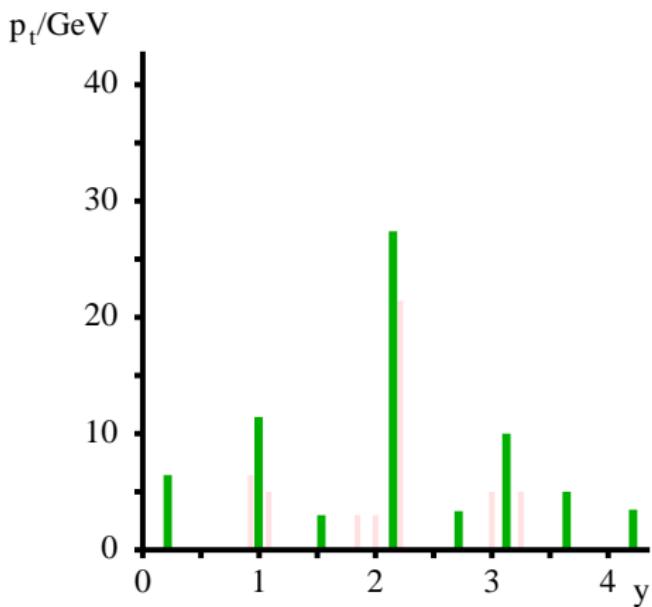
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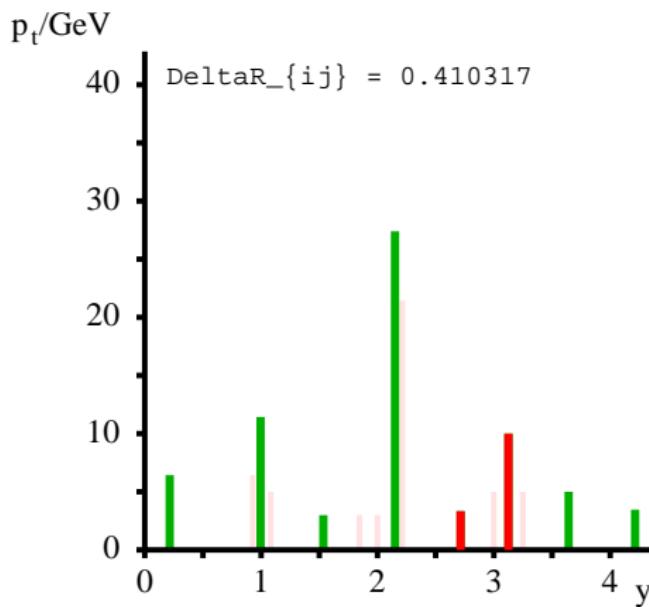
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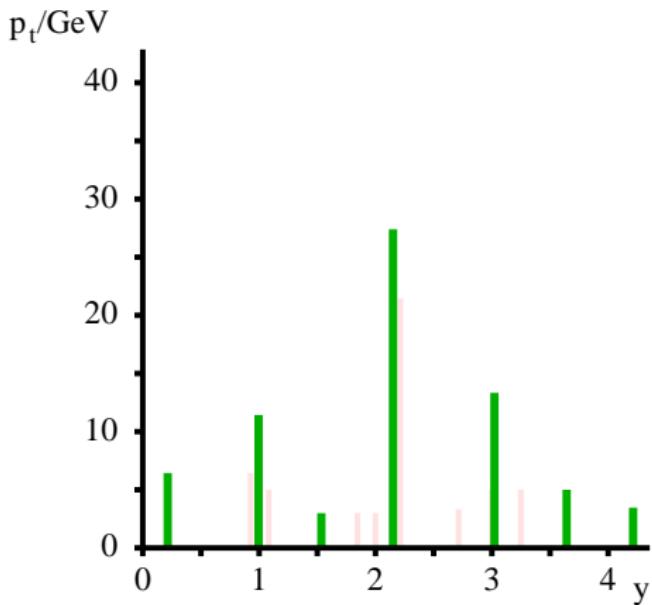
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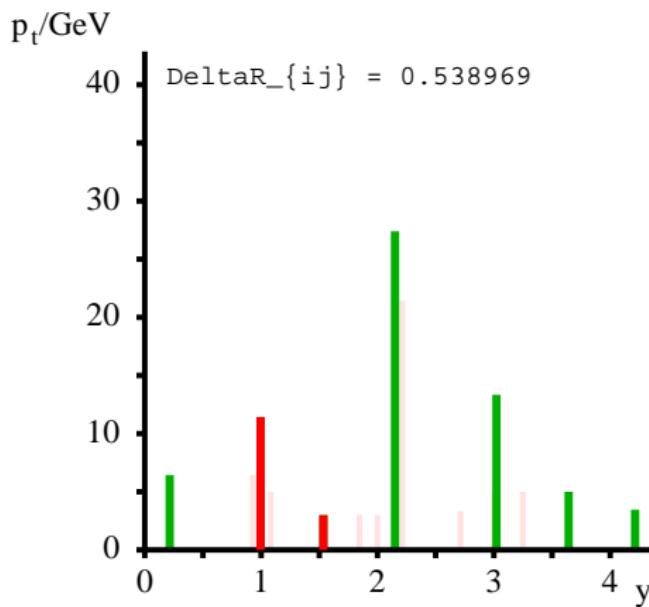
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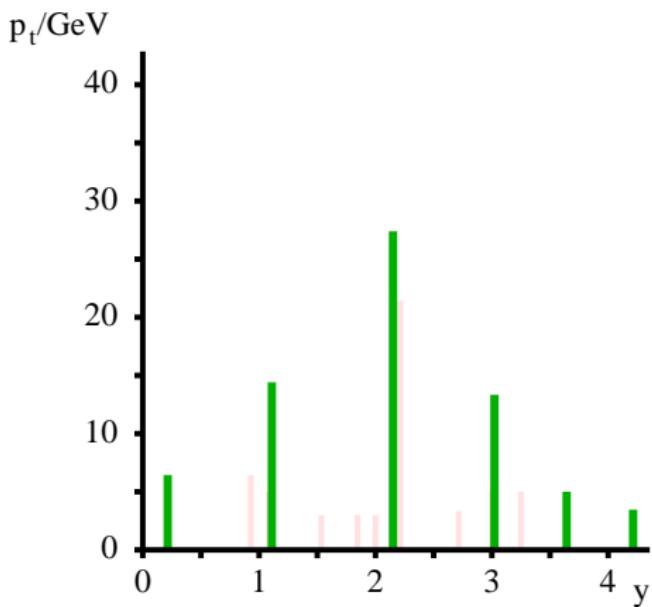
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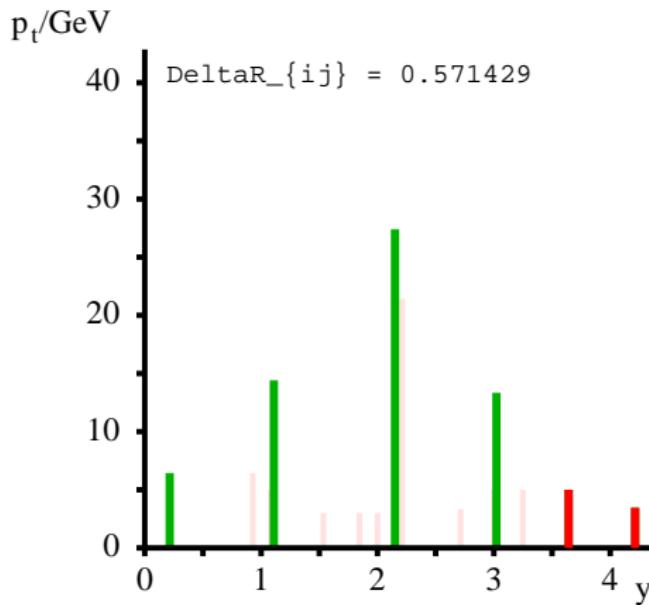
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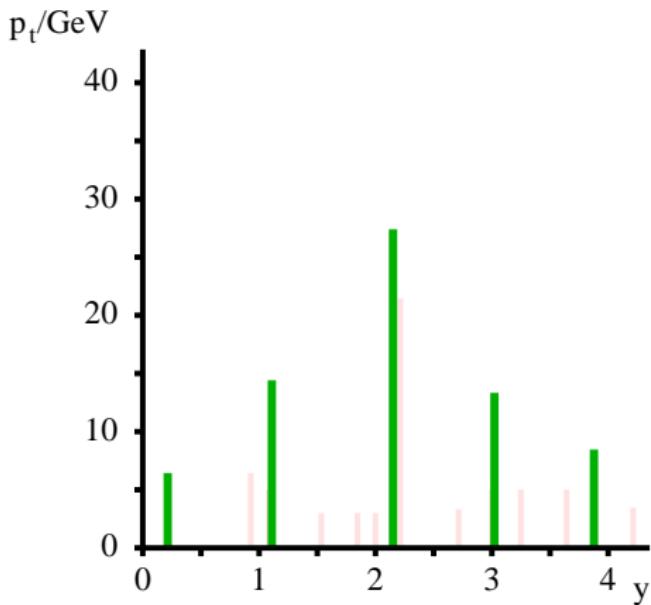
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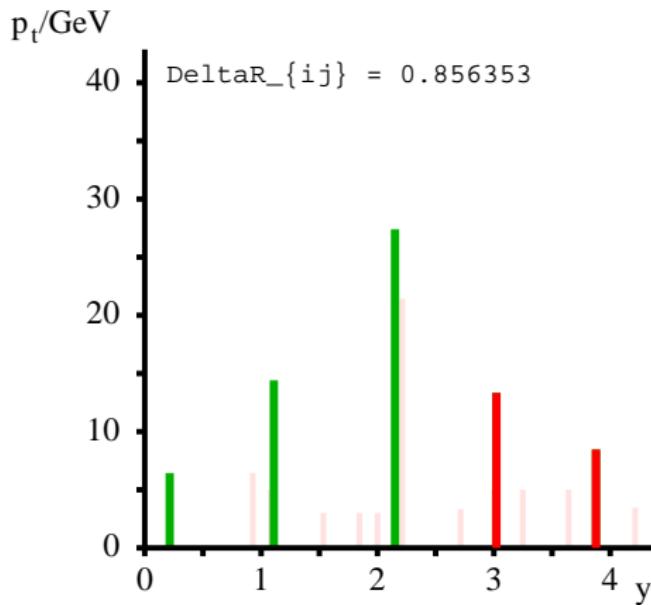
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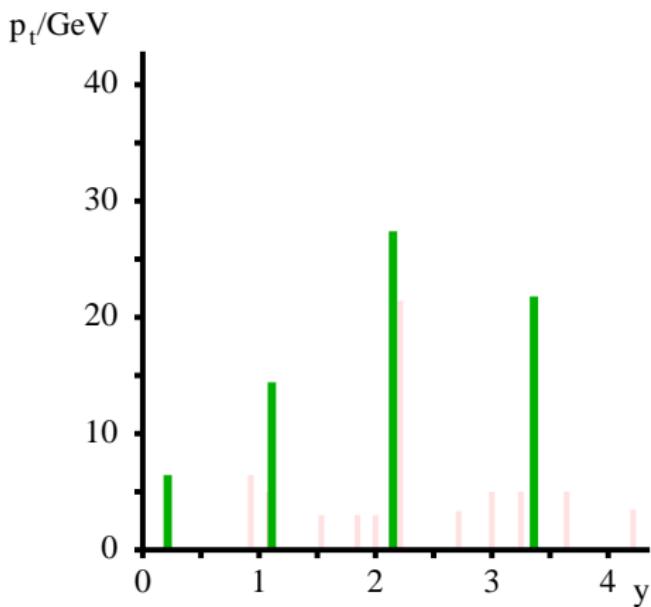
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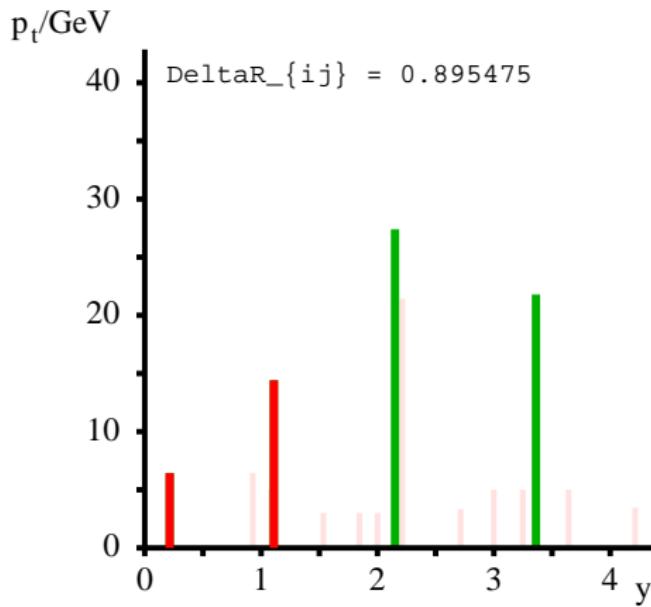
