

# **PROTON STRUCTURE**

## **THE LAST LIGHT PARTON**

**Gavin Salam, CERN**

**with Aneesh Manohar, Paolo Nason and Giulia Zanderighi**

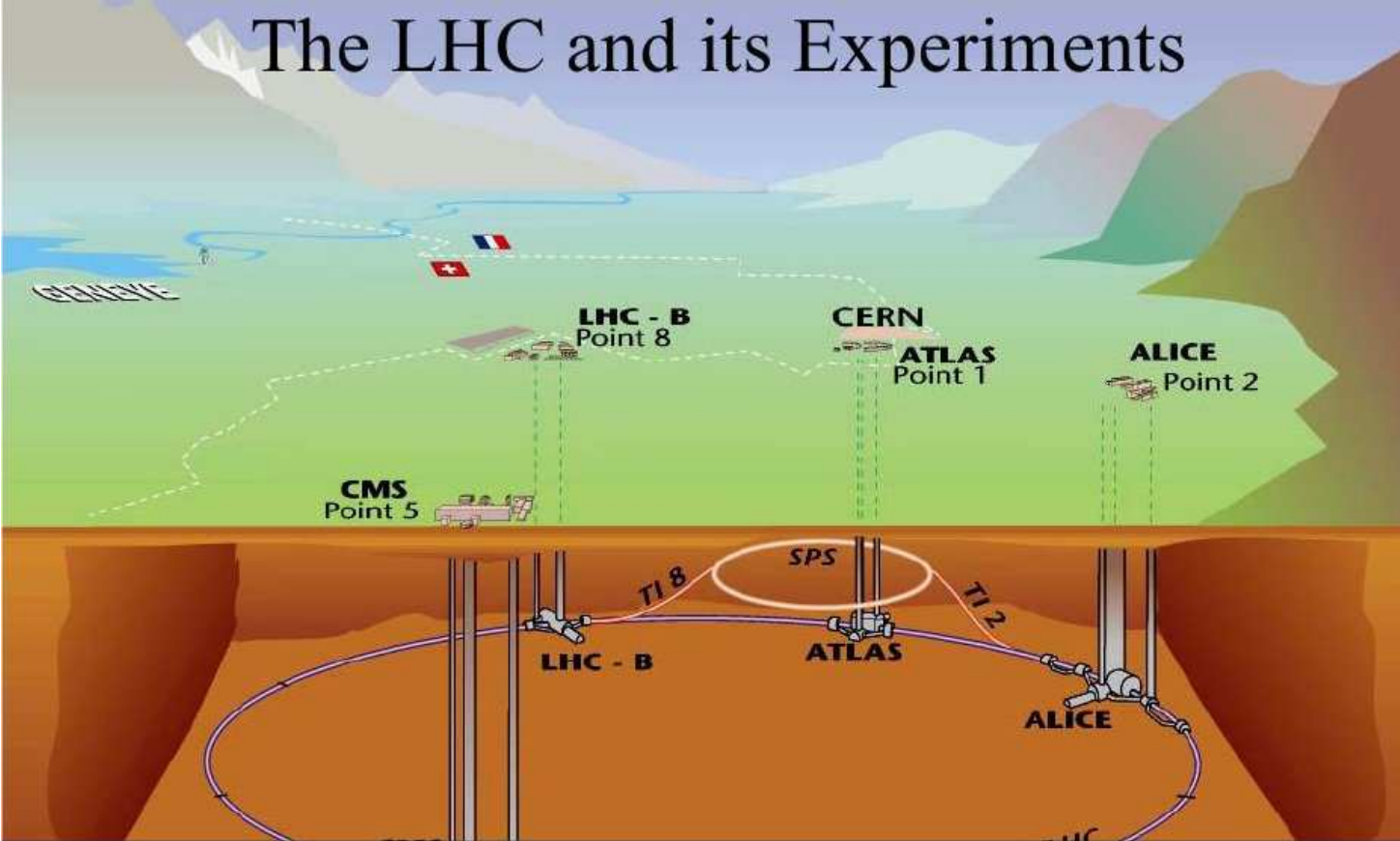
**PRL 117(2016)242002 & work in progress**

**Elementary Particle Physics Seminar**

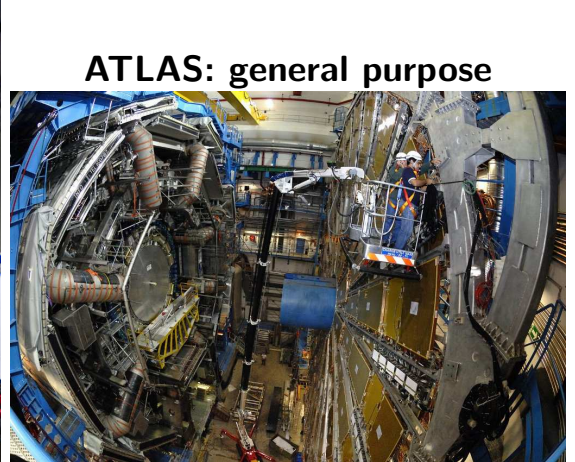
**Universität Würzburg**

**27 April 2017**

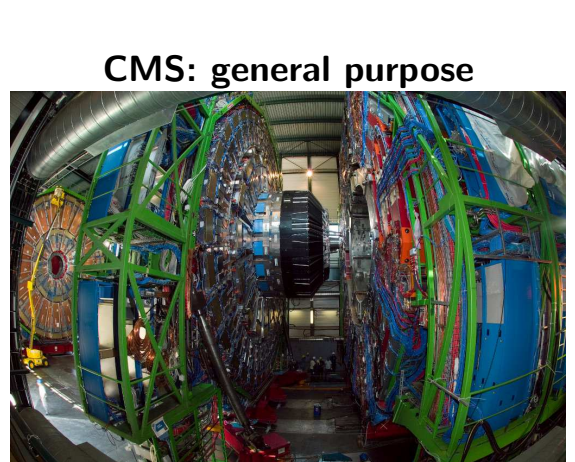
# The LHC and its Experiments



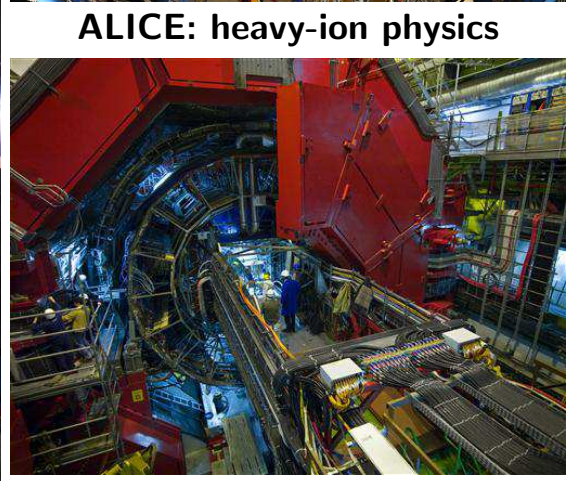
- ~16.5 mi circumference, ~300 feet underground
- 1232 superconducting twin-bore Dipoles (49 ft, 35 t each)
- Dipole Field Strength 8.4 T (13 kA current), Operating Temperature 1.9K
- Beam intensity 0.5 A ( $2.2 \cdot 10^{-6}$  loss causes quench), 362 MJ stored energy



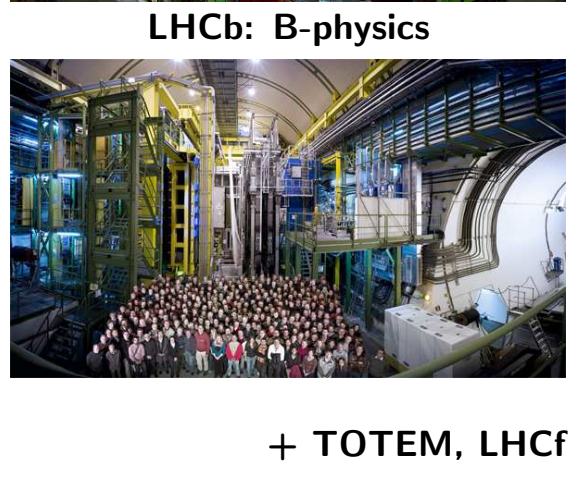
ATLAS: general purpose



CMS: general purpose



ALICE: heavy-ion physics

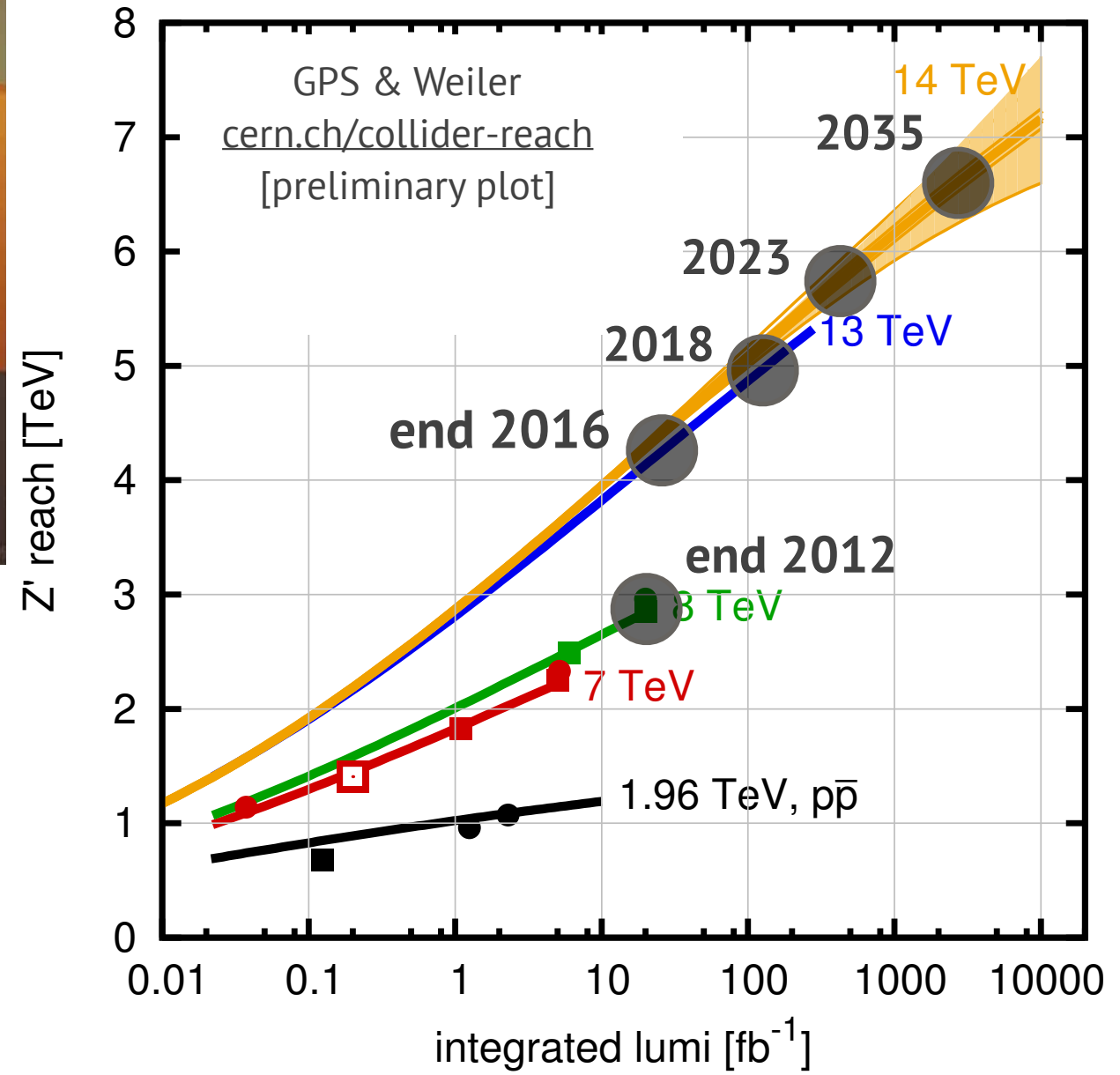


LHCb: B-physics

+ TOTEM, LHCf

# LHC – TWO ROLES – A DISCOVERY MACHINE AND A PRECISION MACHINE

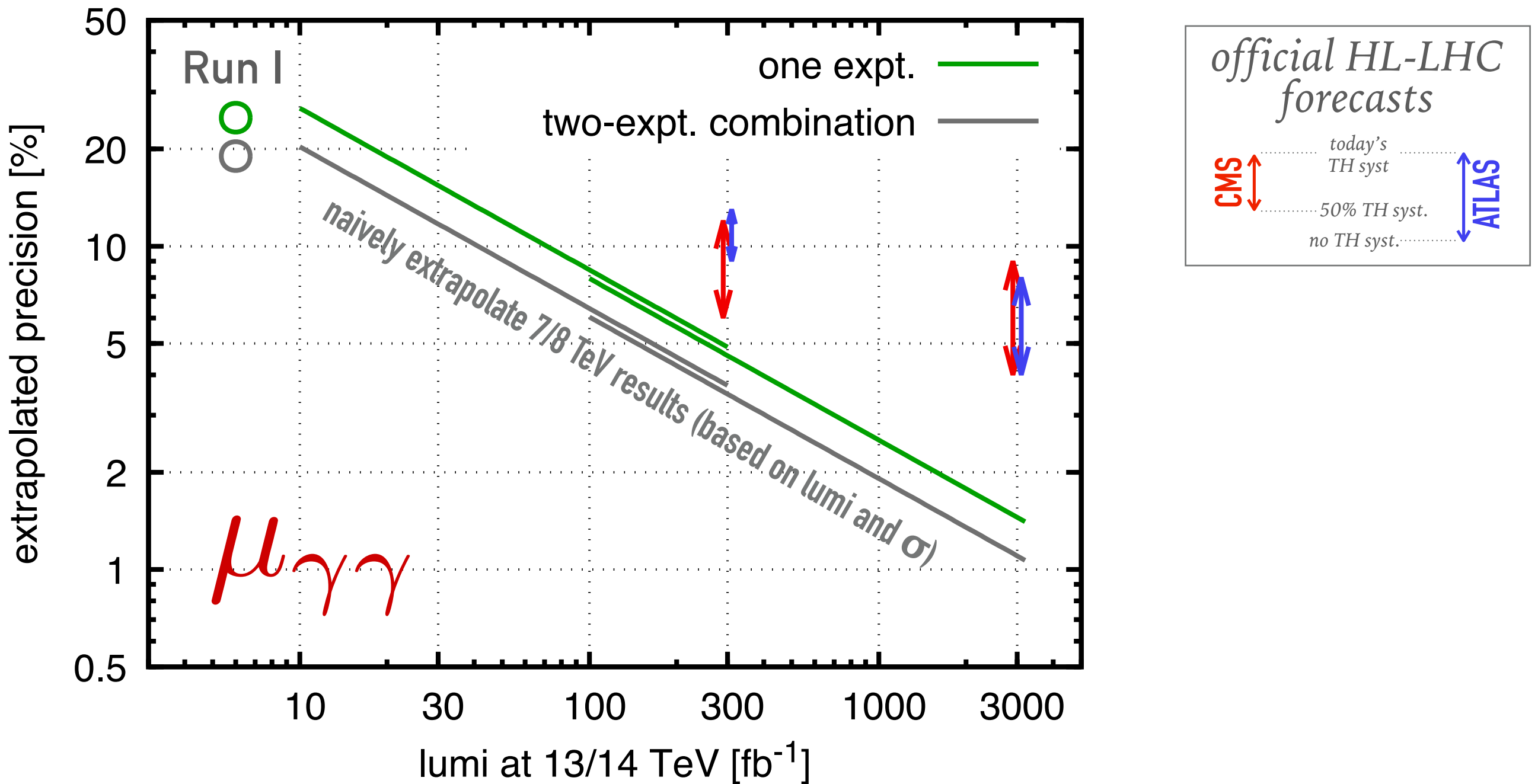
## Z' exclusion reach v. lumi



Increase in luminosity brings discovery reach and precision

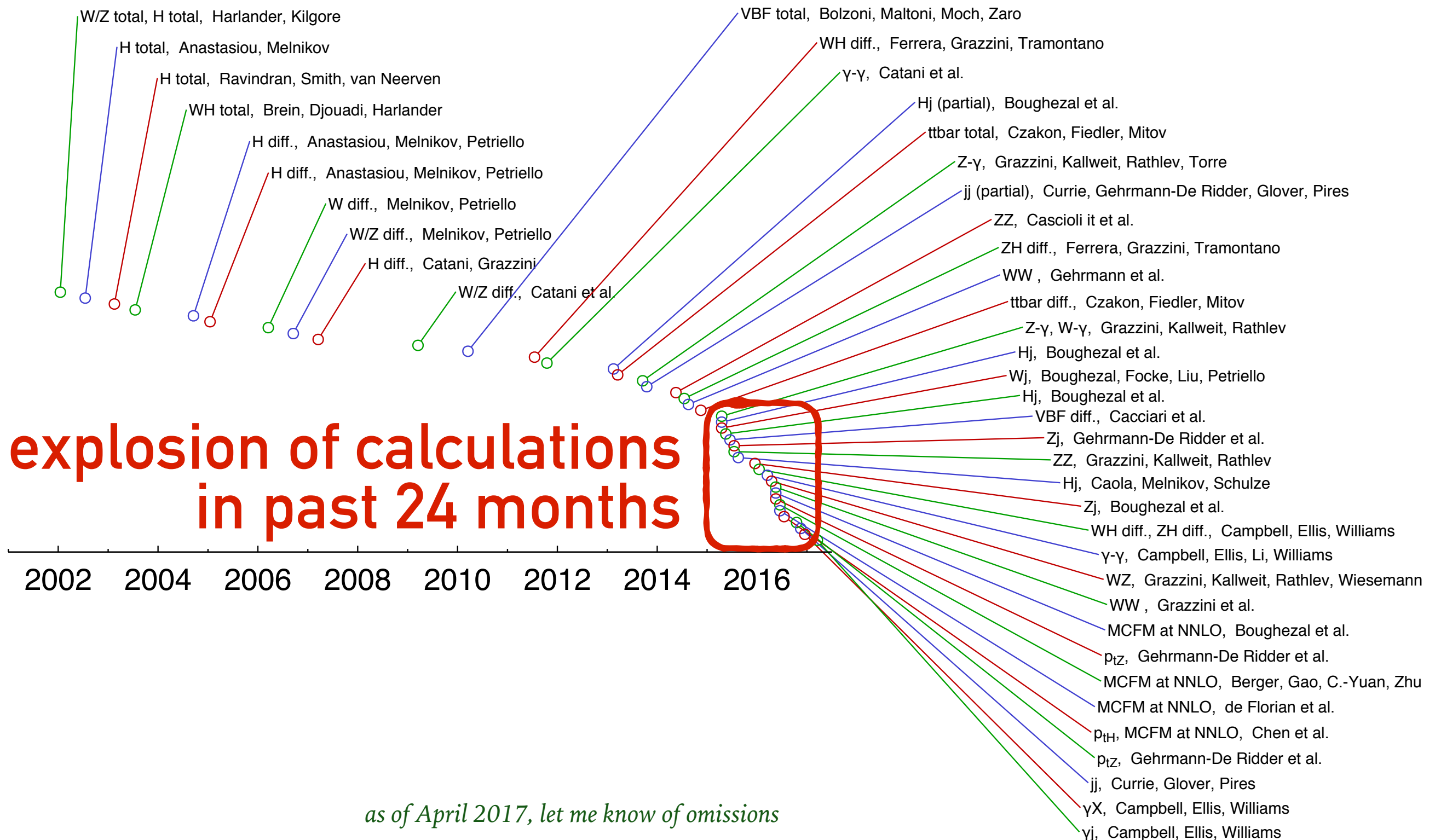


# LONG-TERM HIGGS PRECISION?



Naive extrapolation suggests LHC has long-term potential to do Higgs physics at **1% accuracy**

