

PROTON STRUCTURE

THE LAST LIGHT PARTON

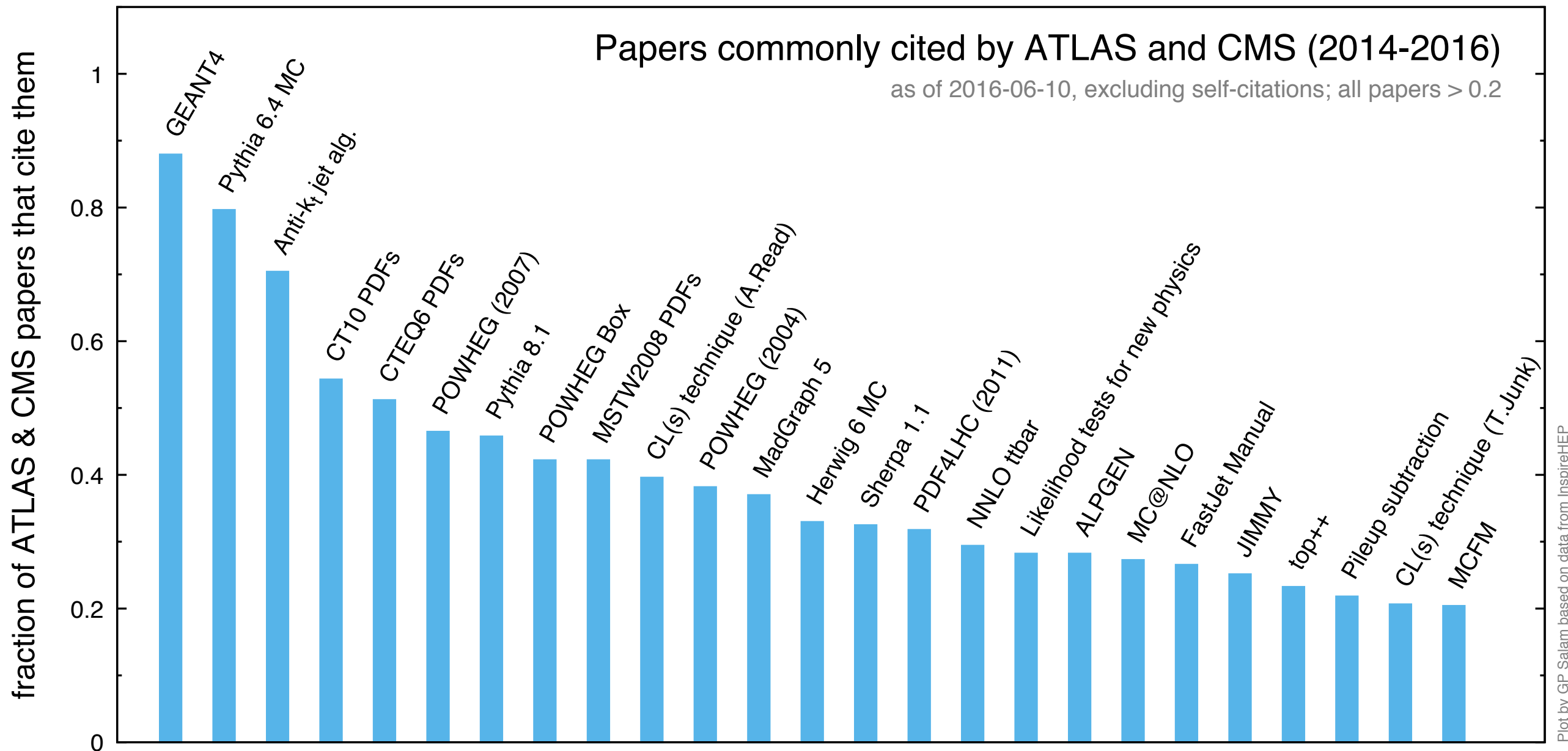
Gavin Salam, CERN

including work with Aneesh Manohar, Paolo Nason and Giulia Zanderighi

Guido Altarelli Memorial Session

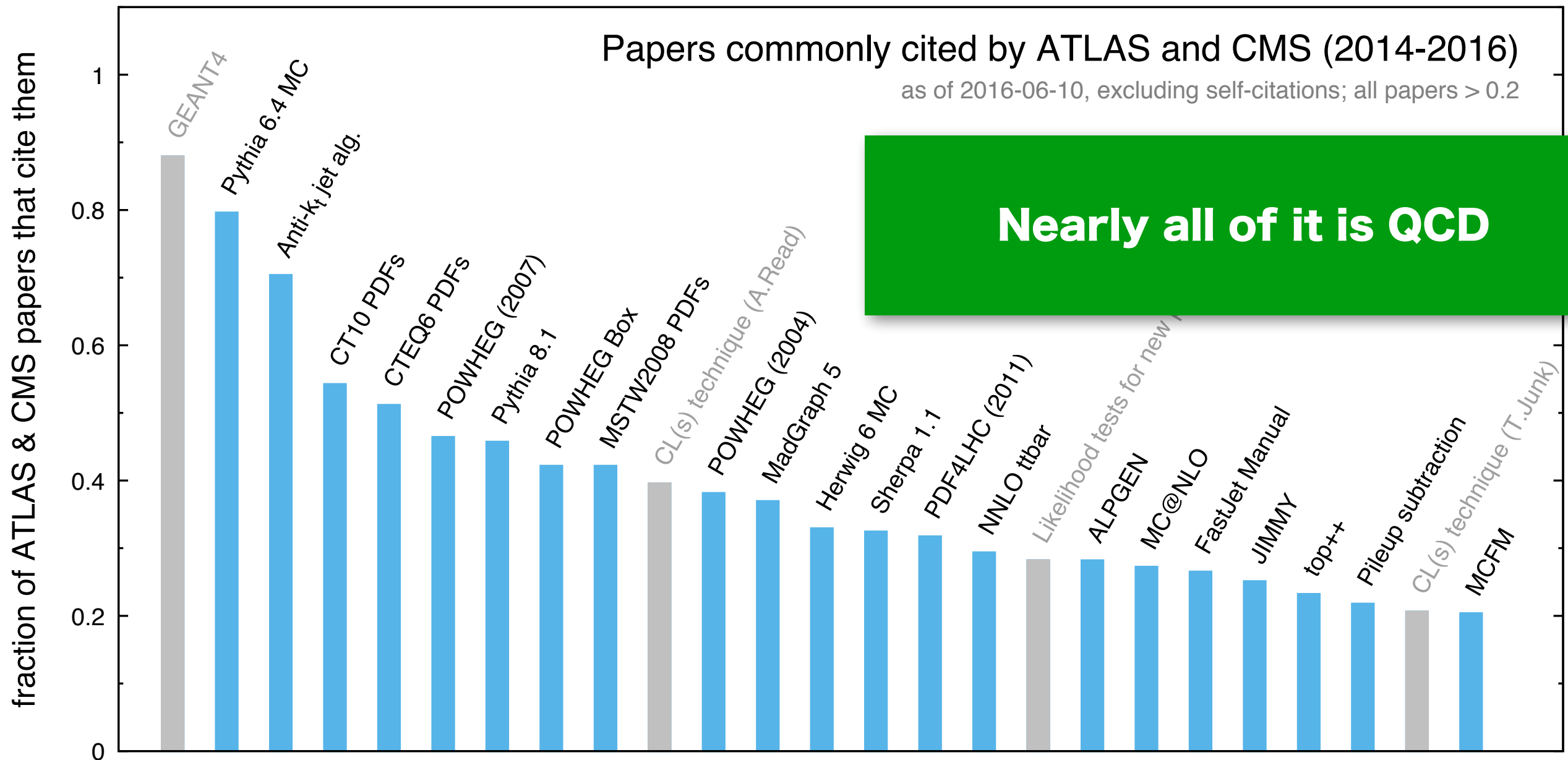
5th International Conference on New Frontiers in Physics, Crete, July 2016

WHAT DO ATLAS & CMS USE MOST FREQUENTLY?

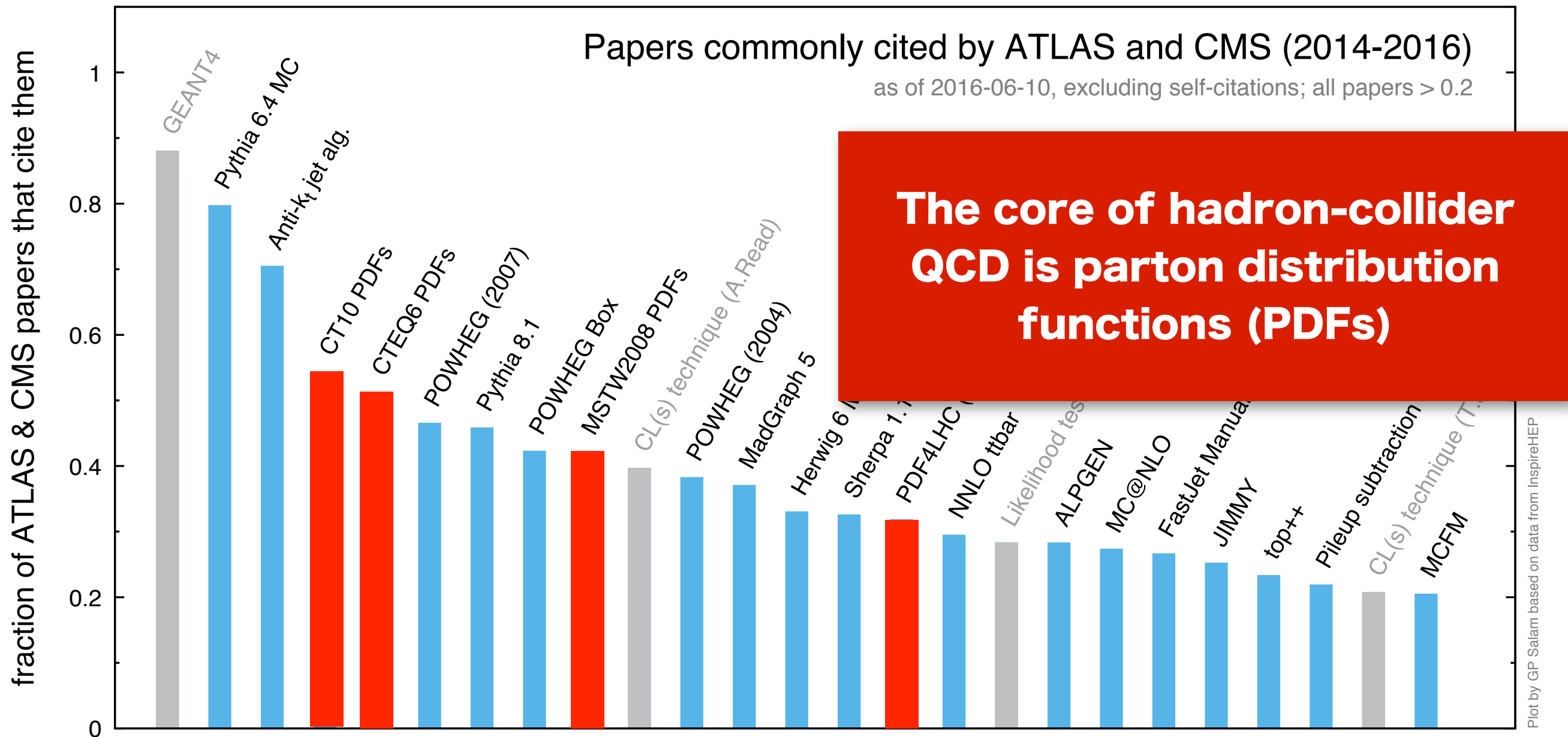


Plot by GP Salam based on data from InspireHEP

WHAT DO ATLAS & CMS USE MOST FREQUENTLY?



WHAT DO ATLAS & CMS USE MOST FREQUENTLY?



Dokshitzer–Gribov–Lipatov–Altarelli–Parisi (DGLAP) equations

ASYMPTOTIC FREEDOM IN PARTON LANGUAGE

G. ALTARELLI *

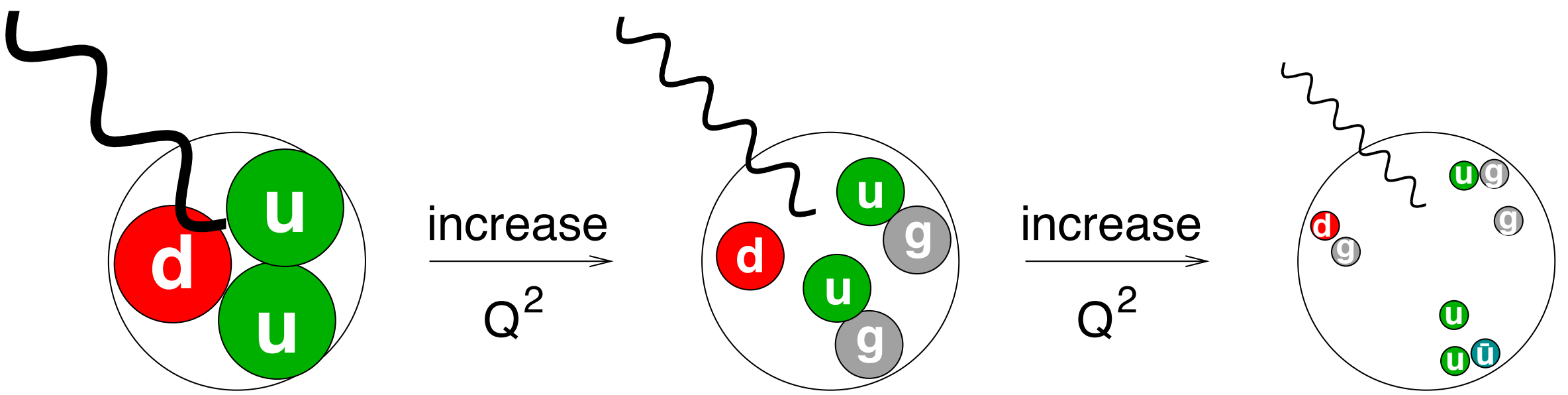
*Laboratoire de Physique Théorique de l'École Normale Supérieure ** , Paris, France*

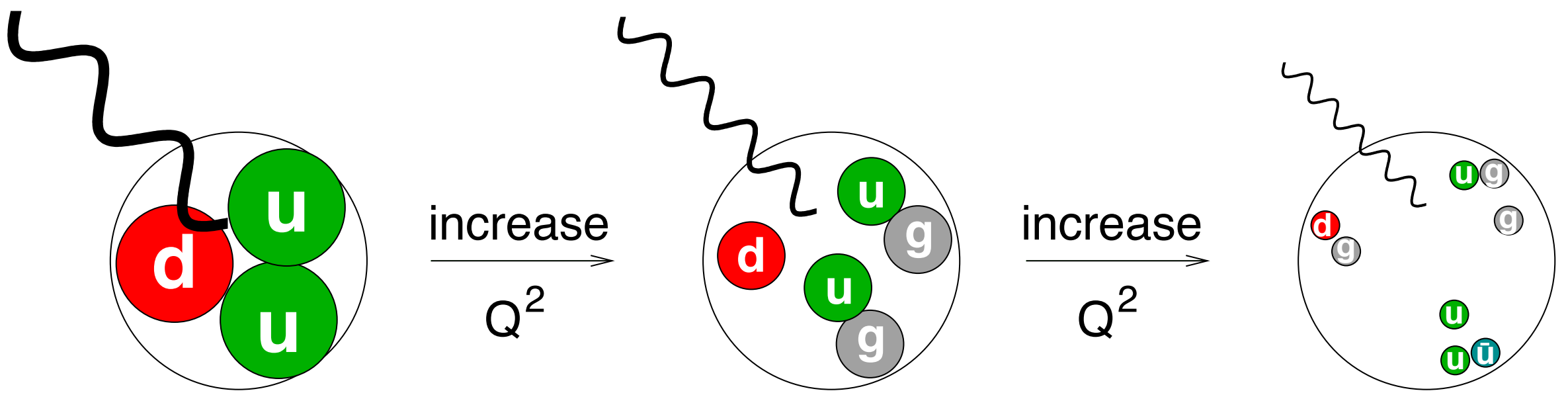
G. PARISI ***

Institut des Hautes Etudes Scientifiques, Bures-sur-Yvette, France

Received 12 April 1977

A novel derivation of the Q^2 dependence of quark and gluon densities (of given helicity) as predicted by quantum chromodynamics is presented. The main body of predictions of the theory for deep-inelastic scattering on either unpolarized or polarized targets is re-obtained by a method which only makes use of the simplest tree diagrams and is entirely phrased in parton language with no reference to the conventional operator formalism.