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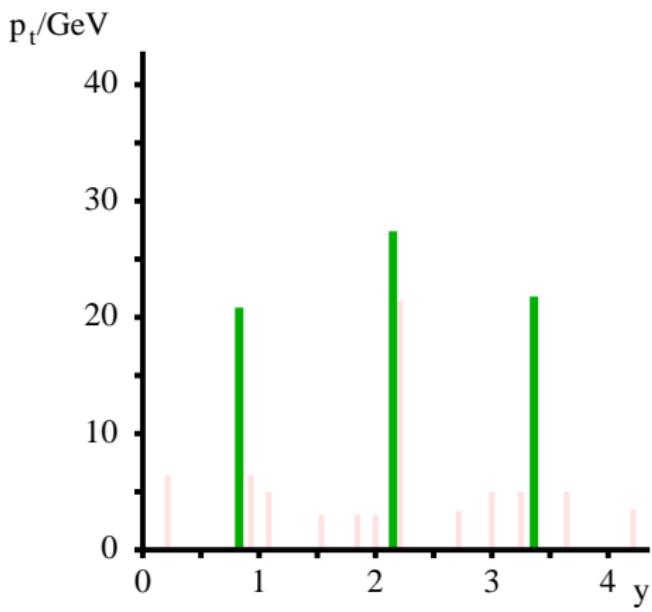
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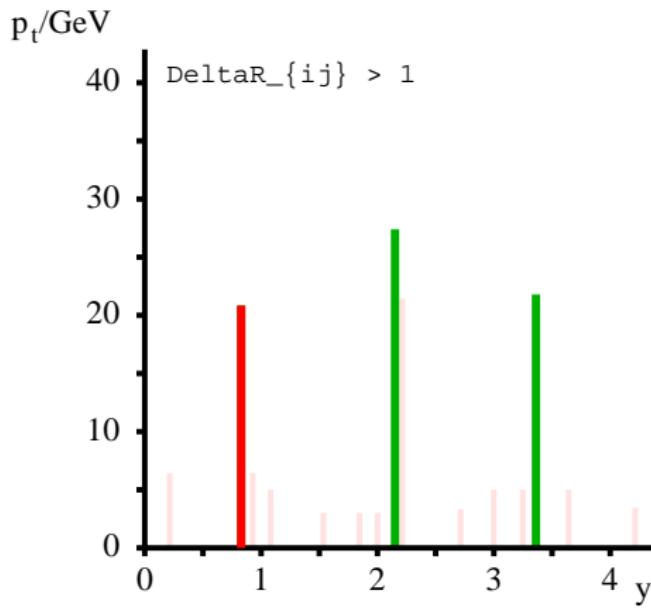
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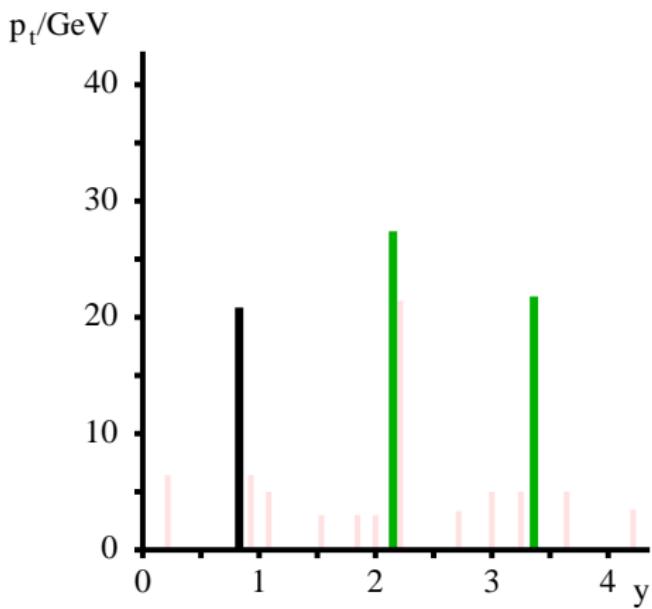
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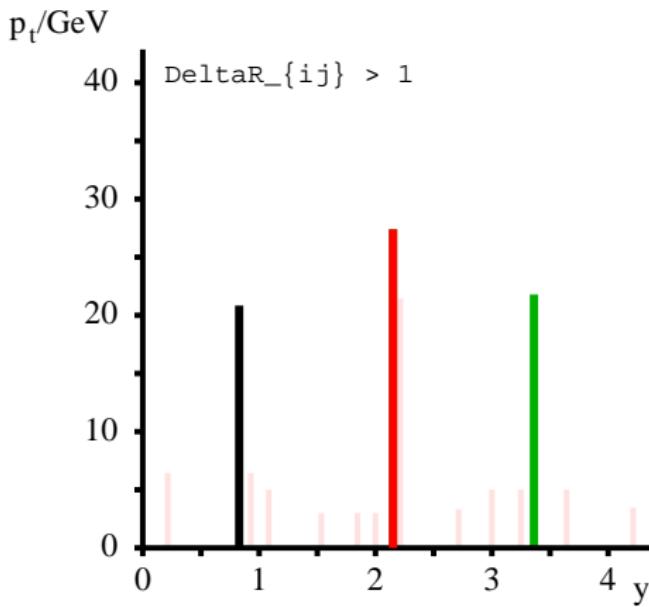
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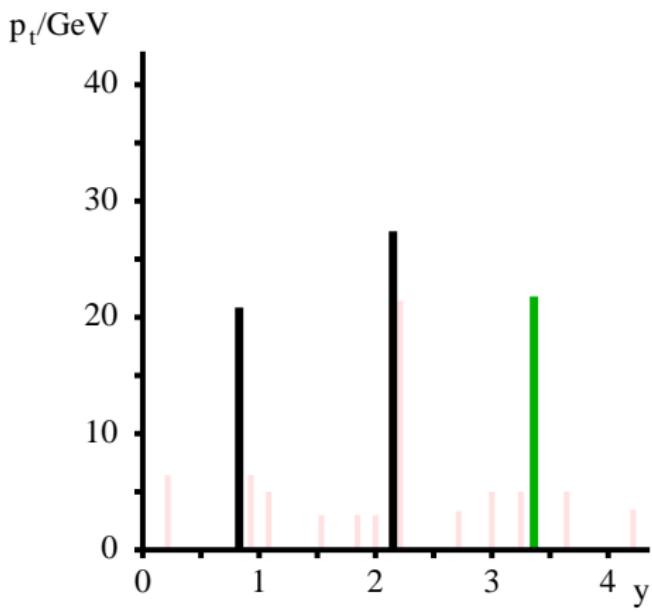
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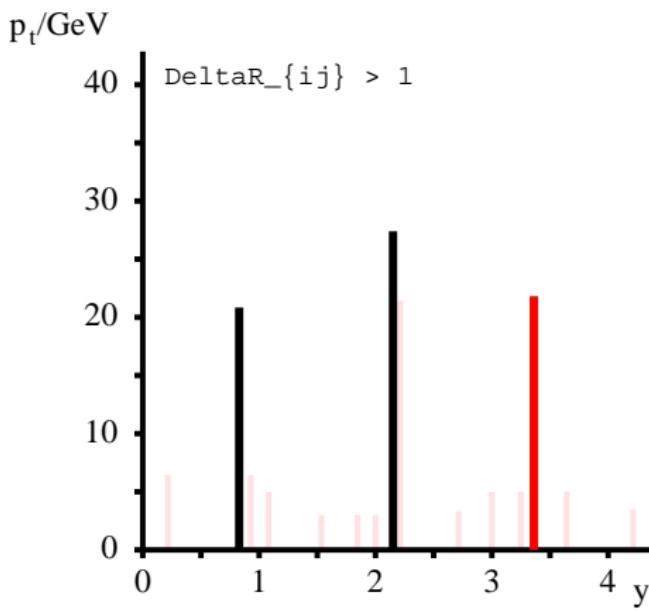
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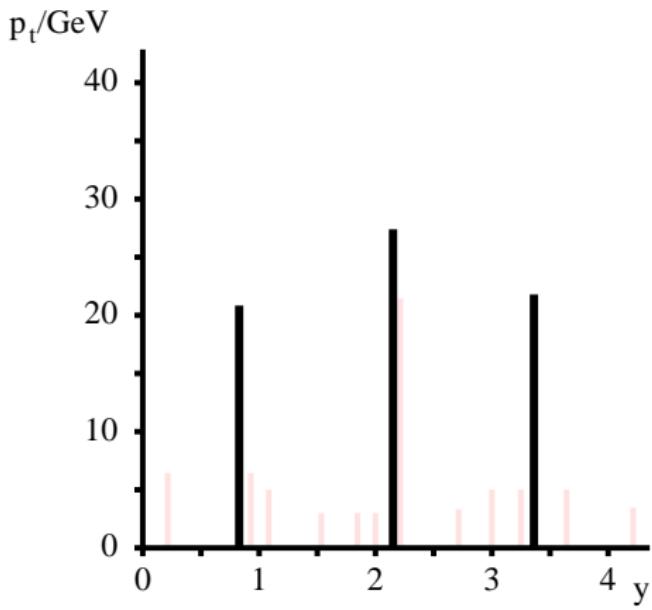
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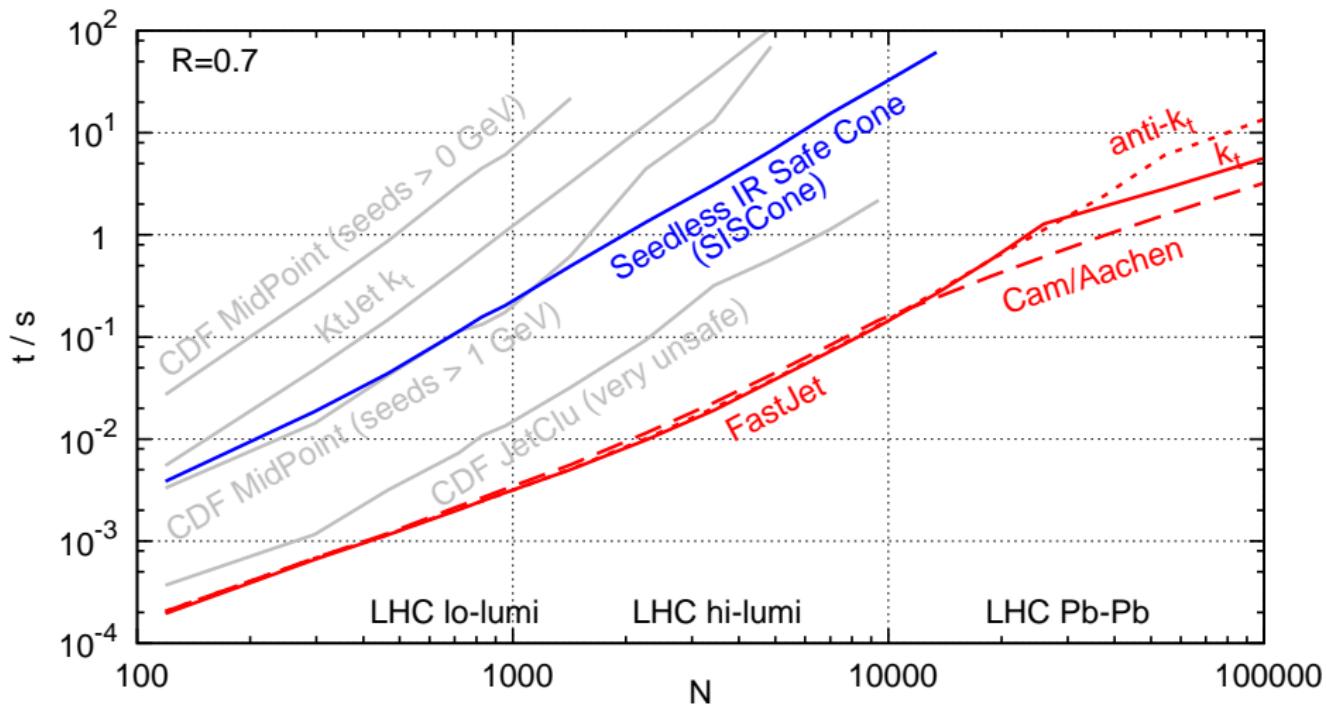
Generalise inclusive-type sequential recombination with