# NNLO hadron-collider calculations v. time

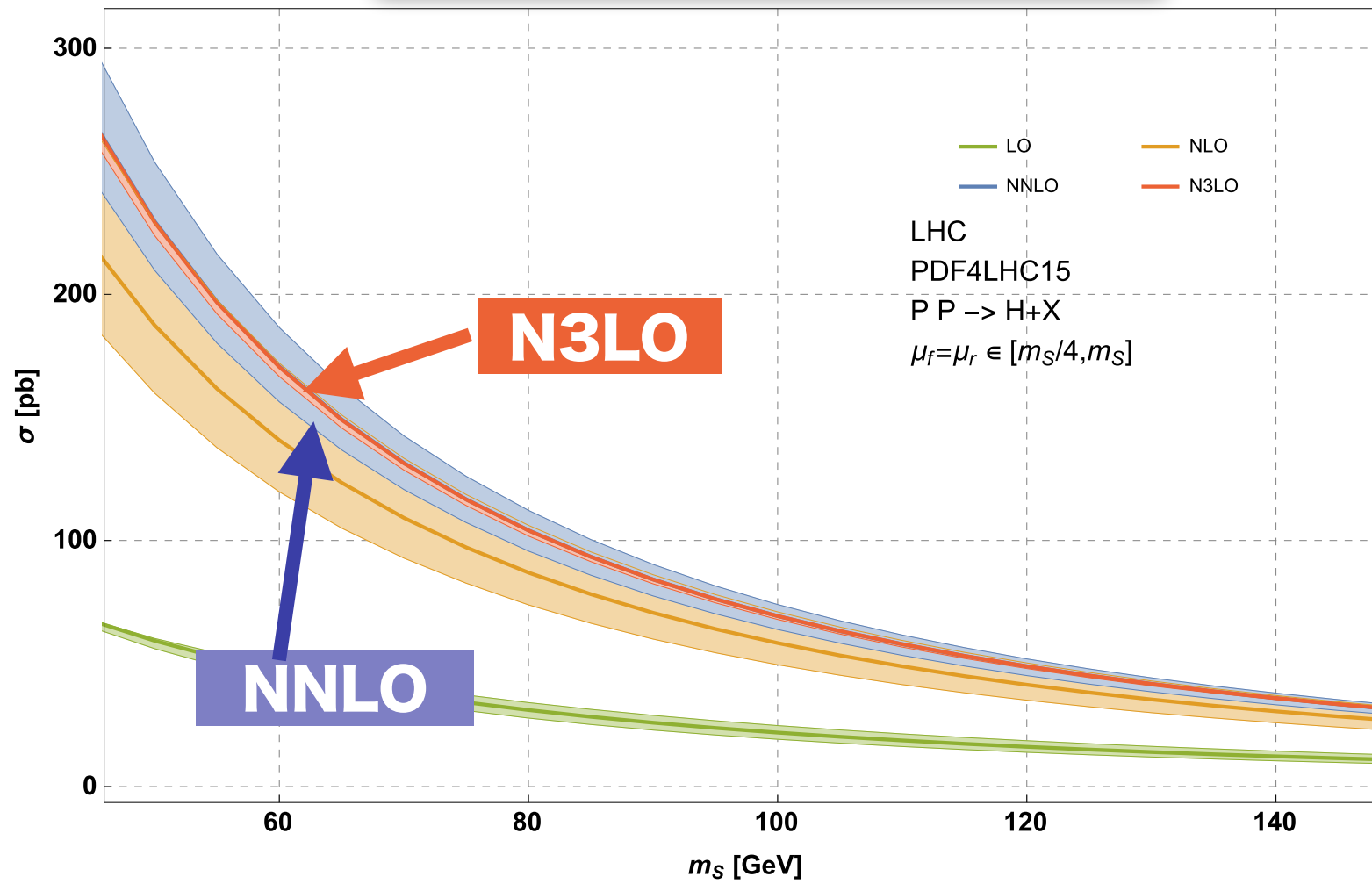




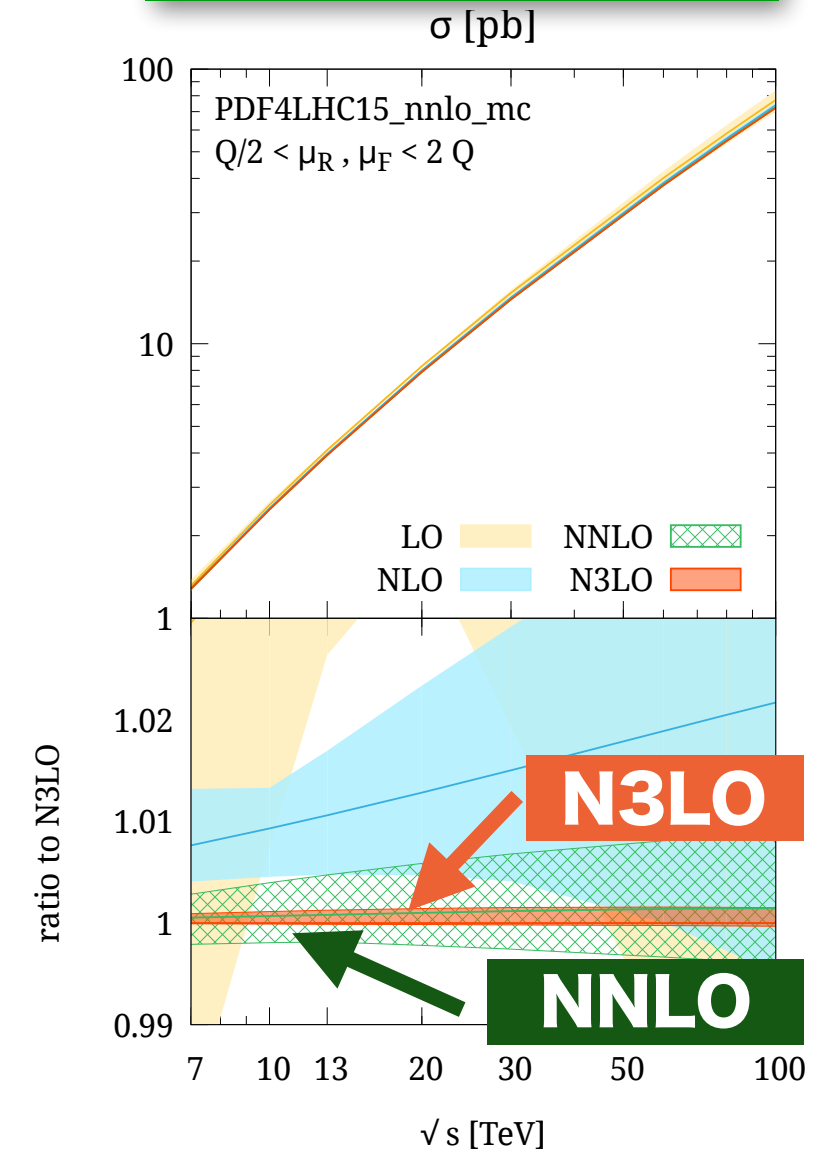
Anastasiou et al, 1602.00695

Dreyer & Karlberg, 1606.00840

## N3LO ggF Higgs

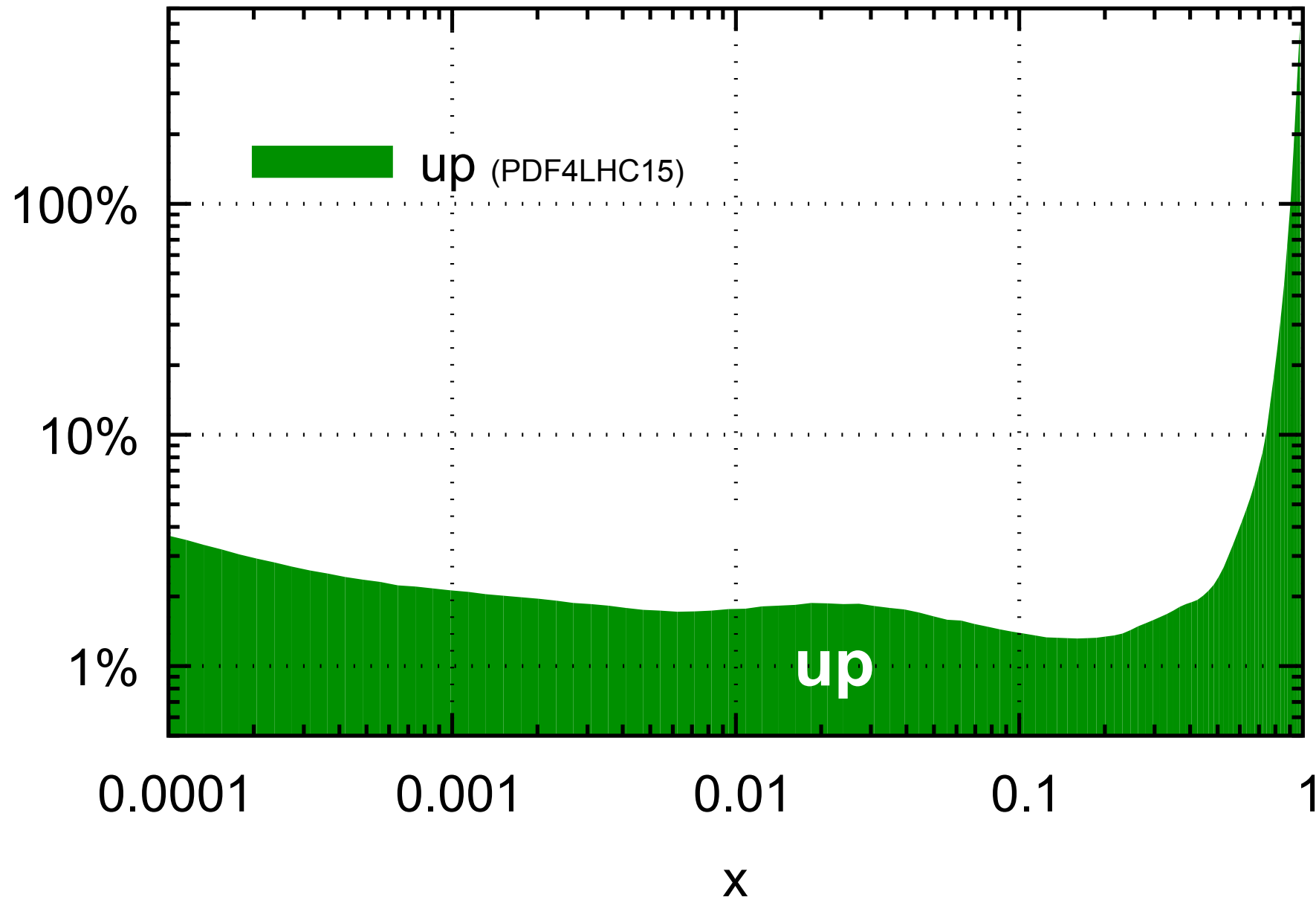


## N3LO VBF Higgs



**how well do we know  
the parton distributions?**

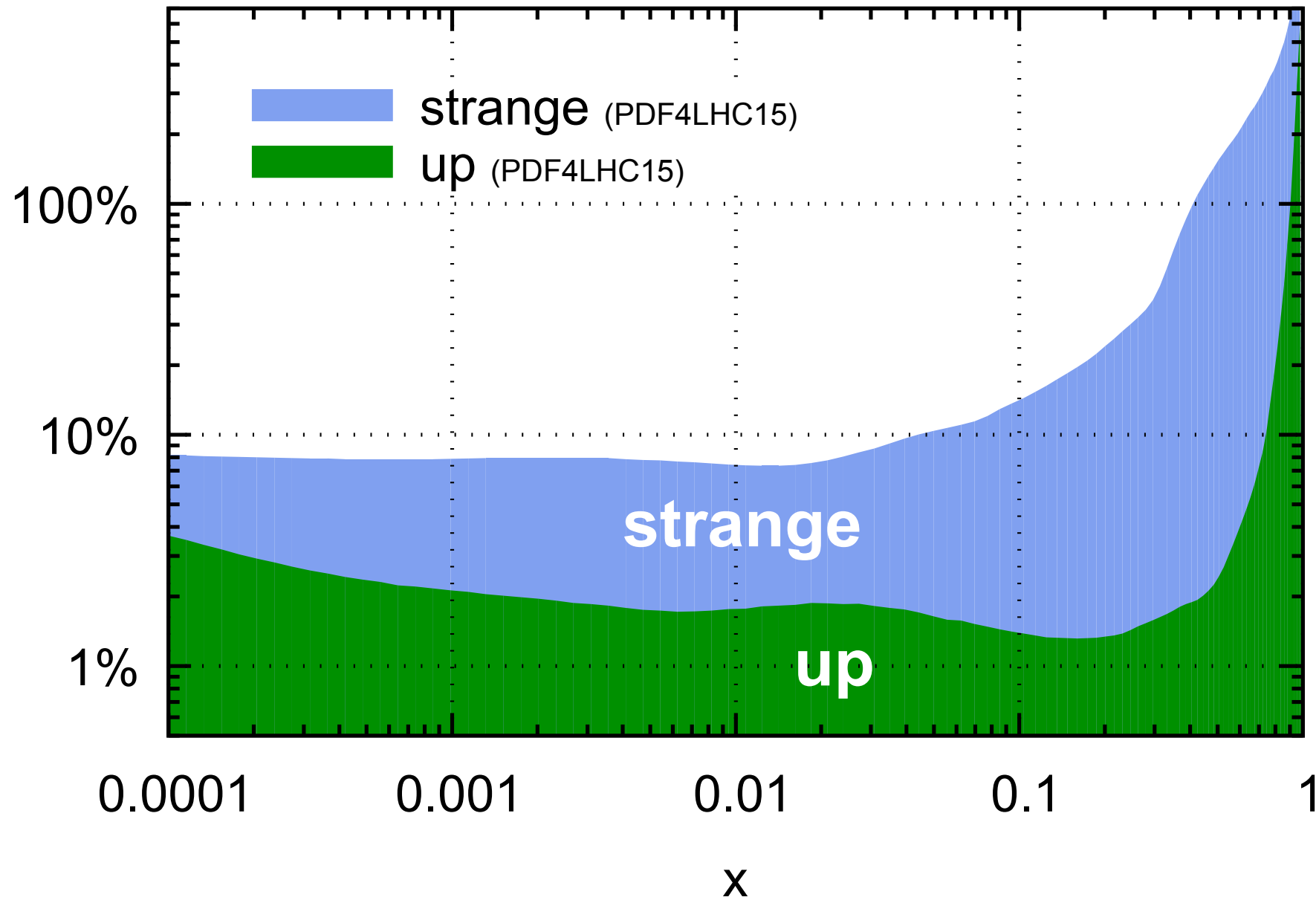
PDF uncertainties (Q = 100 GeV)



➤ core partons (up, down, gluon) are quite well known

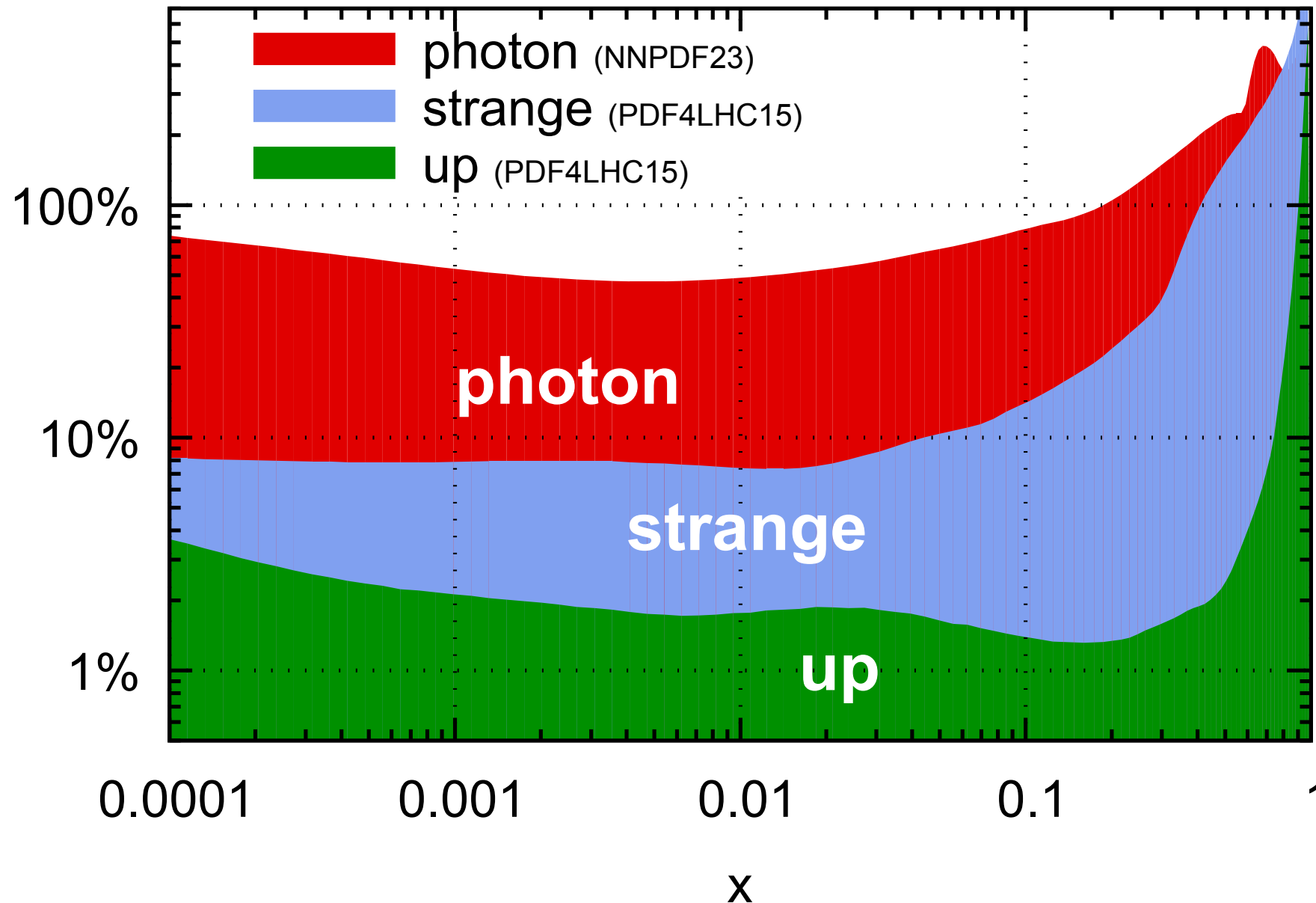


# PDF uncertainties (Q = 100 GeV)



- core partons (up, down, gluon) are quite well known ~2%
- strangeness ~10%

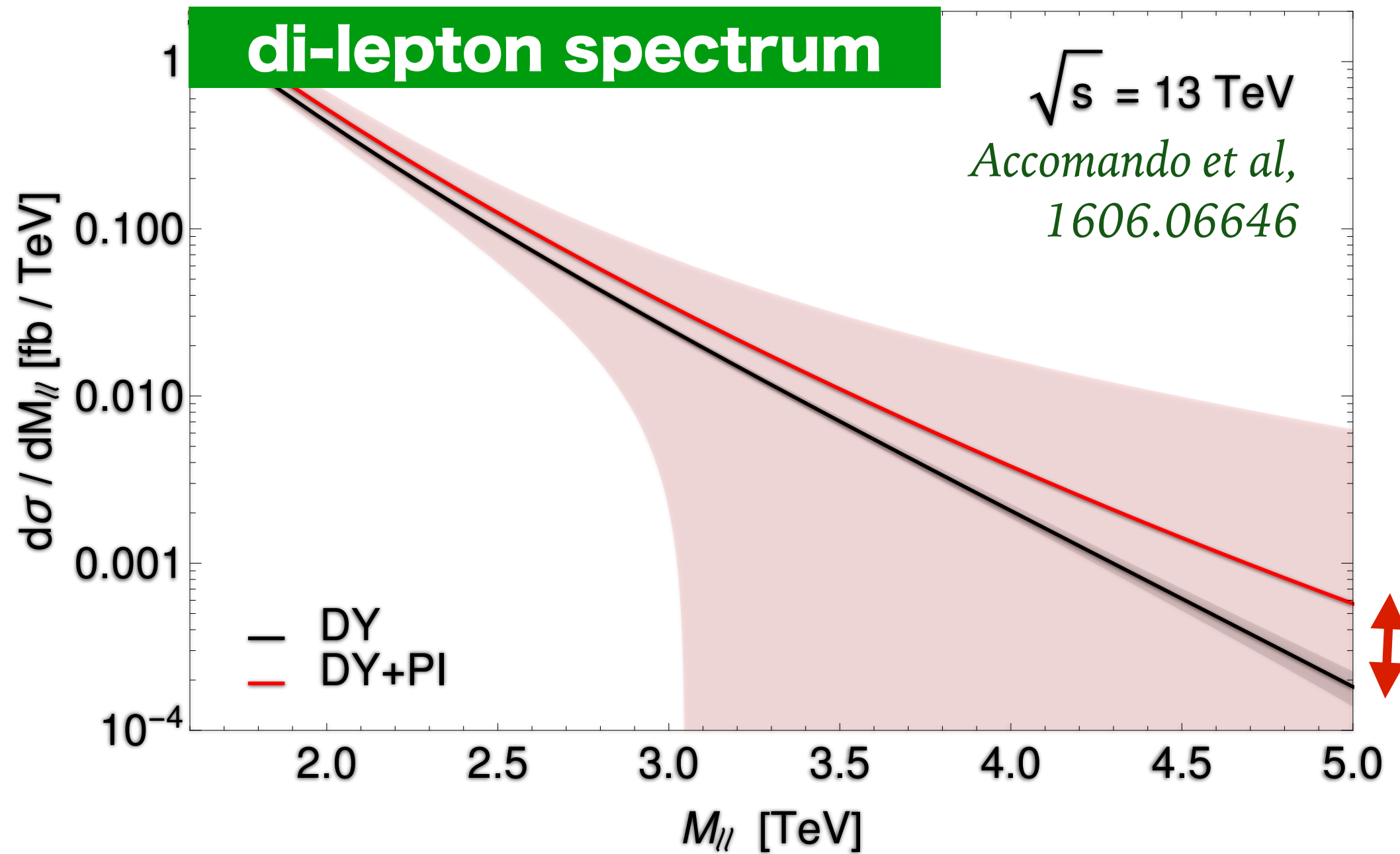
## PDF uncertainties (Q = 100 GeV)



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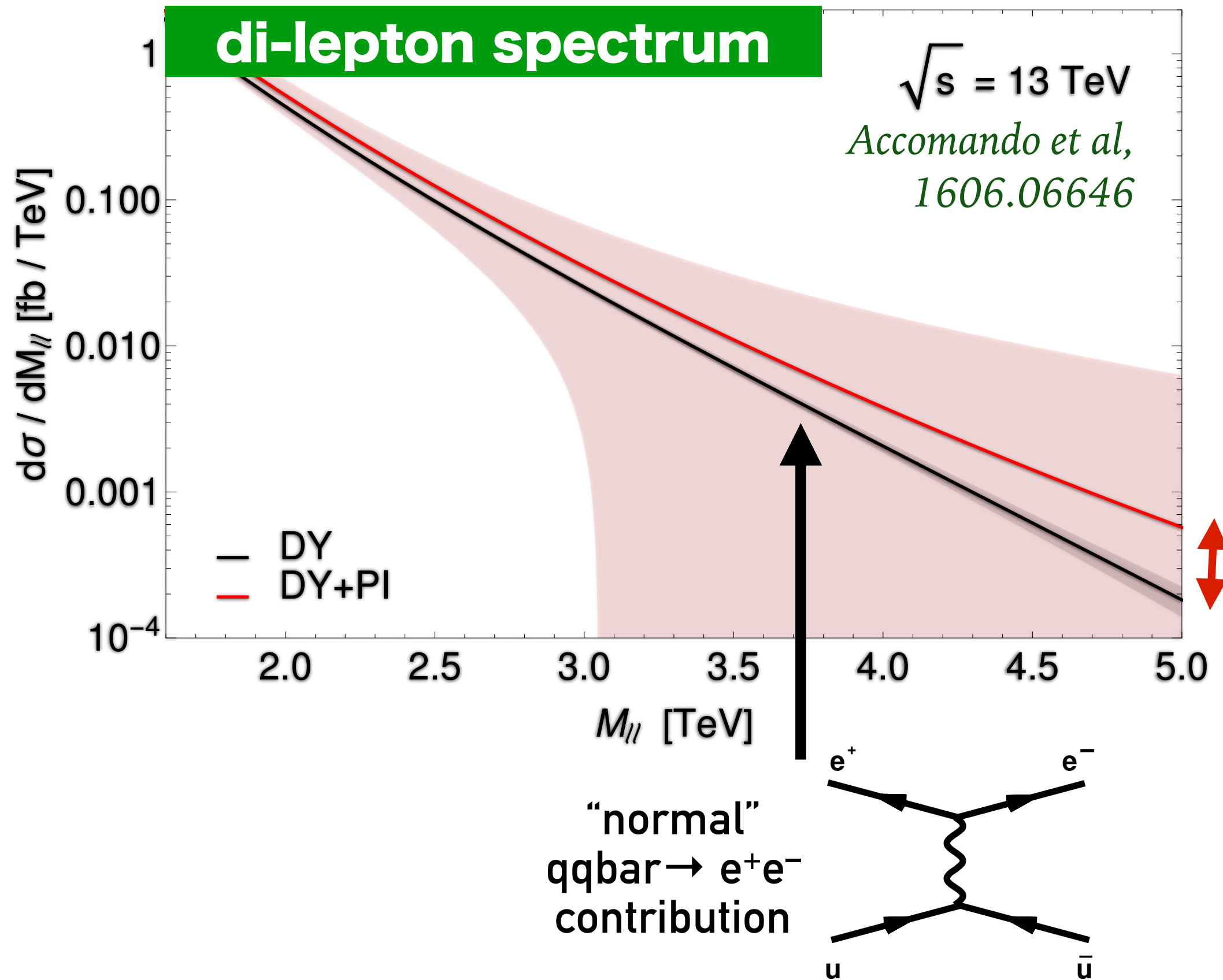
- one other parton, the **photon**, has been debated. Until recently the only model-independent determination (NNPDF23qed) had **0(100%) uncertainty**

# IT MATTERS FOR DI-LEPTON, DI-BOSON, $\tau\tau$ BAR, EW HIGGS, ETC.

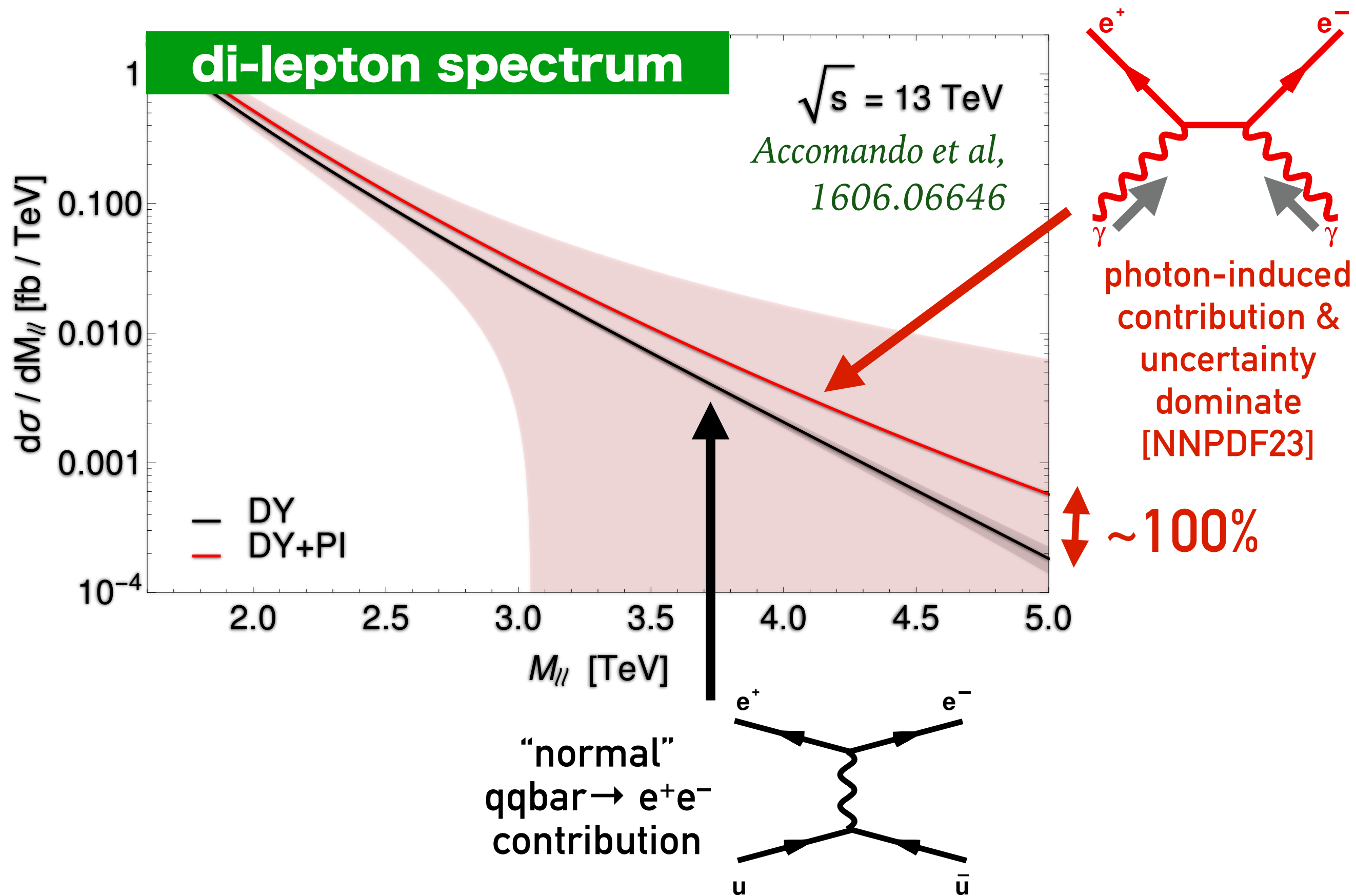




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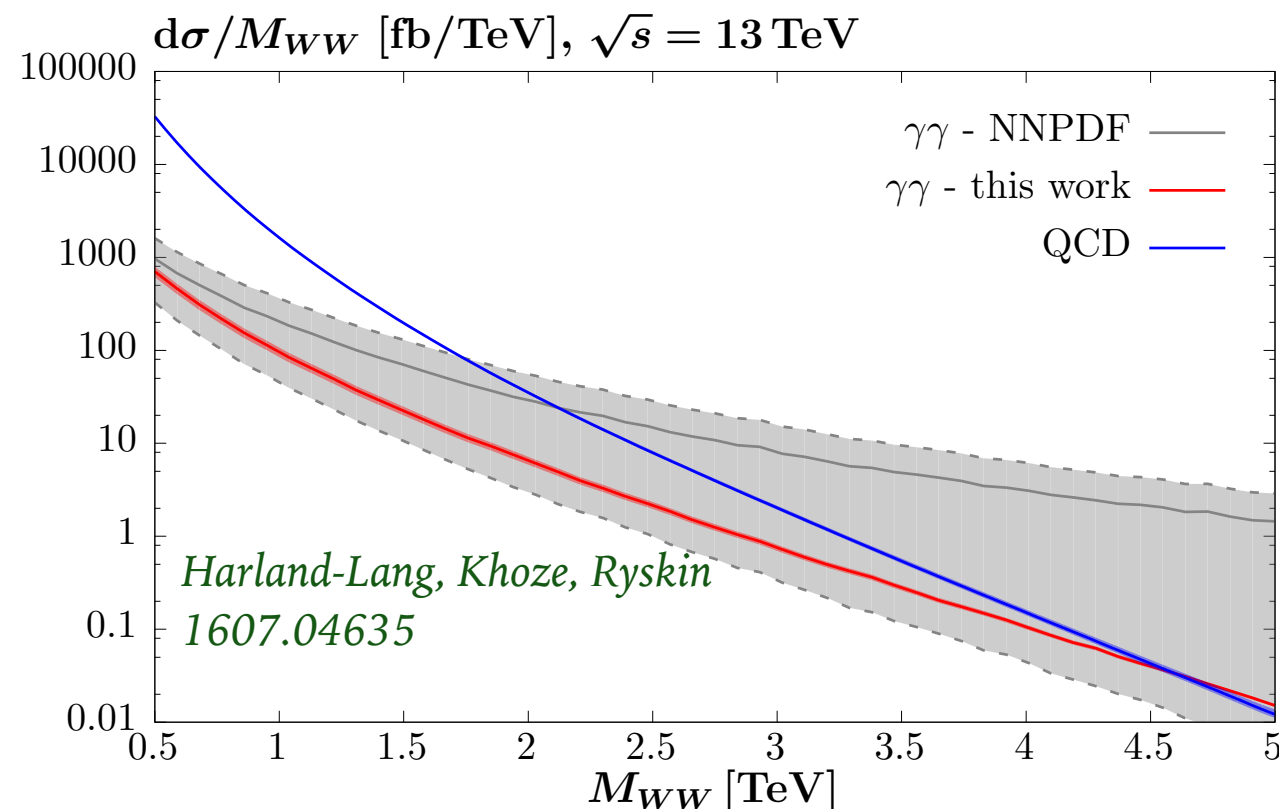


# IT MATTERS FOR DI-LEPTON, DI-BOSON, $T\bar{T}$ , EW HIGGS, ETC.



# where else does the photon come in?

- Electroweak corrections to almost any process
- Largest uncertainty on VBF Higgs and WH ( $\pm$  few %)  
*LHC-HXSWG YR4*
- top production  
*Pagani, Tsinikos, Zaro, arXiv:1606.01915*
- constraints on  $tq\gamma$  coupling  
*Goldouzian & Clerbaux, 1609.04838*
- VV production  
*1409.1803, 1510.08742, 1603.04874, 1601.07787,  
1605.03419, 1604.04080, 1607.04635, ...*



$\gamma\gamma$  (NNPDF)  $100\times$  larger than  $qq$



# photon-induced corrections to $pp \rightarrow HW^+$

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$pp \rightarrow H W^+ (\rightarrow l^+ \nu) + X$  at 13 TeV

non-photon induced contributions

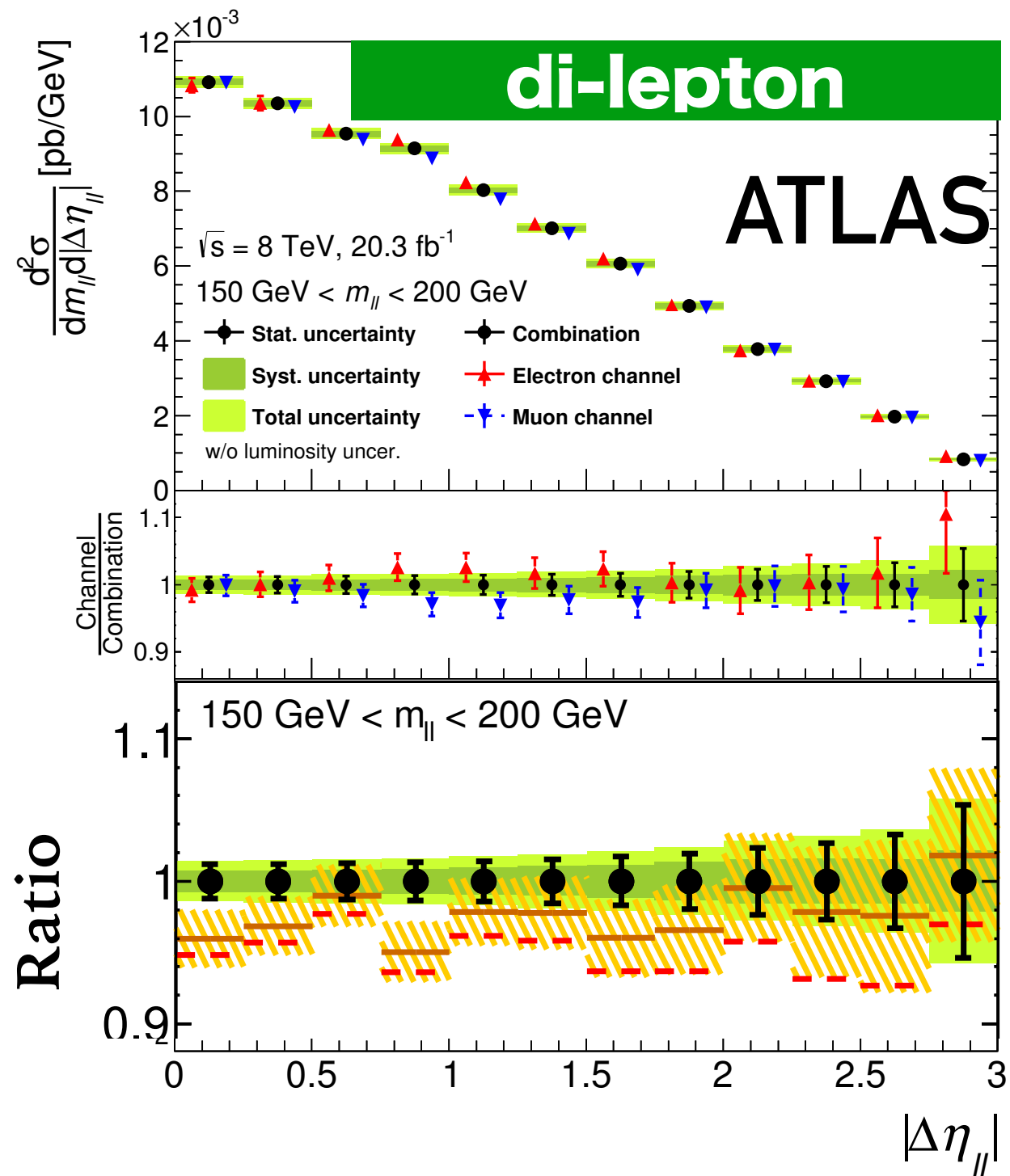
$91.2 \pm 1.8$  fb

photon-induced contriibs (NNPDF23)