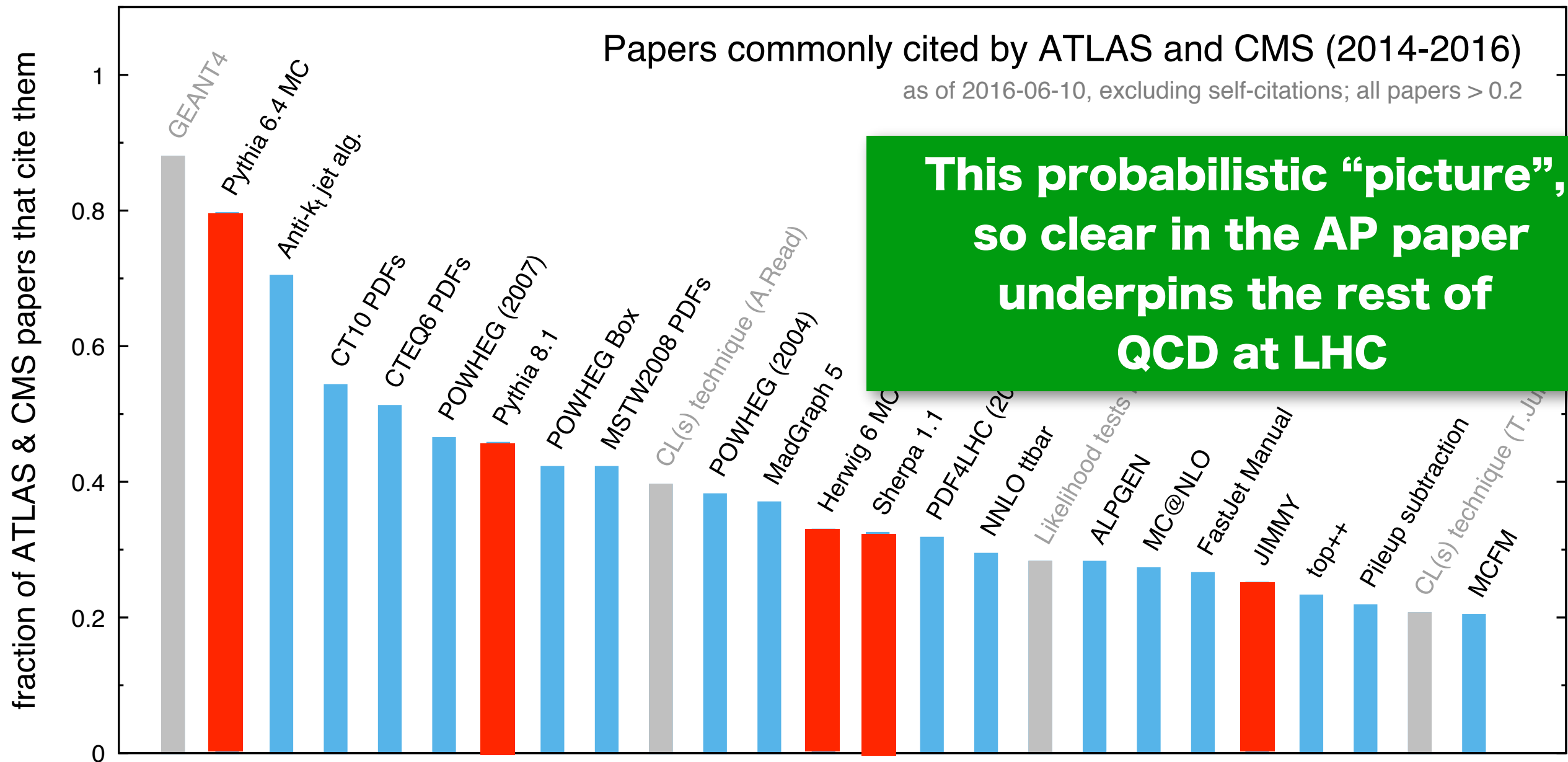


$$\frac{dq^i(x, t)}{dt} = \frac{\alpha(t)}{2\pi} \int_x^1 \frac{dy}{y} \left[\sum_{j=1}^{2f} q^j(y, t) P_{q^i q^j} \left(\frac{x}{y} \right) + G(y, t) P_{q^i G} \left(\frac{x}{y} \right) \right],$$

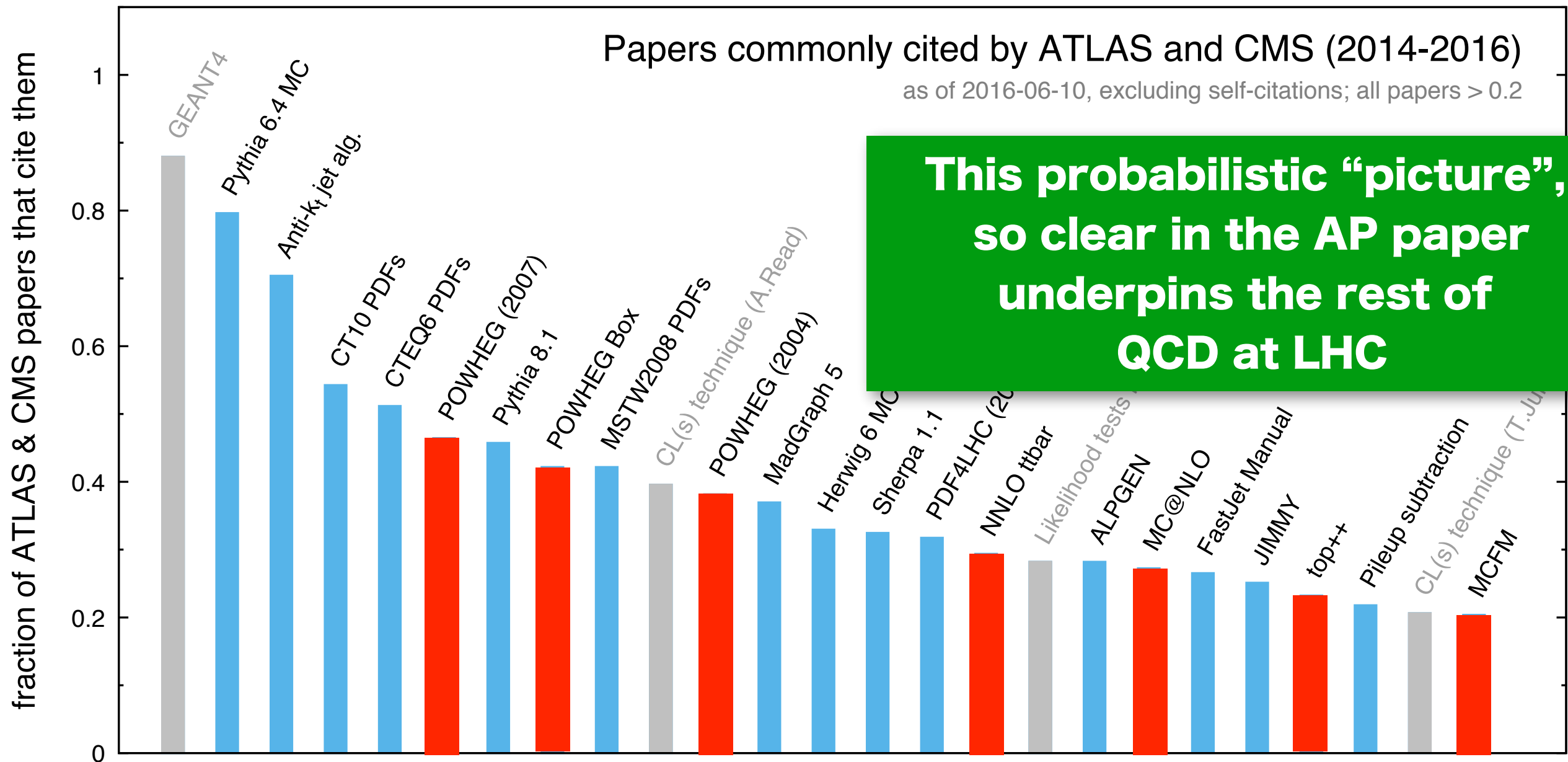
$$\frac{dG(x, t)}{dt} = \frac{\alpha(t)}{2\pi} \int_x^1 \frac{dy}{y} \left[\sum_{j=1}^{2f} q^j(y, t) P_{G q^j} \left(\frac{x}{y} \right) + G(y, t) P_{G G} \left(\frac{x}{y} \right) \right].$$

Quantum mechanics
made probabilistic

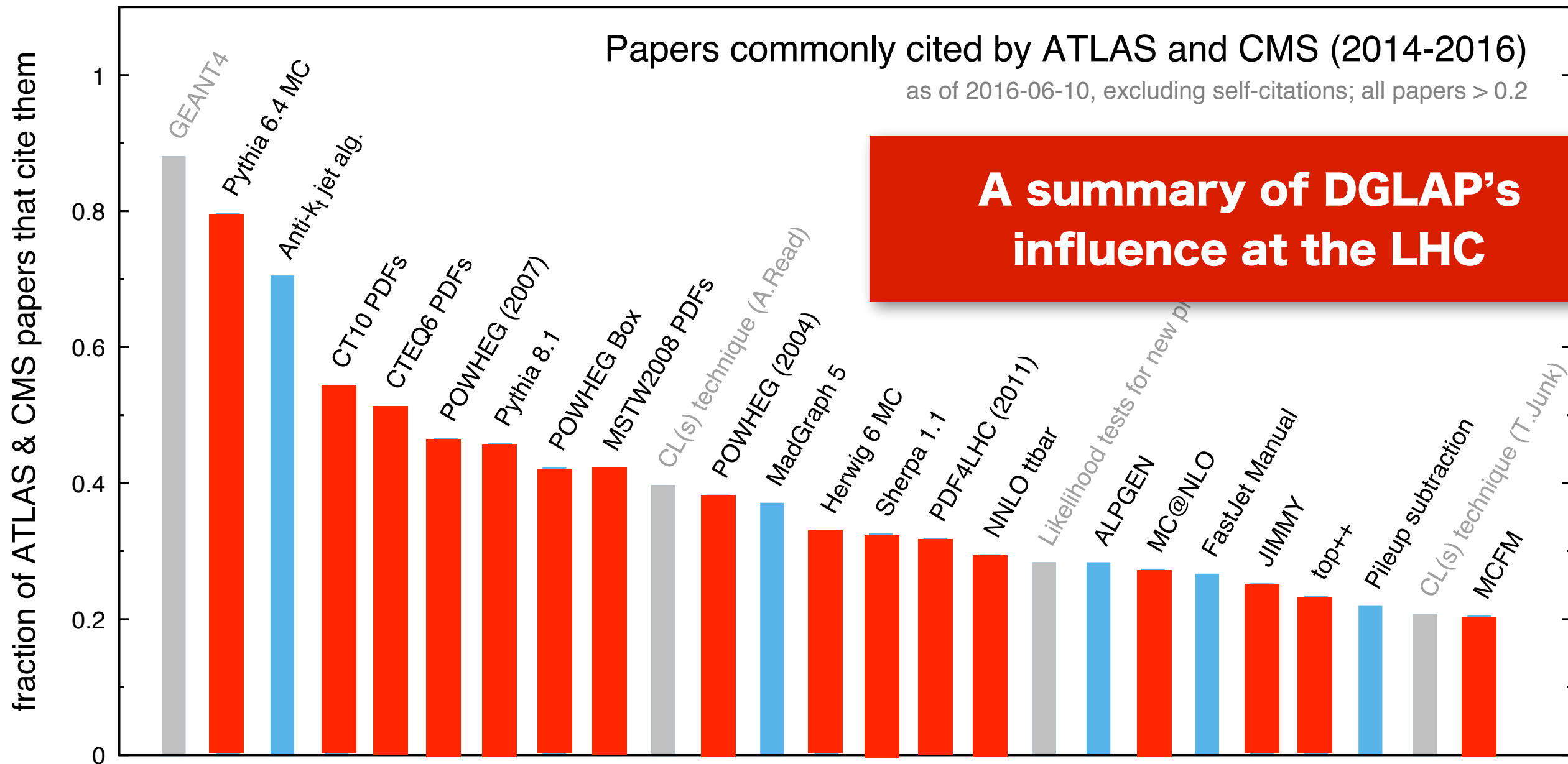
WHAT DO ATLAS & CMS USE MOST FREQUENTLY?



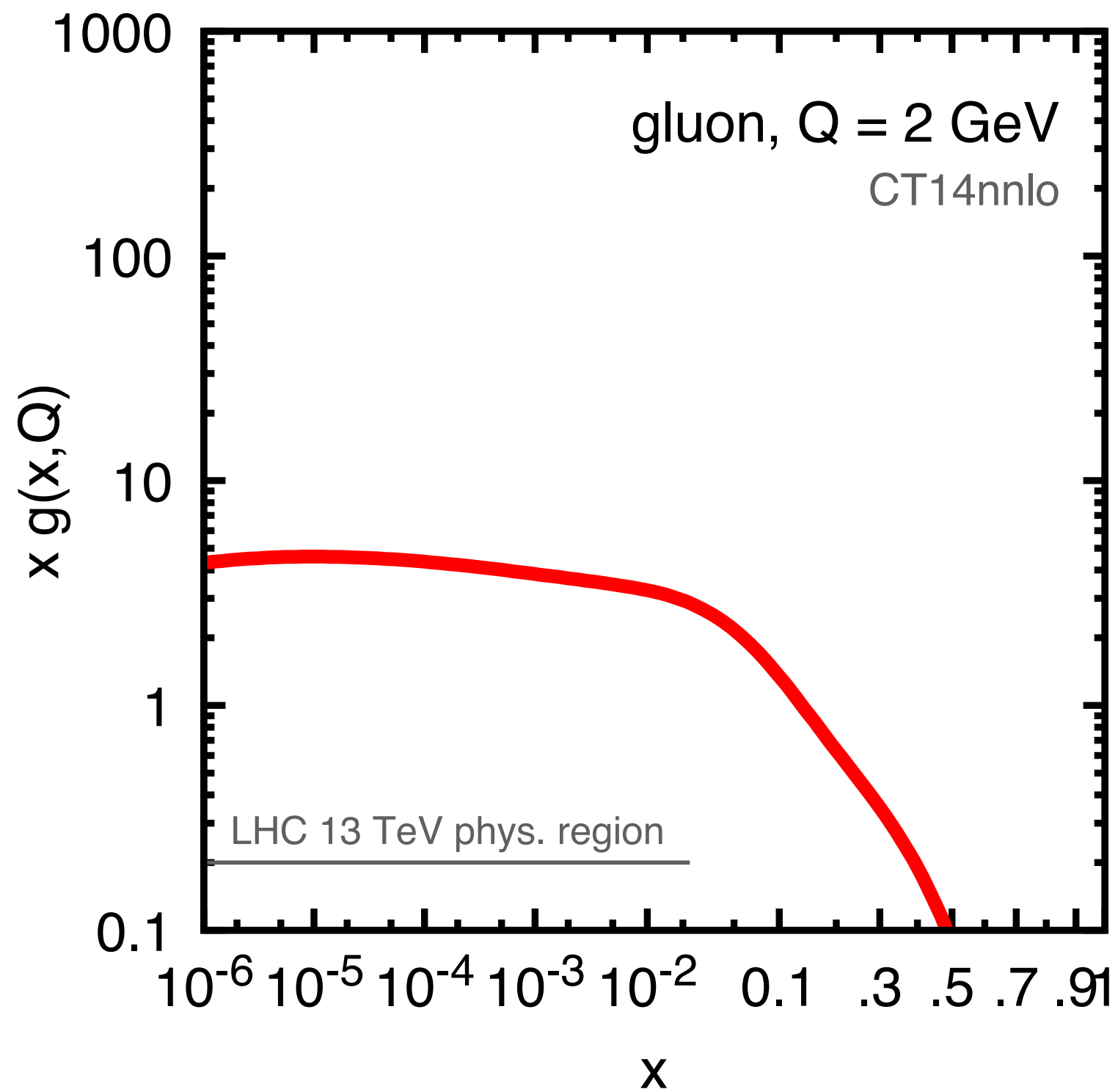
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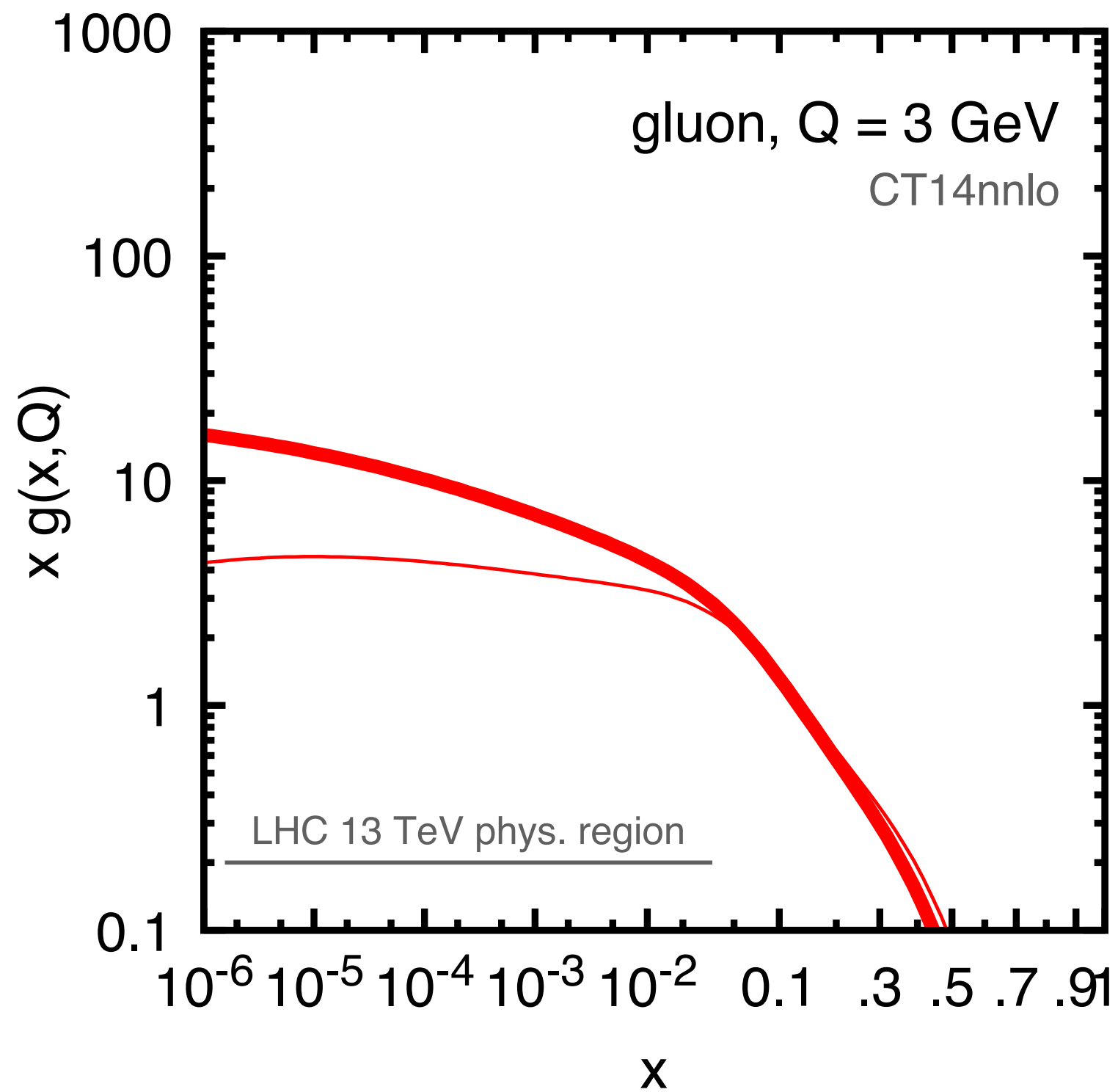
WHAT DO ATLAS & CMS USE MOST FREQUENTLY?