$$d_{ij} = \min(k_{ti}^{2\mathbf{p}}, k_{tj}^{2\mathbf{p}}) \Delta R_{ij}^2 / R^2 \quad d_{iB} = k_{ti}^{2\mathbf{p}}$$

	Alg. name	Comment	time
$p = 1$	k_t <small>CDOSTW '91-93; ES '93</small>	Hierarchical in rel. k_t	$N \ln N$ exp.
$p = 0$	Cambridge/Aachen <small>Dok, Leder, Moretti, Webber '97 Wengler, Wobisch '98</small>	Hierarchical in angle Scan multiple R at once \leftrightarrow QCD angular ordering	$N \ln N$
$p = -1$	anti- k_t Cacciari, GPS, Soyez '08 \sim reverse- k_t Delsart	Hierarchy meaningless, jets like CMS cone (IC-PR)	$N^{3/2}$
SC-SM	SISCone <small>GPS Soyez '07 + Tevatron run II '00</small>	Replaces JetClu, ATLAS MidPoint (xC-SM) cones	$N^2 \ln N$ exp.

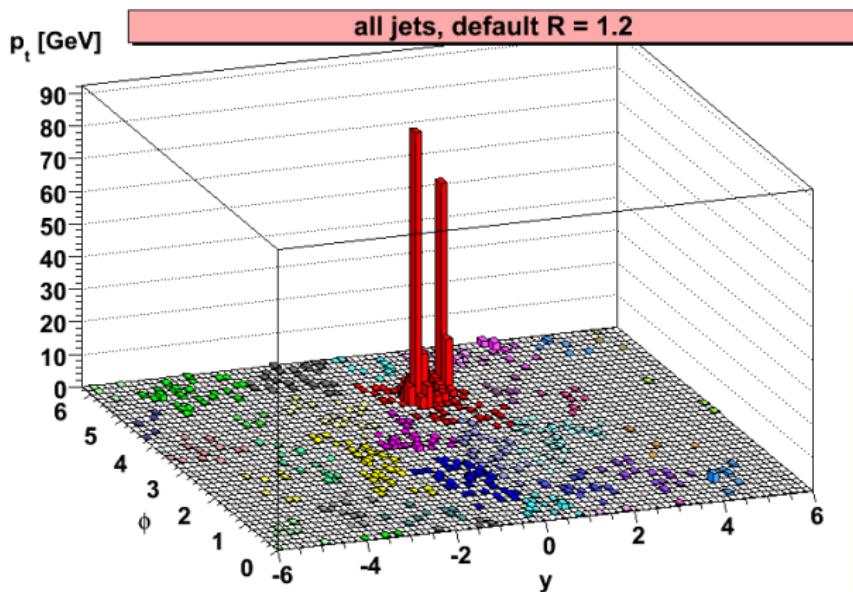
All these algorithms [& much more] coded in (efficient) C++ at
<http://fastjet.fr/> (Cacciari, GPS & Soyez '05-'10)

FastJet: time to cluster N particles



2. Jet areas

Measure jets' susceptibility to contamination by noise



Jets are made of finite number of pointlike particles.

Area not unambiguous concept

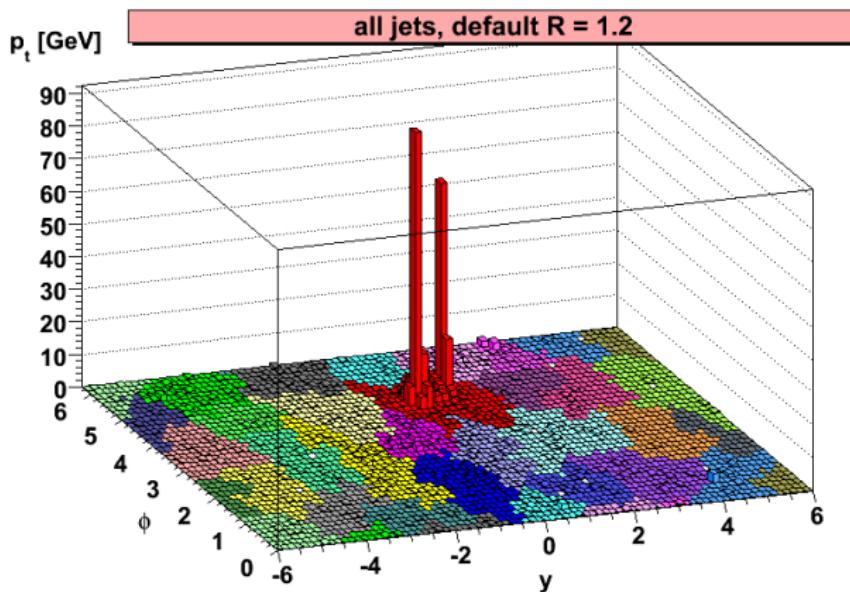
Jet areas must be defined

Add many soft particles to event

10^{-100} GeV each

$A \propto \# \text{ inside jet}$

Cacciari, GPS & Soyez '08
measure of jet's susceptibility to contamination from soft radiation



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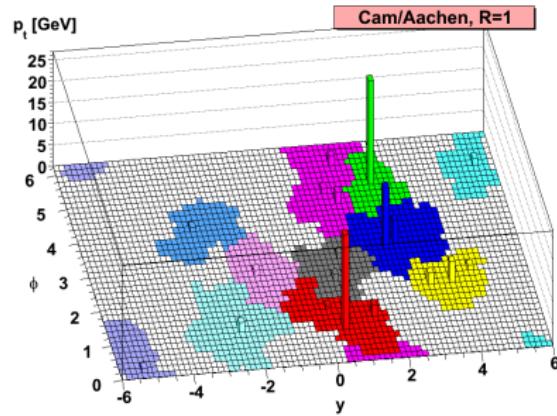
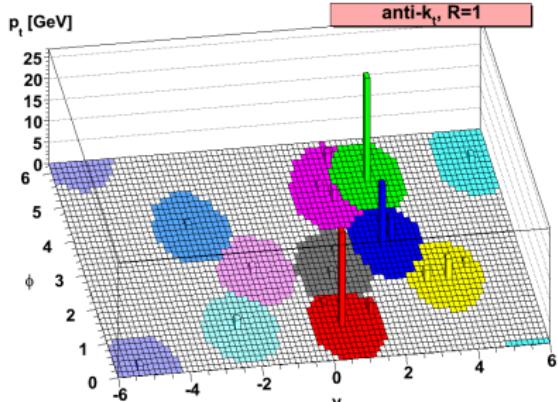
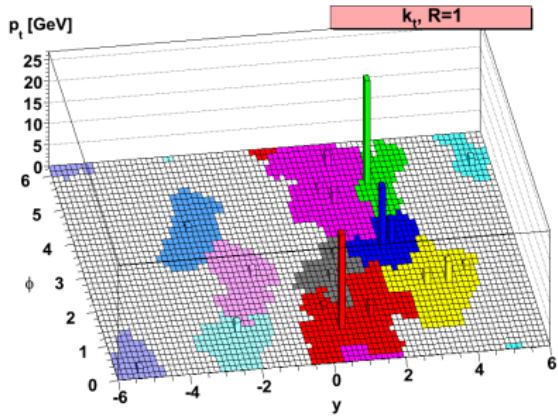
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Areas for 3 jet algorithms



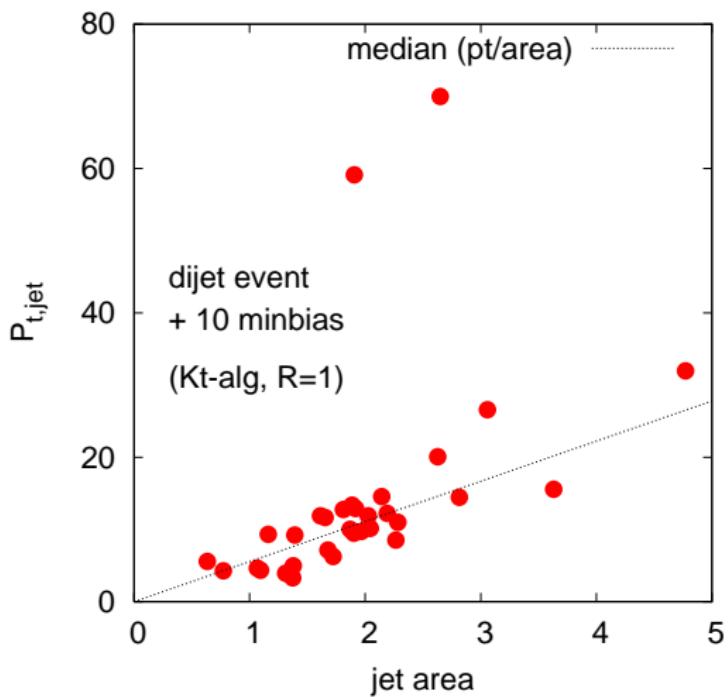
A family of algorithms, all cluster pair with smallest d_{ij} :

$$d_{ij} = \min(p_{ti}^{2p}, p_{tj}^{2p}) \frac{\Delta R_{ij}^2}{R^2}$$

$$p = \begin{cases} 1 & k_t \\ 0 & C/A \\ -1 & \text{anti-}k_t \end{cases}$$

3. Noise estimation

Estimating $\rho \equiv$ background noise level



Most jets in event are “background”

Their p_t is correlated with their area.