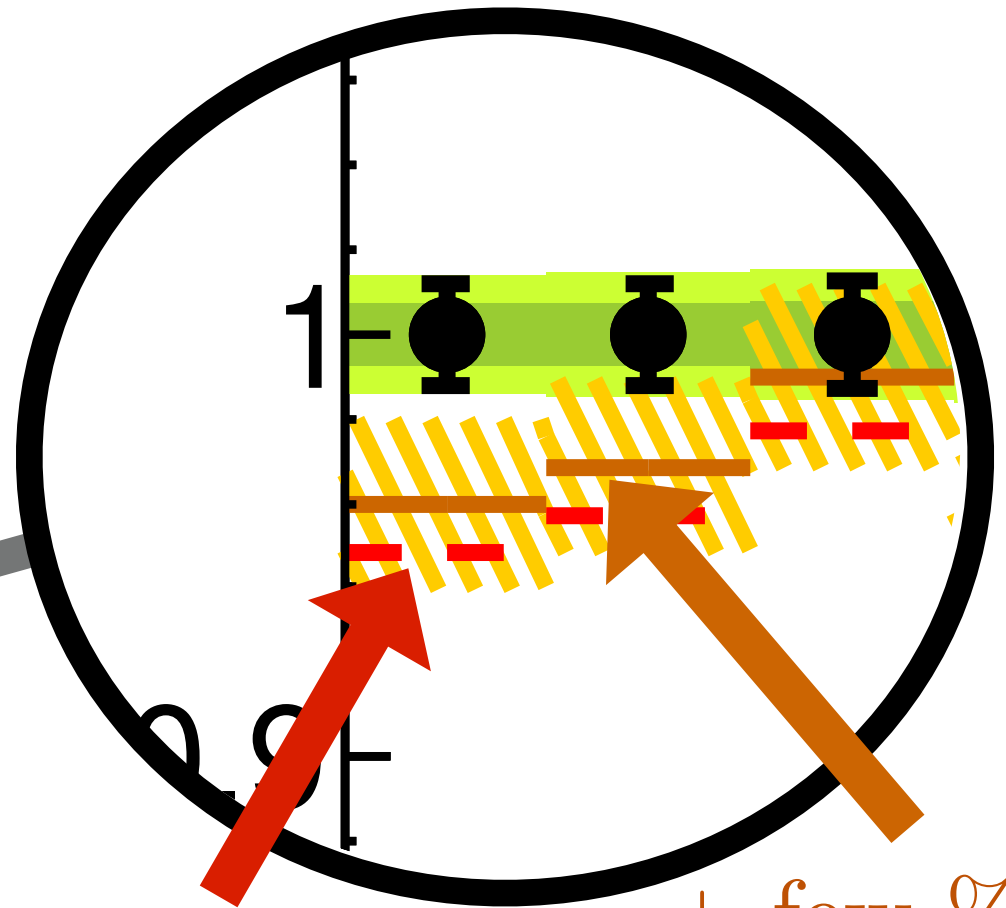
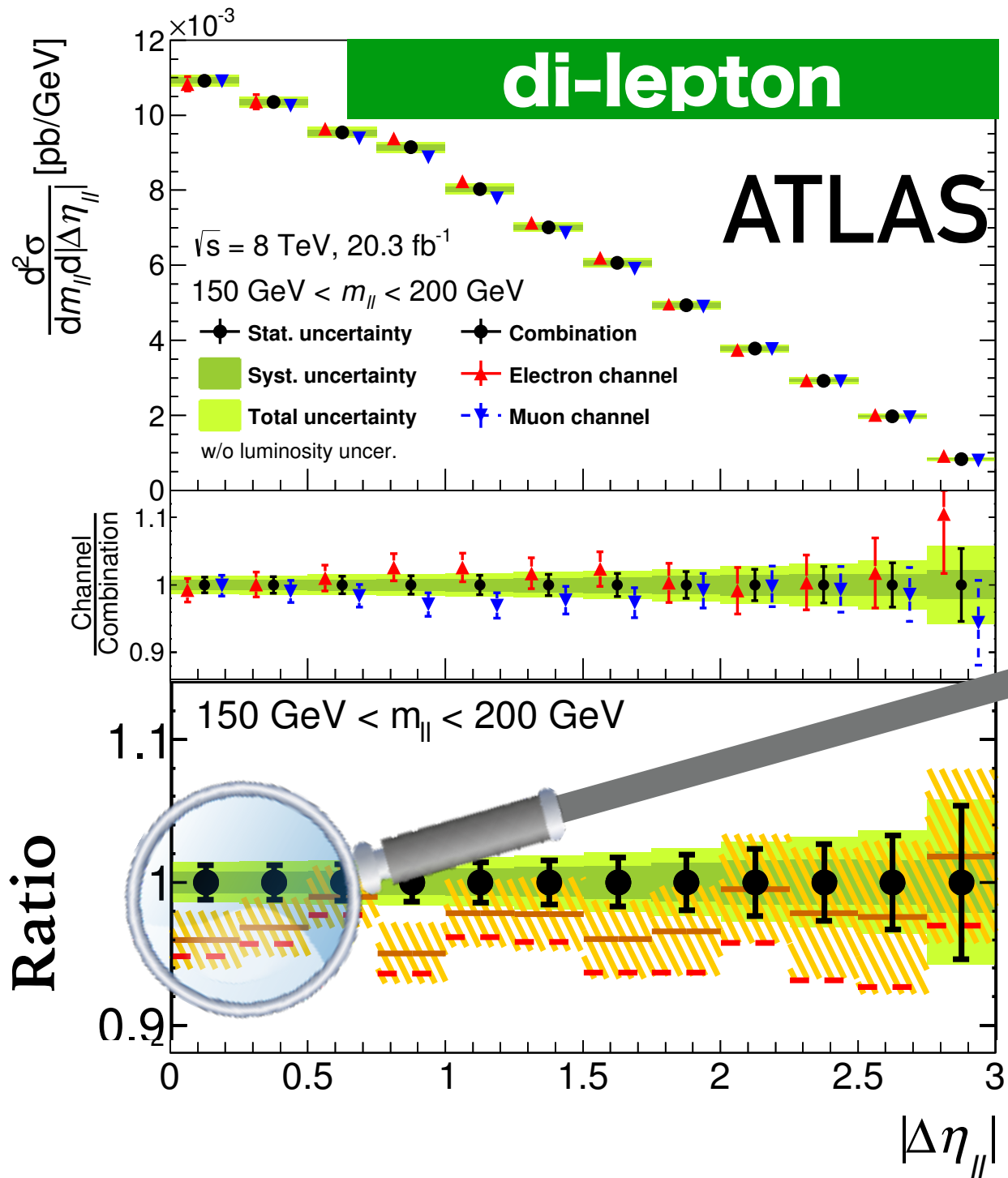
$6.0^{+4.4}_{-2.9}$  fb

*non-photon numbers from LHCHSWG (YR4)  
including PDF uncertainties*

# model-independent $\gamma$ PDF fit (c. 2013)



# model-independent $\gamma$ PDF fit (c. 2013)



95-99% from  $q\bar{q} \rightarrow e^+e^-$

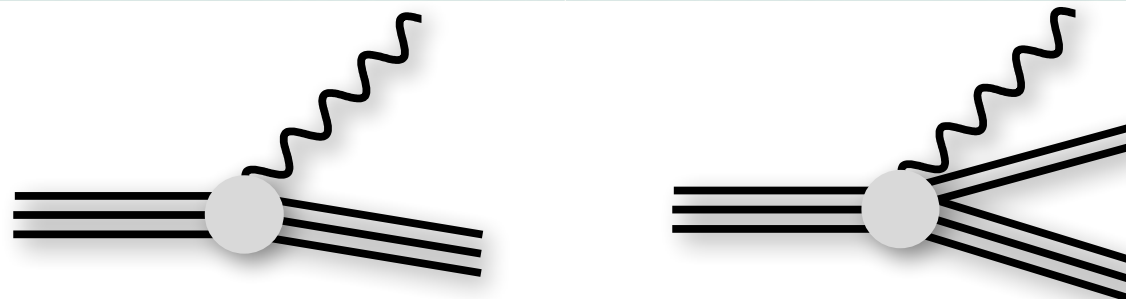
+ few % from  $\gamma\gamma \rightarrow e^+e^-$



# PHOTON PDF ESTIMATES (not exhaustive)

	elastic	inelastic	in LHAPDF?
Gluck Pisano Reya 2002	dipole	model	✗
MRST2004qed	✗	model	✓
NNPDF23qed	no separation; fit to data		✓
CT14qed	✗	model (data-constrained)	✓
CT14qed_inc	dipole	model (data-constrained)	✓
Martin Ryskin 2014	dipole (only electric part)	model	✗
Harland-Lang, Khoze Ryskin 2016	dipole	model	✗

*elastic: Budnev, Ginzburg,  
Meledin, Serbo, 1975*



**YOU SHOULDN'T NEED A MODEL**

**ep scattering (i.e. structure functions) contains all info about proton's EM field**

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**study hypothetical ("BSM") heavy-neutral lepton production process**

**Calculate it in two ways**

**(1) in terms of structure functions (known)**

**(2) in terms of photon distribution (unknown)**

**Equivalence gives us photon distribution**

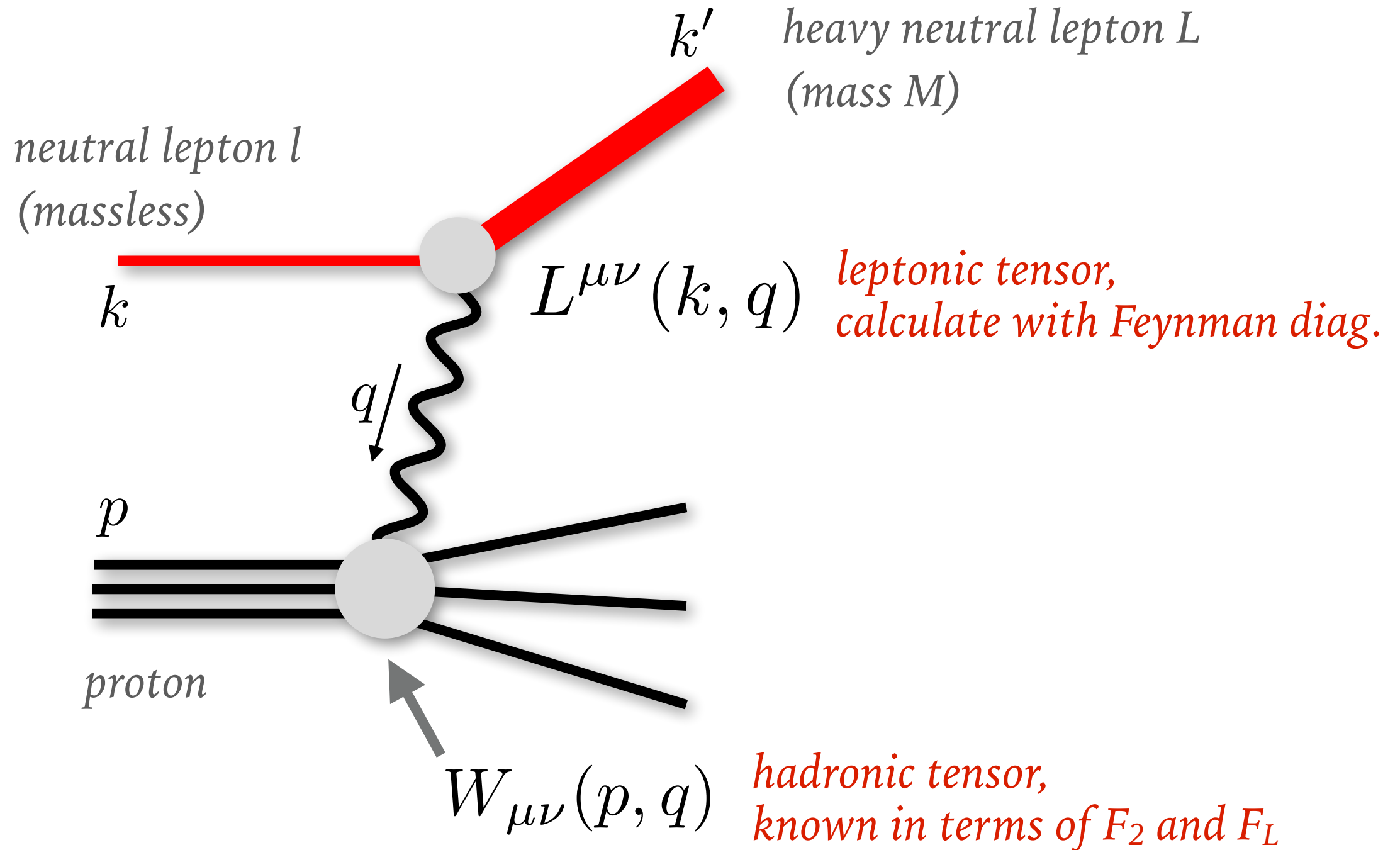
Manohar, Nason, GPS & Zanderighi, arXiv:1607.04266  
(use of BSM inspired by Drees & Zeppenfeld, PRD39(1989)2536)

# **calculation**

## **(one approach)**

# STEP 1

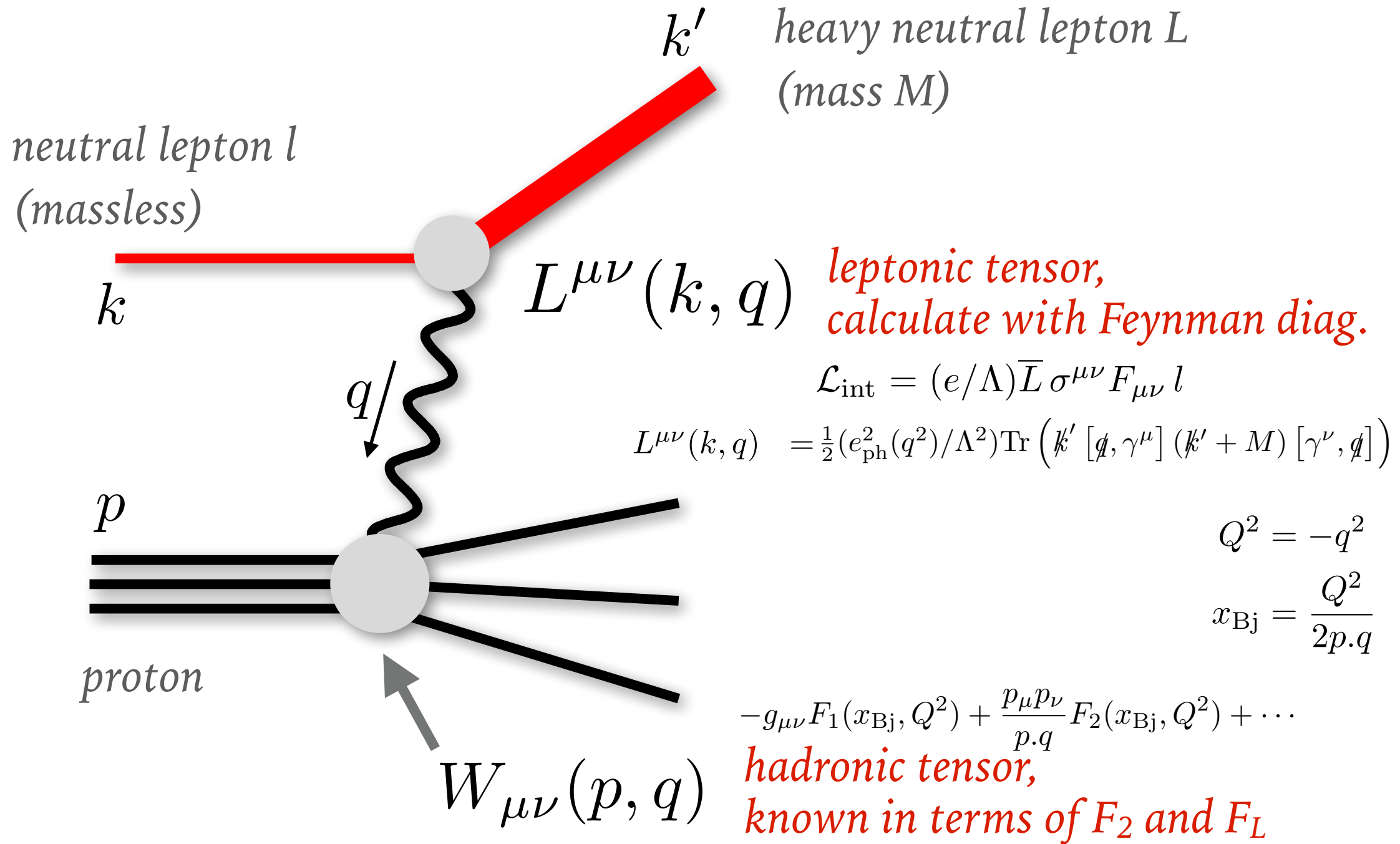
work out a cross section (exact) in terms of  $F_2$  and  $F_L$  struct. fns.



$$\sigma = \frac{1}{4p \cdot k} \int \frac{d^4 q}{(2\pi)^4 q^4} e_{\text{ph}}^2(q^2) [4\pi W_{\mu\nu} L^{\mu\nu}(k, q)] \times 2\pi \delta((k - q)^2 - M^2)$$

# STEP 1

work out a cross section (exact) in terms of  $F_2$  and  $F_L$  struct. fns.



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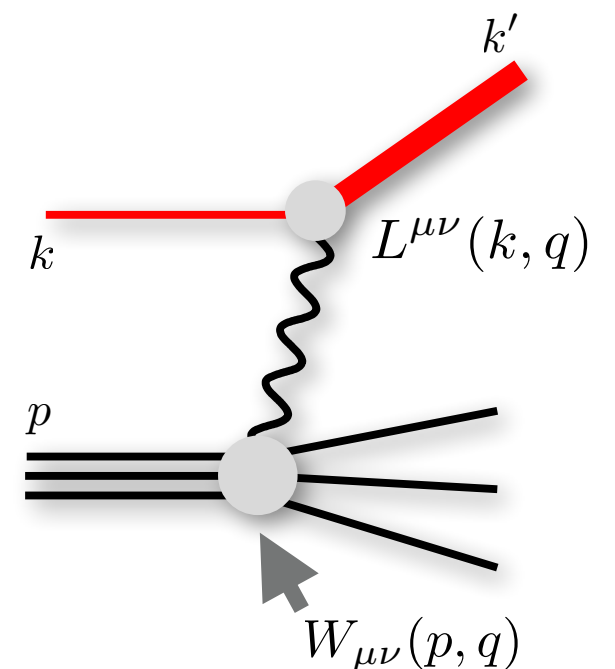


# Cross section in terms of structure functions

- Lagrangian of interaction:  $\mathcal{L}_{\text{int}} = (e/\Lambda)\bar{L}\sigma^{\mu\nu}F_{\mu\nu}l$   
(magnetic moment coupling)
- Using leptons neutral and taking  $\Lambda$  large, ensure that only single-photon exchange is relevant
- **Answer is exact up to  $1/\Lambda$  corrections**

$$\sigma = \frac{c_0}{2\pi} \int_x^{1-\frac{2xm_p}{M}} \frac{dz}{z} \int_{Q_{\text{min}}^2}^{Q_{\text{max}}^2} \frac{dQ^2}{Q^2} \alpha_{\text{ph}}^2(-Q^2) \left[ \left( 2 - 2z + z^2 \right. \right. \\ \left. \left. + \frac{2x^2m_p^2}{Q^2} + \frac{z^2Q^2}{M^2} - \frac{2zQ^2}{M^2} - \frac{2x^2Q^2m_p^2}{M^4} \right) F_2(x/z, Q^2) \right. \\ \left. \left. + \left( -z^2 - \frac{z^2Q^2}{2M^2} + \frac{z^2Q^4}{2M^4} \right) F_L(x/z, Q^2) \right]$$

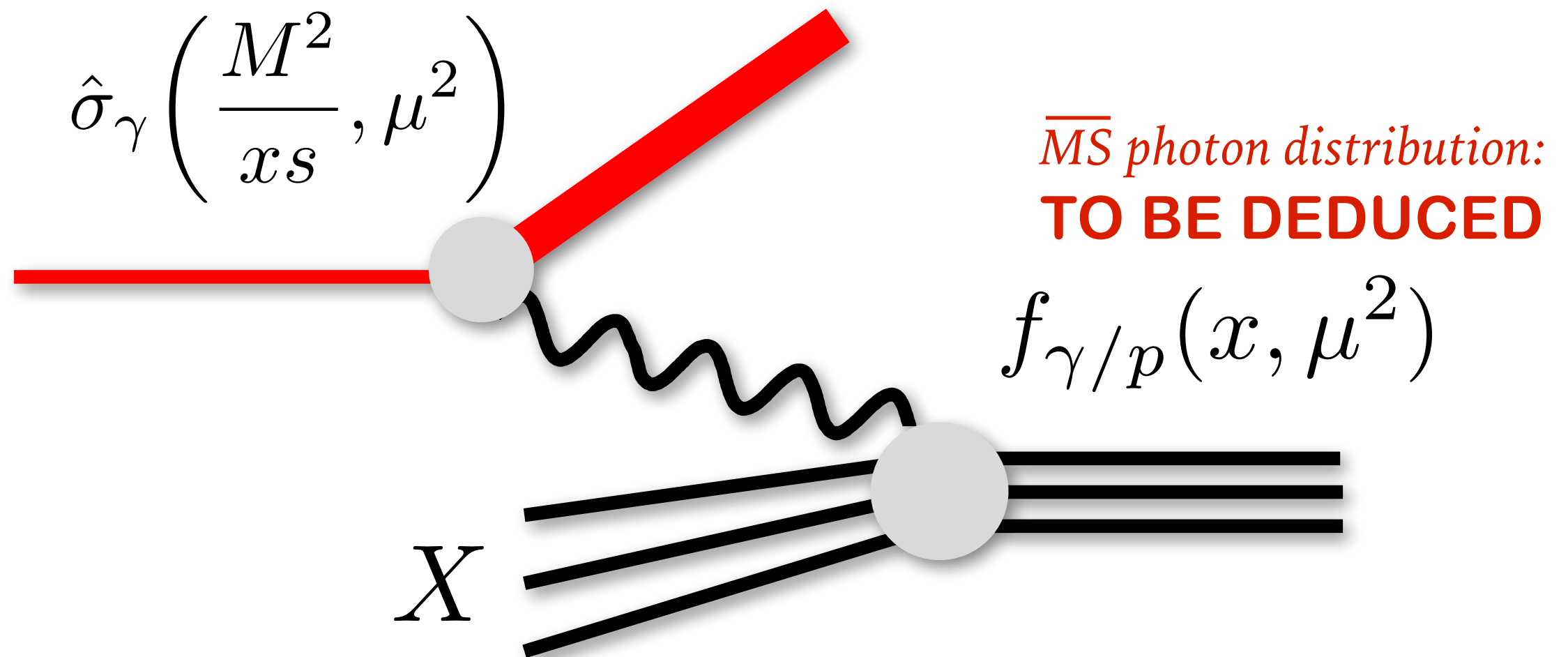
$$c_0 = 16\pi^2/\Lambda^2$$



# STEP 2

work out same cross section in terms of a photon distribution

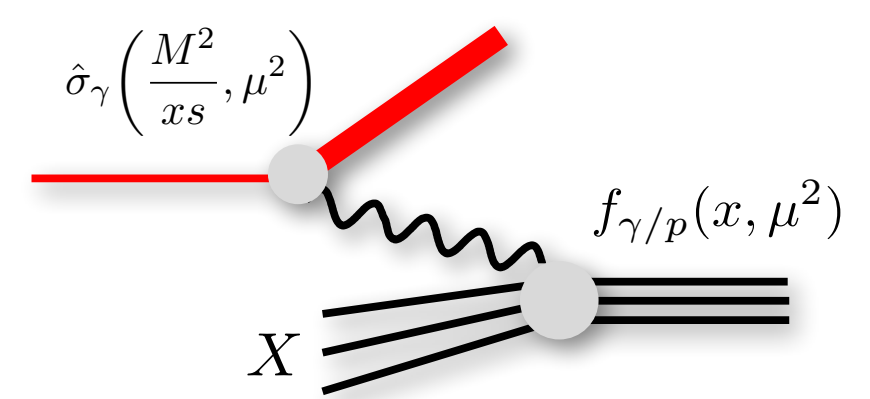
*hard-scattering cross section  
calculate in collinear factorisation*



$$\sigma = c_0 \sum_a \int \frac{dx}{x} \hat{\sigma}_a \left( \frac{M^2}{xS}, \mu^2 \right) x f_{a/p}(x, \mu^2)$$

# Cross section in terms of structure functions

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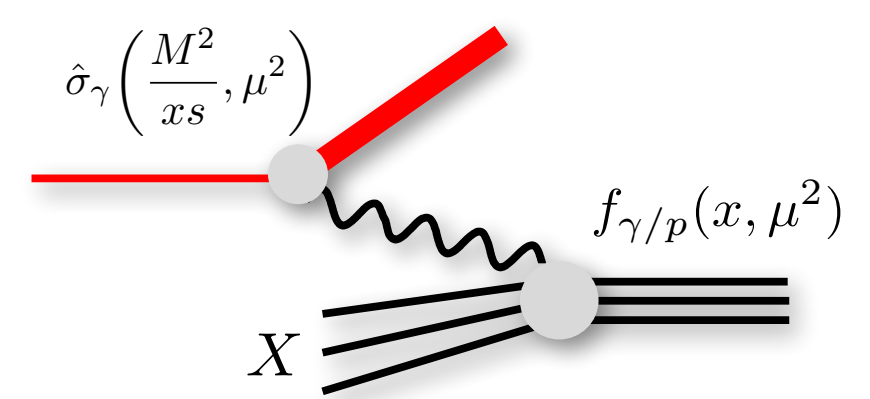


- ▶ Hard cross section driven by the photon distribution at LO



$$\hat{\sigma}_a(z, \mu^2) = \alpha(\mu^2) \delta(1 - z) \delta_{a\gamma}$$

# Cross section in terms of structure functions



- Hard cross section driven by the photon distribution at LO

$$\hat{\sigma}_a(z, \mu^2) = \alpha(\mu^2)\delta(1-z)\delta_{a\gamma} + \frac{\alpha^2(\mu^2)}{2\pi} \left[ -2 + 3z + \right.$$

$$\left. + zp_{\gamma q}(z) \ln \frac{M^2(1-z)^2}{z\mu^2} \right] \sum_{i \in \{q, \bar{q}\}} e_i^2 \delta_{ai} + \dots$$

- Quarks and gluons come in at higher orders

# ACCURACY AIM

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- Take quark and gluon distributions  $\sim O(1)$
- $\alpha$  is QED coupling,  $\alpha_s$  is QCD coupling,  $L = \ln \mu^2/m_p^2$ 
  - Take  $L \sim 1/\alpha_s$ , so all  $(\alpha_s L)^n \sim 1$
  - Think of  $\alpha \sim (\alpha_s)^2$
- To first order, photon distribution  $\sim (\alpha L)$
- we aim to control all terms:
  - $\alpha L (\alpha_s L)^n$  [LO]
  - $\alpha_s \alpha L (\alpha_s L)^n \equiv \alpha (\alpha_s L)^n$  [NLO — extra  $\alpha_s$  or  $1/L$ ]
  - $\alpha^2 L^2 (\alpha_s L)^n$  [NLO — extra  $\alpha L$ ]
- Matching done at large  $M^2$  and  $\mu^2$  to eliminate higher twists

## STEP 3

equate them to deduce the photon distribution (LUXqed)

$$x f_{\gamma/p}(x, \mu^2) = \frac{1}{2\pi\alpha(\mu^2)} \int_x^1 \frac{dz}{z} \left\{ \int_{\frac{x^2 m_p^2}{1-z}}^{\frac{\mu^2}{1-z}} \frac{dQ^2}{Q^2} \alpha^2(Q^2) \right. \\ \left. \left[ \left( z p_{\gamma q}(z) + \frac{2x^2 m_p^2}{Q^2} \right) F_2(x/z, Q^2) - z^2 F_L\left(\frac{x}{z}, Q^2\right) \right] \right. \\ \left. - \alpha^2(\mu^2) z^2 F_2\left(\frac{x}{z}, \mu^2\right) \right\}$$



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with  $F_2 \sim \sum_q e_q^2 x q(x)$  this is just (LO) DGLAP-like piece

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At low  $Q^2$ ,  $F_2$  and  $F_L$  come directly from data (non.pert.)