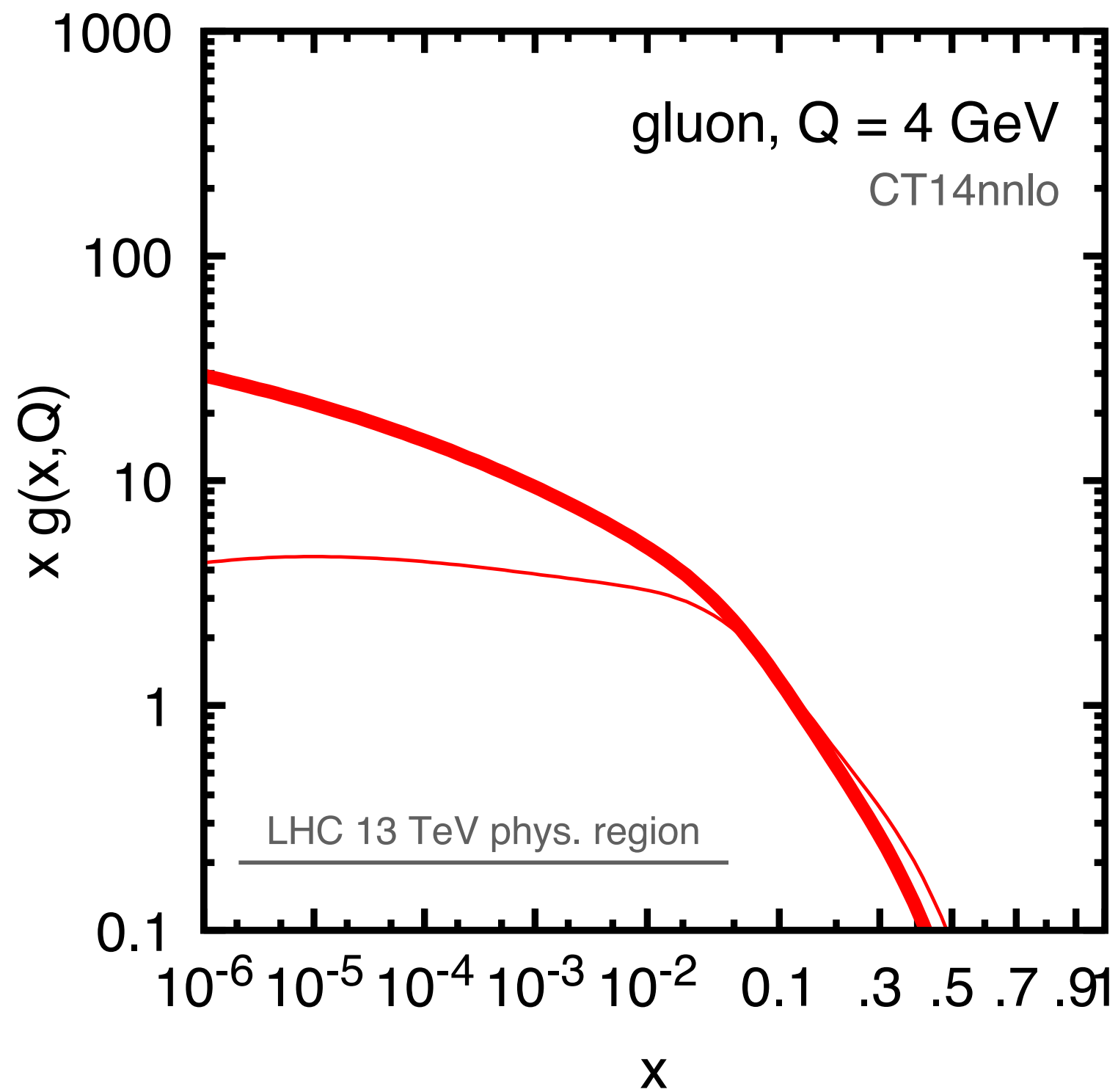
impact of DGLAP evolution from $Q_0 = 2$ GeV



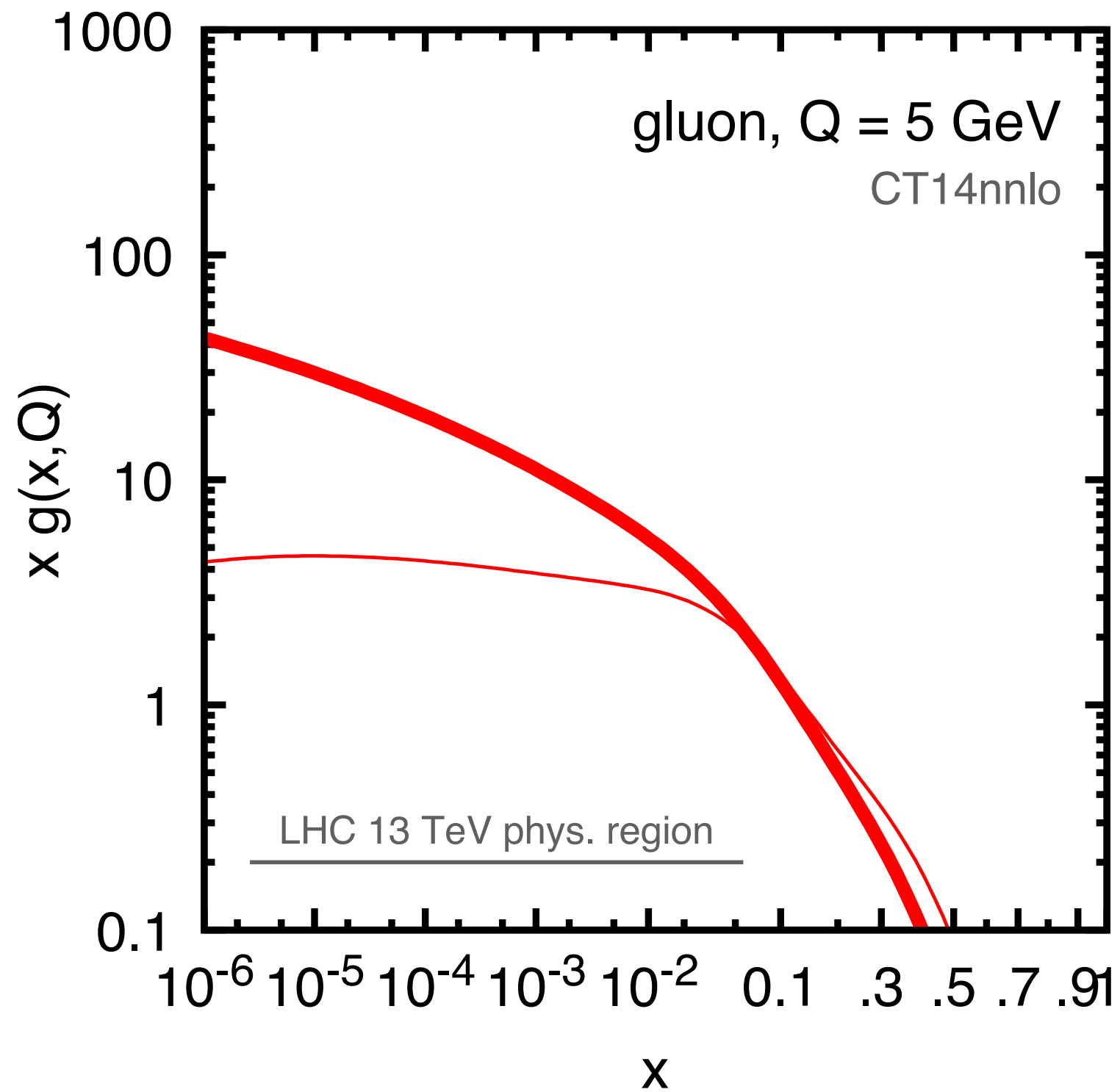
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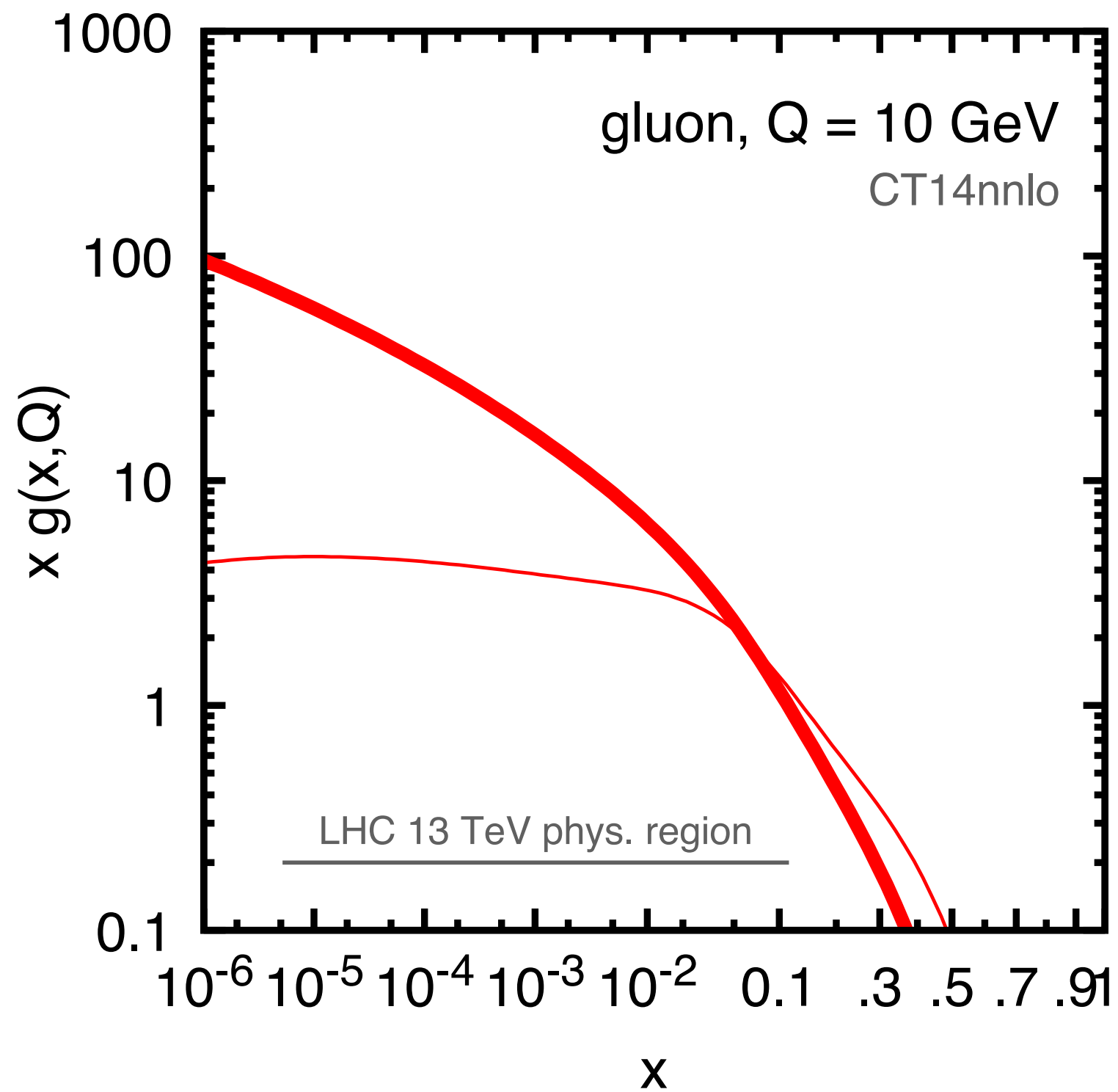
impact of DGLAP evolution from $Q_0 = 2$ GeV



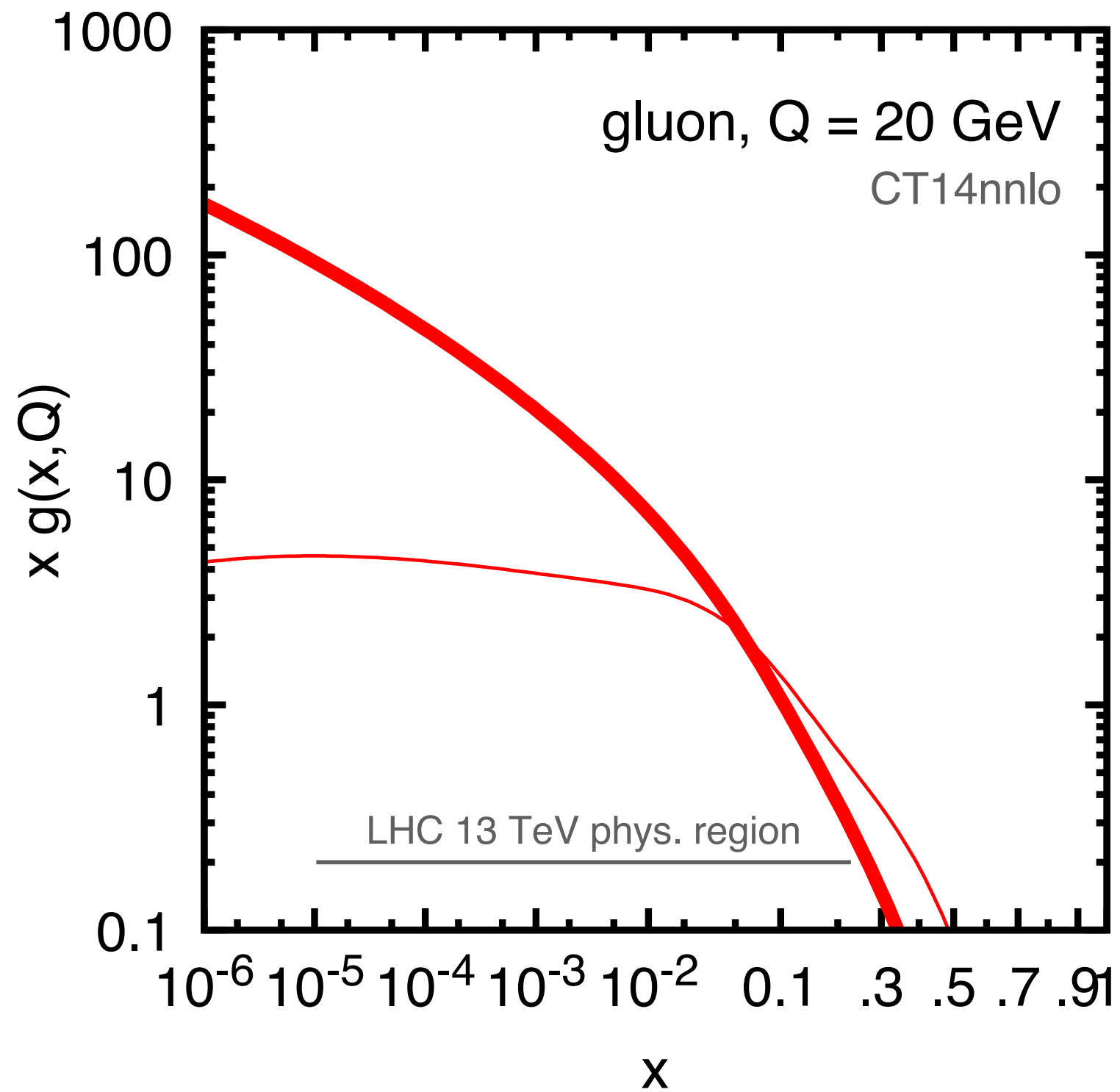
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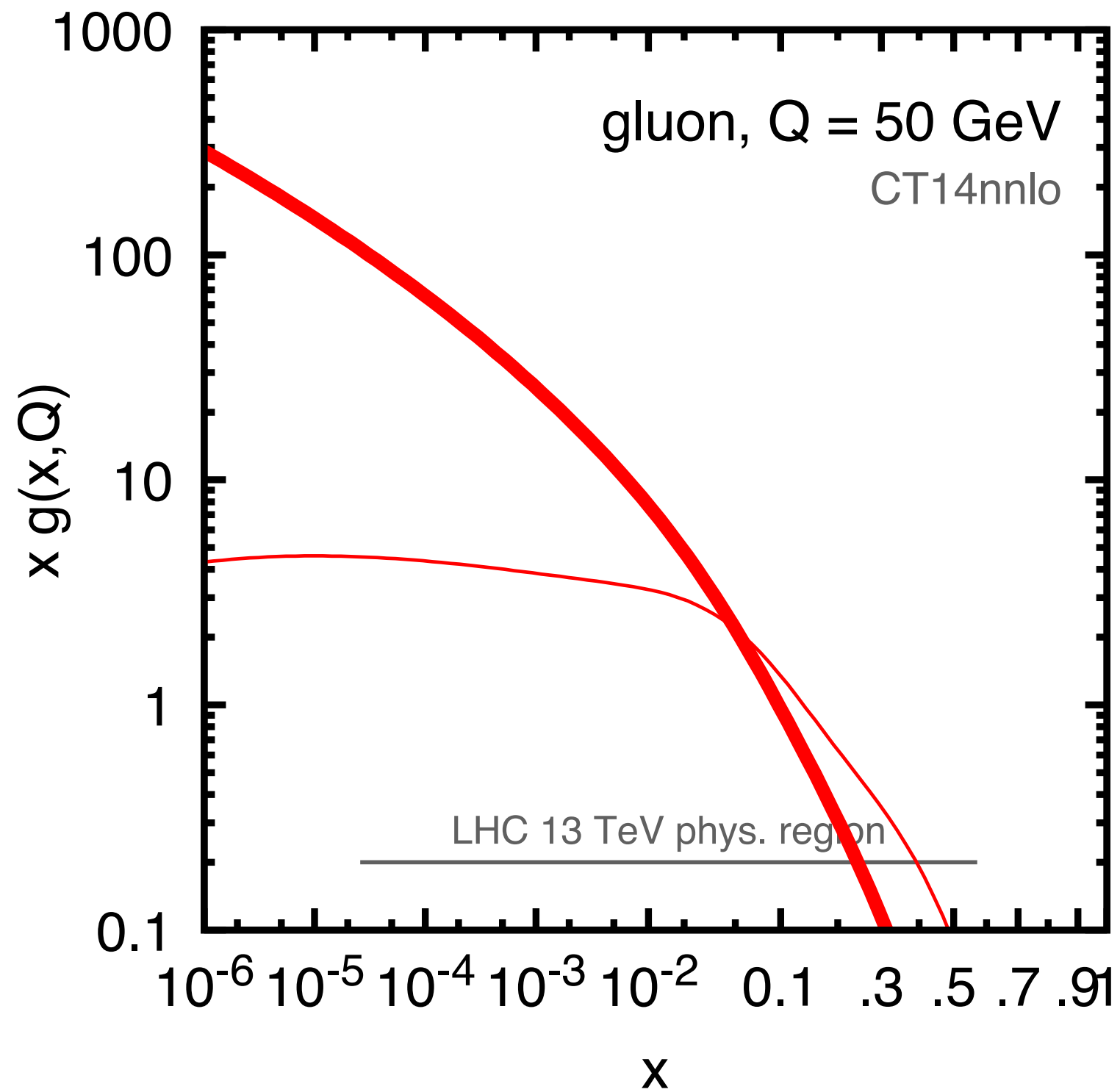
impact of DGLAP evolution from $Q_0 = 2$ GeV