Estimate ρ :

$$\rho \simeq \text{median}_{\{j\}} \left[\frac{p_{t,jet}}{A_{jet}} \right]$$

Median limits bias
from hard jets
Cacciari & GPS '07

4. Noise subtraction

$$p_{t,jet}^{\text{subtracted}} = p_{t,jet} - \rho \times A_{jet}$$

$$A_{jet} = \text{jet area}$$

$\rho = p_t$ per unit area from underlying event
(or “background”)

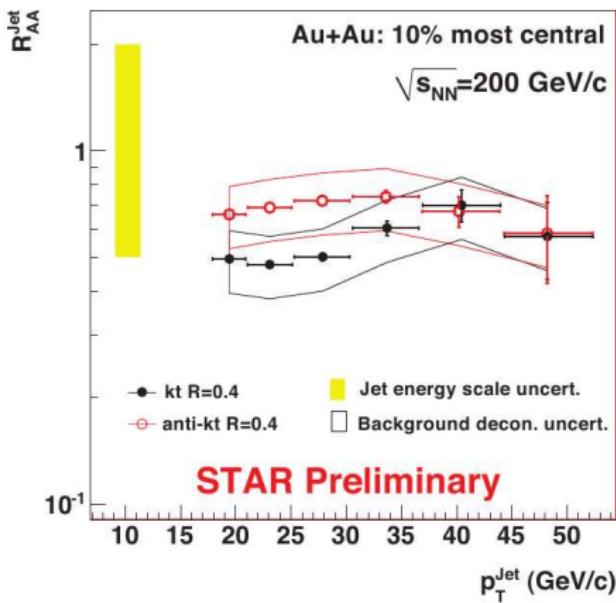
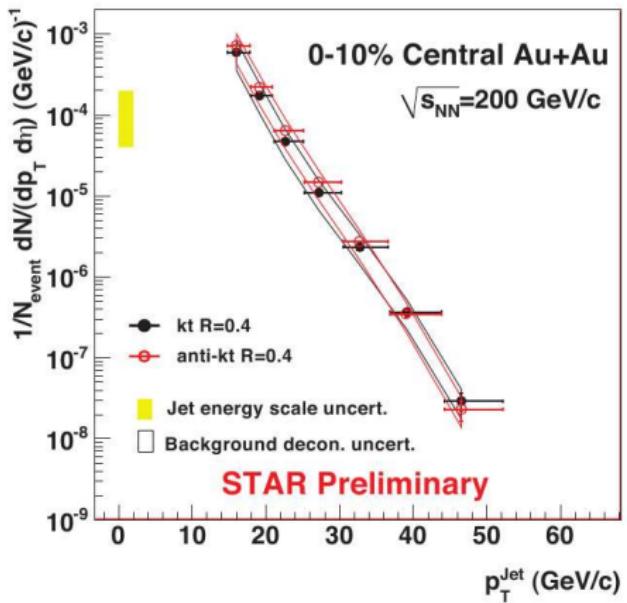
This procedure is intended to be common to pp, pp with pileup (multiple simultaneous minbias) and HIC

NB in AuAu at RHIC: $p_{t,jet}^{\text{subtracted}} = 20 - 50 \text{ GeV}$, $\rho \simeq 80 \text{ GeV}$ and $A_{jet} \simeq 0.5$

Use at RHIC?

Let's examine some of the issues

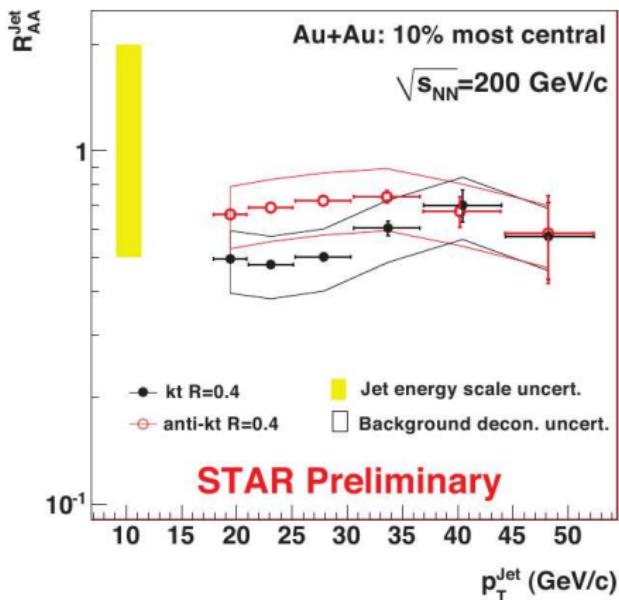
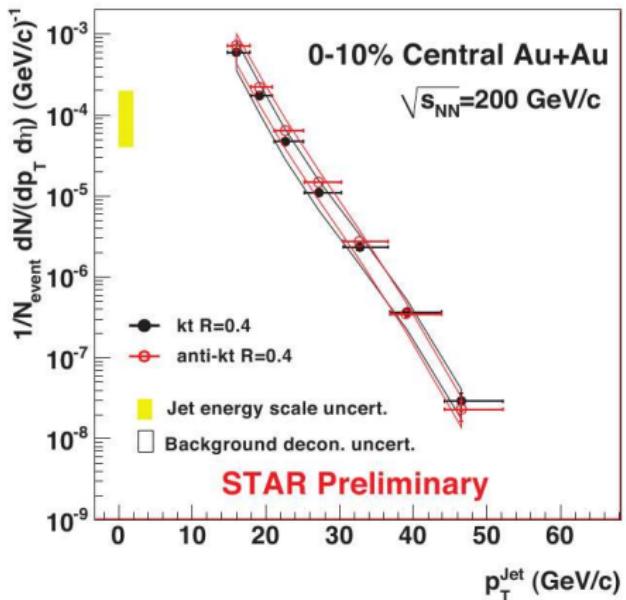
Area/median method → STAR jet results



Method designed to minimise biases, but some still persist.

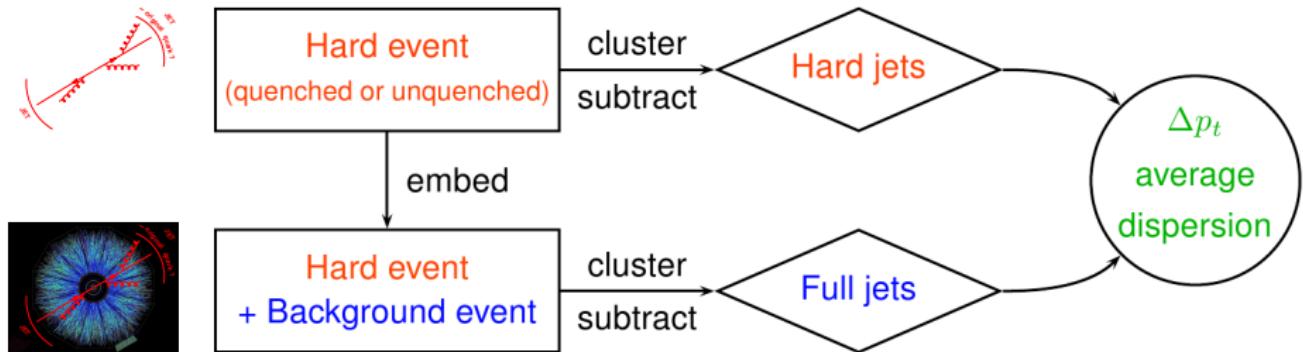
Question: can we calculate size of biases? Can we further reduce them?

Area/median method → STAR jet results



Method designed to minimise biases, but some still persist.

Question: can we calculate size of biases? Can we further reduce them?



Define:

$$\Delta p_t = p_t^{\text{full(subtracted)}} - p_t^{\text{hard}}$$

Study:

average of Δp_t , $\langle \Delta p_t \rangle$
dispersion of Δp_t , $\sigma_{\Delta p_t}$

We don't have real background (or hard!) events, so use

- ▶ Hard event: Pythia 6.4, and optionally PyQyen 1.5 / QPythia
- ▶ Background: Hydjet 1.6 (Hydjet++ 2.1 for cross-checks)
- ▶ Analysis: FastJet 2.4 (& 2.5-devel), $R = 0.4$ for all main jet finders (k_t with $R = 0.5$ for bkgd estimation).

Cacciari, Rojo, GPS & Soyez, to appear soon...

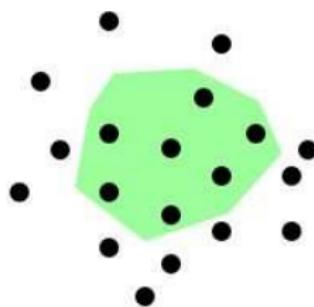
Example #1: non-zero $\langle \Delta p_t \rangle$

(background does not *just* linearly add noise to jet)

BACK REACTION

“How (much) a jet changes when immersed in a background”

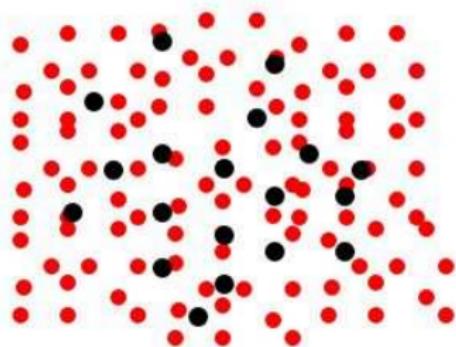
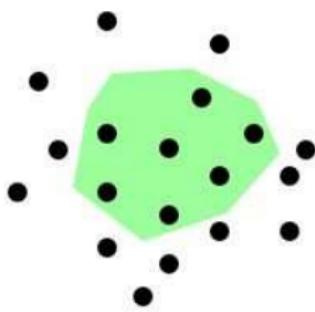
Without
background



BACK REACTION

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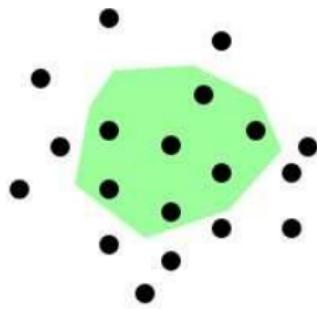
Without
background



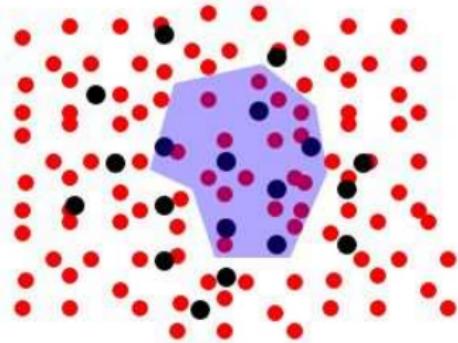
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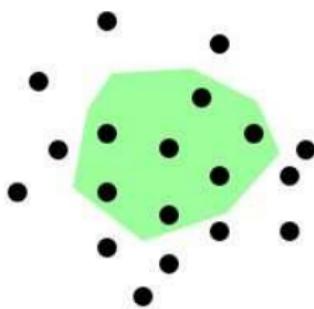
With
background



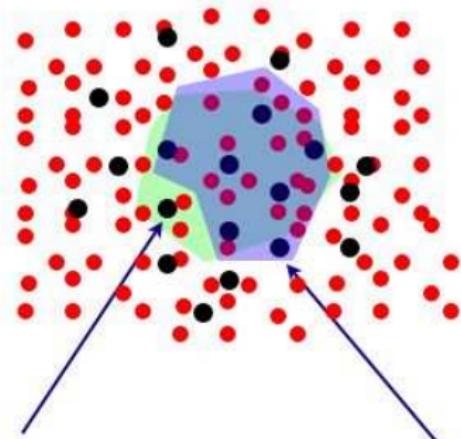
BACK REACTION

“How (much) a jet changes when immersed in a background”

Without
background



With
background



Backreaction **loss**

Backreaction **gain**

Backreaction can be calculated (sort of...)