At high  $Q^2$ , get them from PDFs, including up to  $O(\alpha_s^2)$   
(NNLO) terms

## STEP 3

equate them to deduce the photon distribution (LUXqed)

$$x f_{\gamma/p}(x, \mu^2) = \frac{1}{2\pi\alpha(\mu^2)} \int_x^1 \frac{dz}{z} \left\{ \int_{\frac{x^2 m_p^2}{1-z}}^{\frac{\mu^2}{1-z}} \frac{dQ^2}{Q^2} \alpha^2(Q^2) \right. \\ \left[ \left( z p_{\gamma q}(z) + \frac{2x^2 m_p^2}{Q^2} \right) F_2(x/z, Q^2) - z^2 F_L\left(\frac{x}{z}, Q^2\right) \right] \\ \left. - \alpha^2(\mu^2) z^2 F_2\left(\frac{x}{z}, \mu^2\right) \right\}$$

Terms at boundaries are suppressed by  $1/L$  (NLO)

# STEP 3

equate them to deduce the photon distribution (LUXqed)

$$\begin{aligned}
 x f_{\gamma/p}(x, \mu^2) = & \frac{1}{2\pi\alpha(\mu^2)} \int_x^1 \frac{dz}{z} \left\{ \int_{\frac{x^2 m_p^2}{1-z}}^{\frac{\mu^2}{1-z}} \frac{dQ^2}{Q^2} \alpha^2(Q^2) \right. \\
 & \left[ \left( z p_{\gamma q}(z) + \frac{2x^2 m_p^2}{Q^2} \right) F_2(x/z, Q^2) - z^2 F_L\left(\frac{x}{z}, Q^2\right) \right] \\
 & \left. - \alpha^2(\mu^2) z^2 F_2\left(\frac{x}{z}, \mu^2\right) \right\}
 \end{aligned}$$

terms at boundary  $\sim \mu^2$  ensure  $\overline{\text{MS}}$  fact. scheme

## STEP 3

equate them to deduce the photon distribution (LUXqed)

$$x f_{\gamma/p}(x, \mu^2) = \frac{1}{2\pi\alpha(\mu^2)} \int_x^1 \frac{dz}{z} \left\{ \int_{\frac{x^2 m_p^2}{1-z}}^{\frac{\mu^2}{1-z}} \frac{dQ^2}{Q^2} \alpha^2(Q^2) \right. \\ \left. \left[ \left( z p_{\gamma q}(z) + \frac{2x^2 m_p^2}{Q^2} \right) F_2(x/z, Q^2) - z^2 F_L\left(\frac{x}{z}, Q^2\right) \right] \right. \\ \left. - \alpha^2(\mu^2) z^2 F_2\left(\frac{x}{z}, \mu^2\right) \right\}$$

**QED running of  $\alpha$  accounts for most  $(\alpha L)^2$  effects (NLO)**  
(others come in the way we match to normal PDFs)

**cross-checks**



## Cross checks & literature comparisons

---

- Repeat calculation for a different process ( $\gamma p \rightarrow H + X$ , via  $\gamma\gamma \rightarrow H$ ). Intermediate results differ, **final photon distribution is identical**.
- Repeat calculation using an **operator definition for the photon PDF** (no intermediate BSM process), **identical answer**
- Substitute elastic-scattering component of  $F_2$  and  $F_L$ :

$$F_2^{\text{el}} = \frac{[G_E(Q^2)]^2 + [G_M(Q^2)]^2 \tau}{1 + \tau} \delta(1 - x),$$
$$F_L^{\text{el}} = \frac{[G_E(Q^2)]^2}{\tau} \delta(1 - x), \quad \tau = Q^2 / (4m_p^2)$$

and reproduce widely-used **Equivalent Photon Approximation** with electric ( $G_E$ ) and magnetic ( $G_M$ ) Sachs proton form factors

*Budnev et al., Phys.Rept.15(1975)181*

## Cross checks & literature comparisons

---

- A core part of our answer

$$\left[ \left( zp_{\gamma q}(z) + \frac{2x^2 m_p^2}{Q^2} \right) F_2(x/z, Q^2) - z^2 F_L\left(\frac{x}{z}, Q^2\right) \right]$$

appears in literature for **QED compton process  $ep \rightarrow e\gamma X$**   
(but with inexact treatment of the upper and lower limits for  $Q^2$  integration)

*Anlauf et. al, CPC70(1992)97*

*Mukherjee & Pisano, hep-ph/0306275*

- [NB other literature has an expression for photon distribution in terms of  $F_2$  and  $F_1$  that doesn't reproduce DGLAP limit]

*Luszczak, Schäfer & Szczurek, arXiv:1510.00294*

## Cross checks & literature comparisons

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- $\mu^2$  derivative of our answer should reproduce known DGLAP QCD-QED splitting functions
- At LO, this is trivial.
- At NLO we get relations between QED-QCD splitting functions ( $P$ ) and DIS coefficient functions ( $C$ )

$$P_{\gamma q}^{(1,1)} = e_q^2 \left[ p_{\gamma q} \otimes C_{2q} - h \otimes C_{Lq} + (\bar{p}_{\gamma q} - h) \otimes P_{qq}^{(1,0)} \right] ,$$

$$P_{\gamma g}^{(1,1)} = \sum_{q, \bar{q}} e_q^2 \left[ p_{\gamma q} \otimes C_{2g} - h \otimes C_{Lg} + (\bar{p}_{\gamma q} - h) \otimes P_{qg}^{(1,0)} \right] ,$$

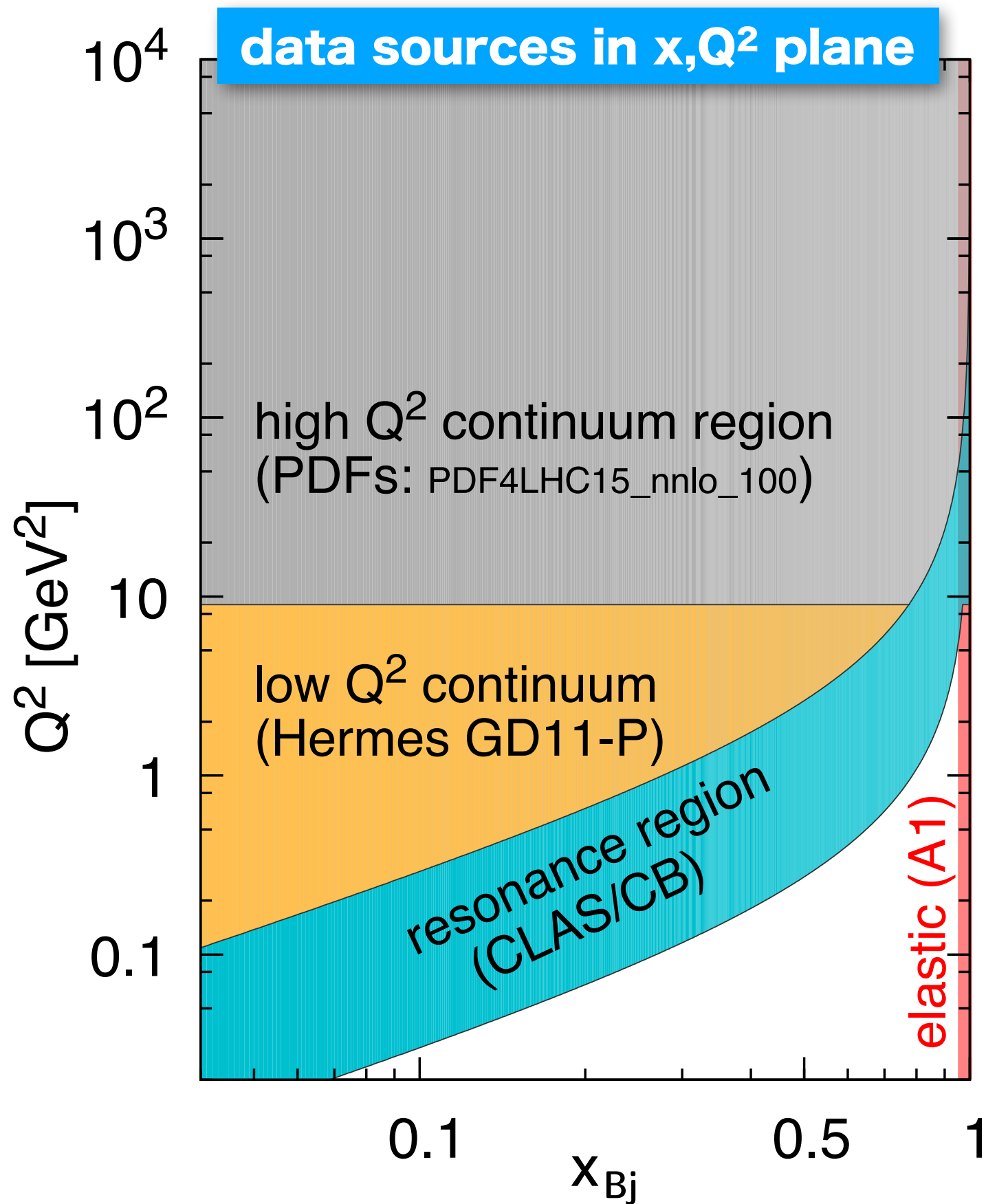
$$P_{\gamma\gamma}^{(1,1)} = (2\pi)^2 b_\alpha^{(1,2)} \delta(1-x) = -C_F N_C \sum_q e_q^2 \delta(1-x)$$

$$h(z) \equiv z \text{ and } \bar{p}_{\gamma q}(z) \equiv p_{\gamma q}(z) \ln \frac{1}{1-z}$$

- These **agree with de Florian, Sborlini & Rodrigo results**

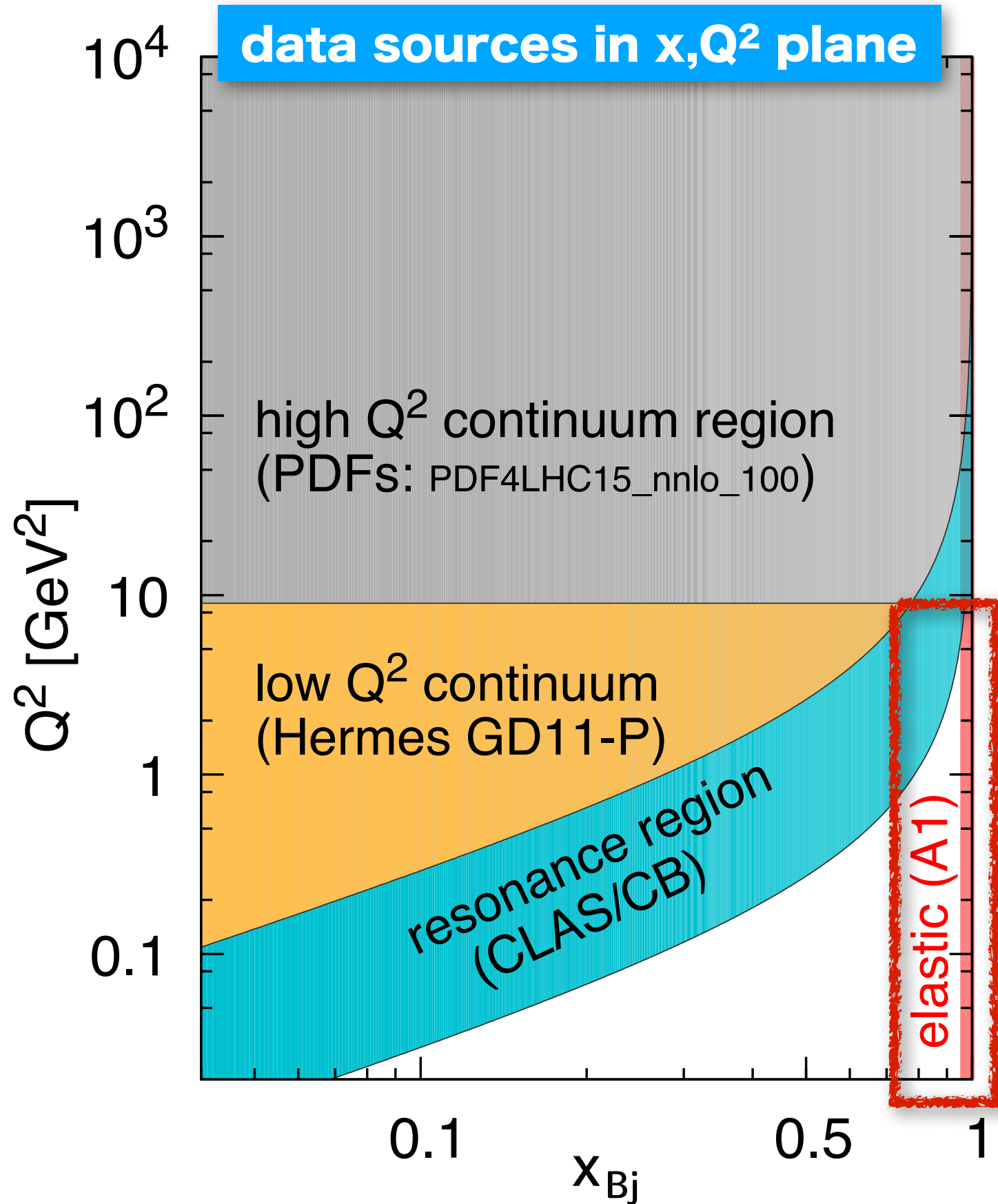
*for  $O(\alpha \alpha_s)$  terms, arXiv:1512.00612*

**data inputs**

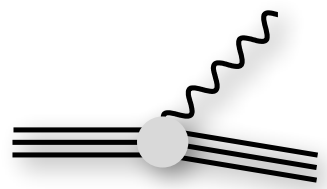


## DATA

- $x, Q^2$  plane naturally breaks up into regions with different physical behaviours and data sources
- We don't use  $F_2$  and  $F_L$  data directly, but rather various fits to data



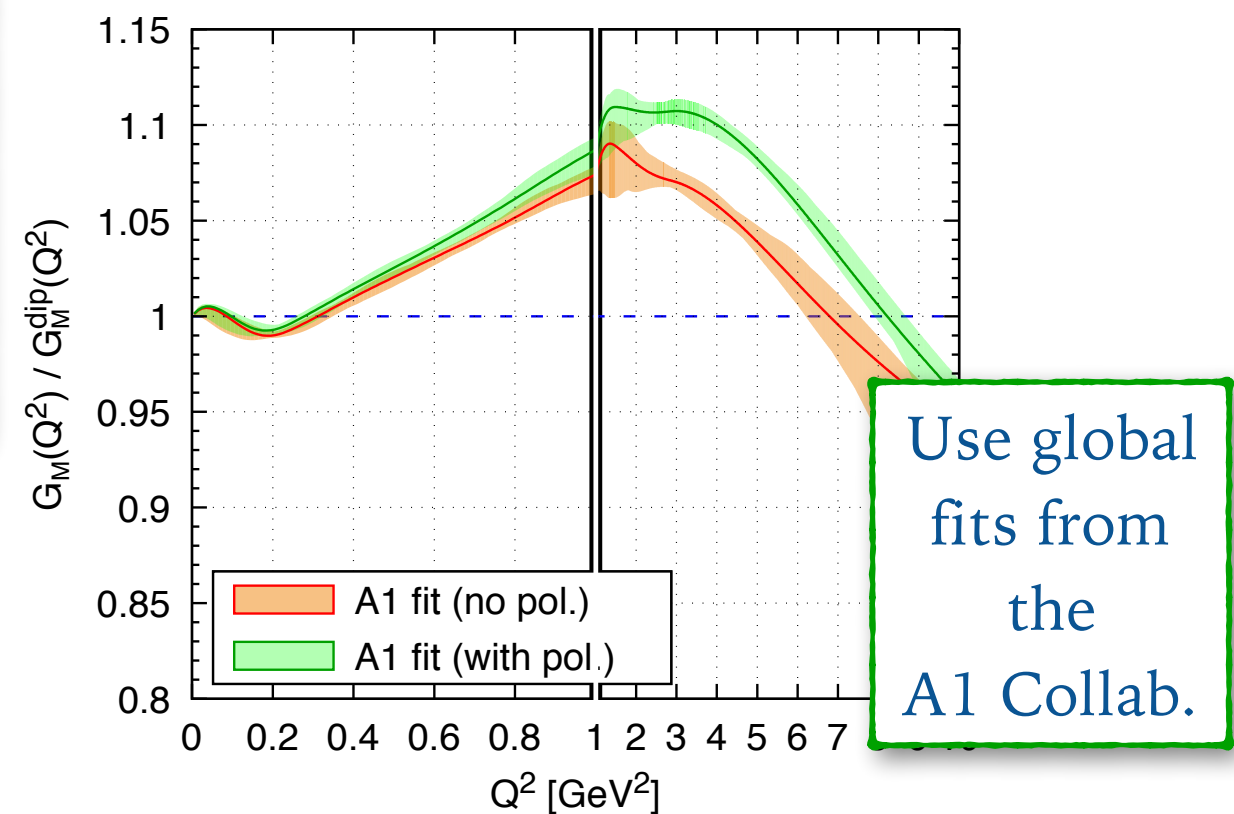
## ELASTIC COMPONENT



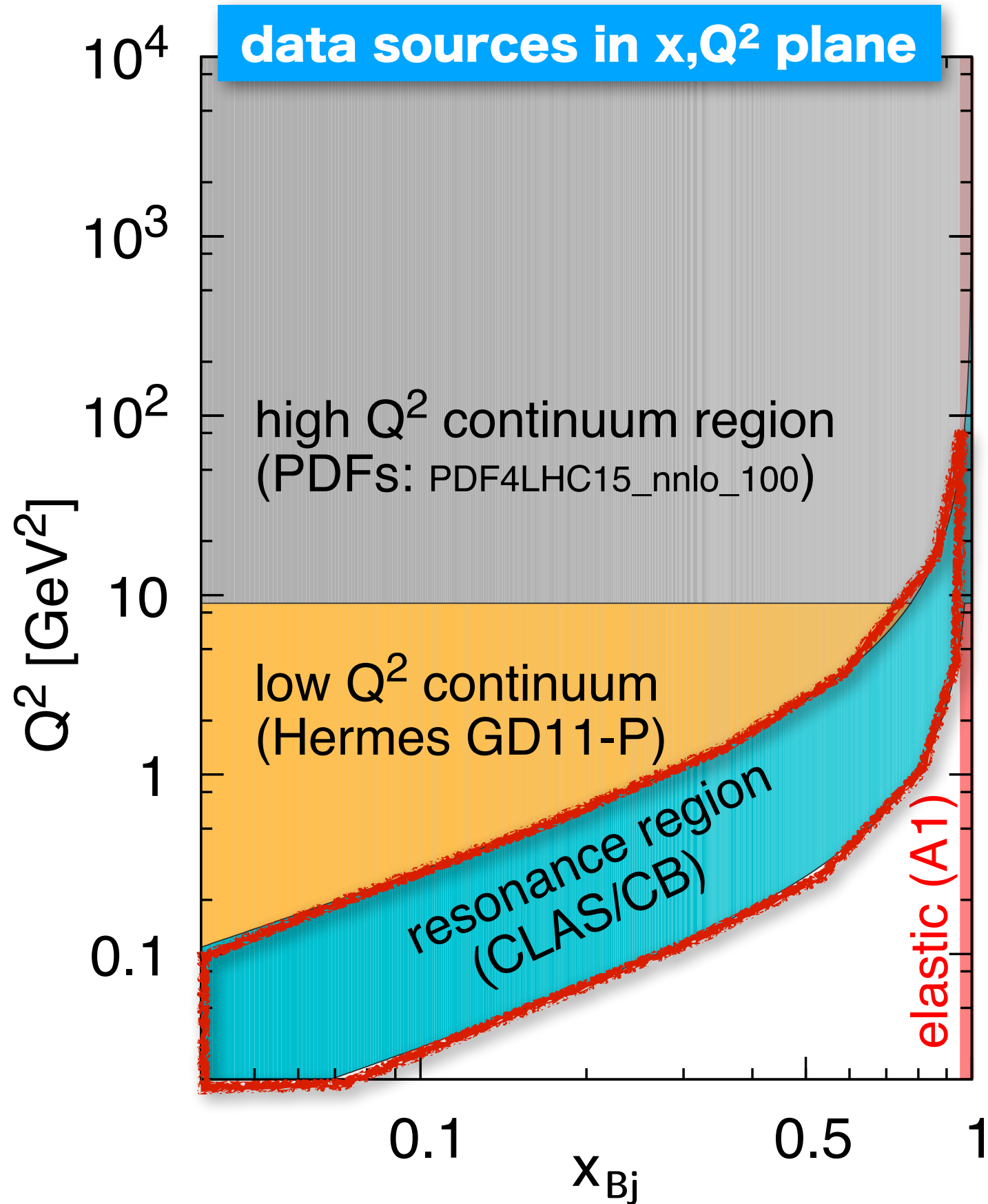
- ▶ Elastic component of  $F_{2/L}$  lives at  $x=1$
- ▶ Express in terms of Sachs Form factors

$$F_2^{\text{el}} = \frac{[G_E(Q^2)]^2 + [G_M(Q^2)]^2 \tau}{1 + \tau} \delta(1 - x),$$

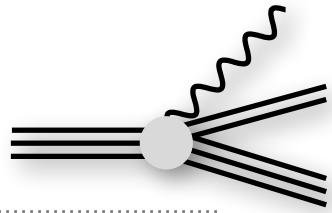
$$F_L^{\text{el}} = \frac{[G_E(Q^2)]^2}{\tau} \delta(1 - x), \quad \tau = Q^2 / (4m_p^2)$$