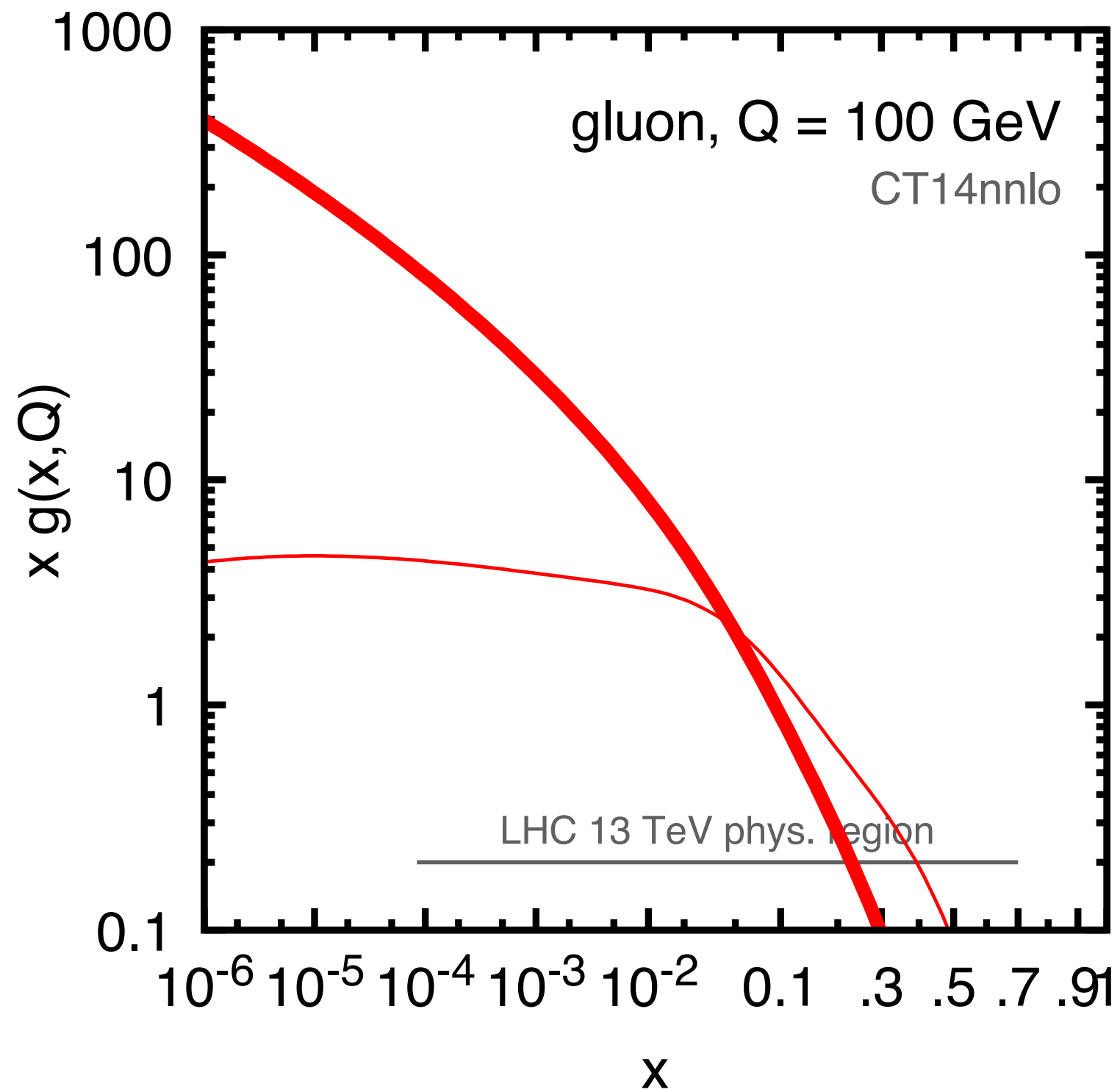
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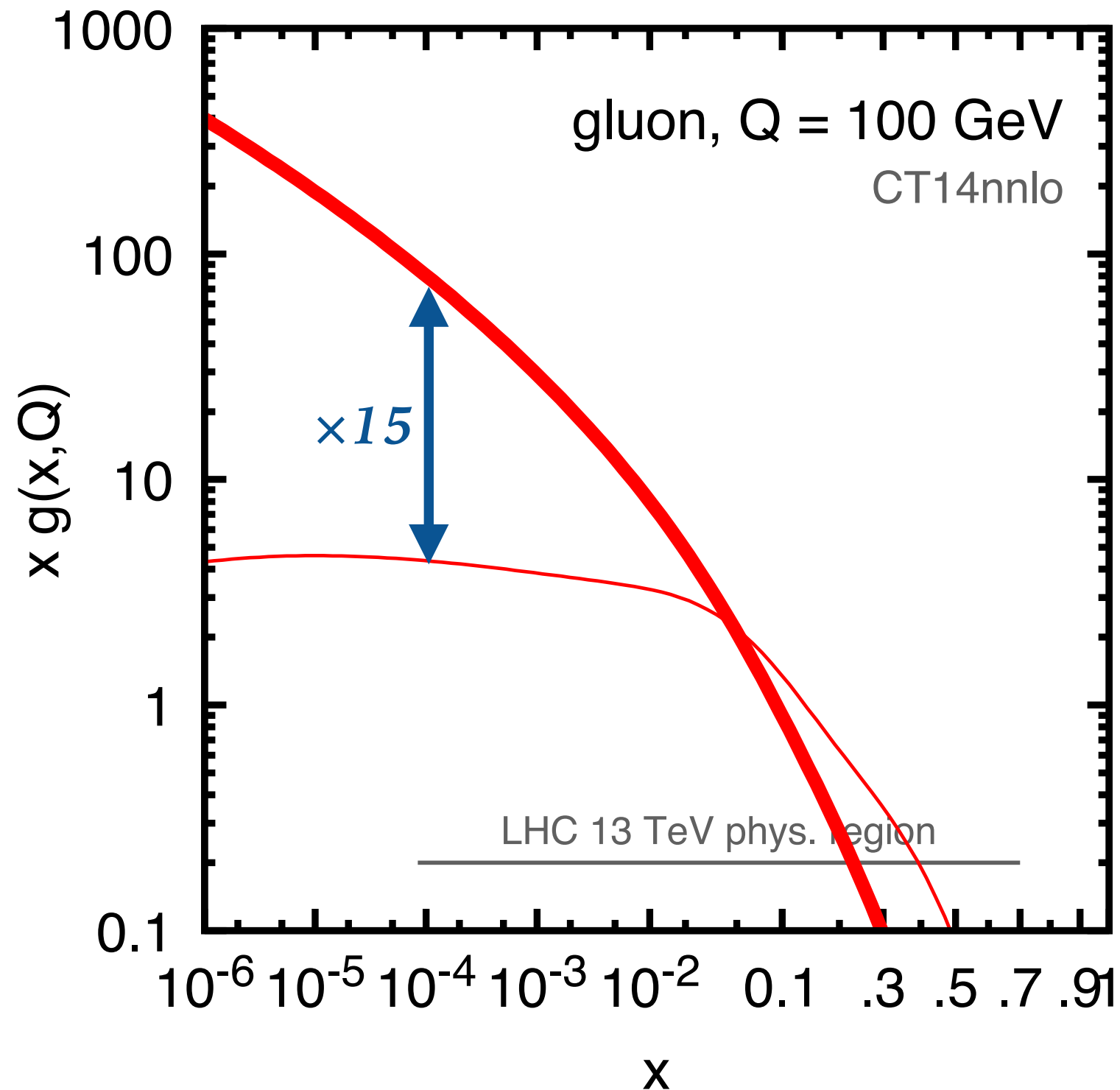
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impact of DGLAP evolution from $Q_0 = 2$ GeV



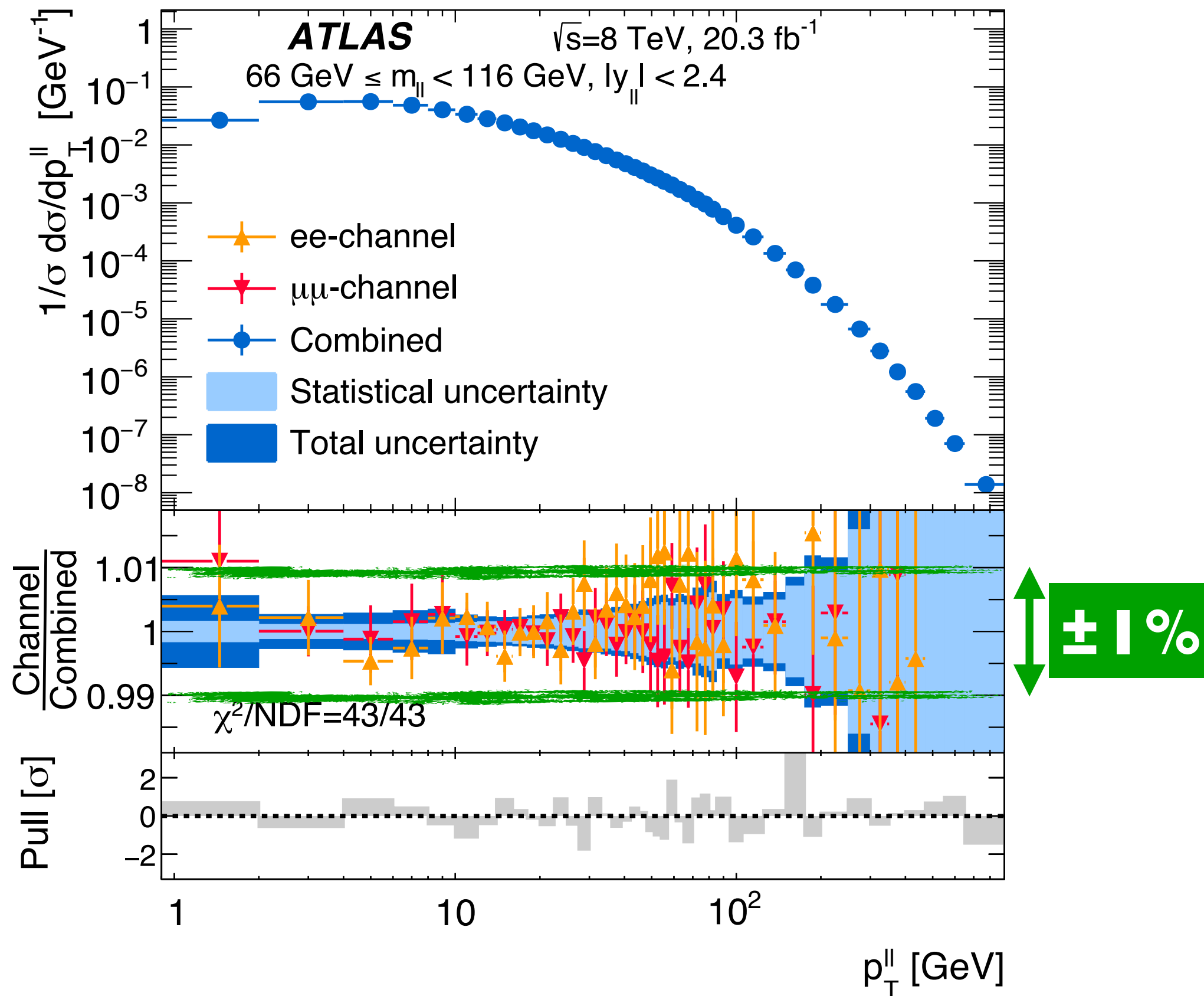
impact of DGLAP evolution from $Q_0 = 2$ GeV



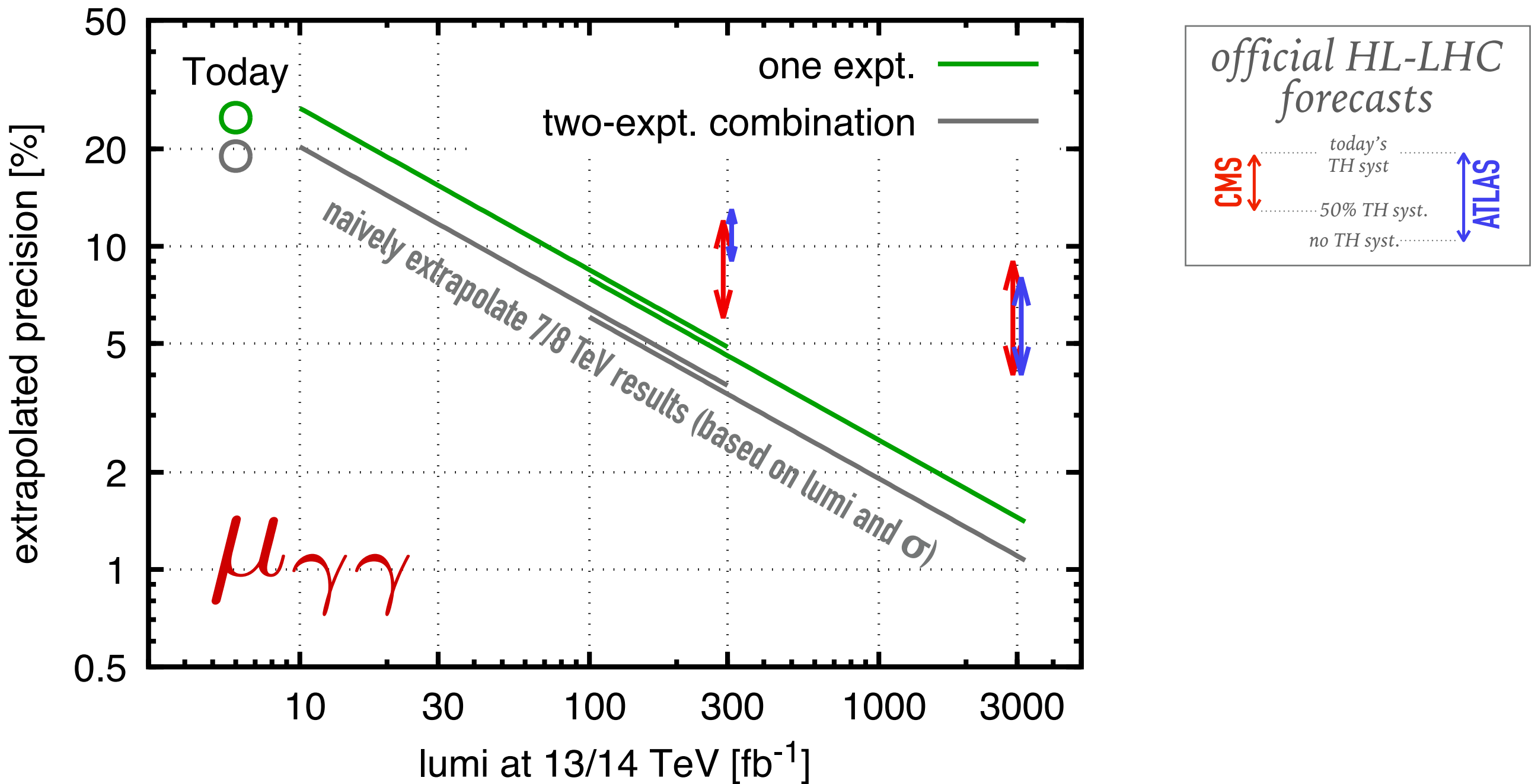
DGLAP
evolution
changes parton
distributions by
factors ~ 10
Higgs cross
section (13 TeV)
would be 6x
smaller without
DGLAP