Soft & collinear approximation:

$$\delta p_t^{BR} = \mathcal{B}_{alg} \cdot \rho R^2 \frac{2C_i}{\pi} \alpha_s \ln \frac{p_t}{\rho R^2}$$

Cacciari, GPS & Soyez '08
+ large corrections

jet alg	\mathcal{B}_{alg}
k_t	-0.3
C/A	-0.3
anti- k_t	0

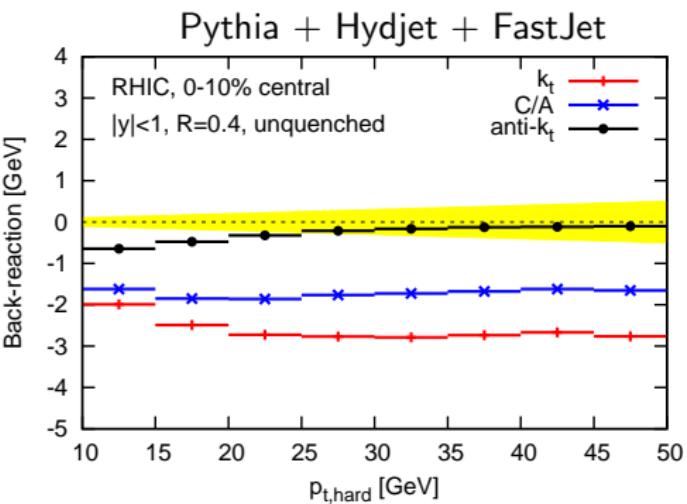
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Cacciari, Rojo, GPS & Soyez, prelim.

anti- k_t bias = 0, as expected

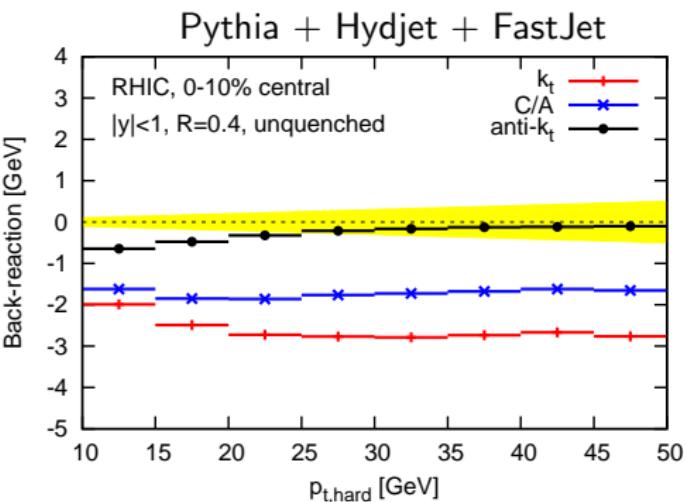
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Cacciari, Rojo, GPS & Soyez, prelim.

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Different jet algorithms have different systematics
Use of more than one provides important cross-checks

Example #2: fluctuations

Fluctuations of amount of background / underlying-event in a square of unit area can be characterised in terms of σ_{UE} , which is $\mathcal{O}(10 \text{ GeV})$ at RHIC.

Dispersion in jet subtraction, σ_{jet} is given by

$$\sigma_{\Delta p_t} = \sigma_{UE} \times \sqrt{A_{jet}}$$

jet alg	$\langle A_{jet} \rangle$
k_t	$0.81\pi R^2$
C/A	$0.81\pi R^2$
anti- k_t	πR^2

+ p_t -dependent scaling
violations for k_t and C/A

Put in numbers:

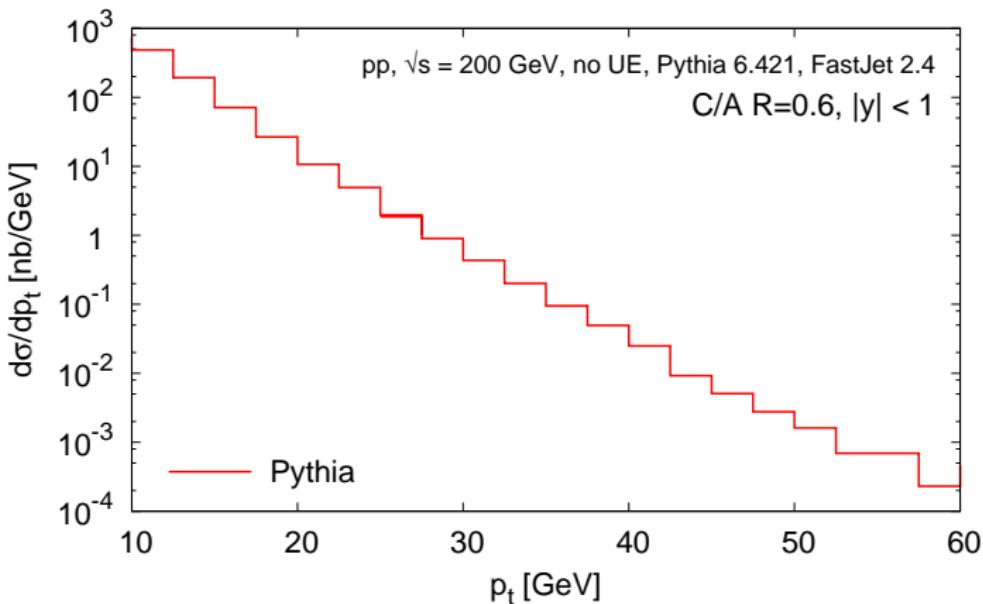
$$\sigma_{UE} \simeq 8 - 10 \text{ GeV}$$

$$\rightarrow \sigma_{\Delta p_t} \sim 6 - 7 \text{ GeV}$$

What impact does this have?

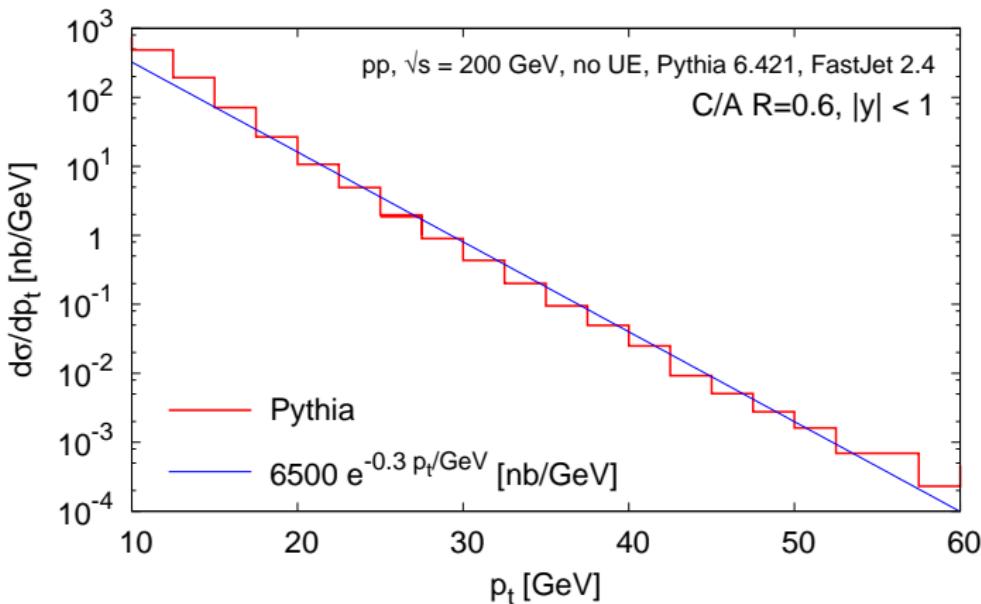
Context: a steeply falling X-section

RHIC Inclusive jet spectrum



Context: a steeply falling X-section

RHIC Inclusive jet spectrum



To help think about impact of falling cross section at RHIC, approximate it as:

$$\frac{d\sigma}{dp_t} \sim \exp(-0.3p_t / \text{GeV})$$

Interplay of PDFs & $1/p_t^4$ matrix element

Exponential spectrum \otimes Gaussian fluctuations

$$\exp(-ap_t) \otimes \frac{1}{\sqrt{2\pi}\sigma} \exp\left(-\frac{\Delta p_t^2}{2\sigma^2}\right)$$

Real fluctuations not quite Gaussian

Real spectrum not quite exponential (especially at low & high p_t !)
 But simple approximations give instructive analytical answers

Convolution rescales spectrum by factor:

$$\exp\left(\frac{1}{2}a^2\sigma^2\right) \sim 10 \text{ for } \sigma \sim 7 \text{ GeV}$$

Convolution migrates p_t 's by

$$a\sigma^2 \sim 15 \text{ GeV for } \sigma \sim 7 \text{ GeV}$$

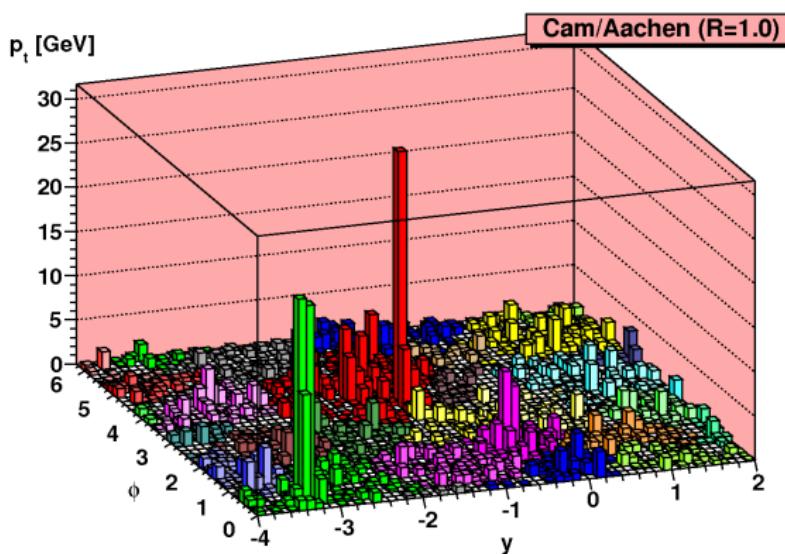
Reducing fluctuations, while
limiting bias:

filtering

Idea to improve resolution for an LHC Higgs search in $H \rightarrow b\bar{b}$ decay mode!