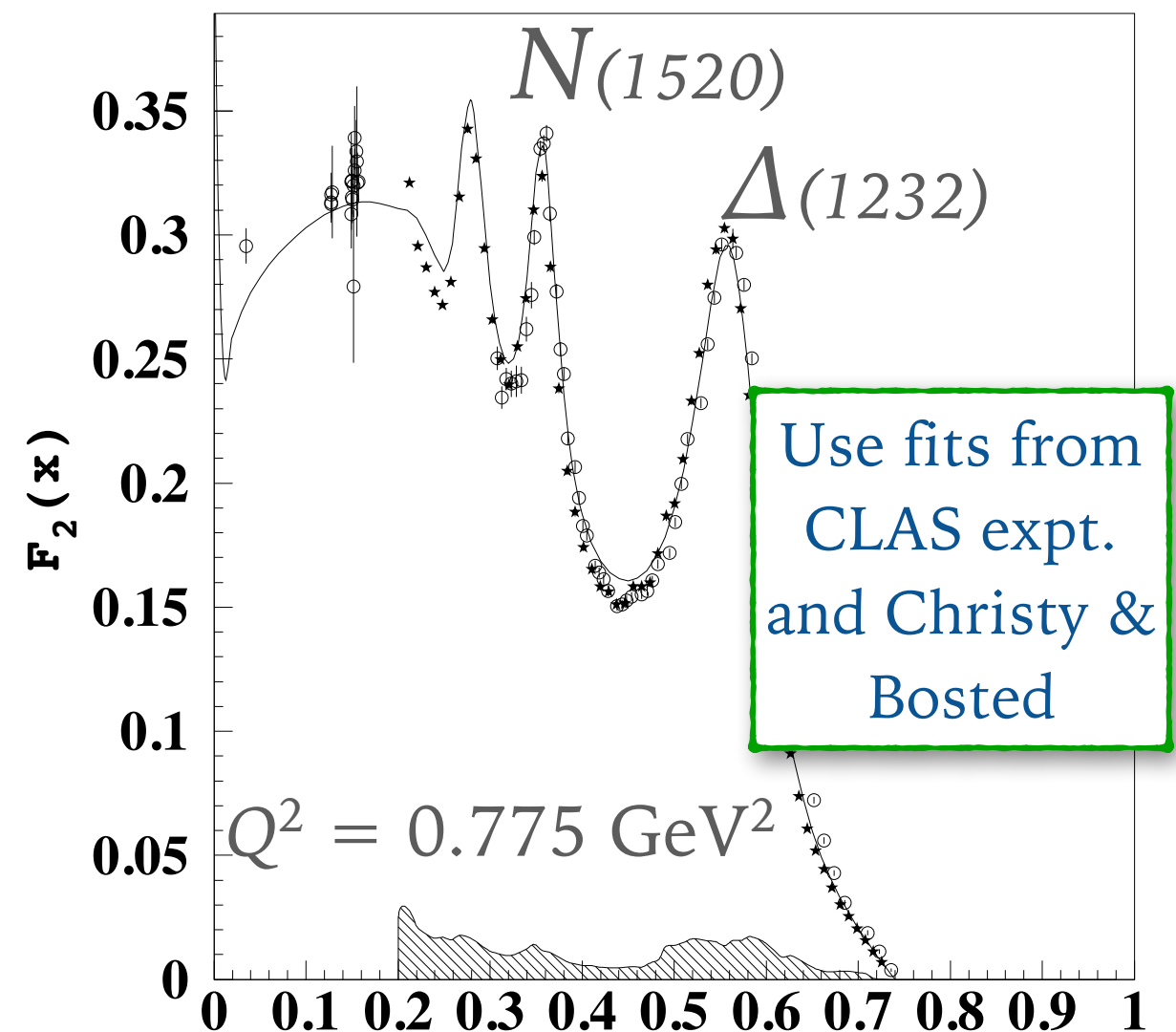


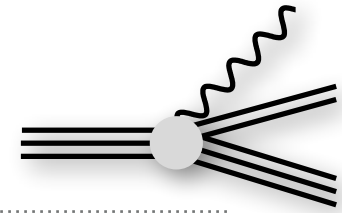


## RESONANCE COMPONENT



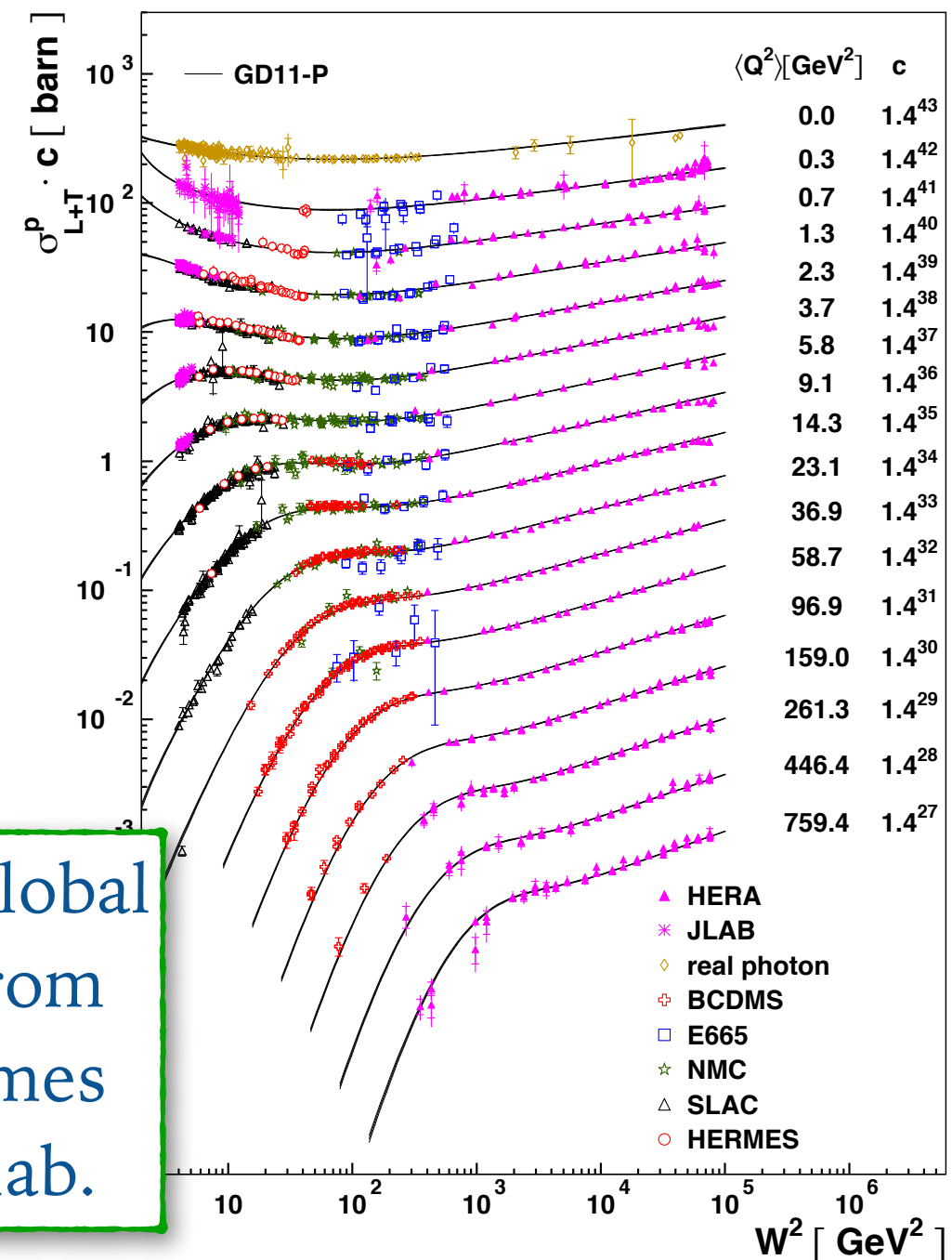
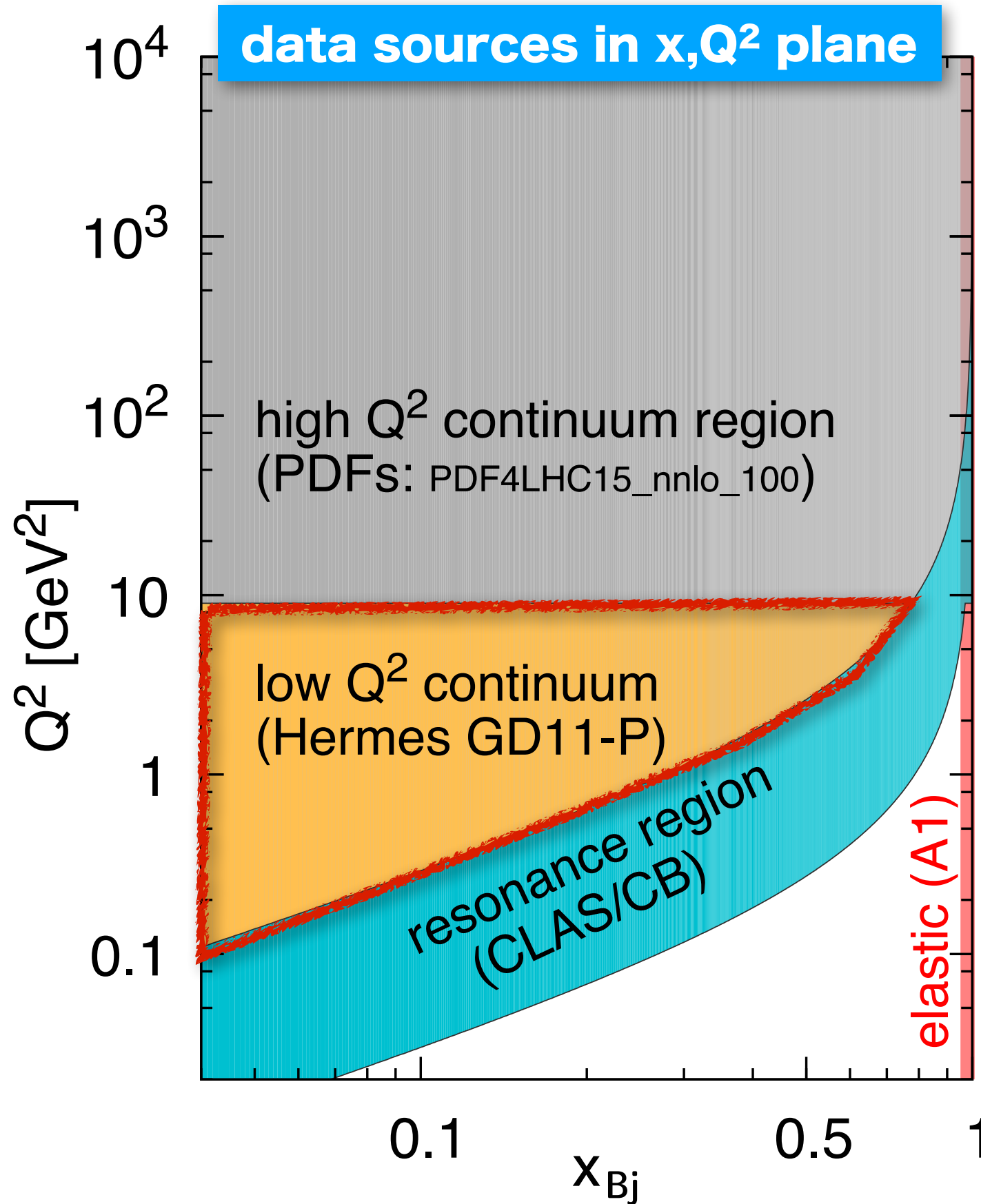
- ▶ proton gets excited, e.g. to  $\Delta \rightarrow p\pi$  and higher resonances
- ▶ relevant for  $(m_p + m_\pi)^2 < W^2 < 3.5 \text{ GeV}^2$



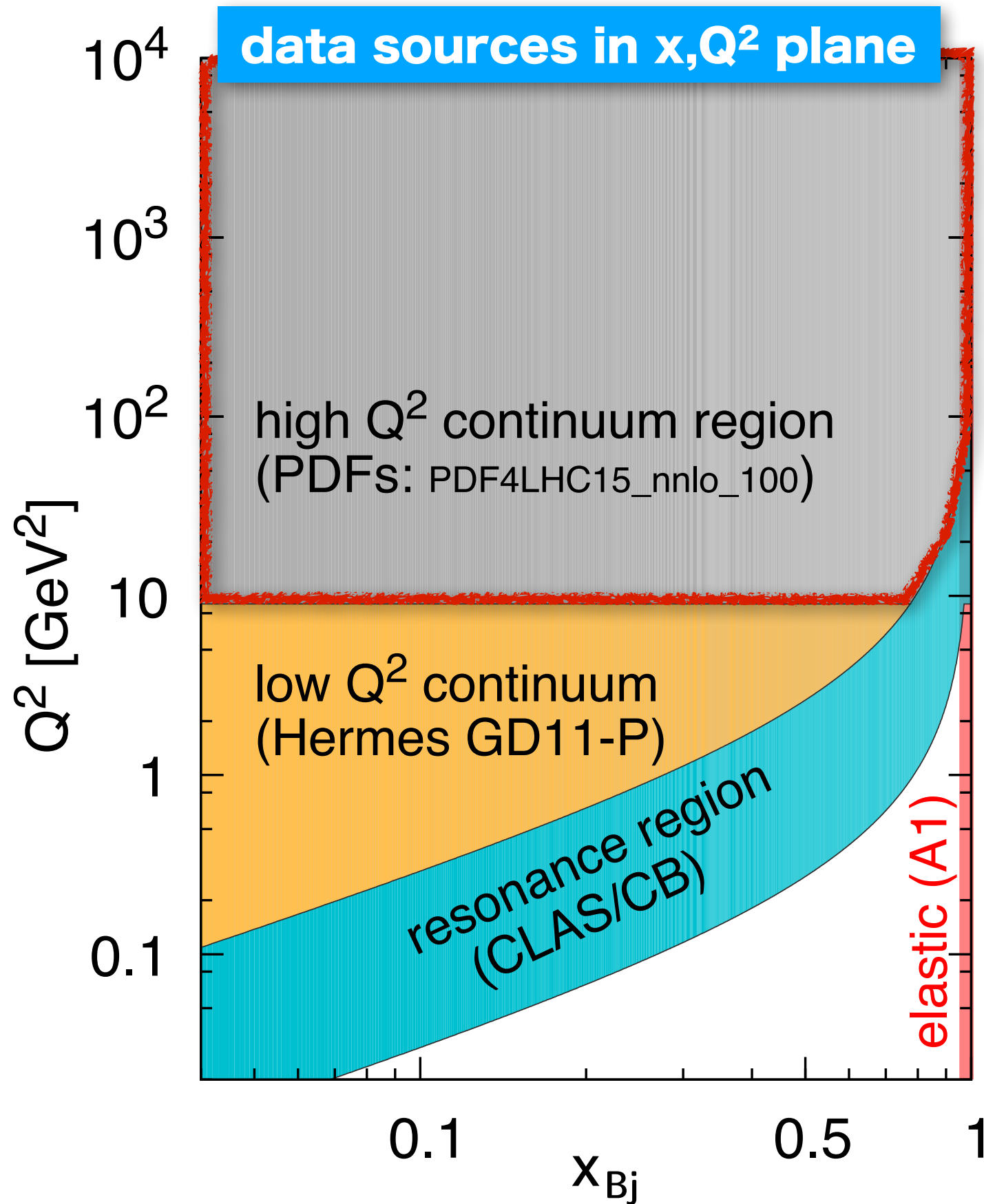


# CONTINUUM COMPONENT

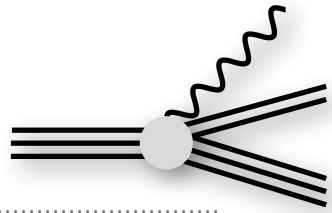
- Much data
- For  $Q^2 \rightarrow 0$ ,  $\sigma_{\text{yp}}$  indep. of  $Q^2$  at fixed  $W^2$



Use global fit from Hermes Collab.

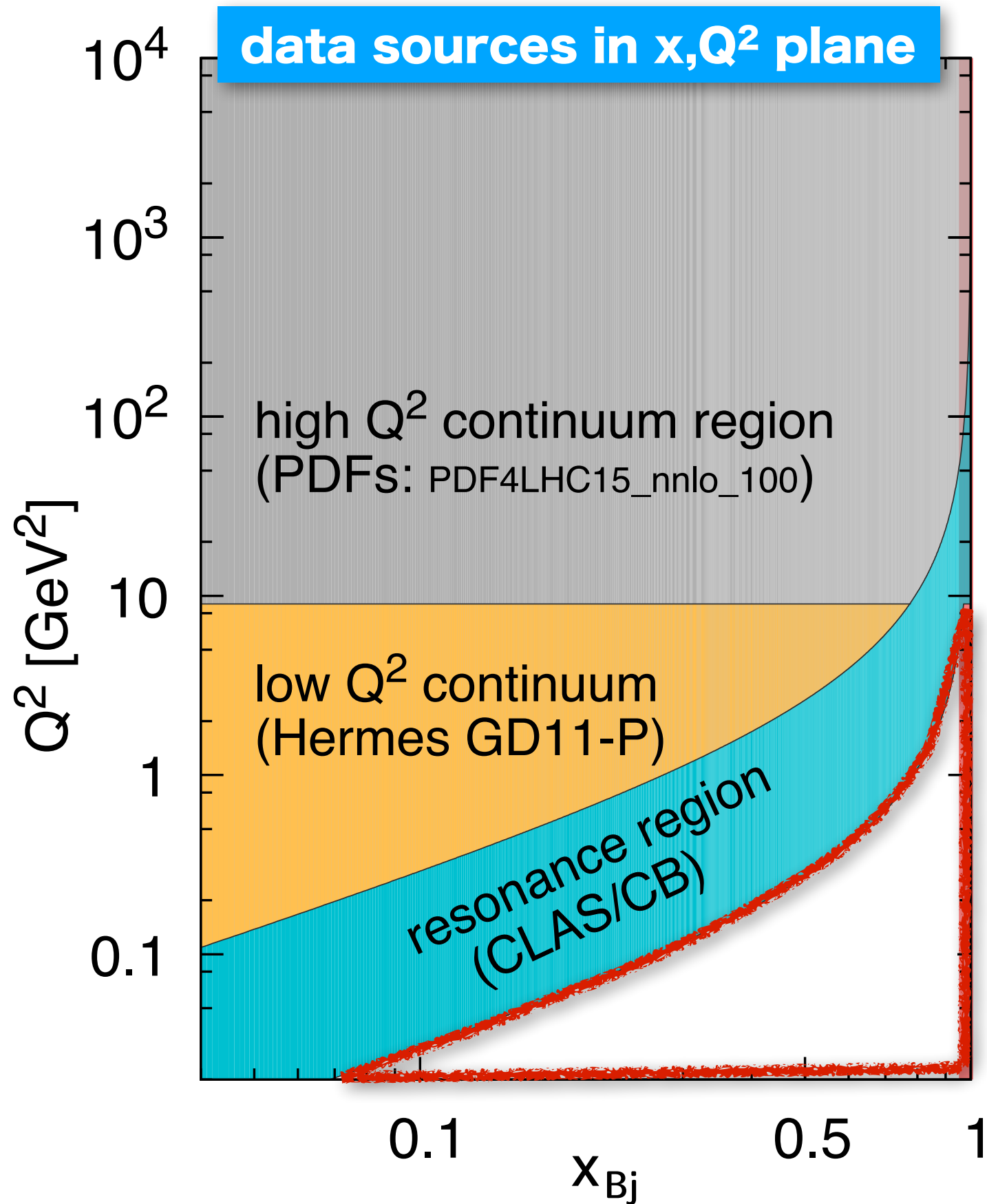


## CONTINUUM COMPONENT

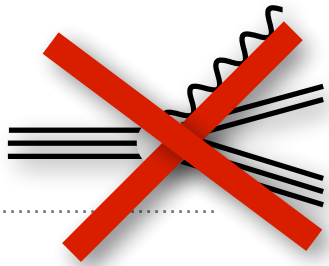


- Less direct data for  $F_2$  and  $F_L$  at high  $Q^2$
- But we can reliably use PDFs and coefficient functions (up to NNLO) to calculate them
- Our default choice is PDF4LHC15\_nnlo\_100 (and zero-mass variable flavour-number scheme)

As a PDF we use  
**PDF4LHC15\_nnlo\_100**  
from LHAPDF



## EMPTY AREA

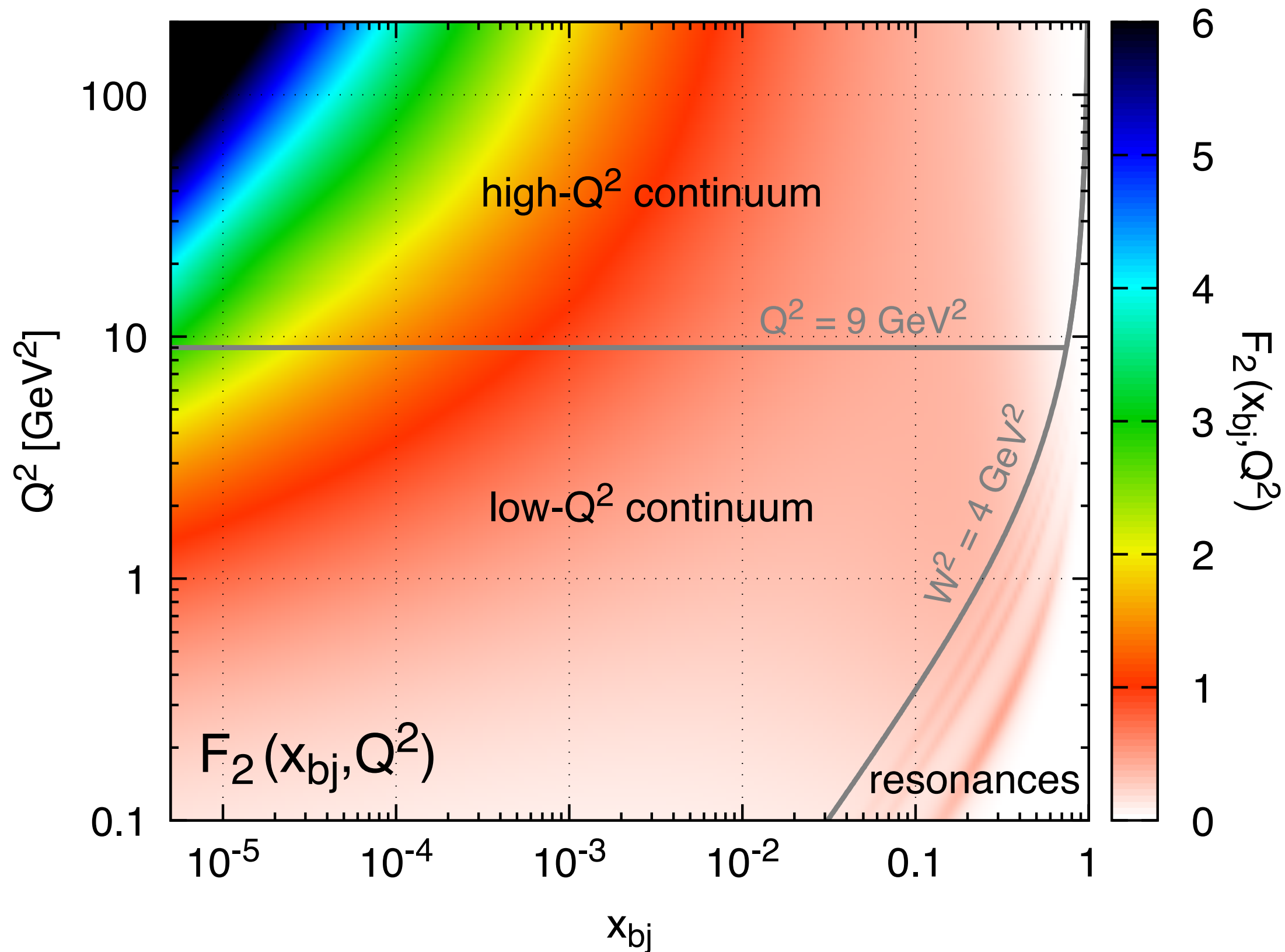


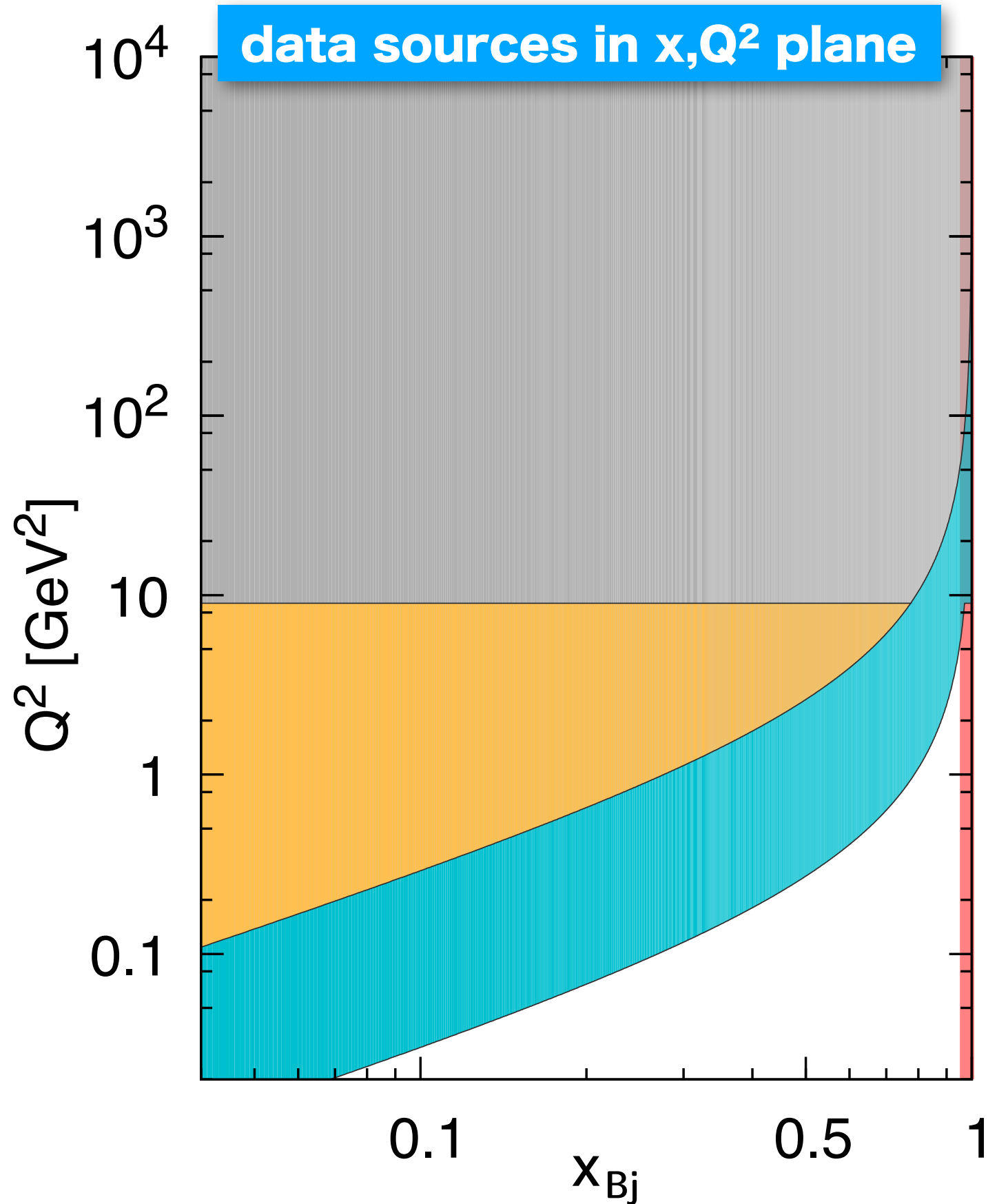
- ▶ kinematically inaccessible region: hadronic final-state mass  $W$  in range

$$m_p < W < m_p + m_\pi$$

- ▶ i.e. the QCD mass gap
- ▶ [at higher order in QED, beyond our accuracy, can be filled with photon radiation]

# Final assembly of $F_2(x, Q^2)$

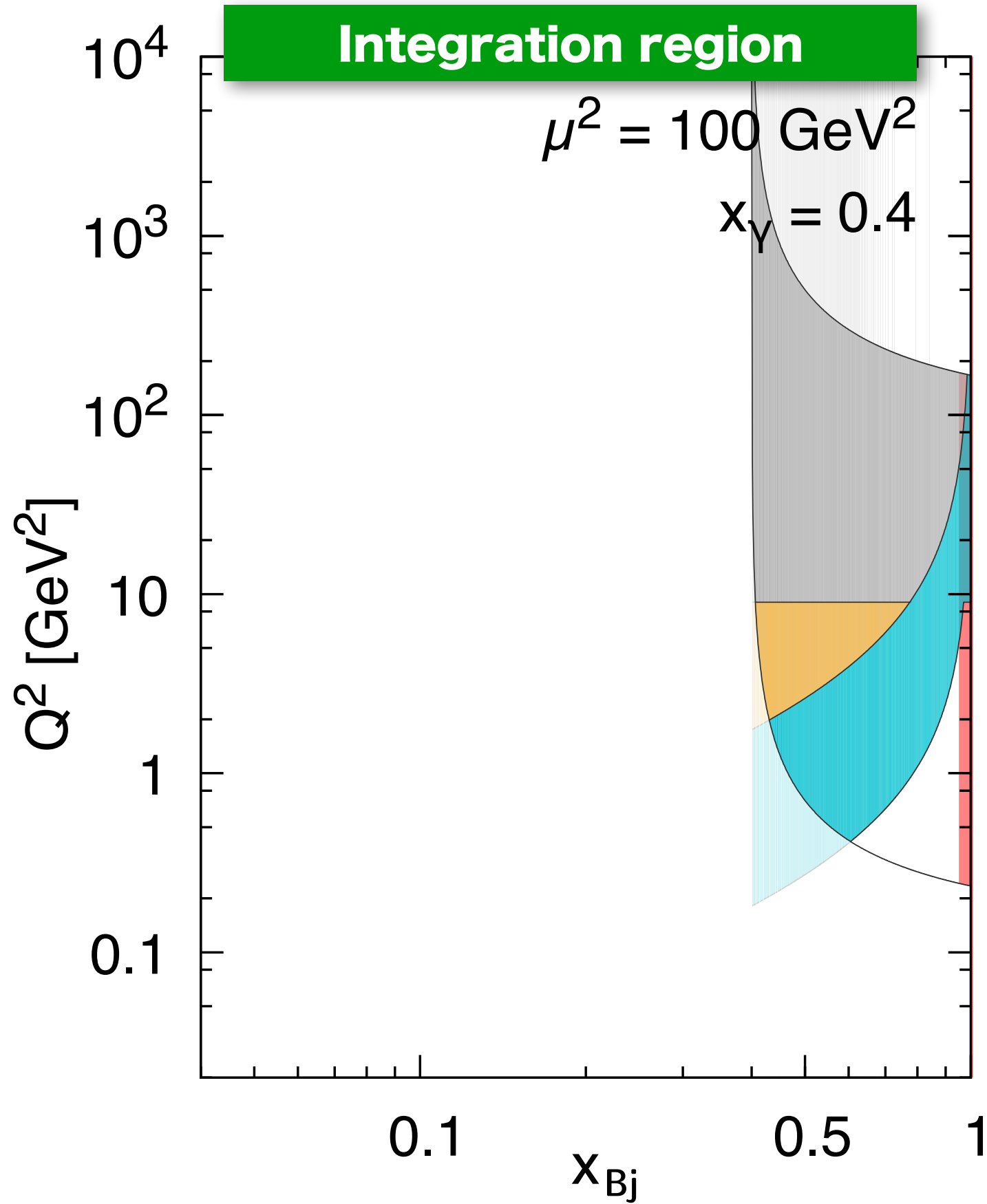




## INTEGRATION REGION

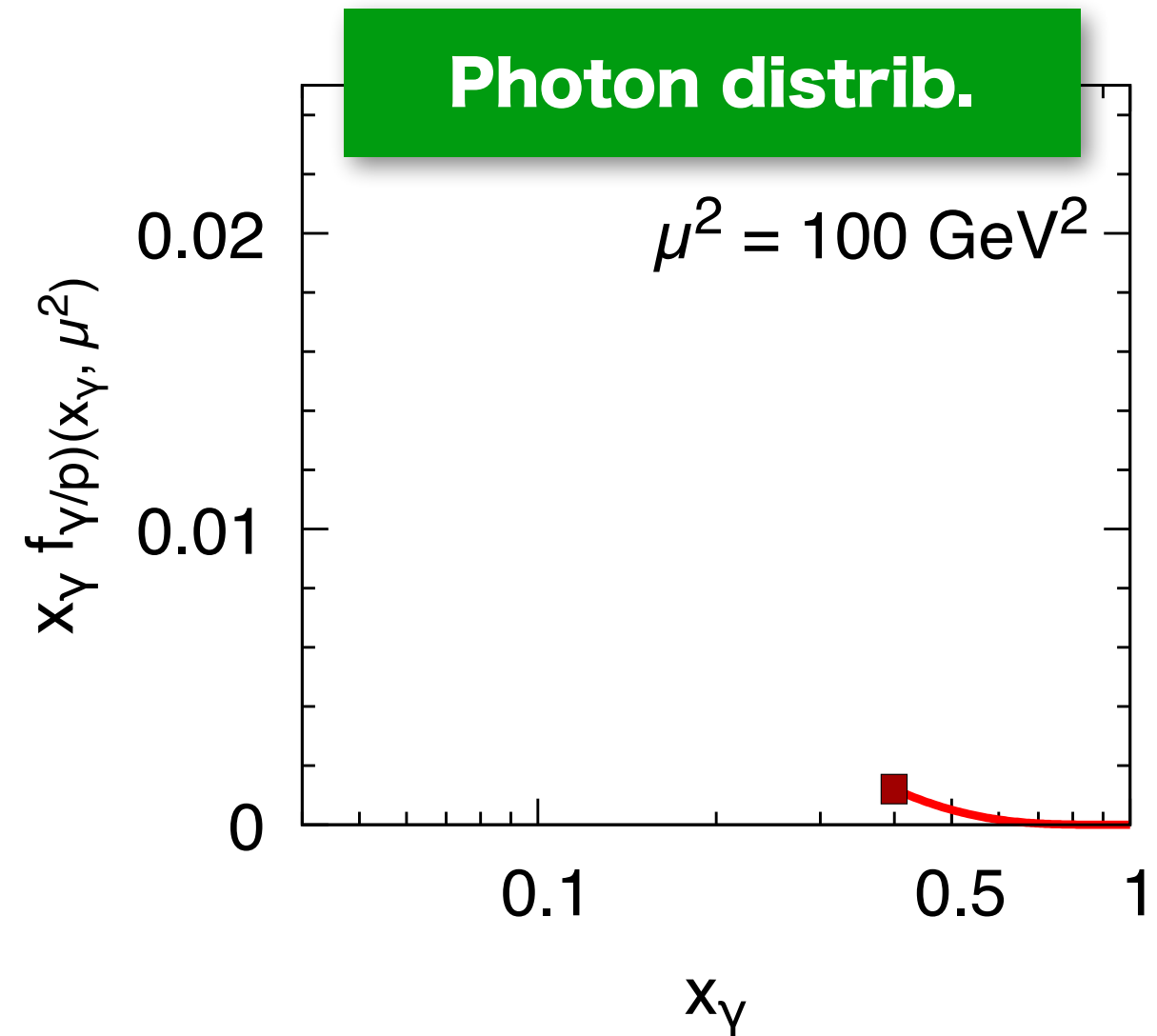
- depends on momentum fraction of the photon ( $x_\gamma$ ) and factorisation scale ( $\mu^2$ )

$$x f_{\gamma/p}(x, \mu^2) = \frac{1}{2\pi\alpha(\mu^2)} \int_x^1 \frac{dz}{z} \left\{ \int_{\frac{x^2 m_p^2}{1-z}}^{\frac{\mu^2}{1-z}} \frac{dQ^2}{Q^2} \alpha^2(Q^2) \right. \\ \left. \left[ \left( z p_{\gamma q}(z) + \frac{2x^2 m_p^2}{Q^2} \right) F_2(x/z, Q^2) - z^2 F_L\left(\frac{x}{z}, Q^2\right) \right] \right. \\ \left. - \alpha^2(\mu^2) z^2 F_2\left(\frac{x}{z}, \mu^2\right) \right\}, \quad (6)$$