*nowadays, used at NNLO, thanks
to Moch, Vermaseren & Vogt*

EXPERIMENTAL PRECISION TODAY CAN REACH 1%



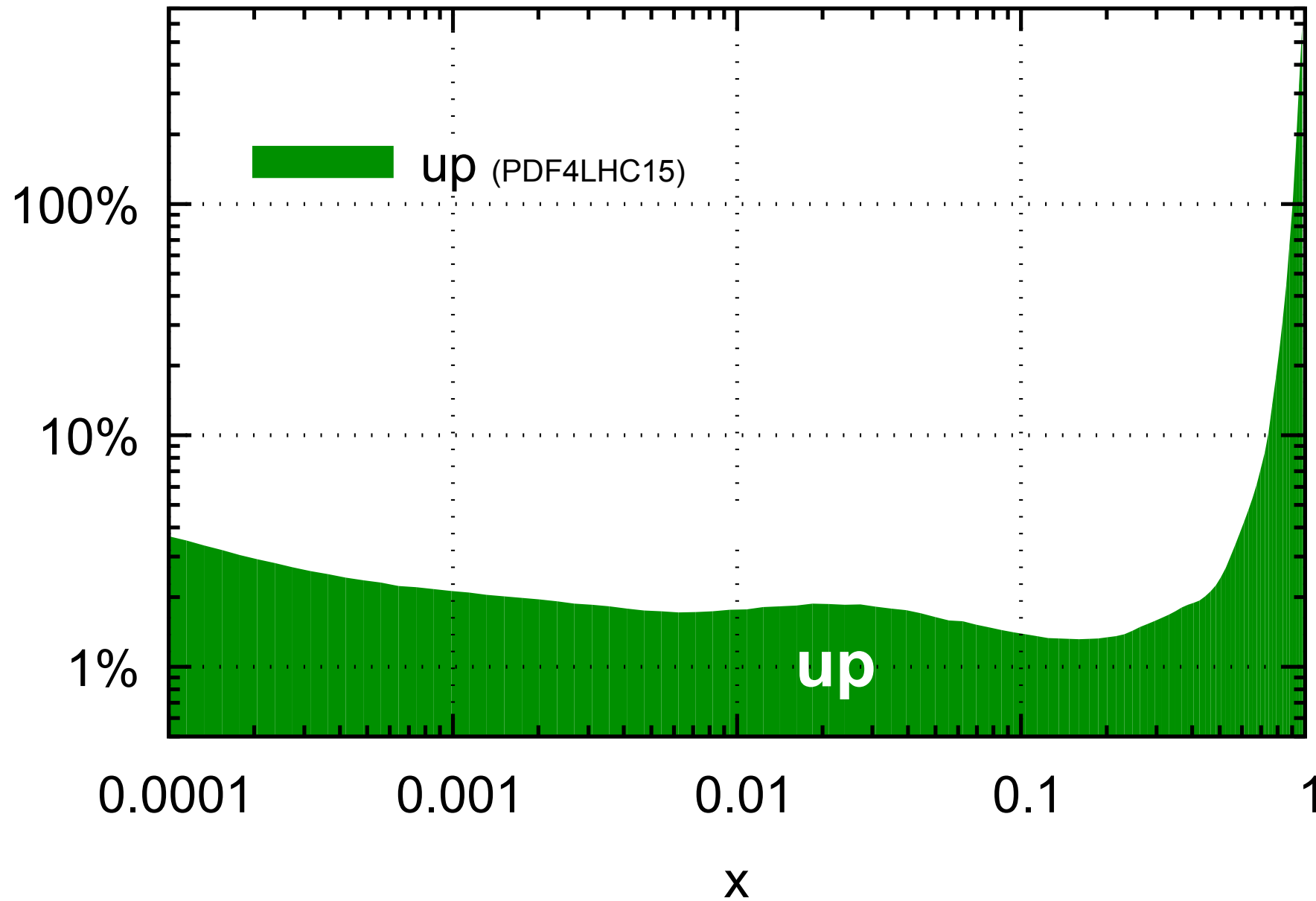
WHAT ACCURACY DO WE NEED? E.G. FOR LONG-TERM HIGGS PRECISION



Naive extrapolation suggests LHC has long-term potential to do Higgs (and much other) physics at **1% accuracy**

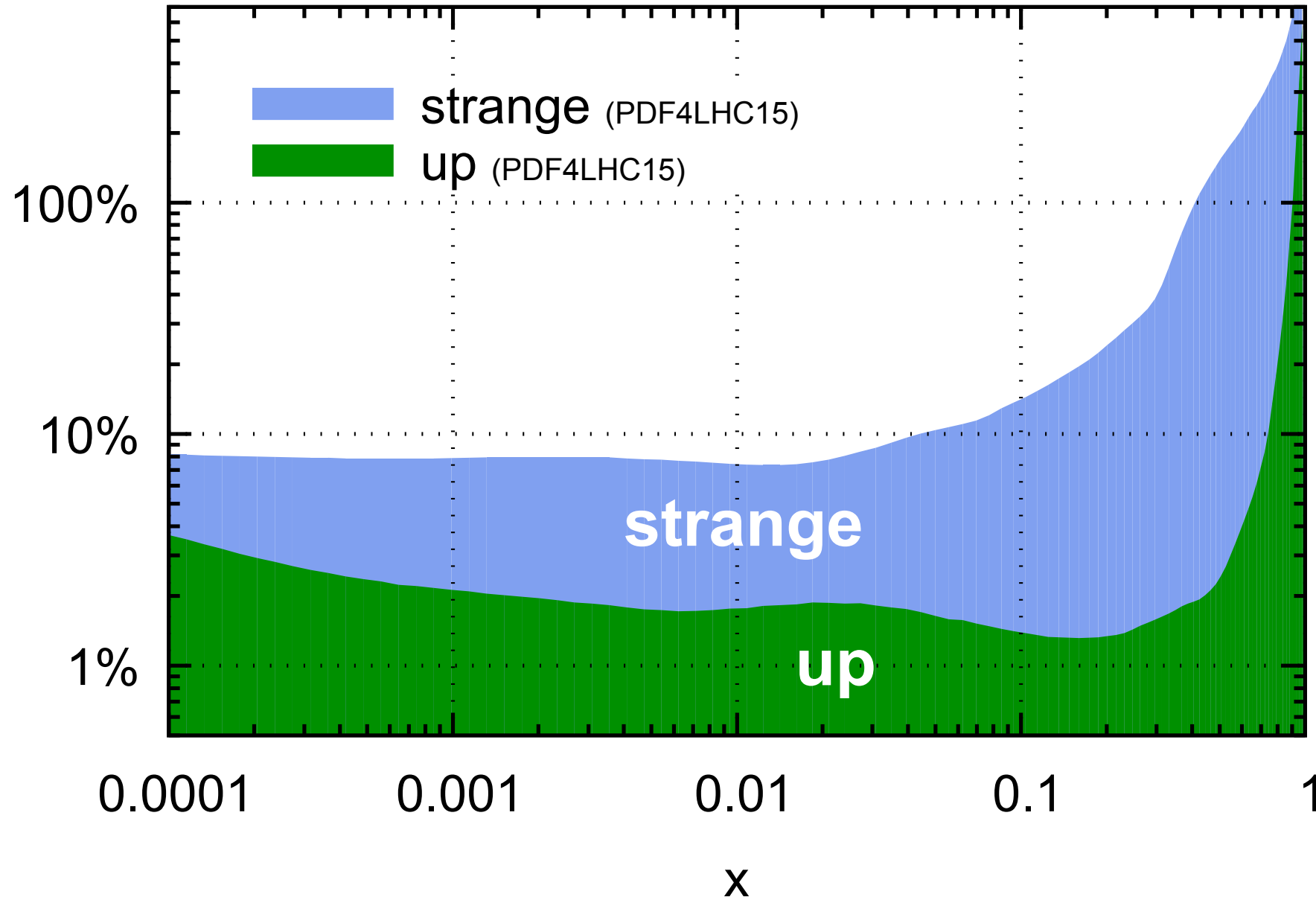
**how well do we know
the parton distributions?**

PDF uncertainties (Q = 100 GeV)



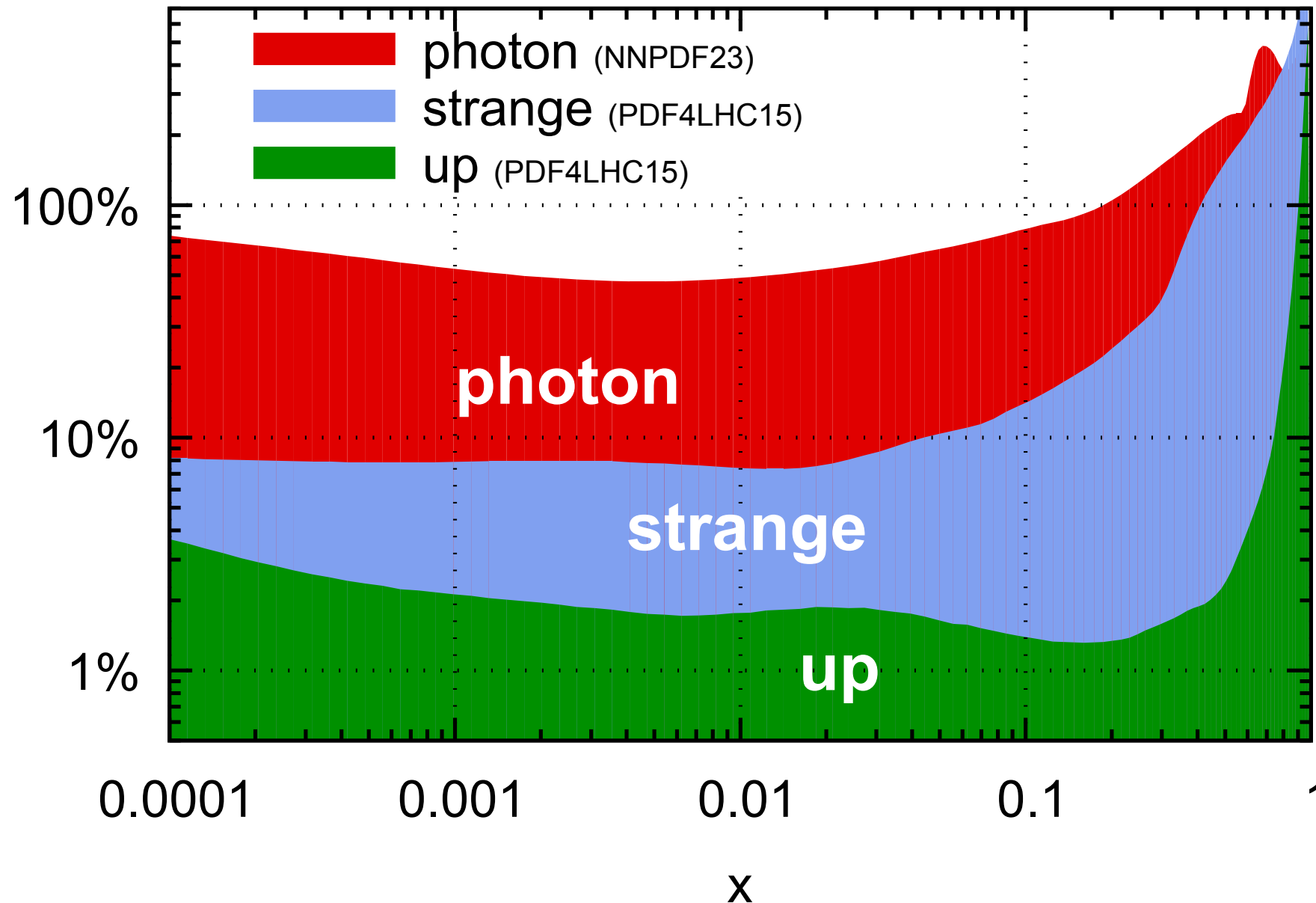
➤ core partons (up, down, gluon) are quite well known

PDF uncertainties (Q = 100 GeV)



- core partons (up, down, gluon) are quite well known ~2%
- strangeness ~10%

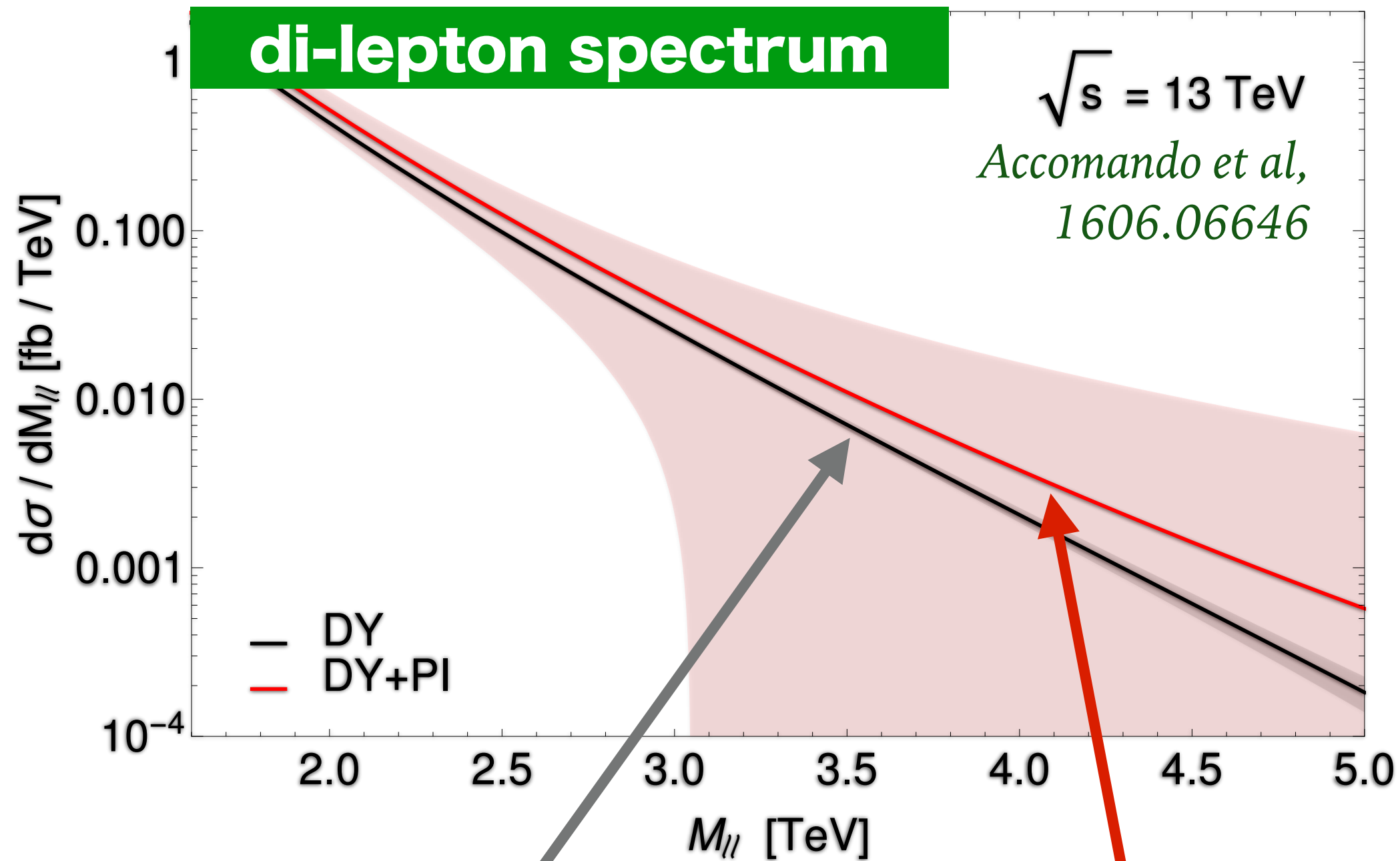
PDF uncertainties (Q = 100 GeV)



- core partons (up, down, gluon) are quite well known $\sim 2\%$
- strangeness $\sim 10\%$

- one other parton, the **photon**, is debated. The only model-independent determination (NNPDF23qed) has **$O(100\%)$ uncertainty**

IT MATTERS FOR DI-LEPTON, DI-BOSON, $T\bar{T}$, EW HIGGS, ETC.