Keep hardest $\mathcal{O}(\alpha_s)$ gluon emission in jet, while throwing out soft “junk”

Butterworth, Davison, Rubin & GPS '08



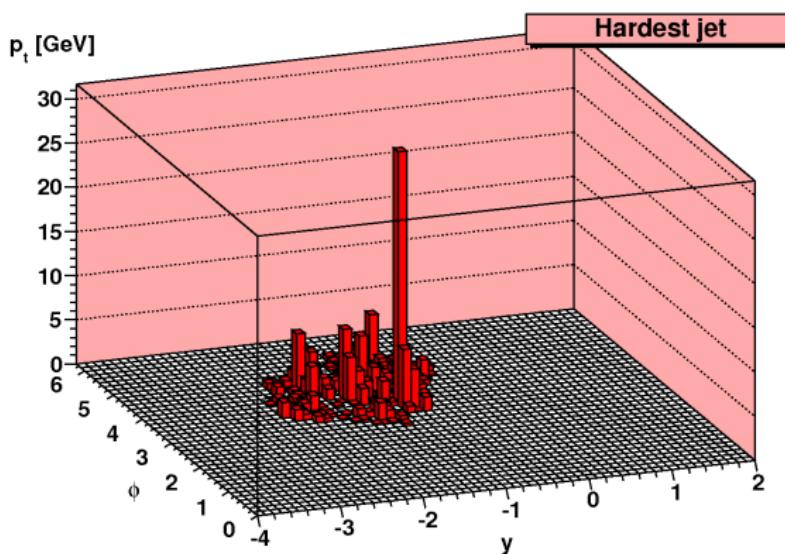
1. Consider a jet
2. View it on smaller angular resolution scale R_{filt}
3. Take (e.g.) 2 hardest “subjets” leading quark + 1 gluon
4. The result is a “filtered” jet

Related ideas by Ellis, Vermillion & Walsh '09 and Krohn, Thaler & Wang '09

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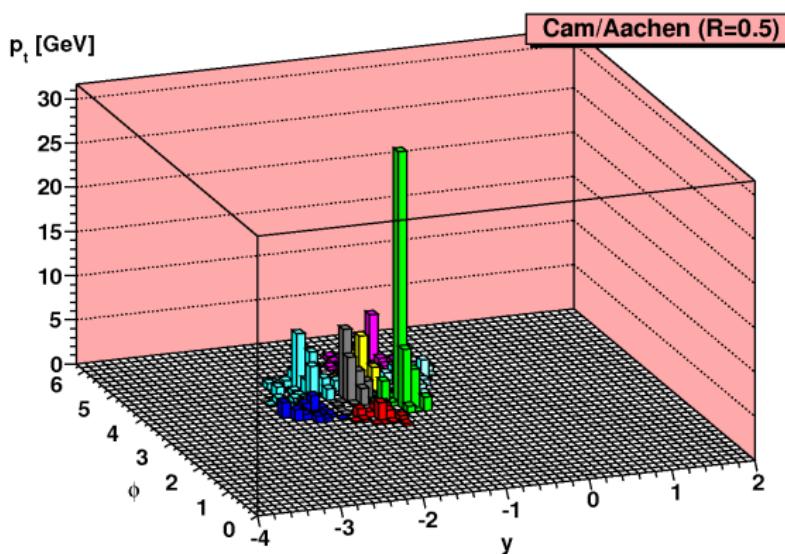
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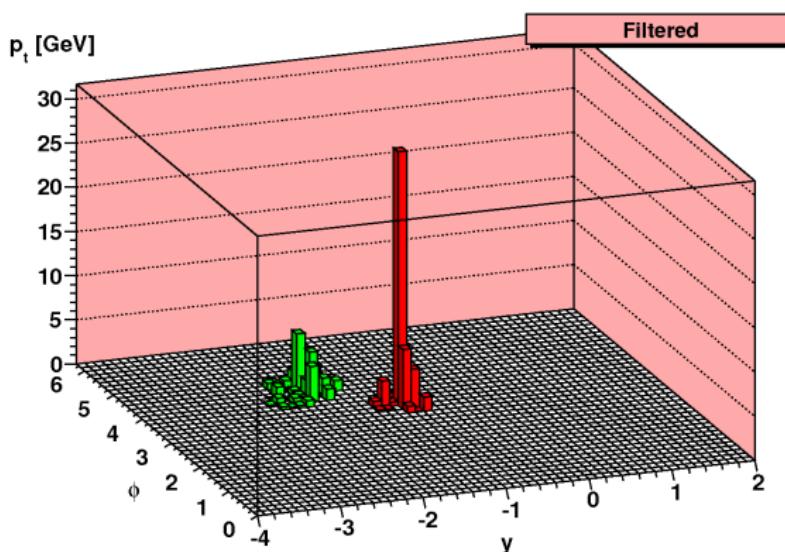
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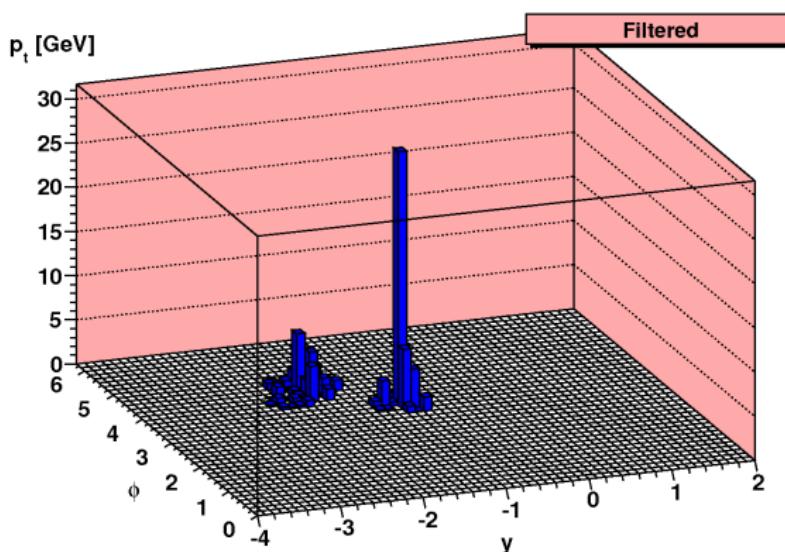
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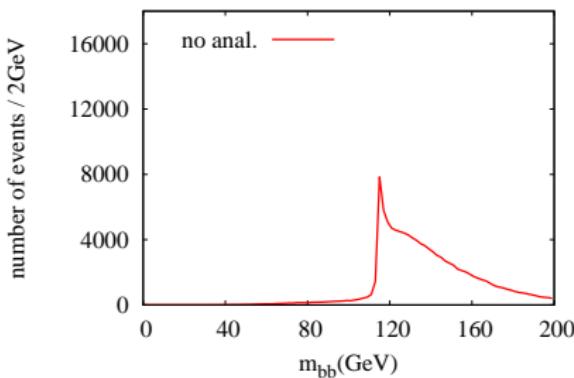


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Related ideas by Ellis, Vermillion & Walsh '09 and Krohn, Thaler & Wang '09

Reconstructed mass for jets from decay of high- p_t Higgs-boson [without pileup]

Without Filtering



With Filtering

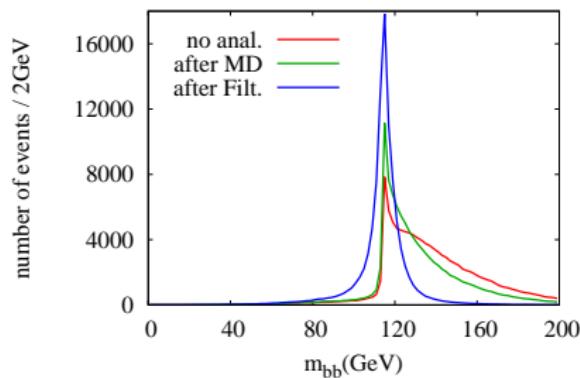
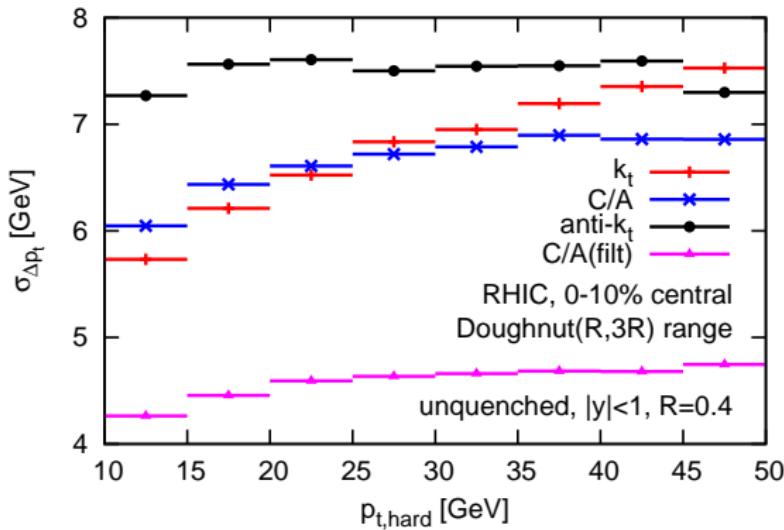