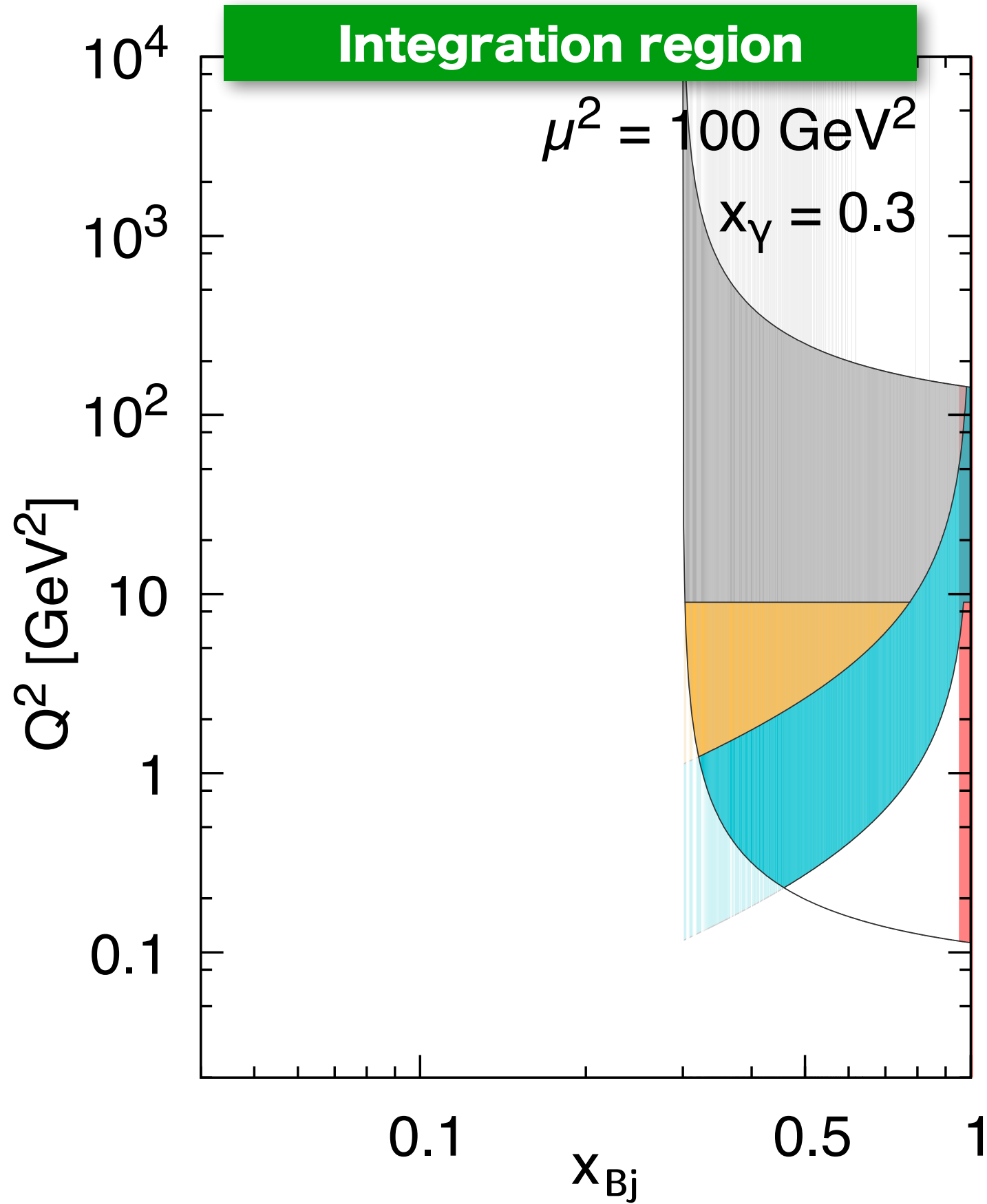


## INTEGRATION REGION

- depends on momentum fraction of the photon ( $x_\gamma$ ) and factorisation scale ( $\mu^2$ )

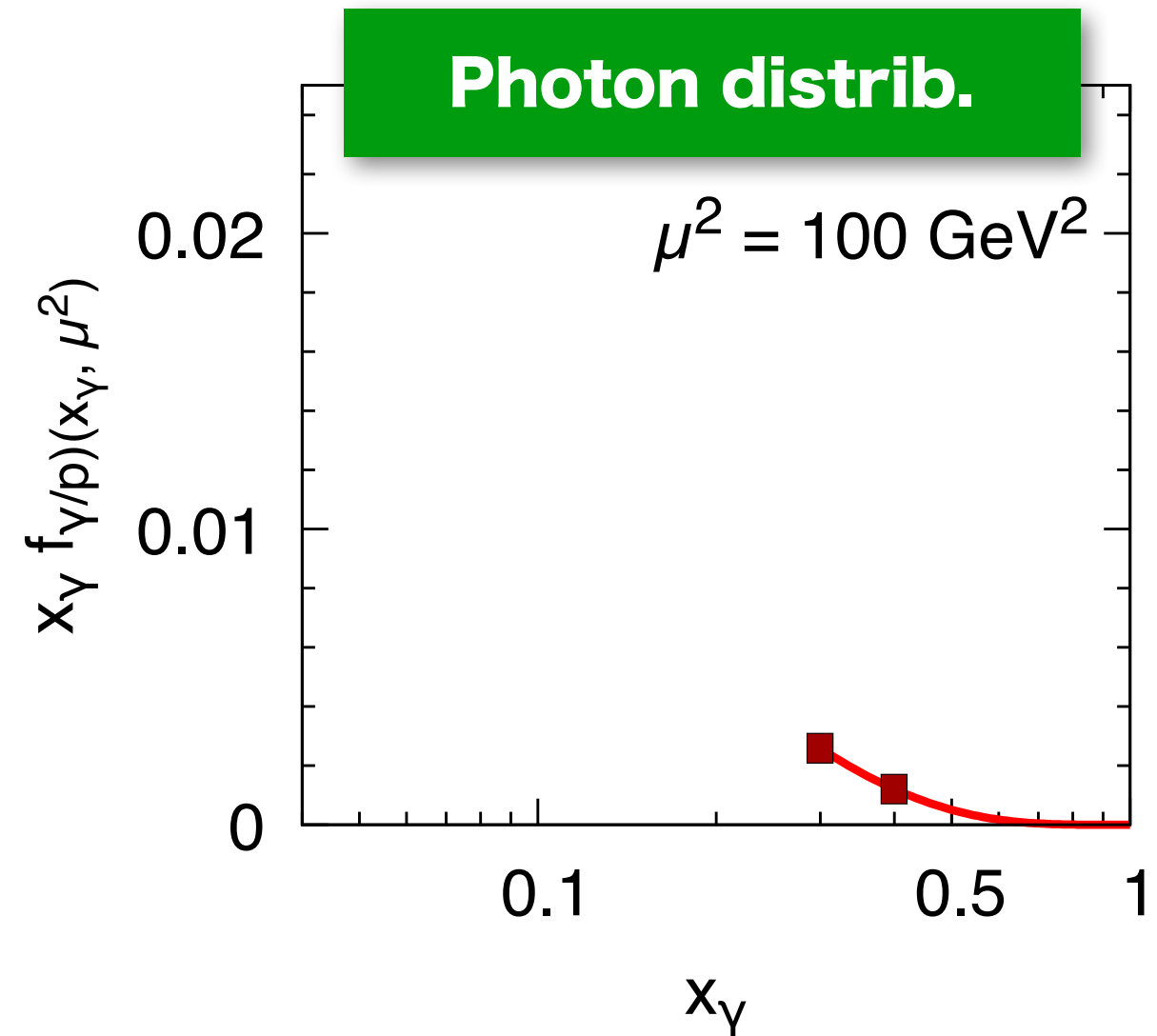




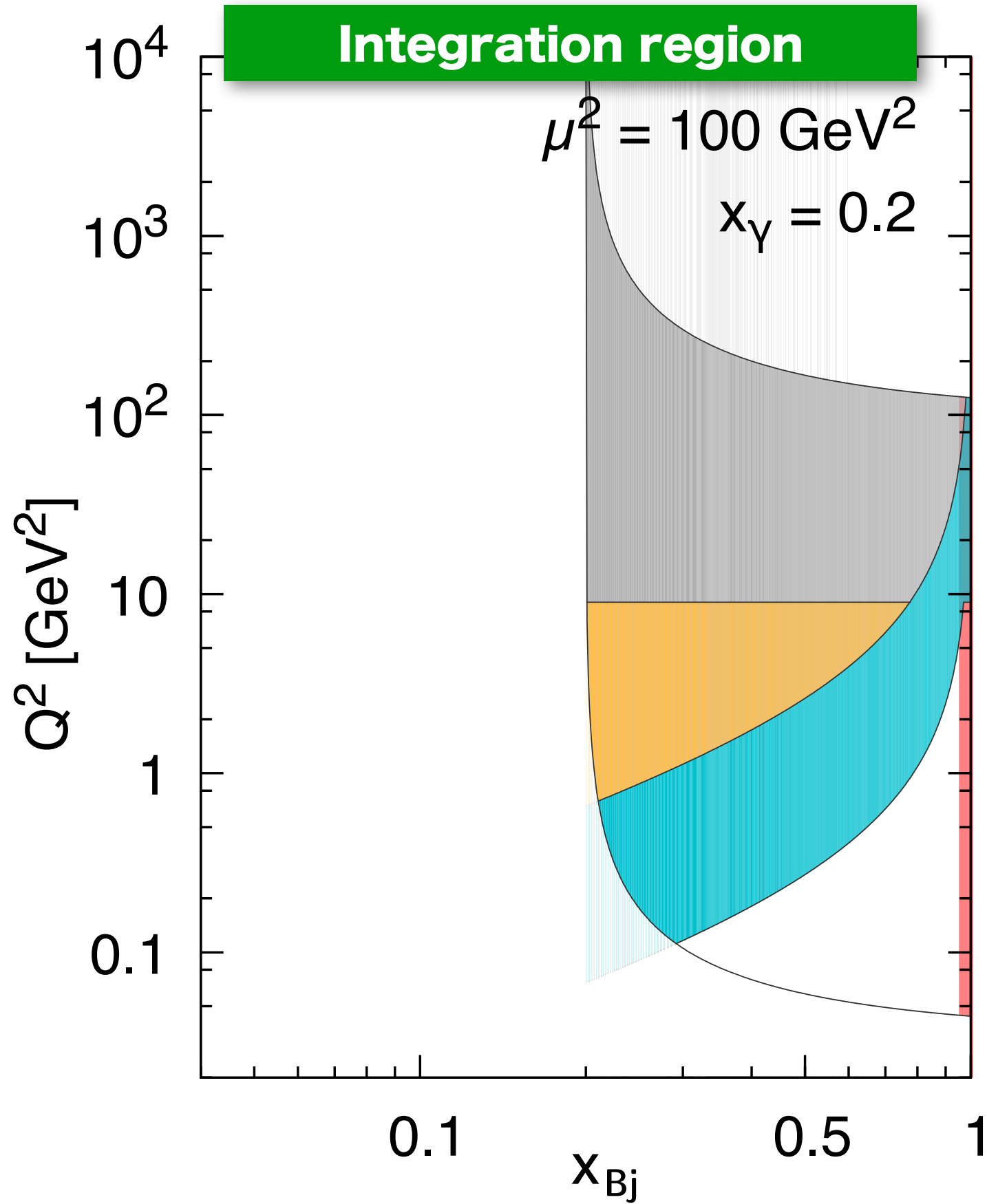


## INTEGRATION REGION

- depends on momentum fraction of the photon ( $x_\gamma$ ) and factorisation scale ( $\mu^2$ )

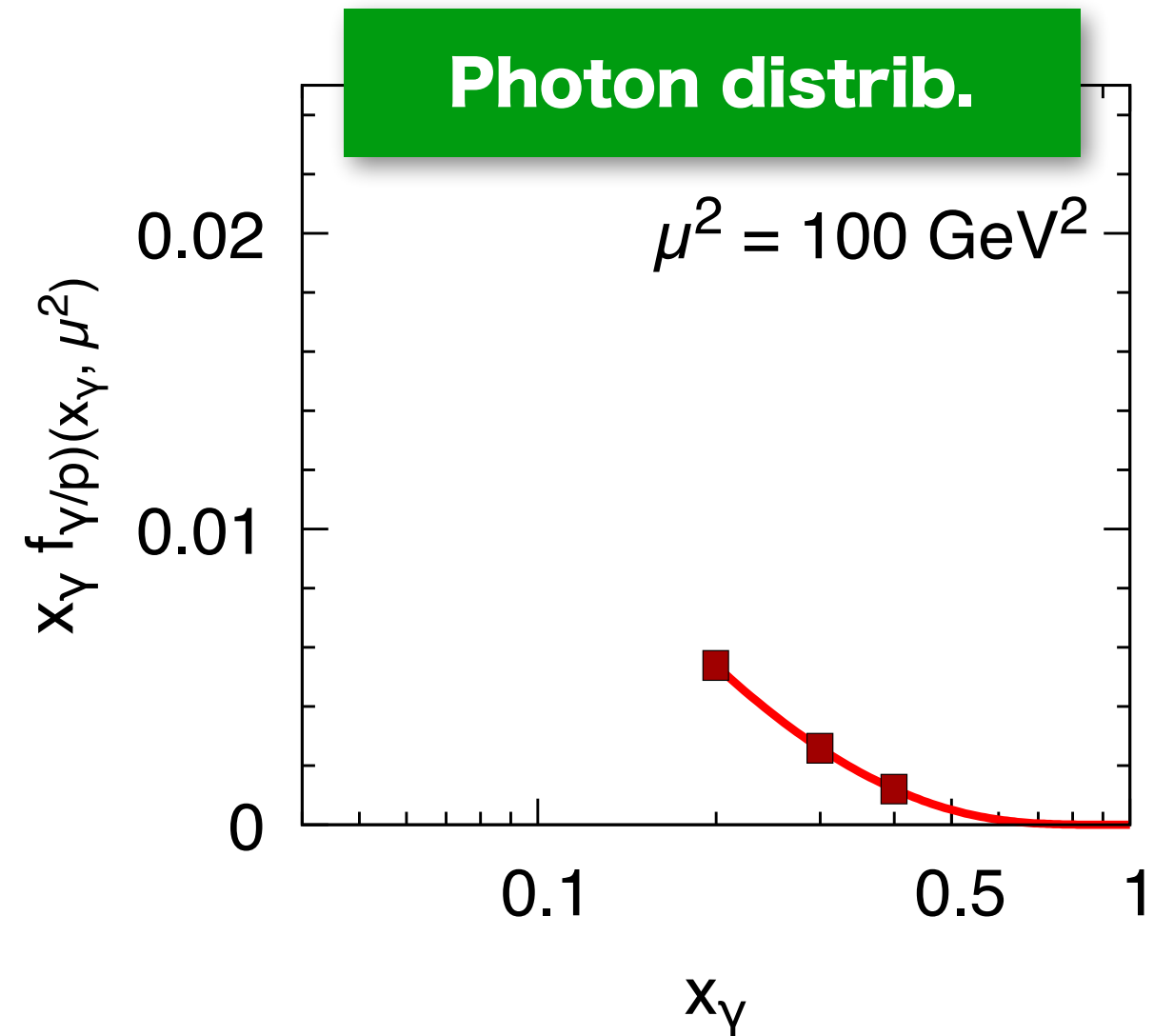


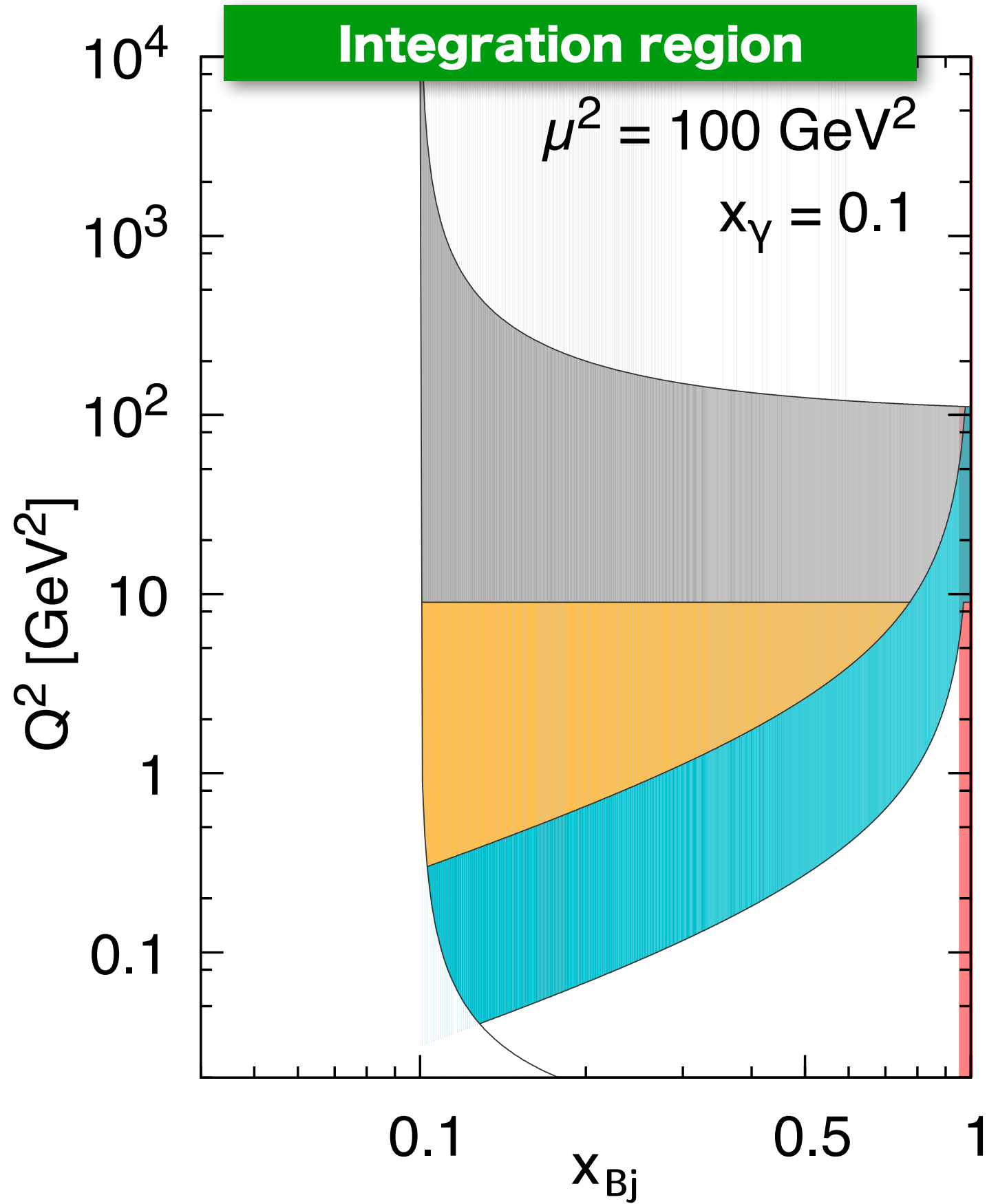




## INTEGRATION REGION

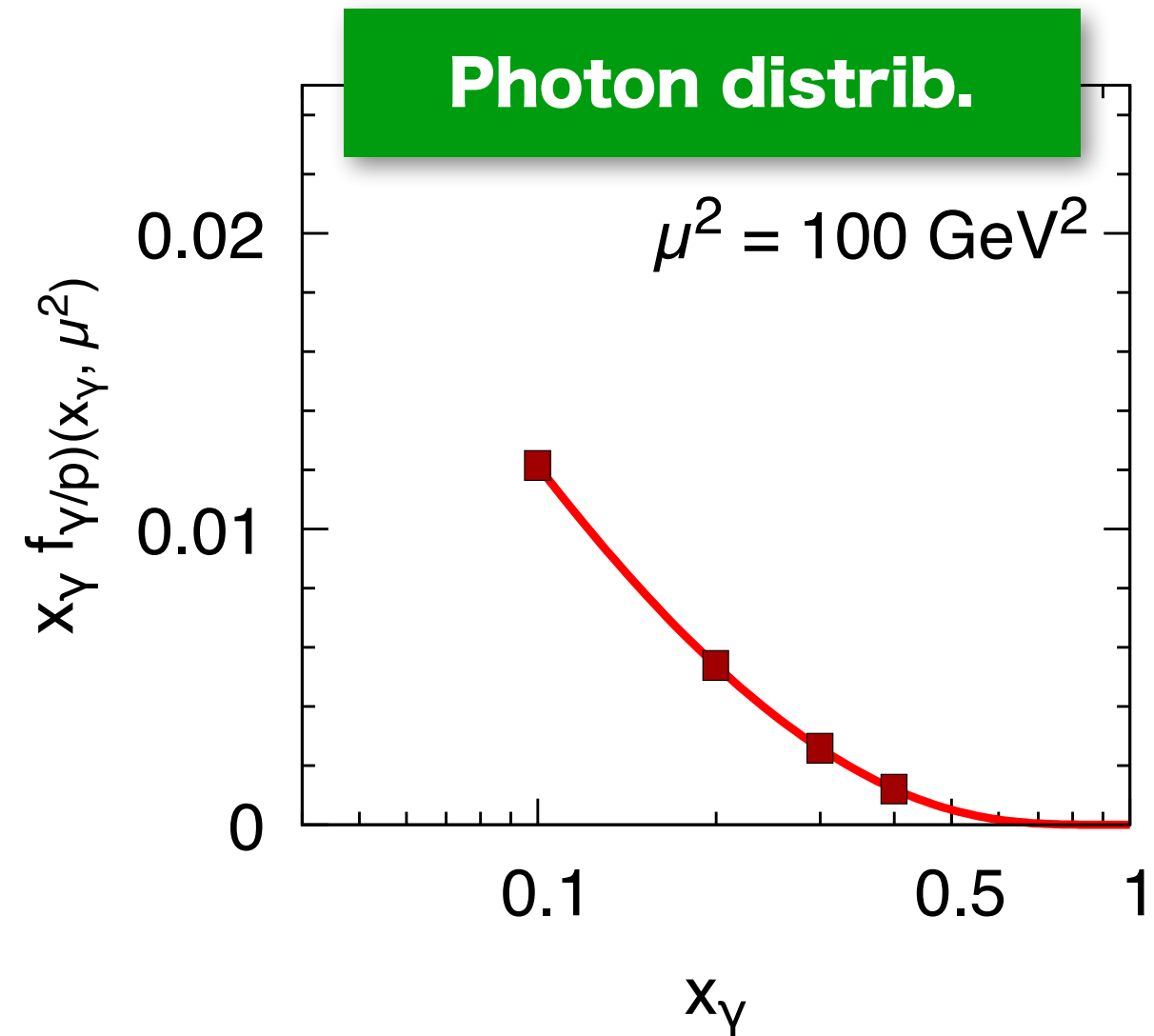
- depends on momentum fraction of the photon ( $x_\gamma$ ) and factorisation scale ( $\mu^2$ )