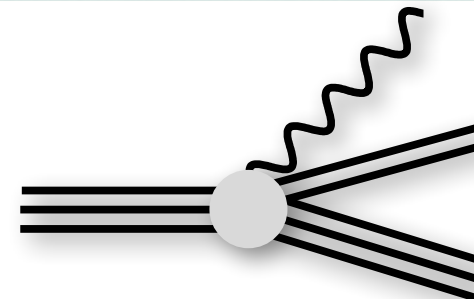
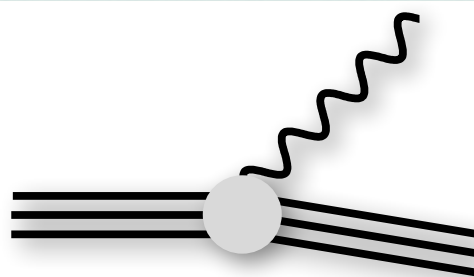
normal DY
contribution

photon-induced contribution
and uncertainty [NNPDF23]

PHOTON PDF ESTIMATES (not exhaustive)

	elastic	inelastic	in LHAPDF?
Gluck Pisano Reya 2002	dipole	model	✗
MRST2004qed	✗	model	✓
NNPDF23qed	no separation; fit to data		✓
CT14qed	✗	model (data-constrained)	✓
CT14qed_inc	dipole	model (data-constrained)	✓
Martin Ryskin 2014	dipole (only electric part)	model	✗
Harland-Lang, Khoze Ryskin 2016	dipole	model	✗

*elastic: Budnev, Ginzburg,
Meledin, Serbo, 1975*



YOU SHOULDN'T NEED A MODEL

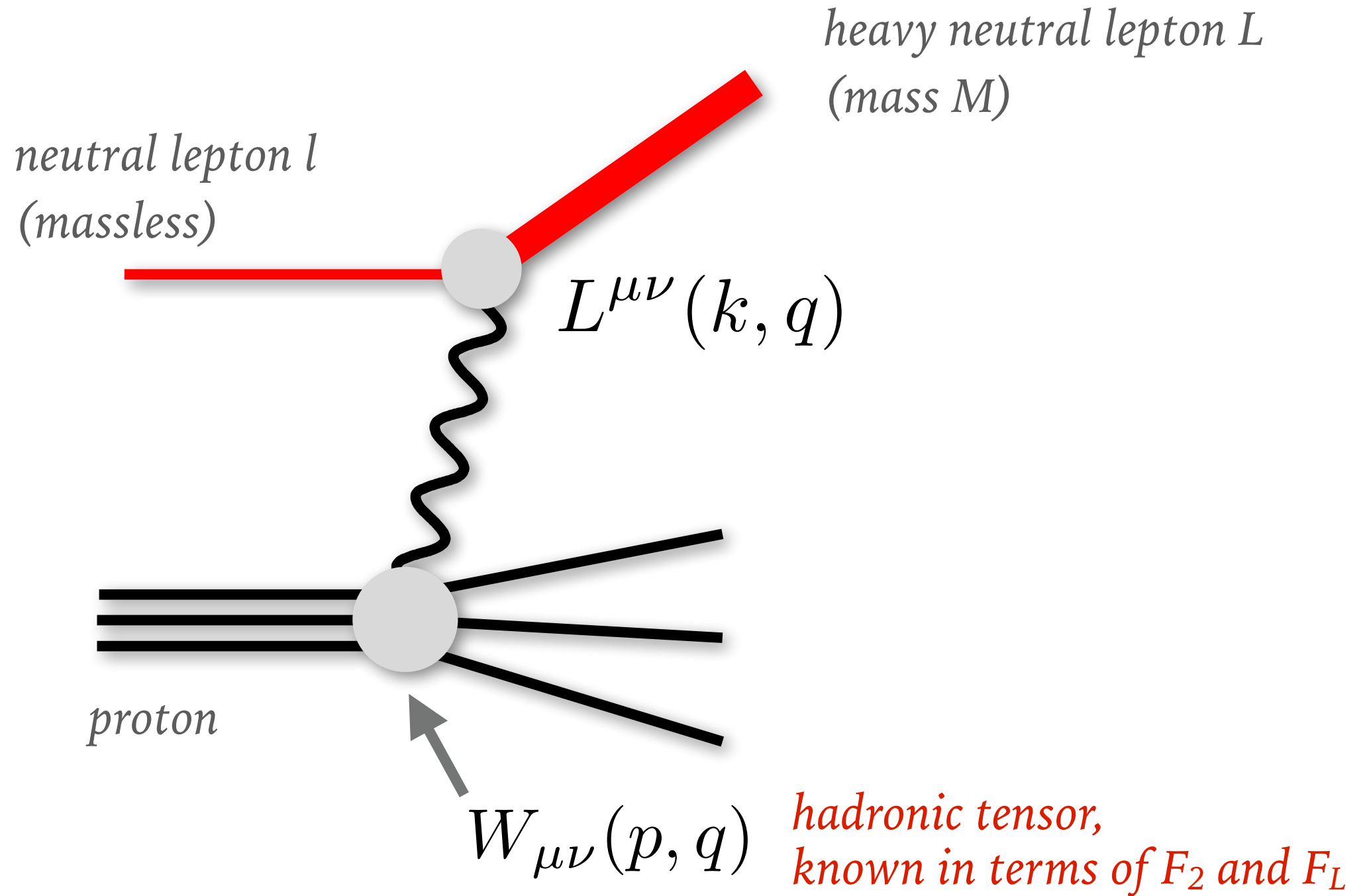
ep scattering (i.e. structure functions) contains all info about proton's EM field

**to extract it, we'll study a hypothetical ("BSM") heavy-neutral lepton
production process**

*Manohar, Nason, GPS
& Zanderighi, to appear*

STEP 1

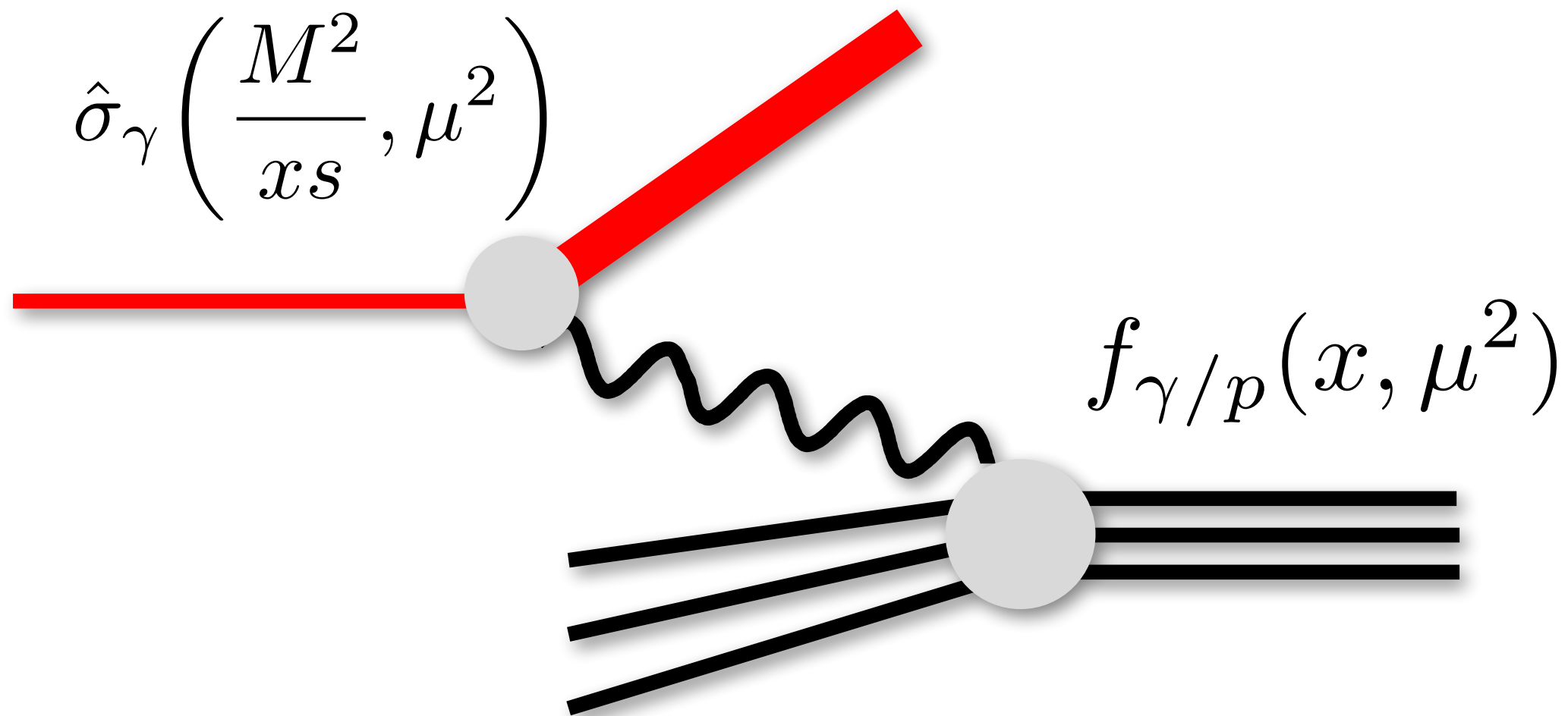
work out a cross section (exact) in terms of F_2 and F_L struct. fns.



$$\sigma = \frac{1}{4p \cdot k} \int \frac{d^4 q}{(2\pi)^4 q^4} e_{\text{ph}}^2(q^2) [4\pi W_{\mu\nu} L^{\mu\nu}(k, q)] \times 2\pi \delta((k - q)^2 - M^2)$$

STEP 2

work out same cross section in terms of a photon distribution



$$\sigma = c_0 \sum_a \int \frac{dx}{x} \hat{\sigma}_a \left(\frac{M^2}{xS}, \mu^2 \right) x f_{a/p}(x, \mu^2)$$

STEP 3

equate them to deduce the photon distribution (LUXqed)

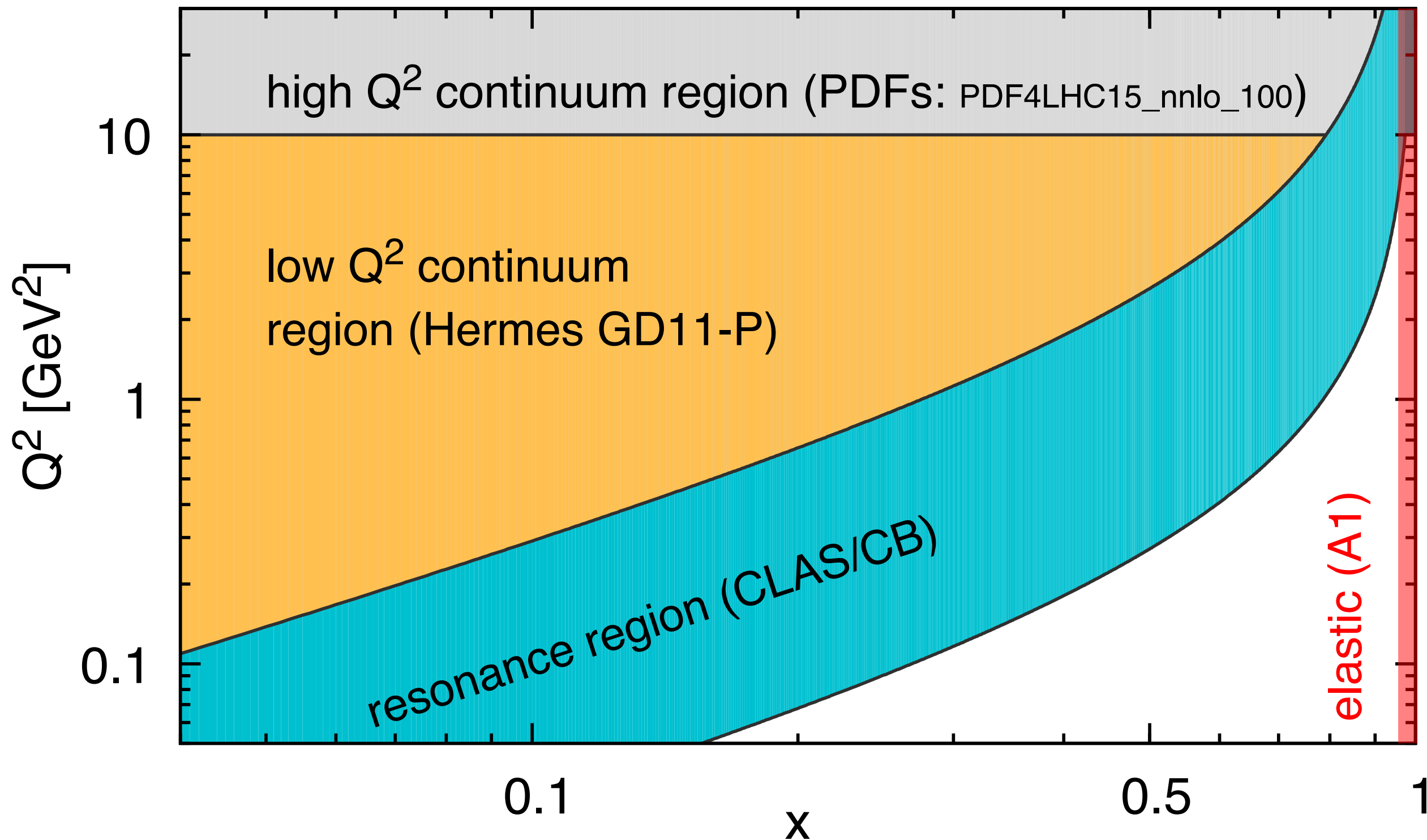
$$x f_{\gamma/p}(x, \mu^2) = \frac{1}{2\pi\alpha(\mu^2)} \int_x^1 \frac{dz}{z} \left\{ \int_{Q_{\min}^2}^{Q_{\max}^2} \frac{dQ^2}{Q^2} \alpha^2(Q^2) \left[\left(2 - 2z + z^2 + \frac{2x^2 m_p^2}{Q^2} \right) F_2(x/z, Q^2) - z^2 F_L\left(\frac{x}{z}, Q^2\right) \right] - \alpha^2(\mu^2) z^2 F_2\left(\frac{x}{z}, \mu^2\right) \right\},$$

Result is in MSbar scheme & consistent with
2015 de Florian, Rodrigo, Sborlini $O(\alpha\alpha_s)$ $P_{\gamma X}$ QED split.fns.