Figure from Rubin

Among the techniques adopted in search for $H \rightarrow b\bar{b}$ at LHC

Impact of filtering on dispersion in HIC

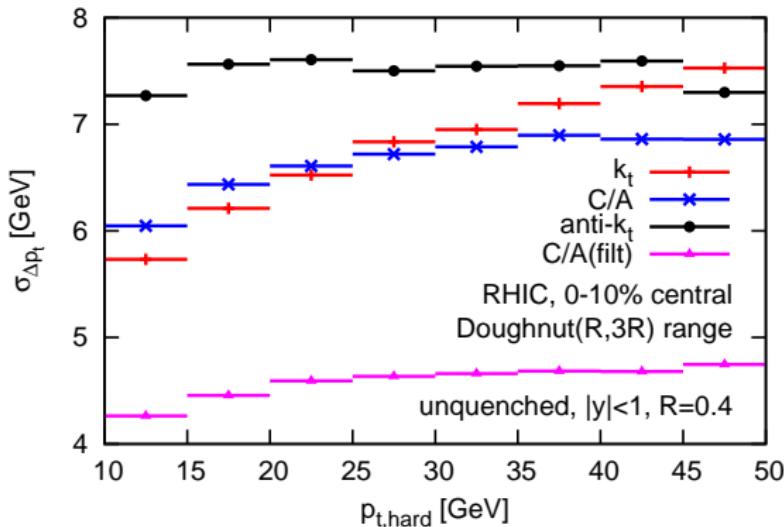


Filtering reduces jet area by $\sim \frac{1}{2}$

Fluctuations $\propto \sqrt{A}$
should go down by $\sim \sqrt{\frac{1}{2}}$

And they do

Impact of filtering on dispersion in HIC



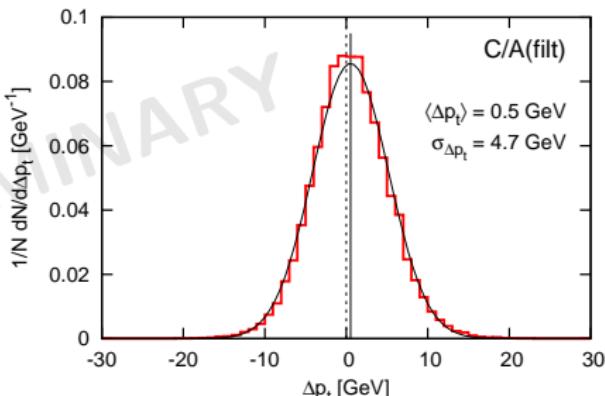
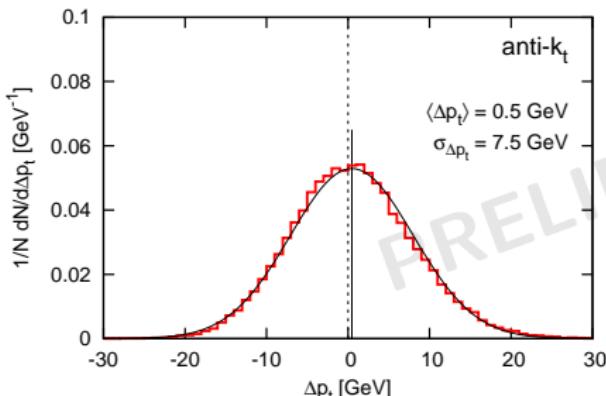
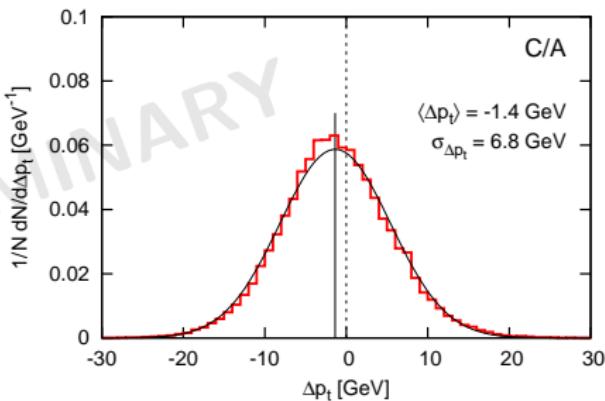
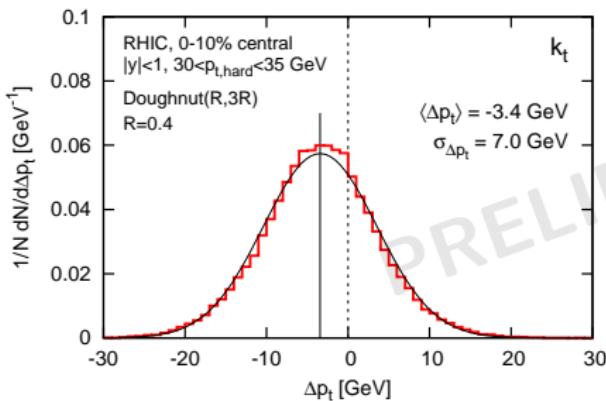
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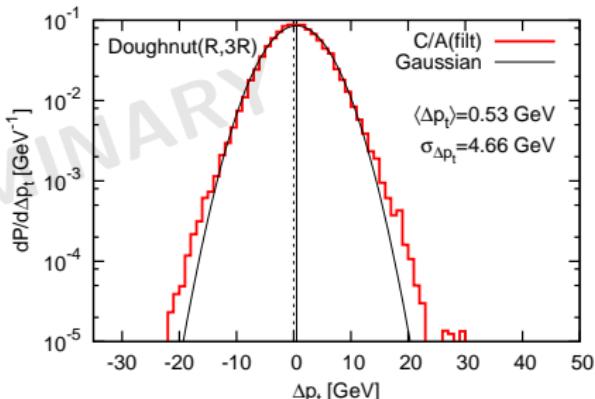
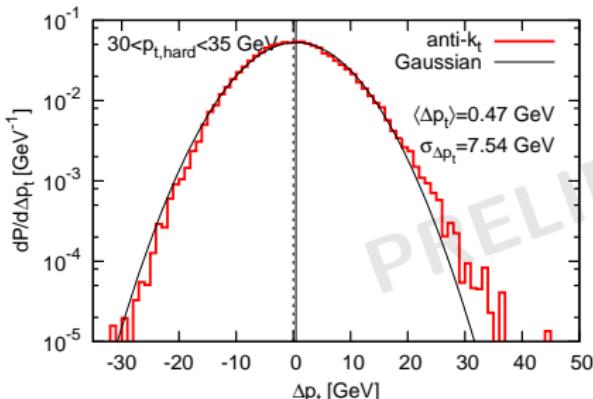
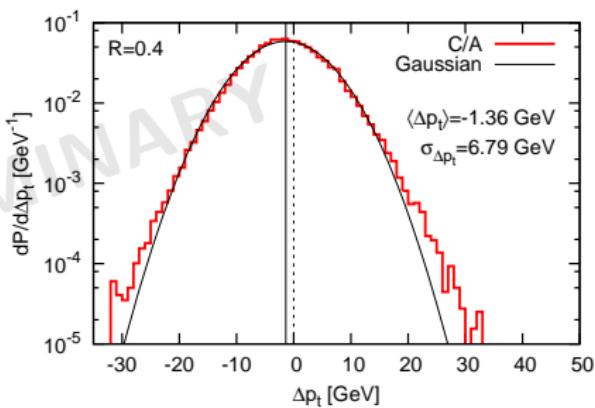
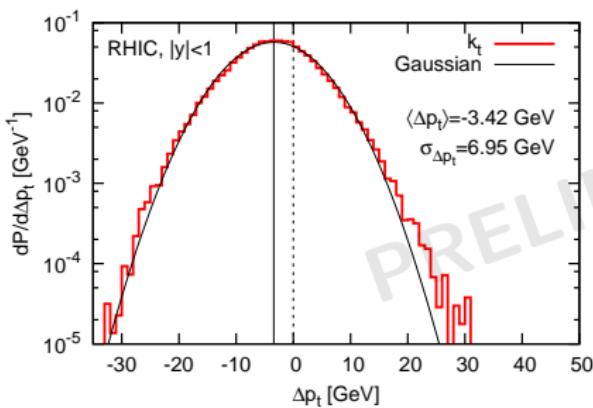
Fluctuations $\propto \sqrt{A}$
should go down by $\sim \sqrt{\frac{1}{2}}$

And they do

Filtering's reduction of dispersion from 7 GeV to 5 GeV means experimental "unfolding" might be factor 3 instead of factor 10

Numbers are rough – intended to give an idea of impact
NB: Gaussian filtering (Cole & Lai '08) not the same thing



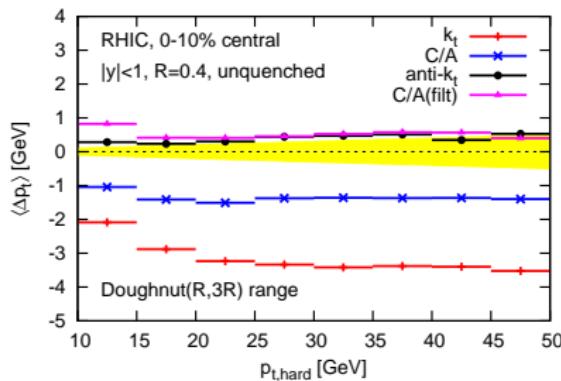


Does filtering introduce new biases in jets in quenched case?

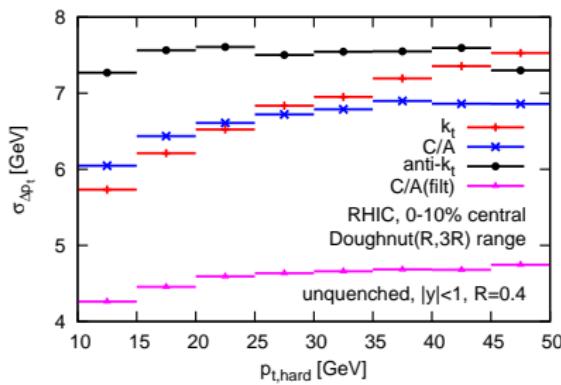
Vacuum QCD: we know how much gluon radiation we lose
QCD in medium: extra medium-induced radiation lost?

UNQUENCHED

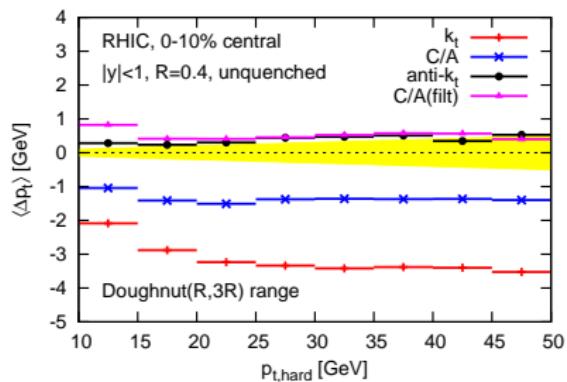
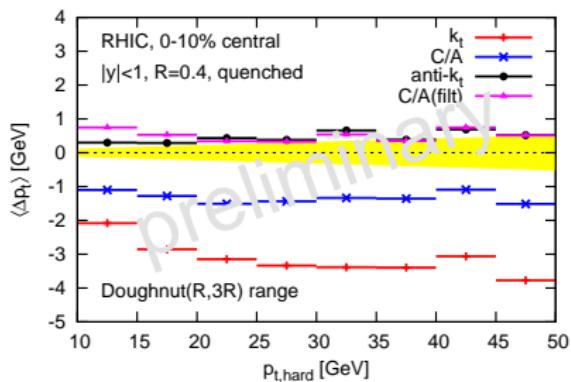
PT SHIFT



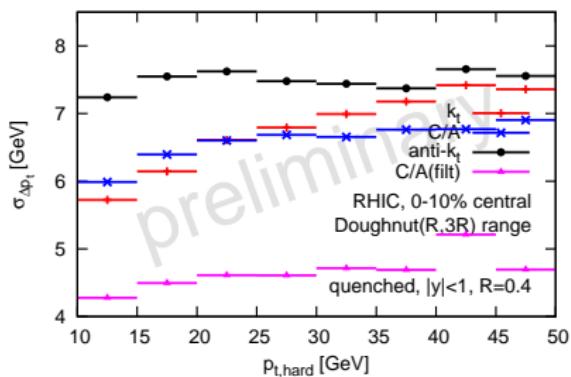
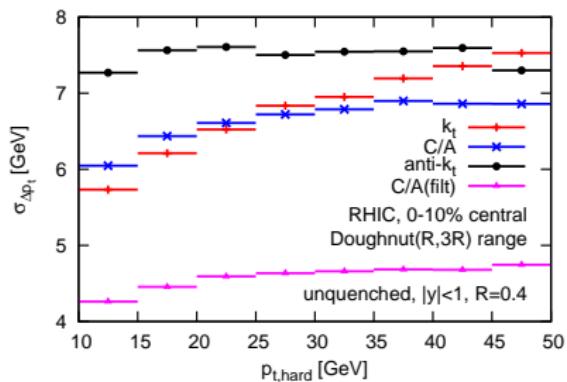
DISPERSION