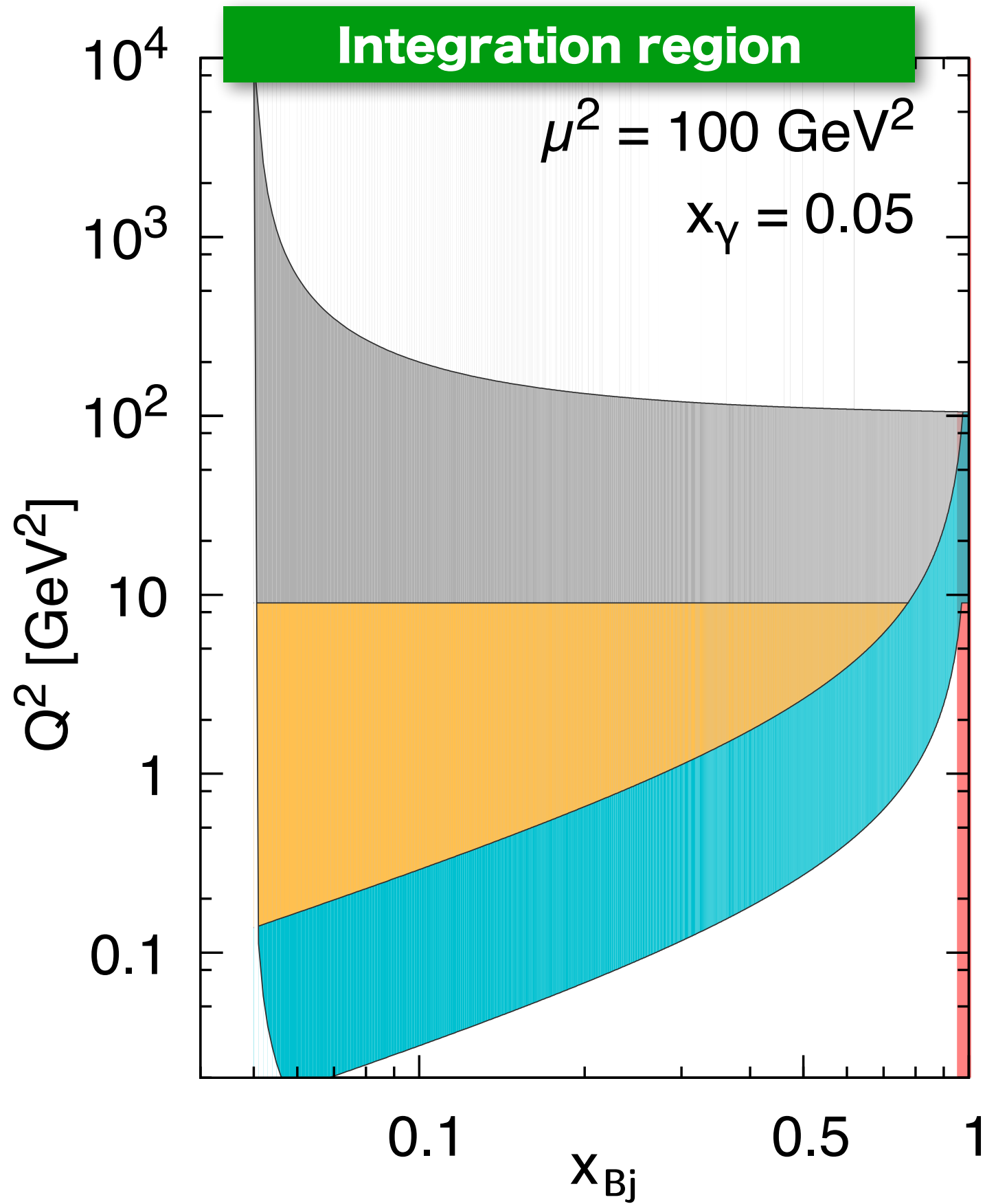




## INTEGRATION REGION

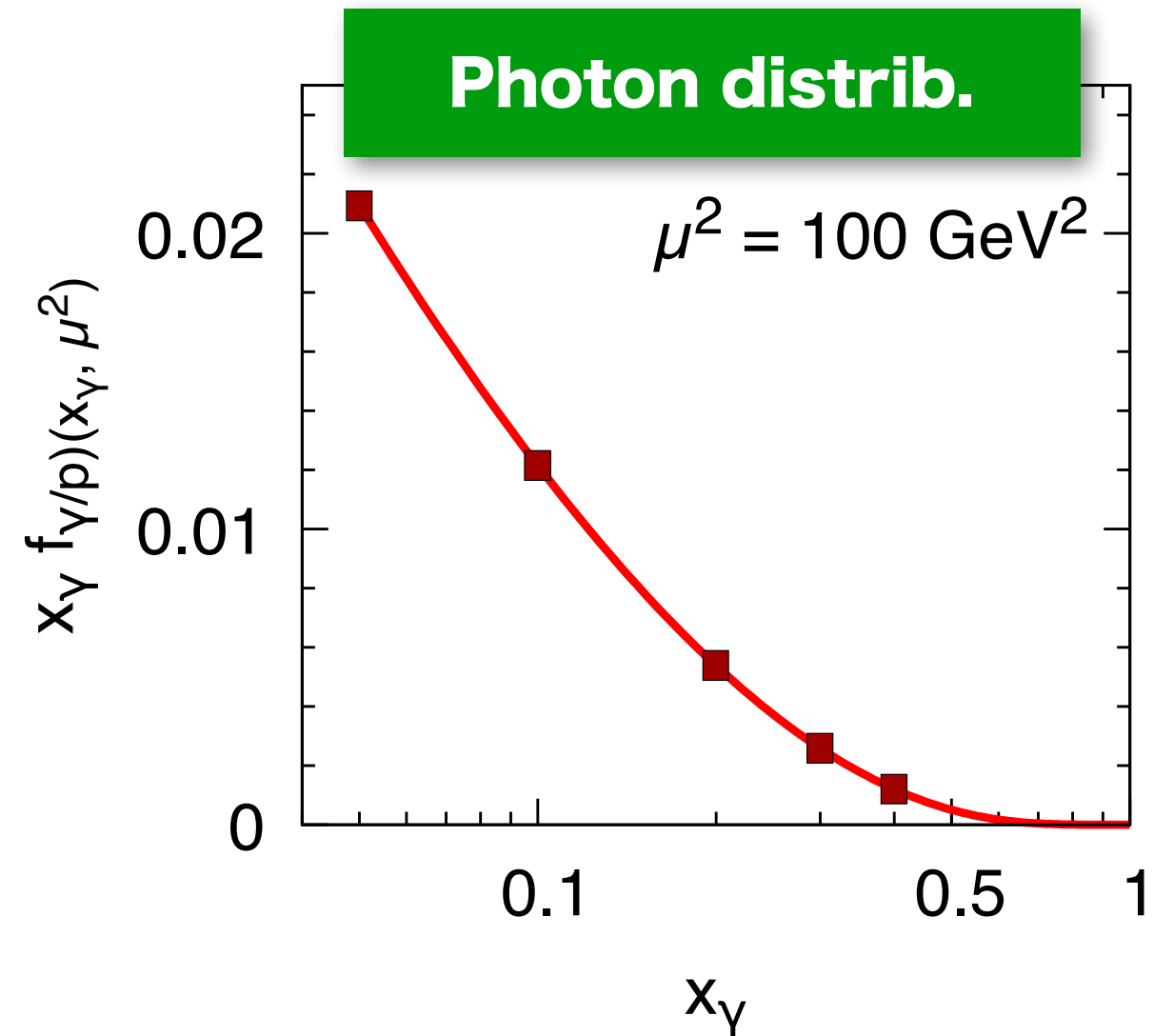
- depends on momentum fraction of the photon ( $x_\gamma$ ) and factorisation scale ( $\mu^2$ )



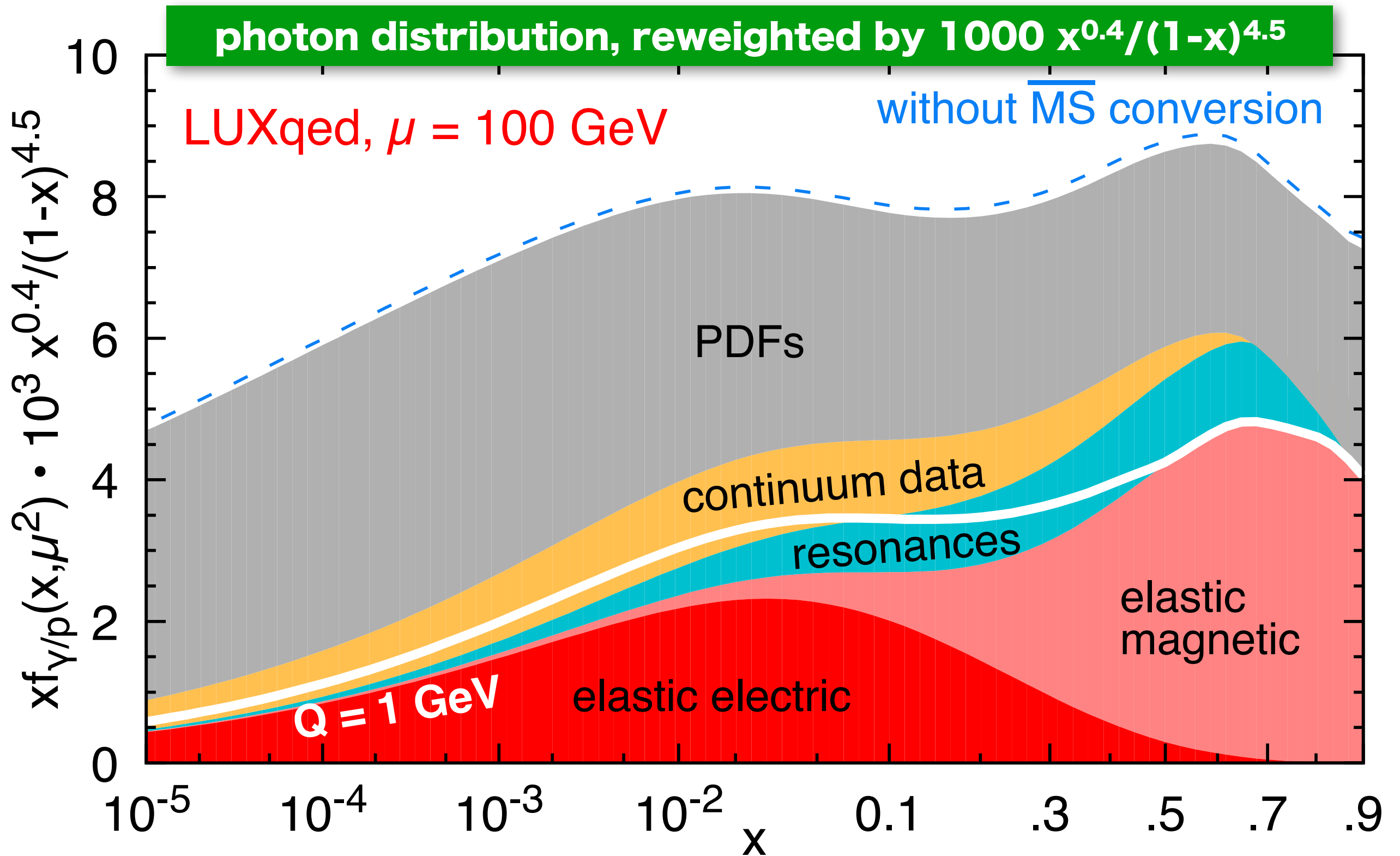


## INTEGRATION REGION

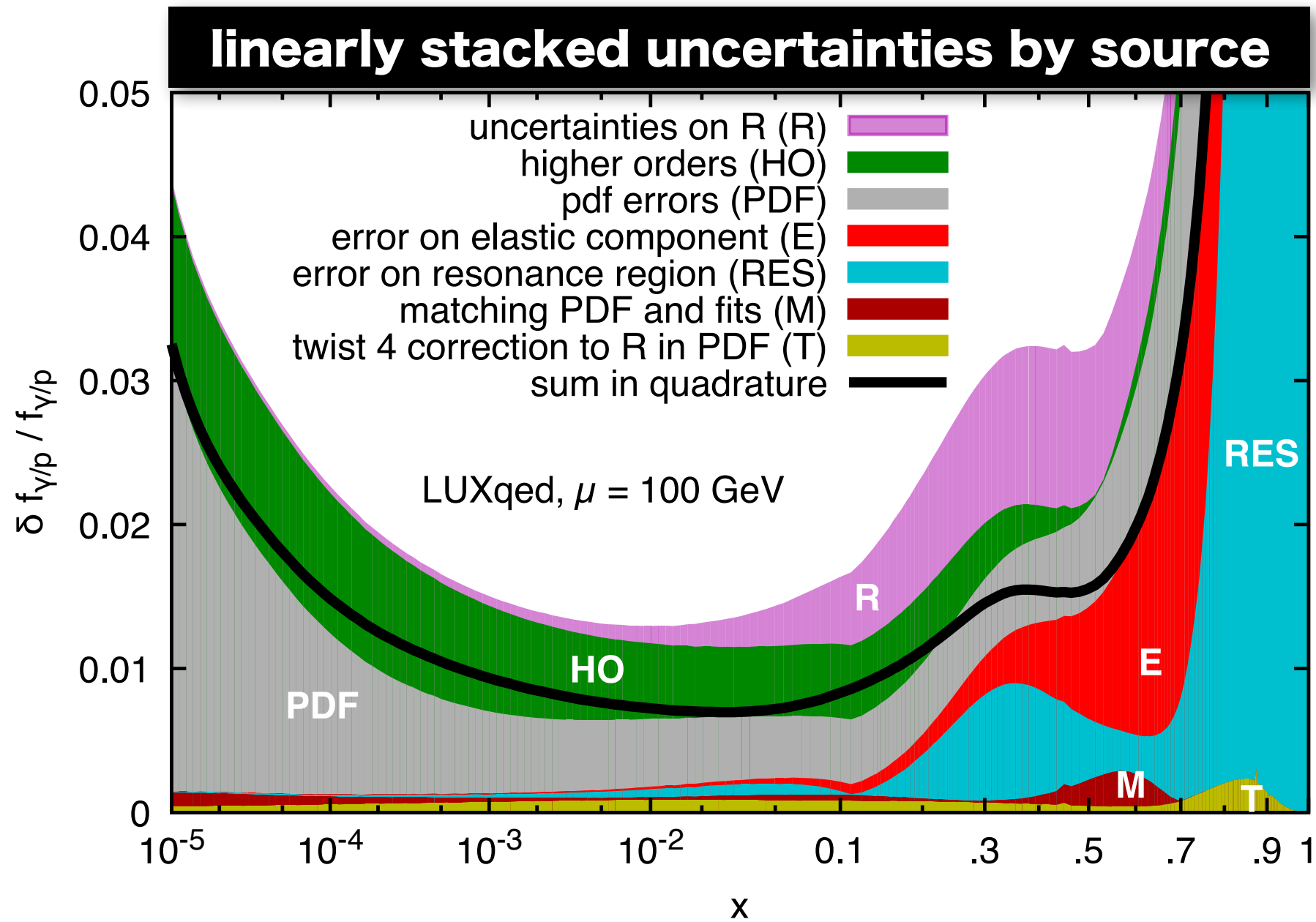
- depends on momentum fraction of the photon ( $x_\gamma$ ) and factorisation scale ( $\mu^2$ )



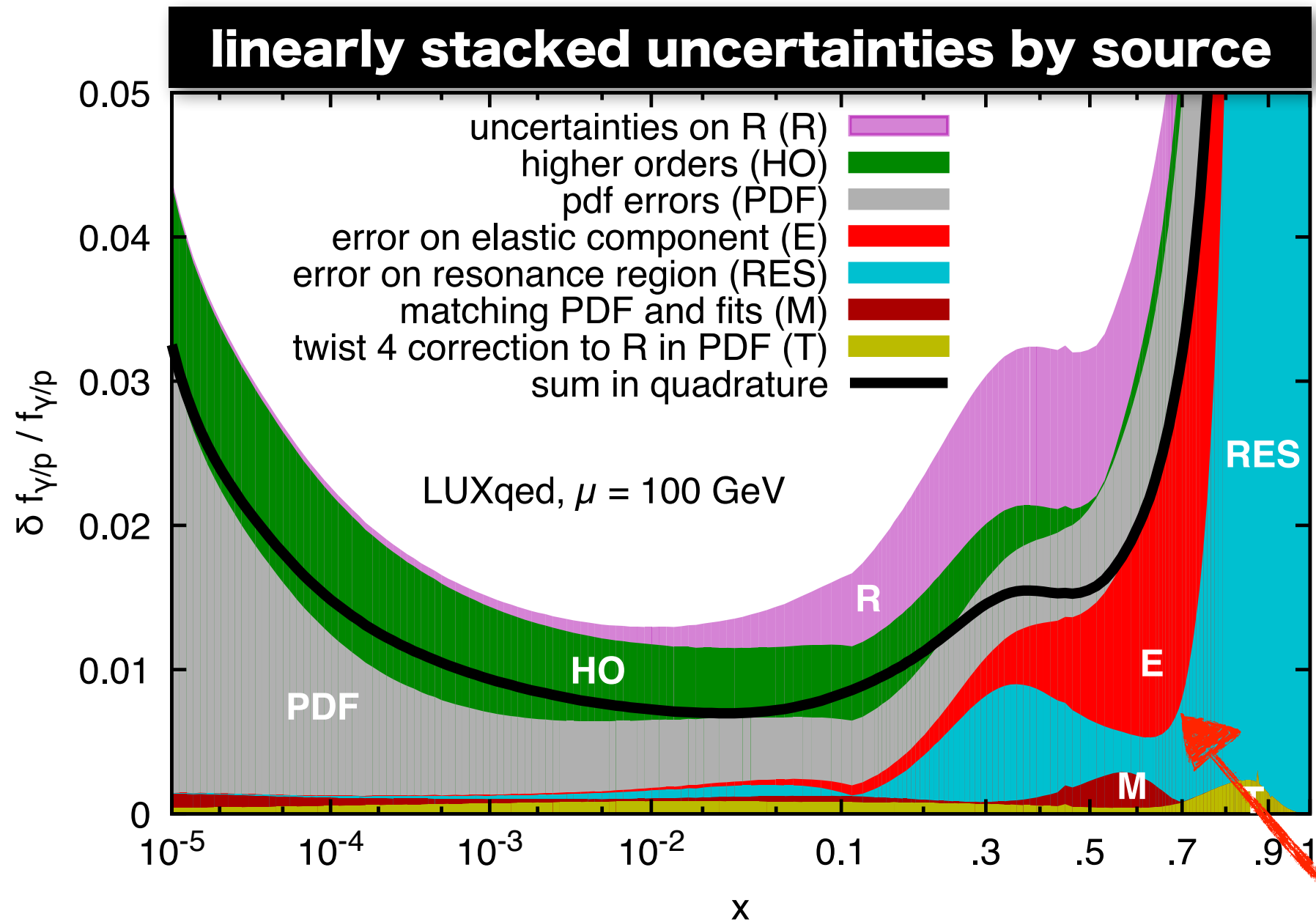
# SEPARATE CONTRIBUTIONS TO PHOTON PDF



# photon uncertainties (aim to be conservative & pragmatic)

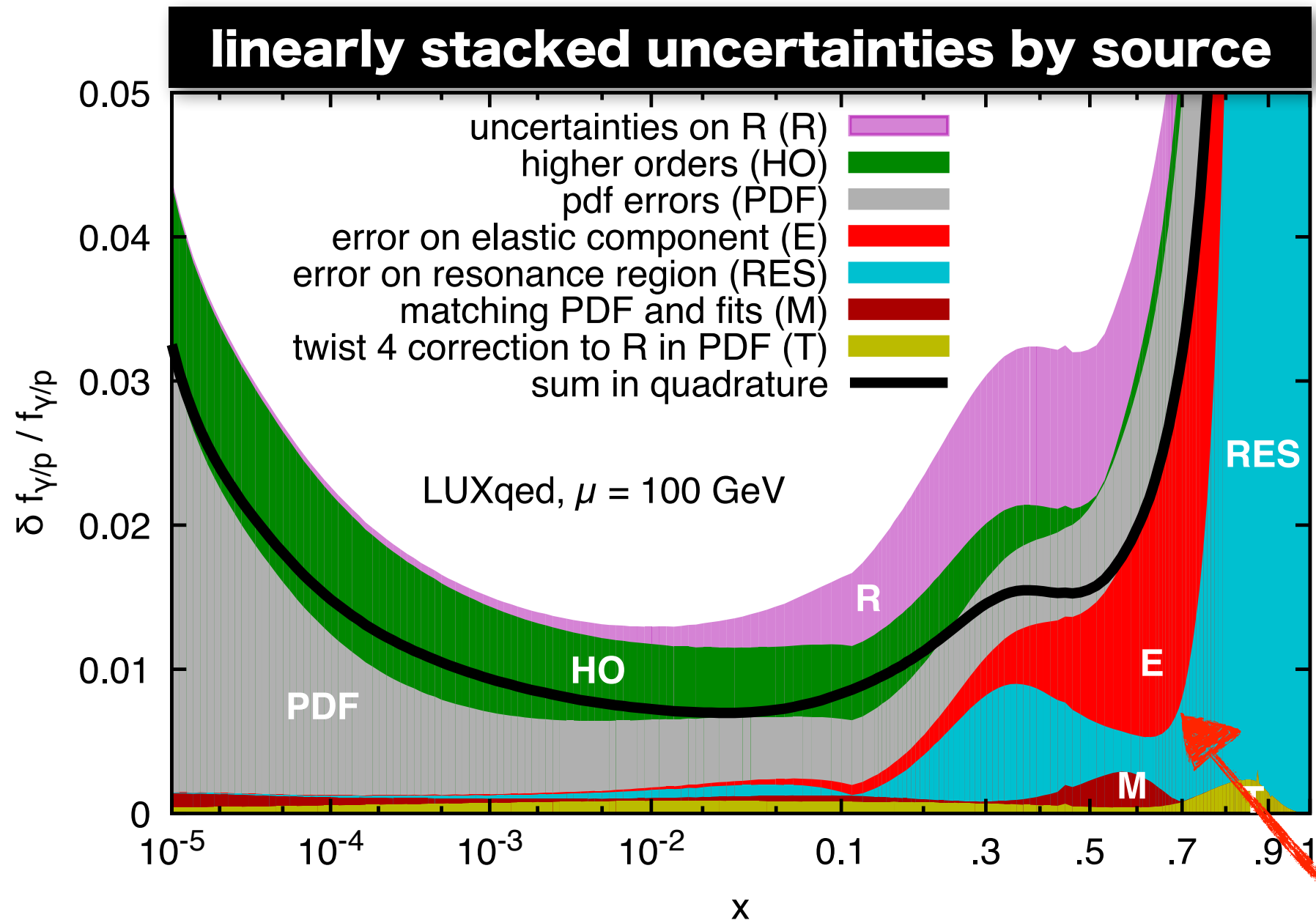


# photon uncertainties (aim to be conservative & pragmatic)



uncertainty on  
elastic component  
(quoted  $\oplus$   
unpol./pol.)

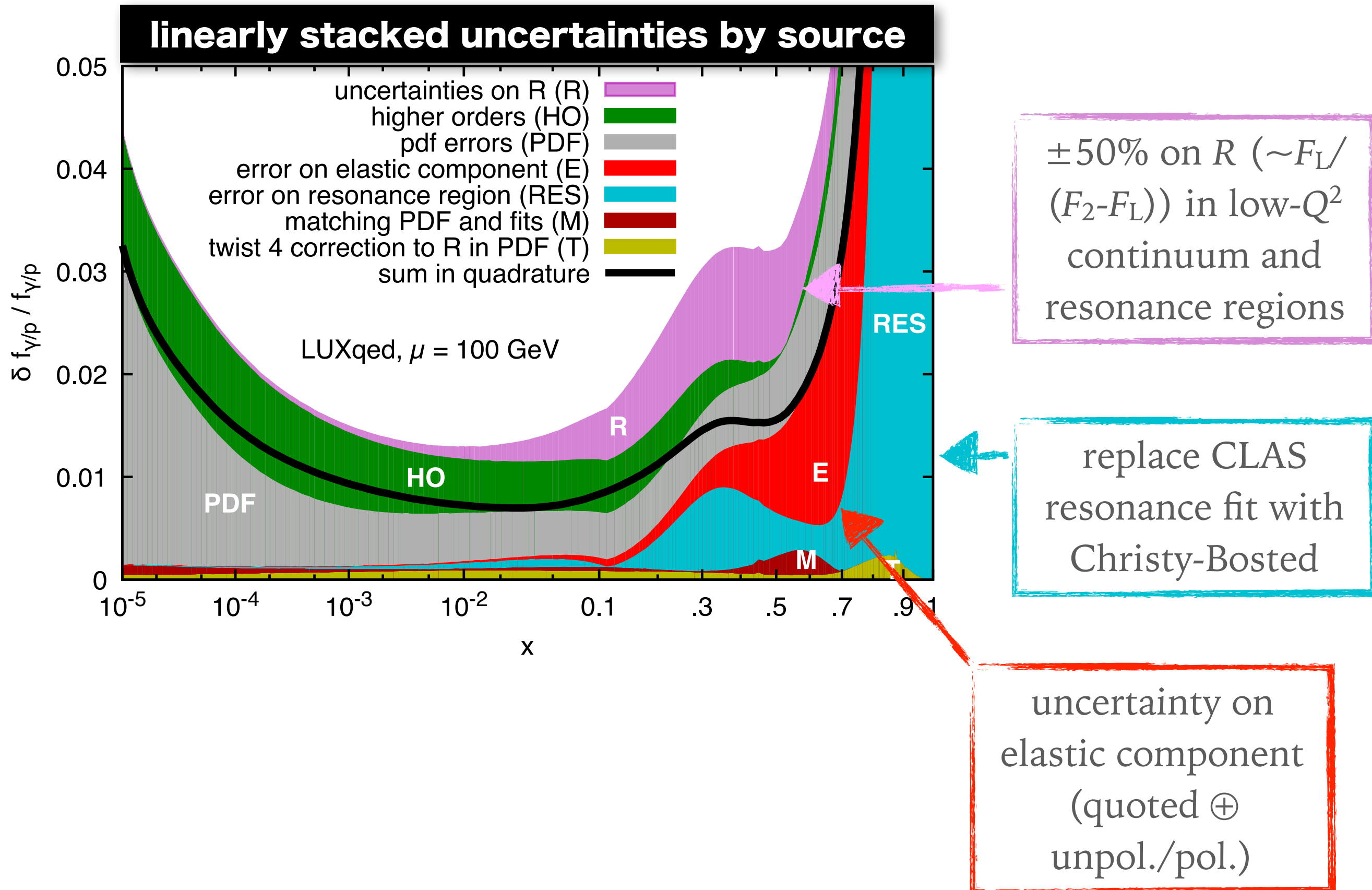
# photon uncertainties (aim to be conservative & pragmatic)



replace CLAS resonance fit with Christy-Bosted

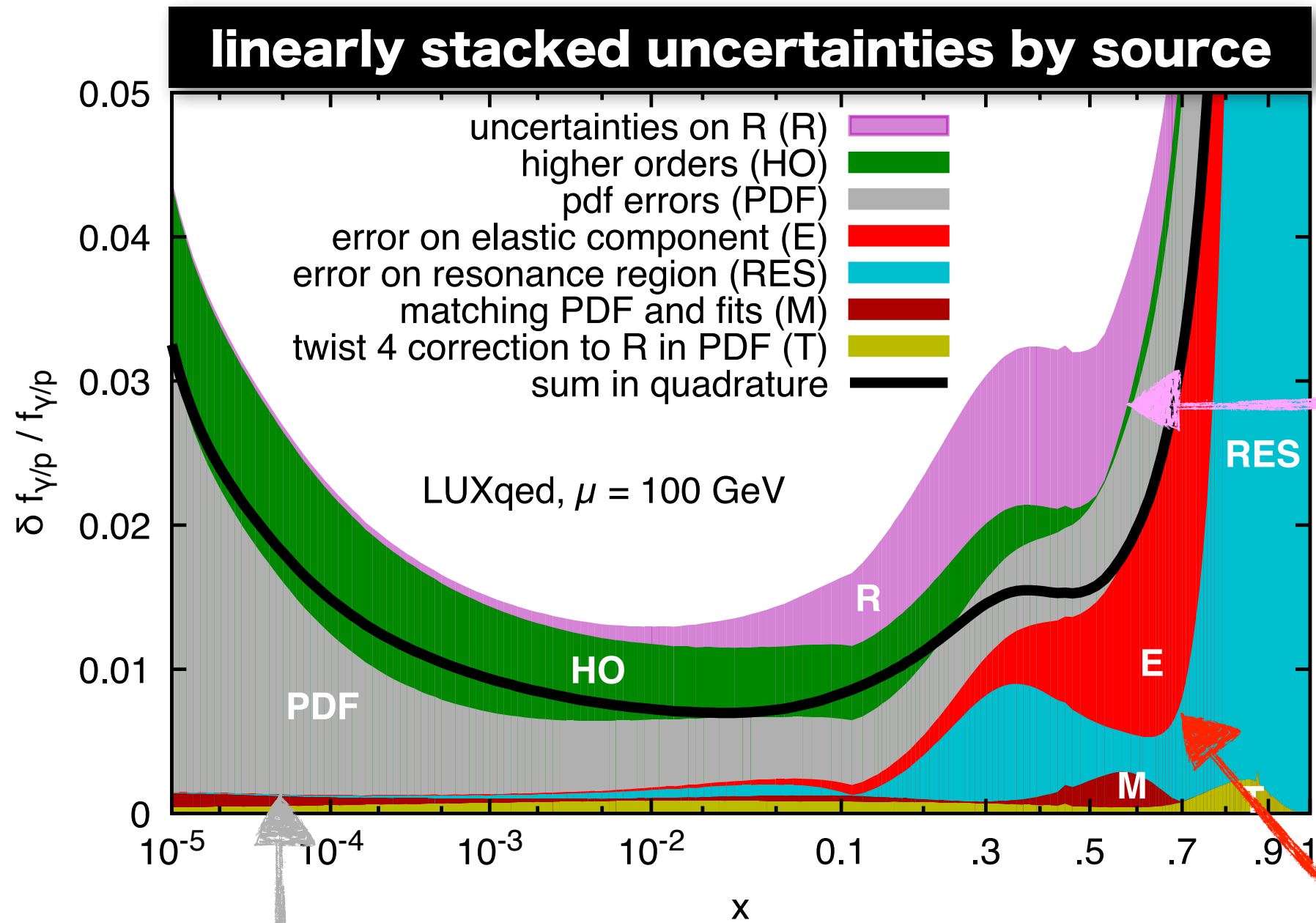
uncertainty on elastic component (quoted  $\oplus$  unpol./pol.)