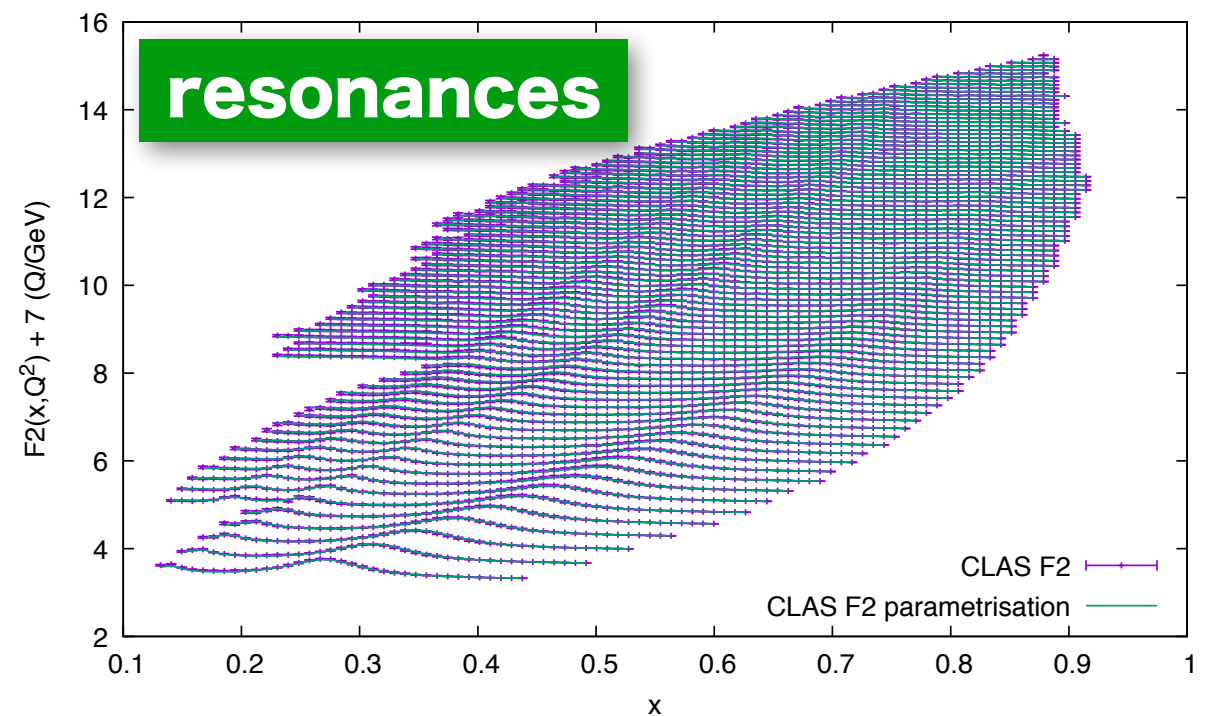
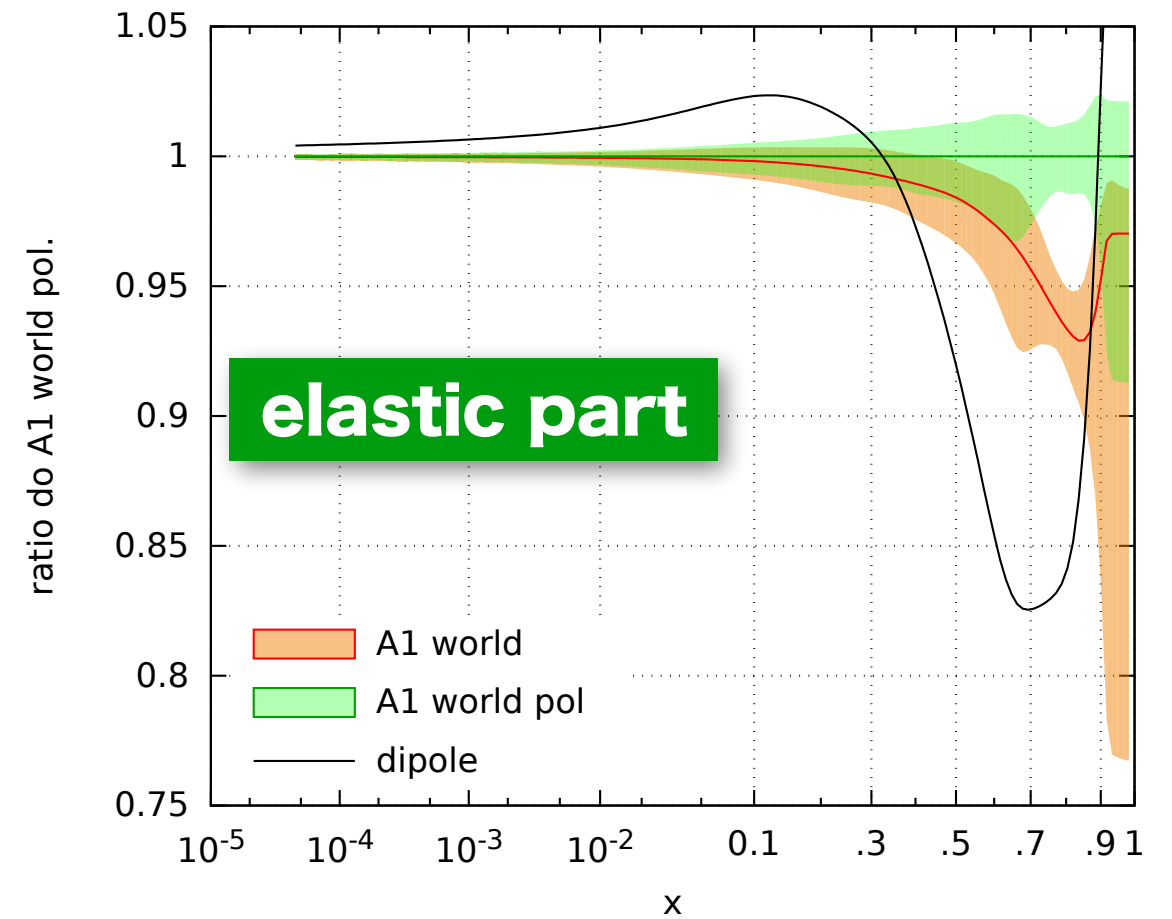
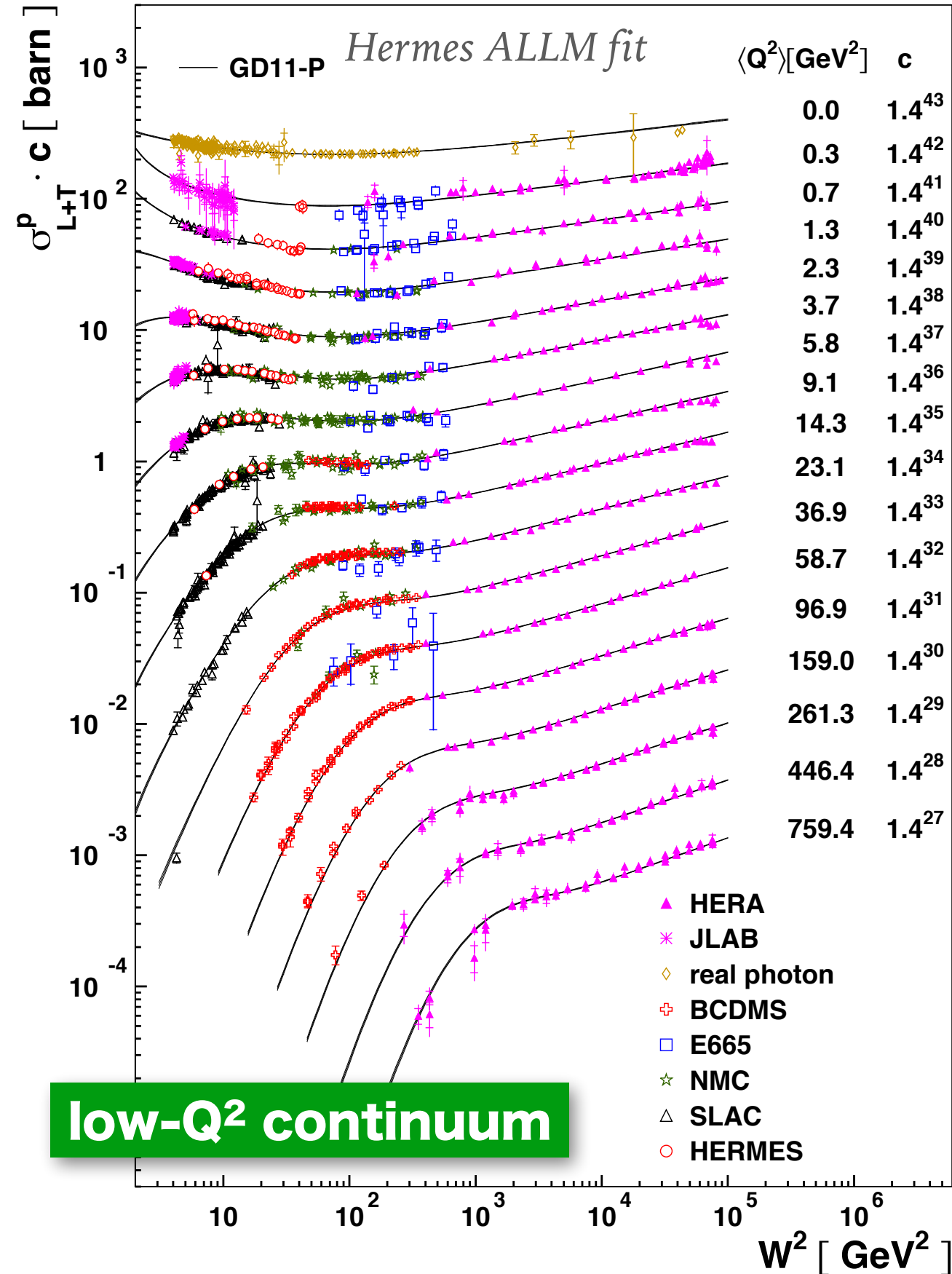
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LUXqed 2016	data	data	soon

DATA SOURCES – various fits to F2, FL & elastic form factors

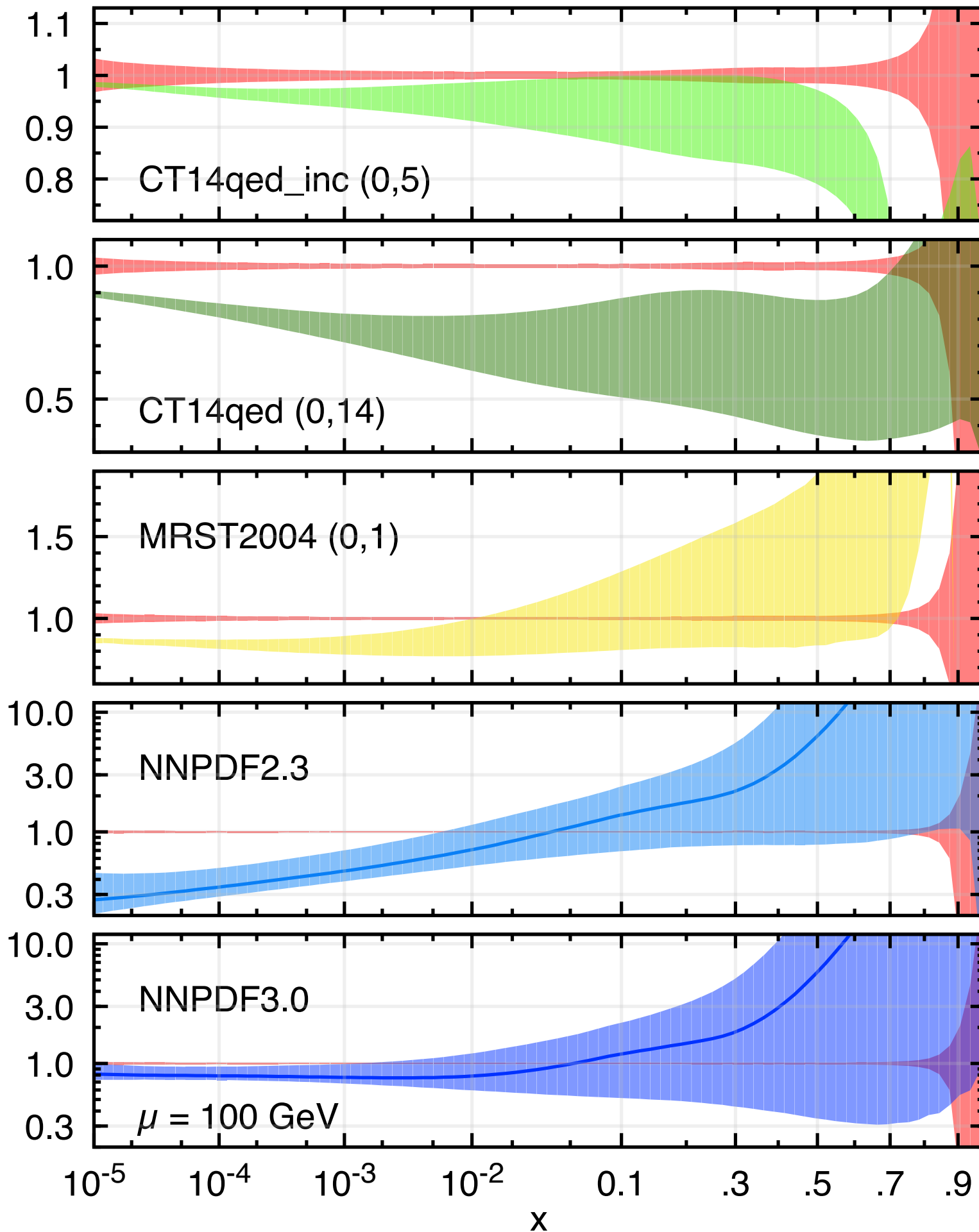


DATA SOURCES – various fits to F2, FL & elastic form factors



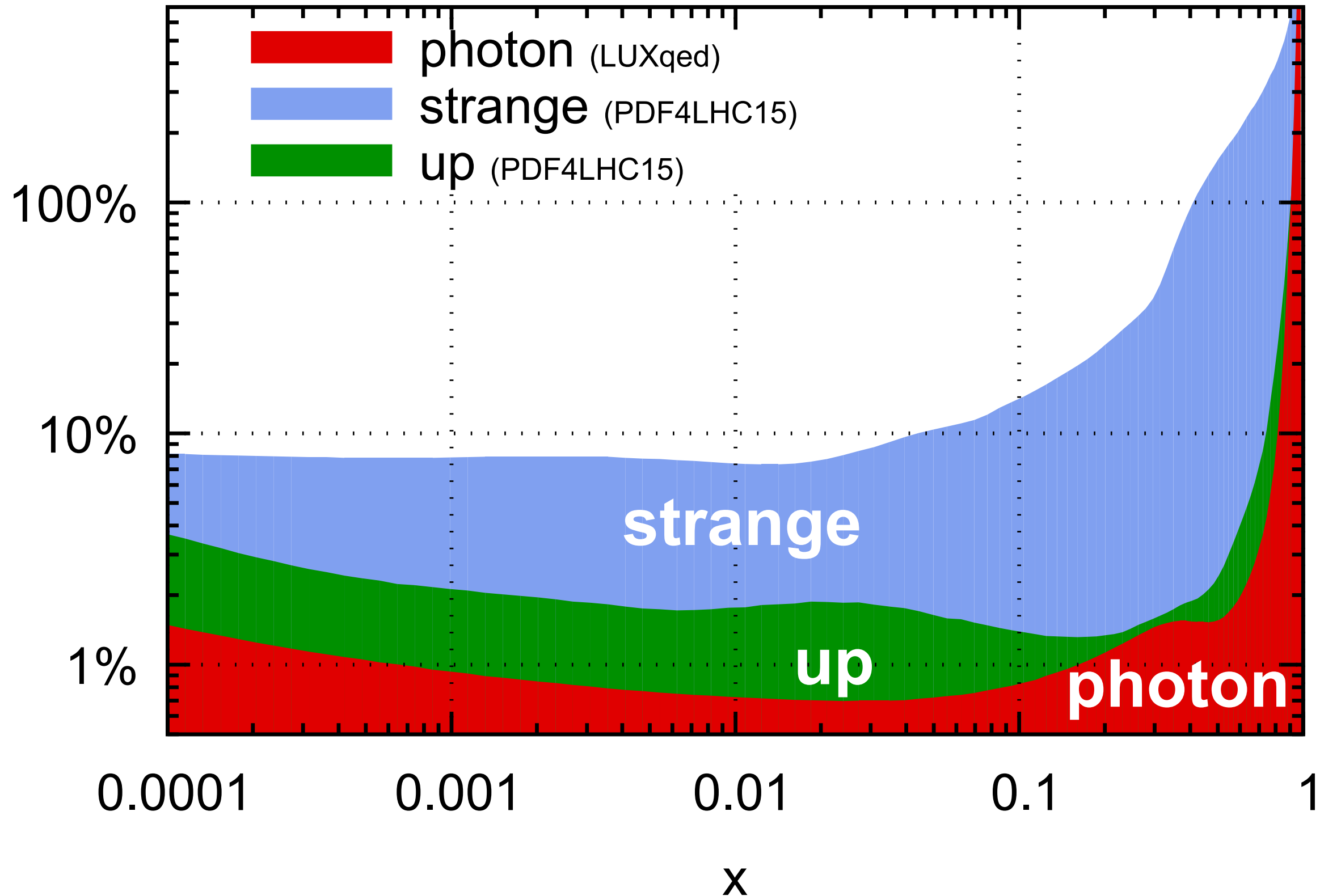
the results

ratio of some
widely used
PDFs to
LUXqed (red)



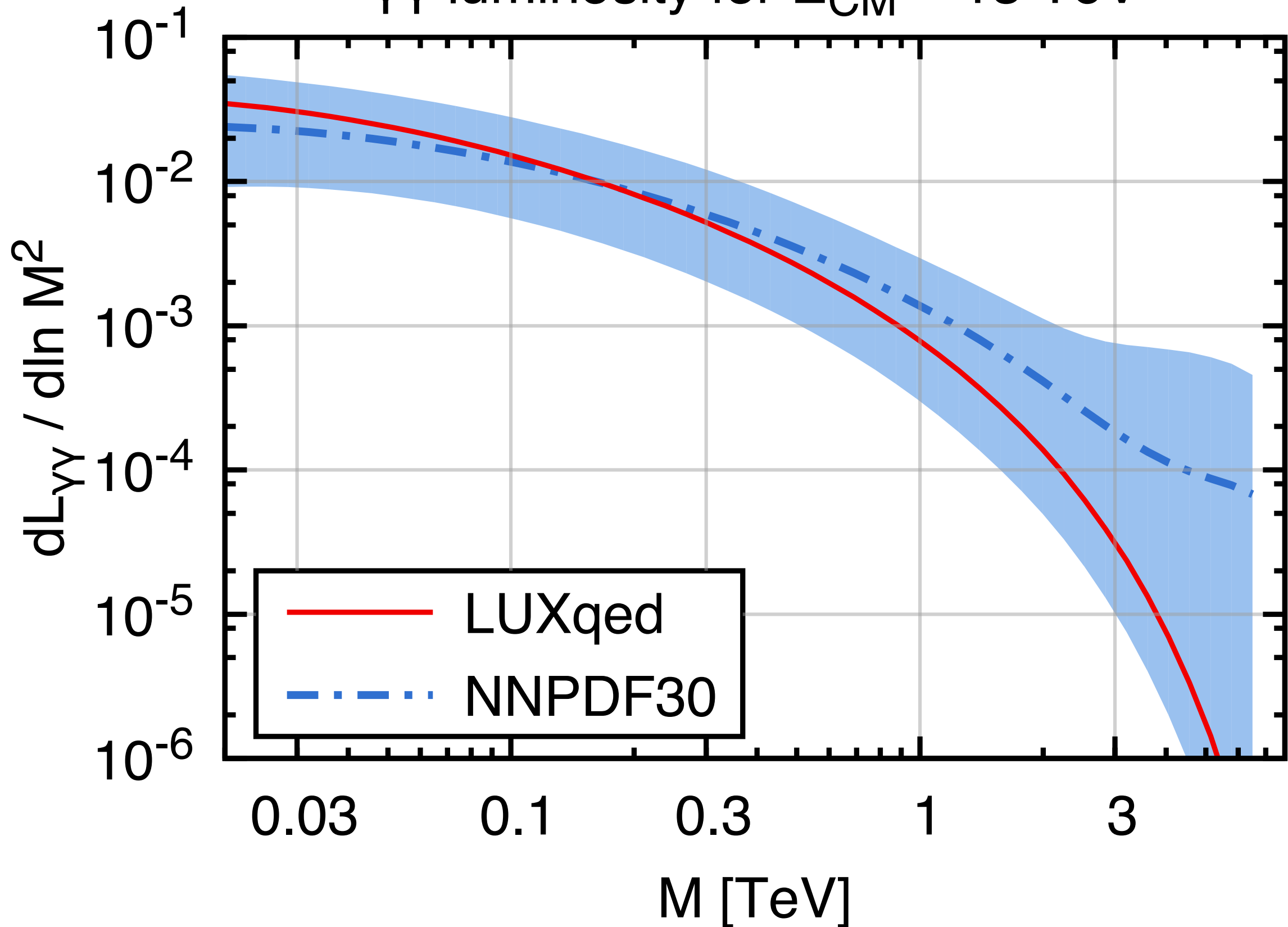
PHOTON UNCERTAINTY (1-2%) COMPARED TO OTHER FLAVOURS