PT SHIFT

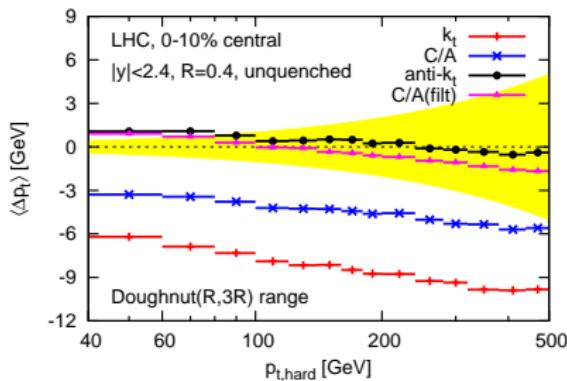
**QUENCHED**

DISPERSION

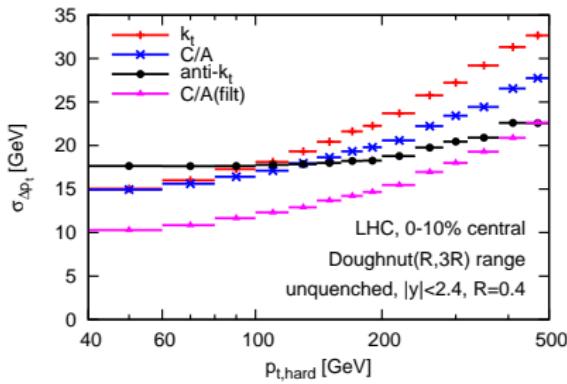


UNQUENCHED

PT SHIFT

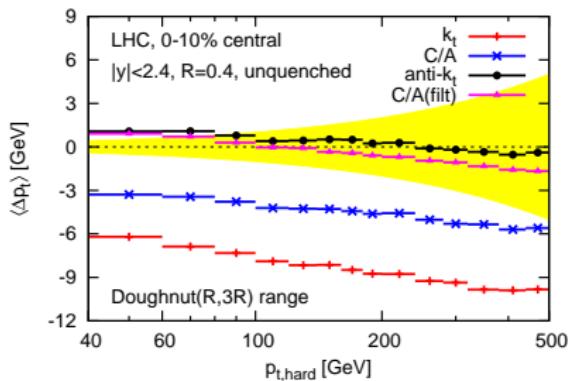


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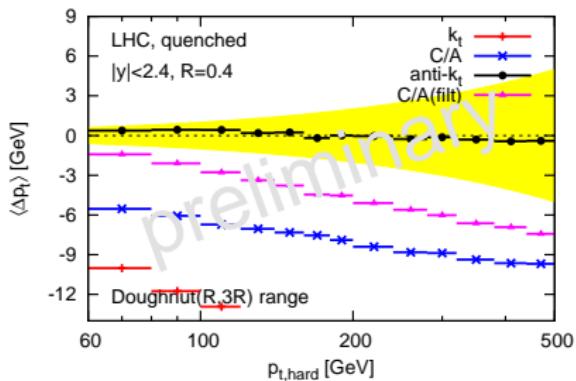


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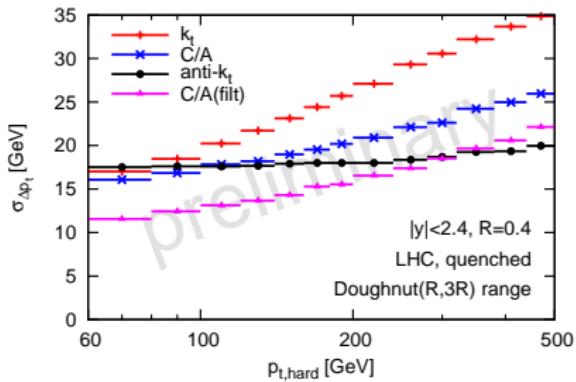
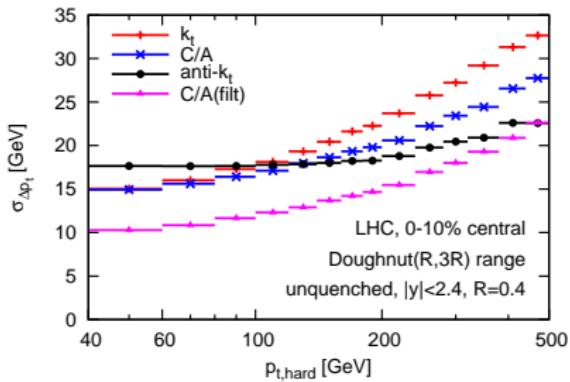
UNQUENCHED



QUENCHED



DISPERSION



It's still early days for jet-finding in HIC (& high-luminosity LHC)

It's a tough job to accurately remove 40 GeV of noise from a 40 GeV hard jet in the context of a steeply falling cross-section.

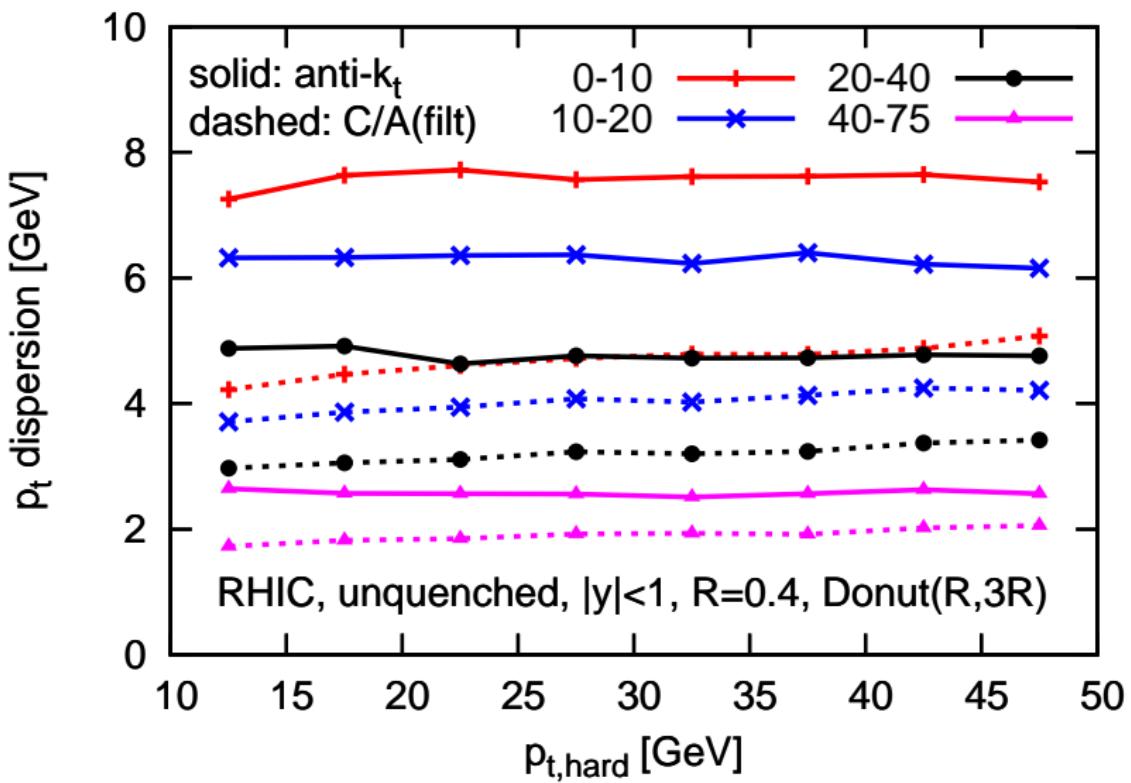
Theory calculations can guide the choices one makes

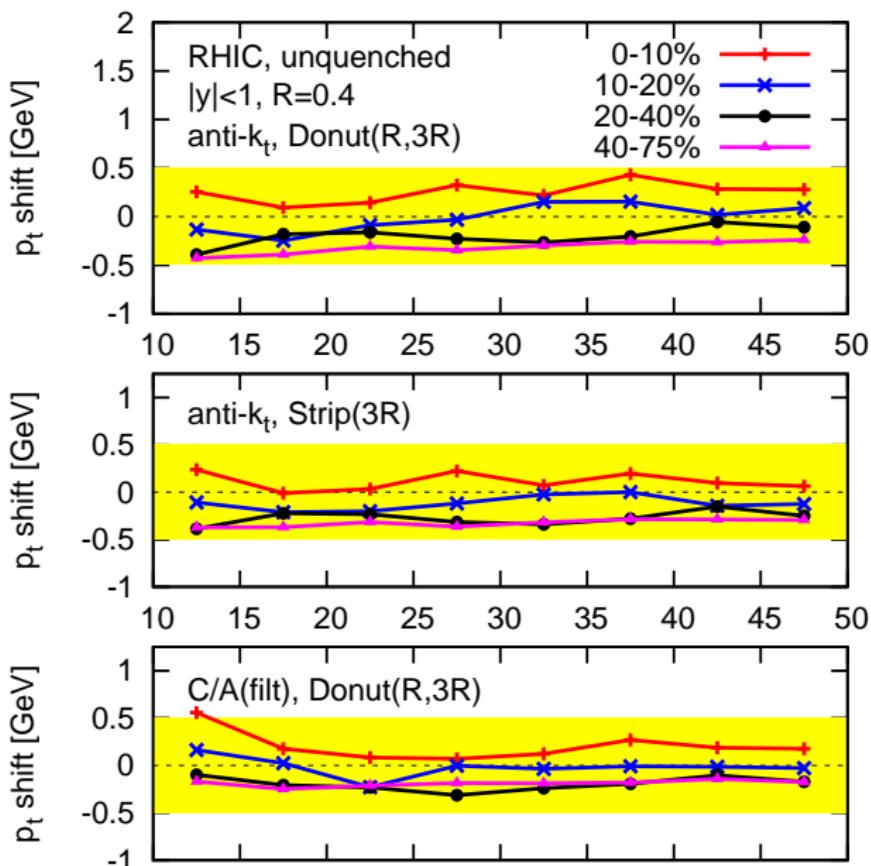
- ▶ Give us an idea of size of corrections semi-independently of Monte Carlo
Some of them are rather large
- ▶ Tell us which approaches are complementary in their systematics
Adding to robustness of experimental measurements, e.g. k_t v. anti- k_t
NB: it's still hard to estimate how quenching affects systematics
- ▶ Guide design of new tools that have smaller systematics
Like filtering, yet to be tried out at RHIC

Important potential for cross-fertilization between ideas in HIC and LHC pp programs.

EXTRAS

Dispersion for non central AuAu



P_t shift for non central AuAu

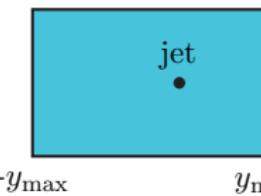
Example #2: another bias

is ρ measured correctly?

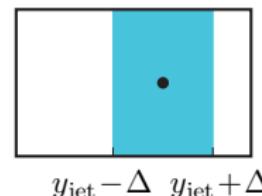
What could go wrong?

- ▶ Rapidity and azimuth dependence of ρ distribution means ρ near jet $\neq \rho$ measured over large region. So try various regions:

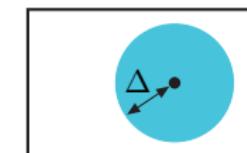
Global



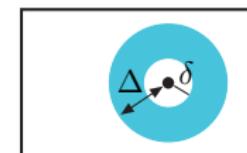
StripRange(Δ)



CircularRange(Δ)



DonutRange(δ, Δ)



- ▶ Median estimate \neq mean contamination. Can be studied in toy models:

$$\rho^{\text{median}} \simeq \rho^{\text{true}} \left(1 - \frac{1}{3\nu R^2} \right)$$

$\nu = \text{number of particles / unit area}$

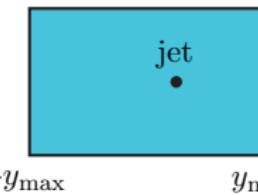
With $\nu = 100$, $R = 0.4$, $\mathcal{O}(2\%) \rightarrow \mathcal{O}(1 \text{ GeV})$ on jet p_t

Cacciari, GPS & Sapeta '09, for measuring $\rho \sim 2 \text{ GeV}$ in pp collisions!

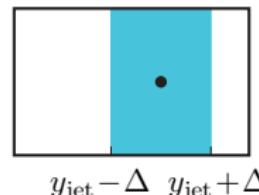
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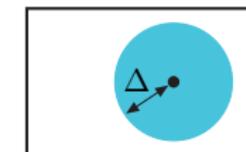
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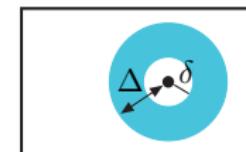
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