PDF uncertainties ($Q = 100 \text{ GeV}$)



$\gamma\gamma$ luminosity

$\gamma\gamma$ luminosity for $E_{\text{CM}} = 13$ TeV



APPLICATION TO HIGGS PHYSICS

$pp \rightarrow H W^+ (\rightarrow l^+ \nu) + X$ at 13 TeV

non-photon induced contributions

91.2 ± 1.8 fb

photon-induced contribs (NNPDF23)

$6.0^{+4.4}_{-2.9}$ fb

photon-induced contribs (LUXqed)

4.4 ± 0.1 fb

non-photon numbers from LHCHSWG

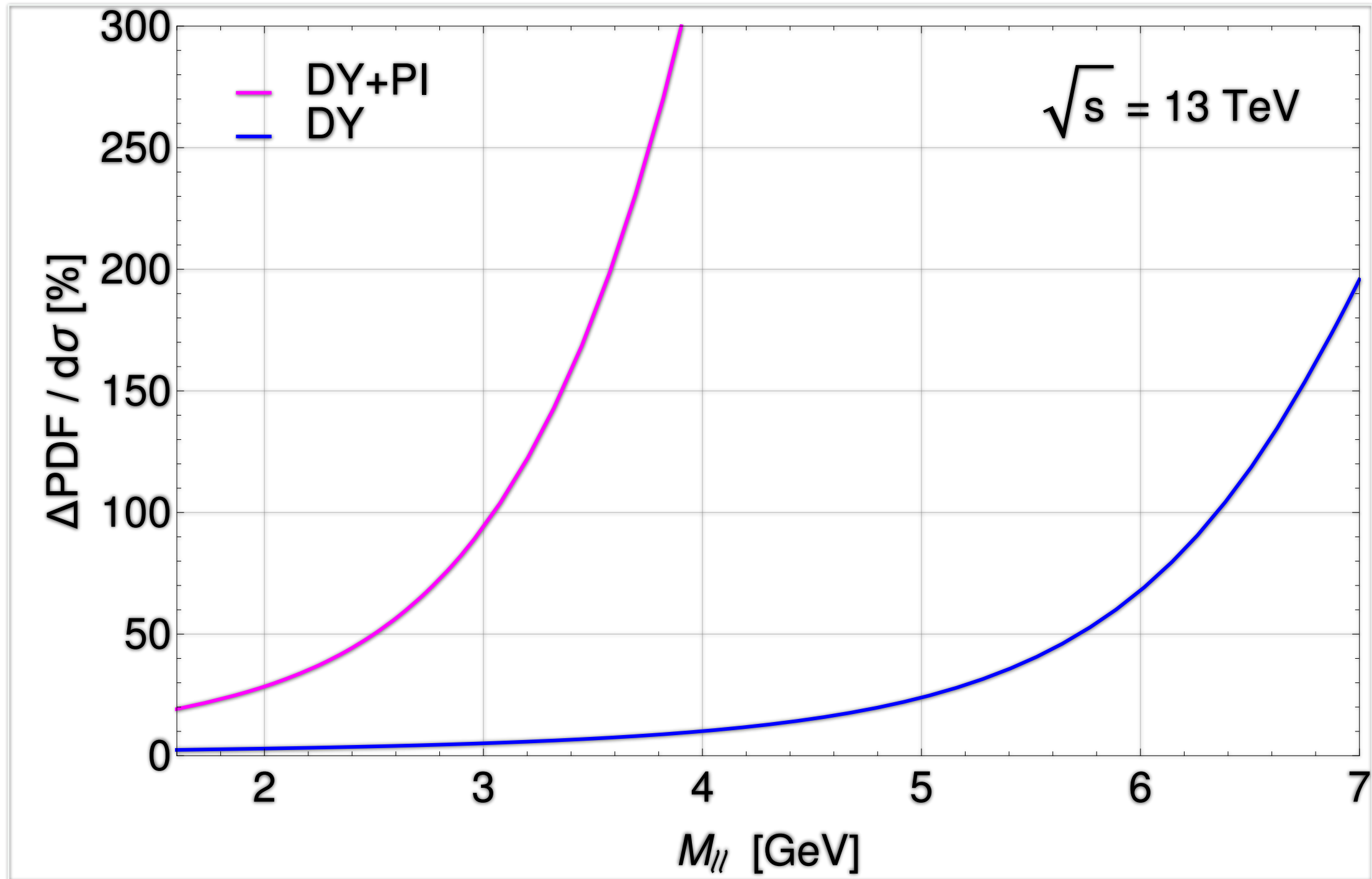
CLOSING REMARKS

- LHC physics would be unrecognisable without Guido's contributions, first and foremost the simple physical picture contained in the DGLAP equations.
- Parton distribution functions are among the crucial inputs to LHC physics, with significant open problems still to solve today.
- More generally, Guido's dedication, his combination of breadth and attention to detail, all serve as a model for what a physicist may aspire to.

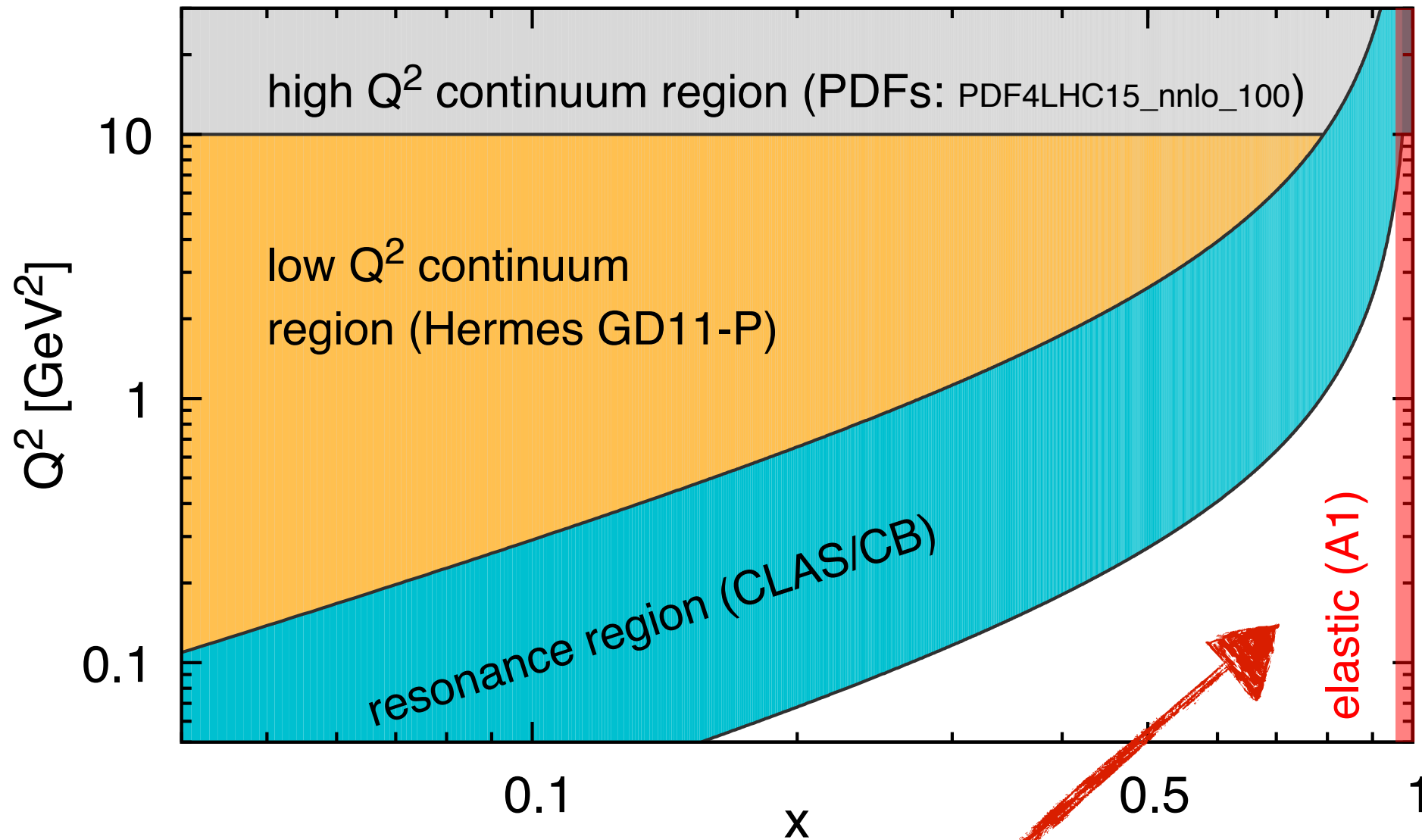
extra slides

1606.06646v1

Elena Accomando,^{1,2,*} Juri Fiaschi,^{1,2,†} Francesco Hautmann,^{2,3,‡}
Stefano Moretti,^{1,2,§} and C.H. Shepherd-Themistocleous^{1,2,¶}



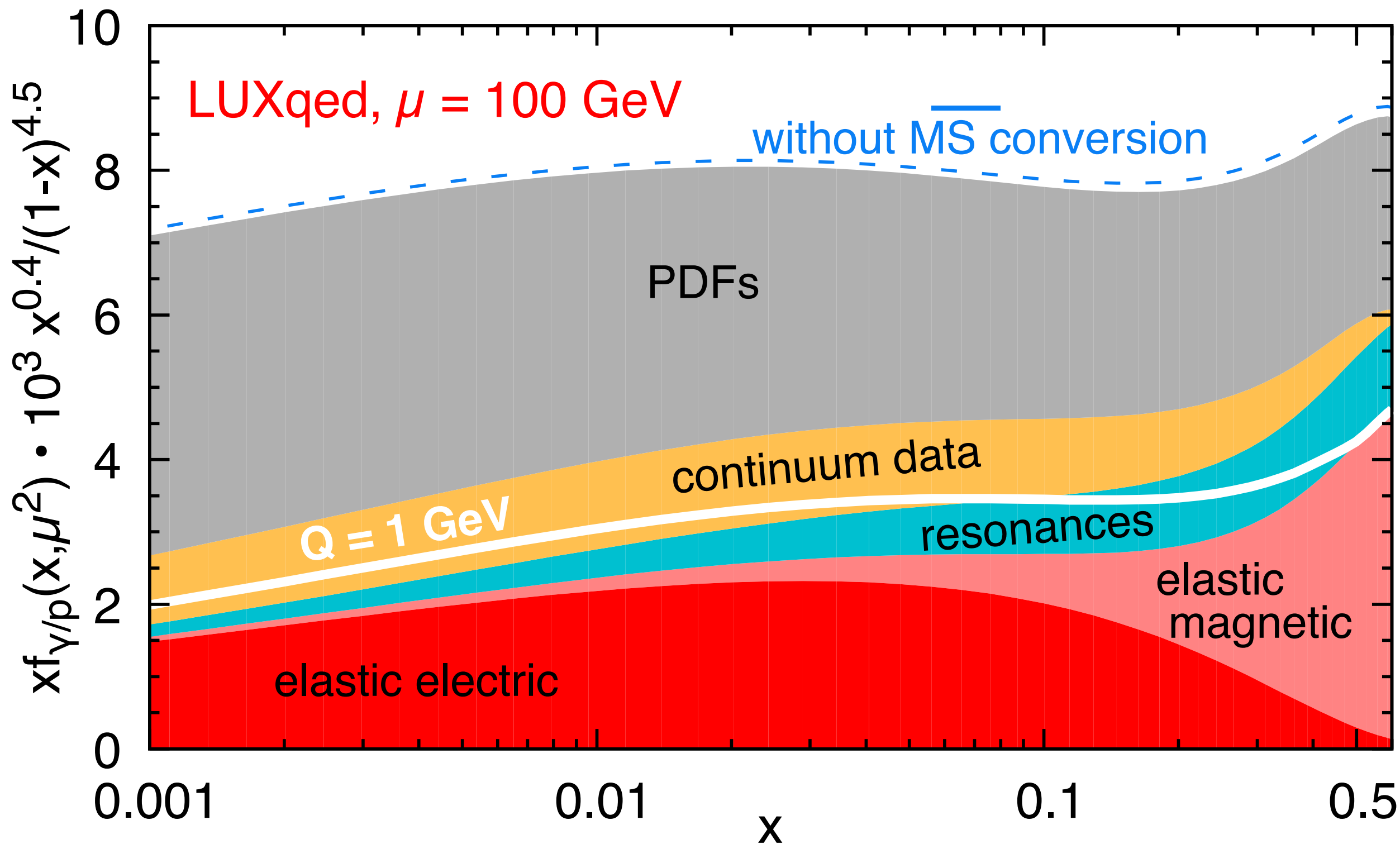
DATA SOURCES – elastic form factors



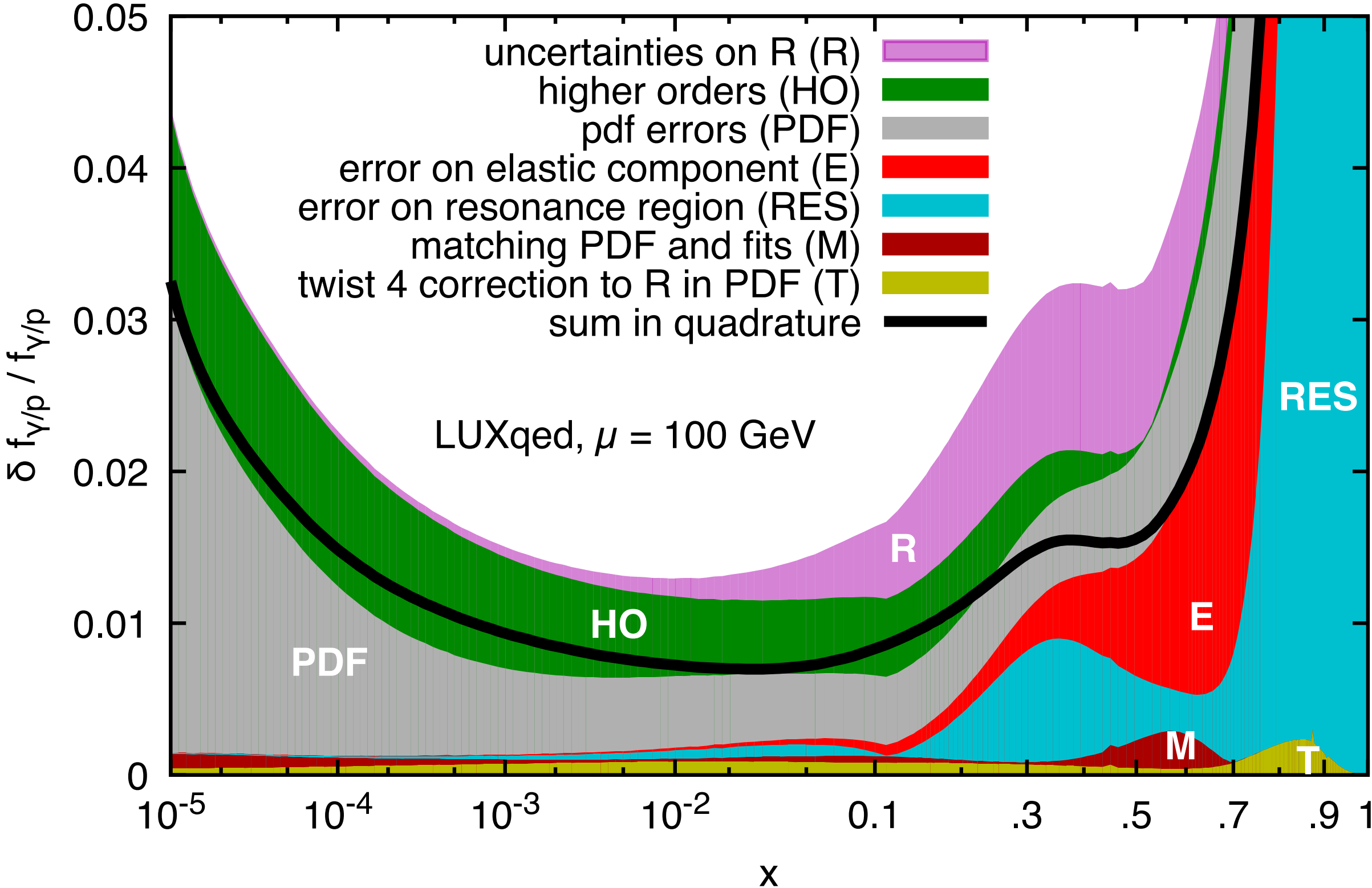
$$F_2^{\text{el}} = \frac{[G_E(Q^2)]^2 + [G_M(Q^2)]^2 \tau}{1 + \tau} \delta(1 - x),$$

$$F_L^{\text{el}} = \frac{[G_E(Q^2)]^2}{\tau} \delta(1 - x),$$

SEPARATE CONTRIBUTIONS TO PHOTON PDF



BREAKDOWN OF UNCERTAINTIES



$\gamma\gamma$ luminosity

$\gamma\gamma$ luminosity for $E_{\text{CM}} = 8$ TeV