# photon uncertainties (aim to be conservative & pragmatic)





# photon uncertainties (aim to be conservative & pragmatic)



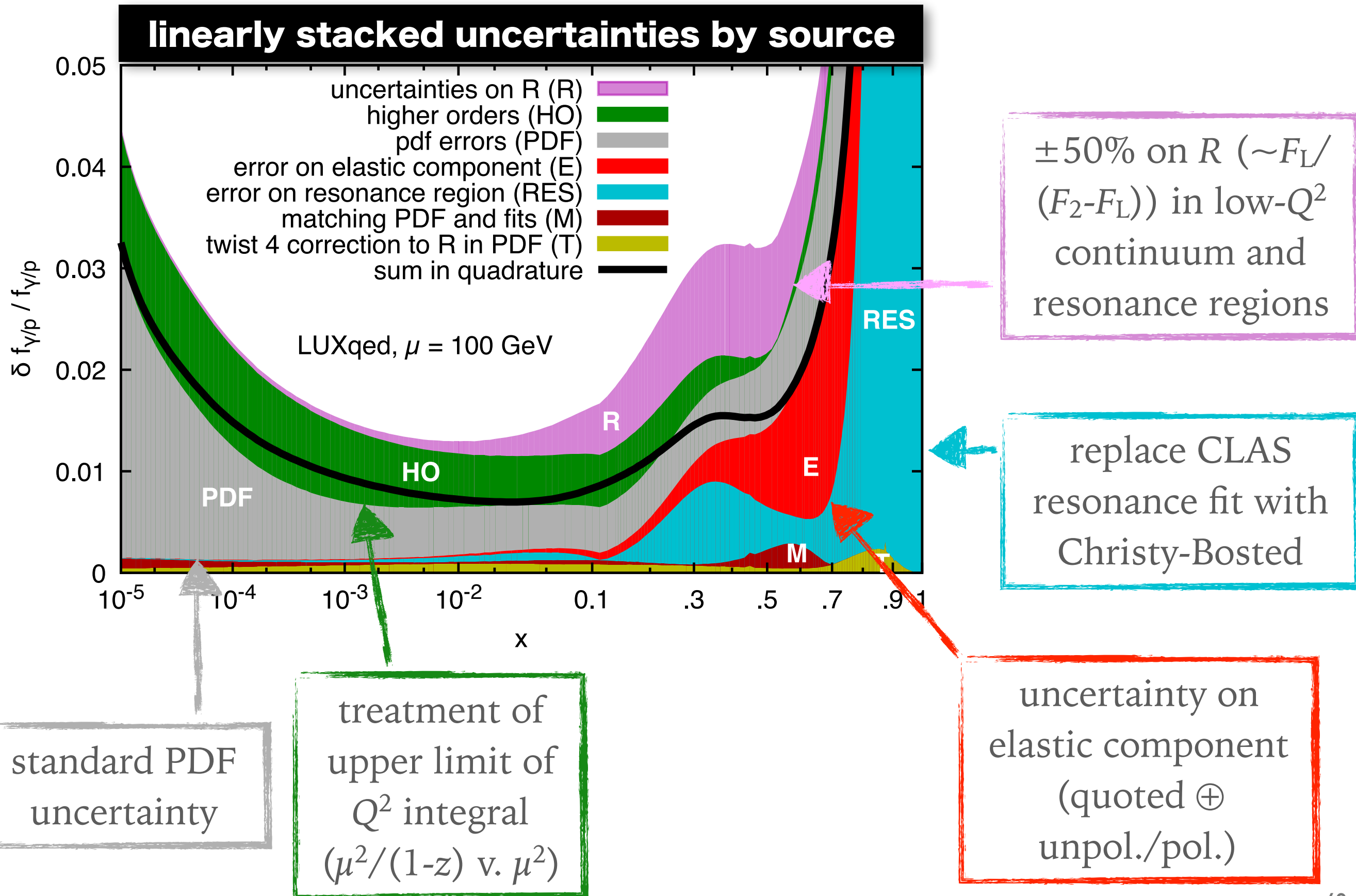
$\pm 50\%$  on  $R$  ( $\sim F_L / (F_2 - F_L)$ ) in low- $Q^2$  continuum and resonance regions

replace CLAS resonance fit with Christy-Bosted

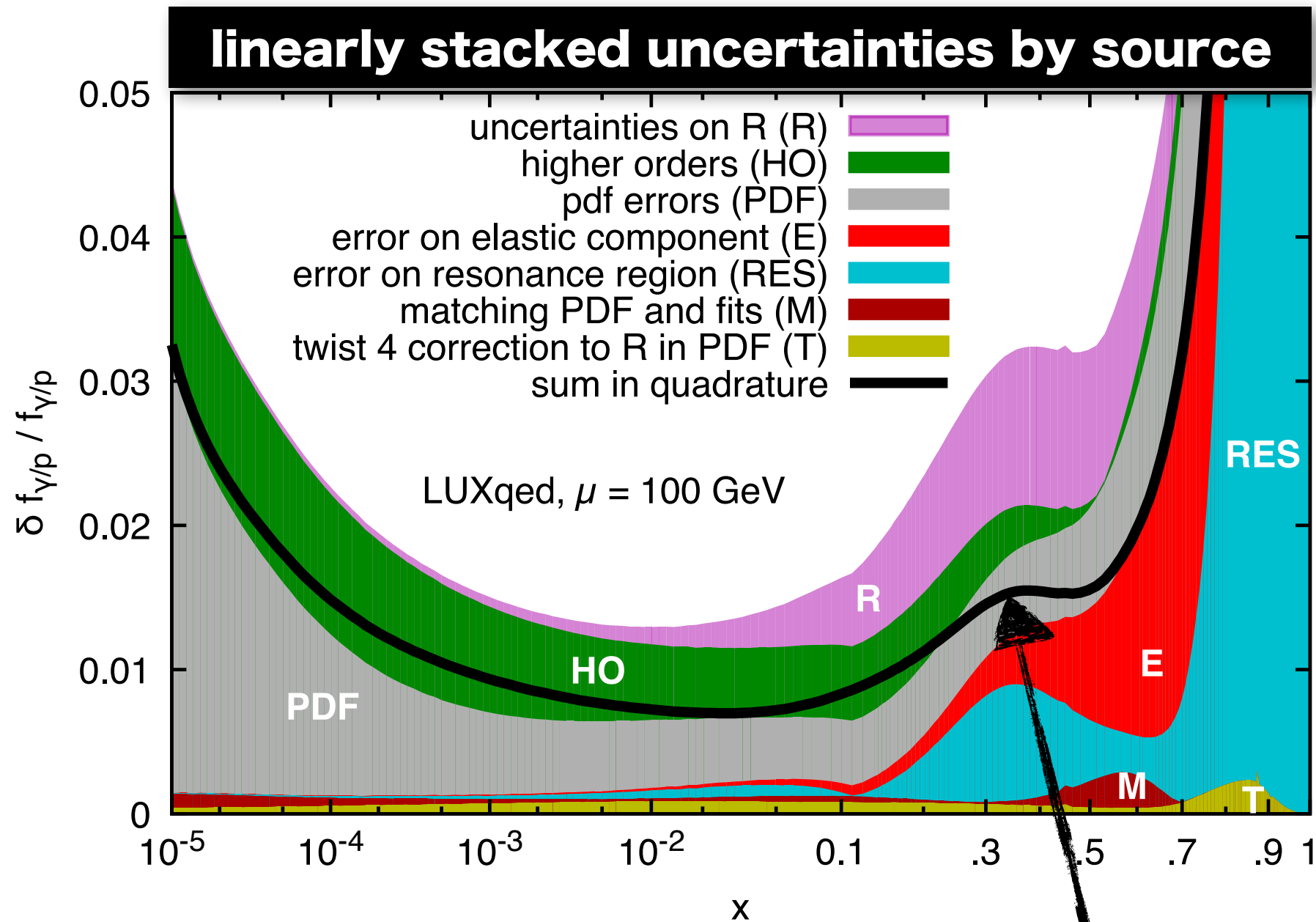
uncertainty on elastic component (quoted  $\oplus$  unpol./pol.)

standard PDF uncertainty

# photon uncertainties (aim to be conservative & pragmatic)



# photon uncertainties (aim to be conservative & pragmatic)



**final total  
uncertainty  
~ 1 – 2%  
(sum in quad. of  
all sources)**



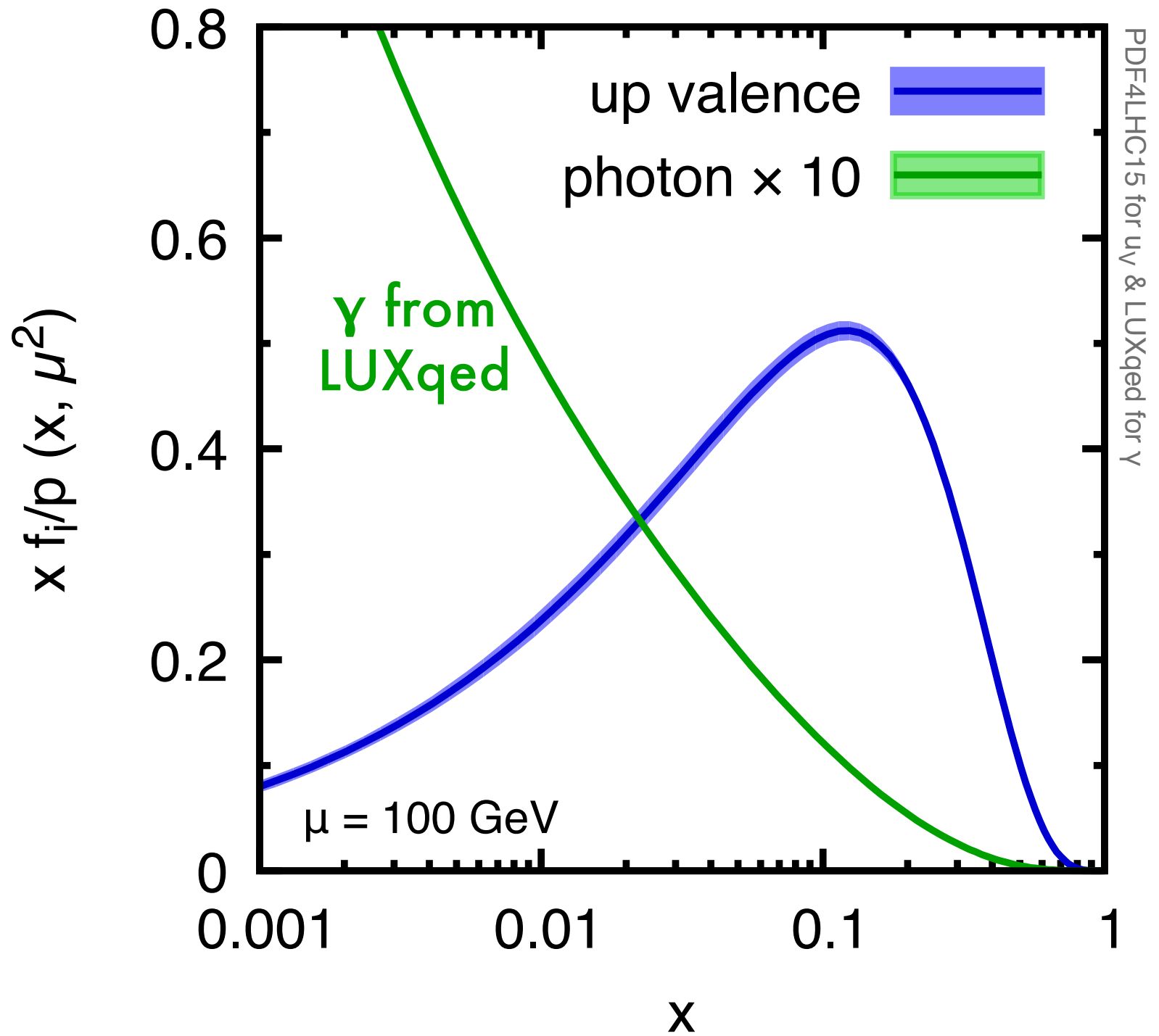
# PHOTON PDF ESTIMATES (not exhaustive)

	elastic	inelastic	in LHAPDF?
Gluck Pisano Reya 2002	dipole	model	✗
MRST2004qed	✗	model	✓
NNPDF23qed	no separation; fit to data		✓
CT14qed	✗	model (data-constrained)	✓
CT14qed_inc	dipole	model (data-constrained)	✓
Martin Ryskin 2014	dipole (only electric part)	model	✗
Harland-Lang, Khoze Ryskin 2016	dipole	model	✗
LUXqed 2016	data	data	✓

**examine result**

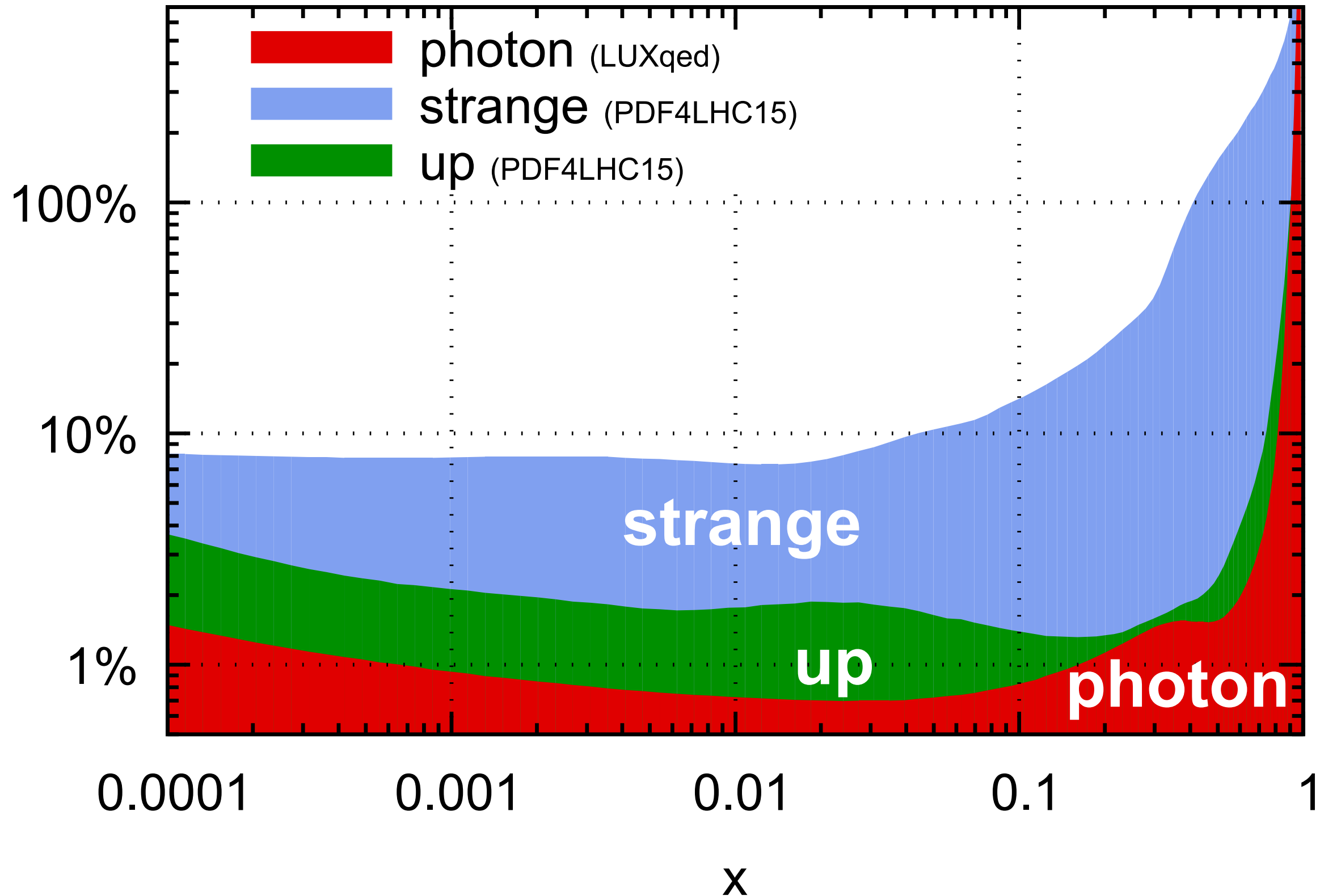
# photon PDF results

- Model-independent uncertainty (NNPDF) was 50–100%
- Goes down to  $O(1\%)$  with LUXqed determination



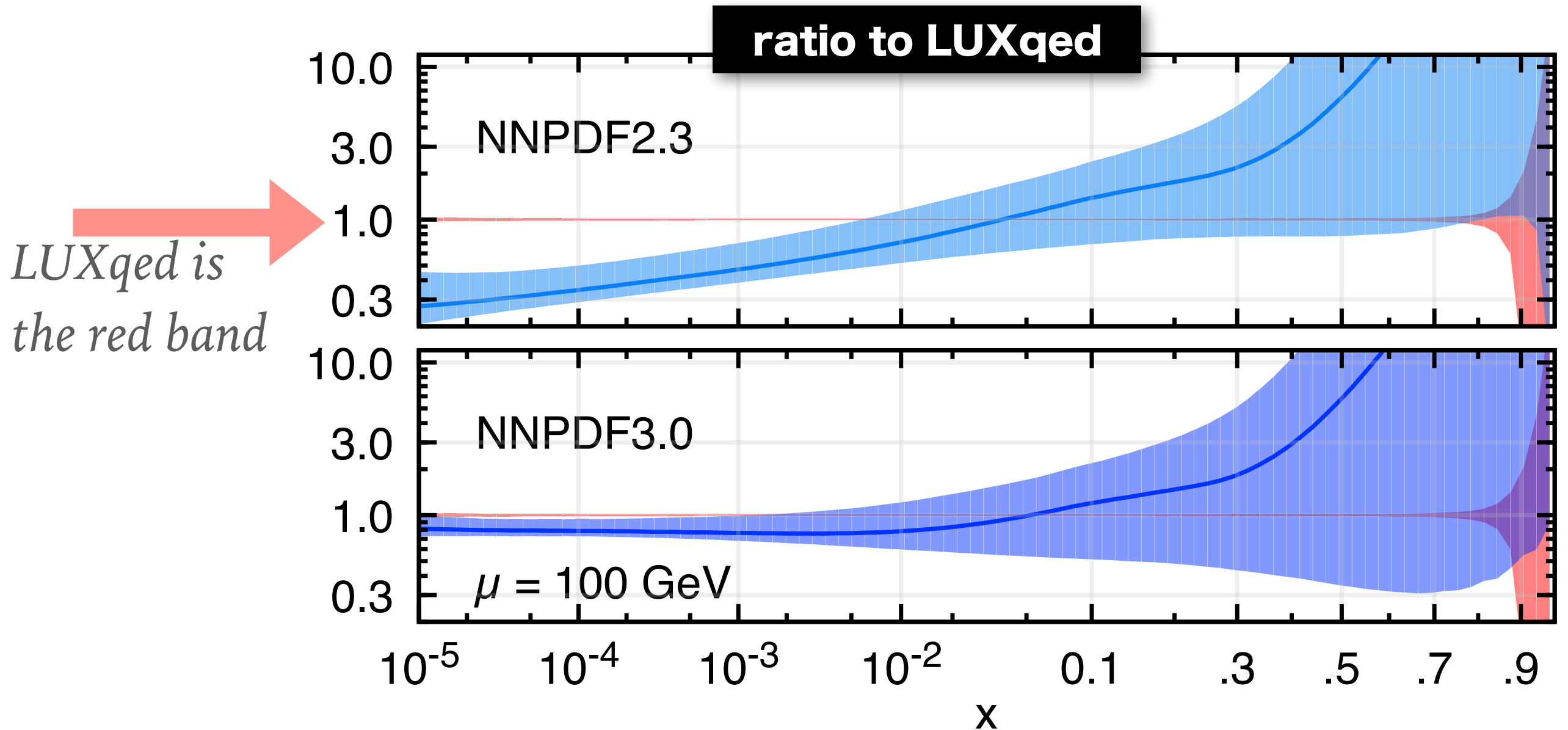
# PHOTON UNCERTAINTY (1-2%) COMPARED TO OTHER FLAVOURS

PDF uncertainties ( $Q = 100 \text{ GeV}$ )





# other PDFs v. LUXqed



**central NNPDF result much higher at large  $x$   
(but consistent within errors)**

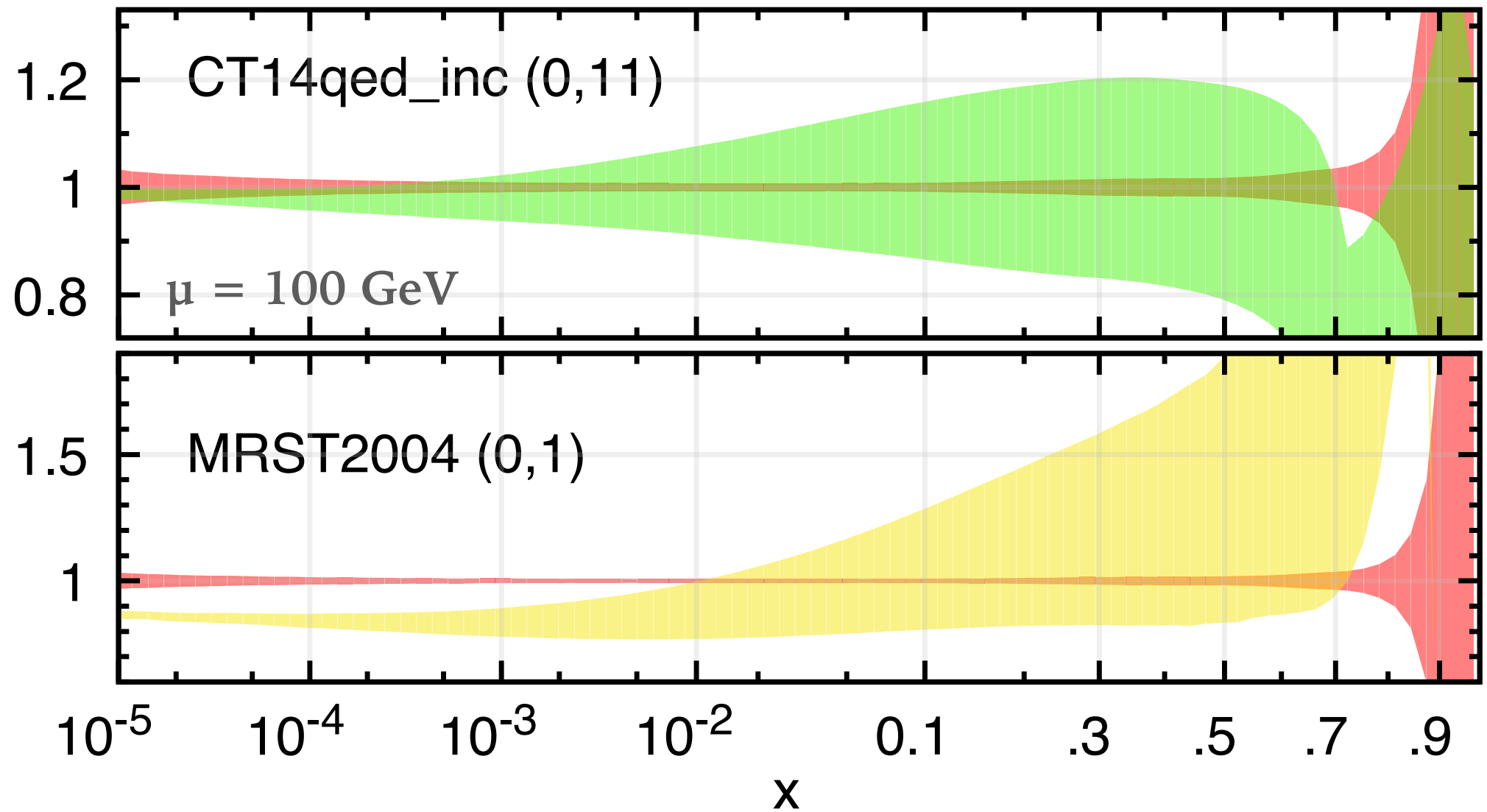
at small  $x$ , with corrected evolution (NNPDF30), about 20% smaller

# other PDFs v. LUXqed

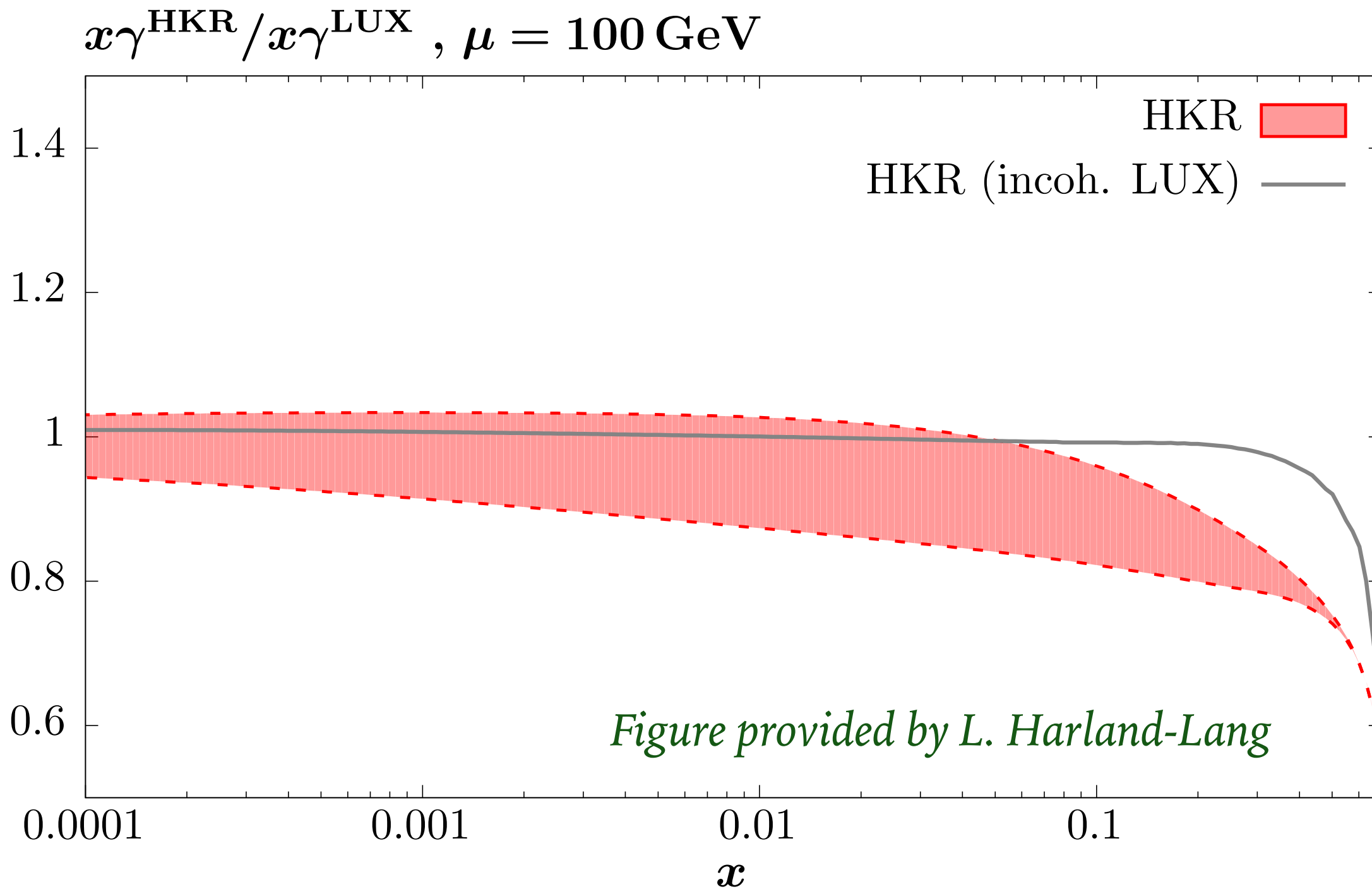
Others are numerically closer

Error bands don't always overlap with LUXqed, but within ~10-20%

ratio to LUXqed

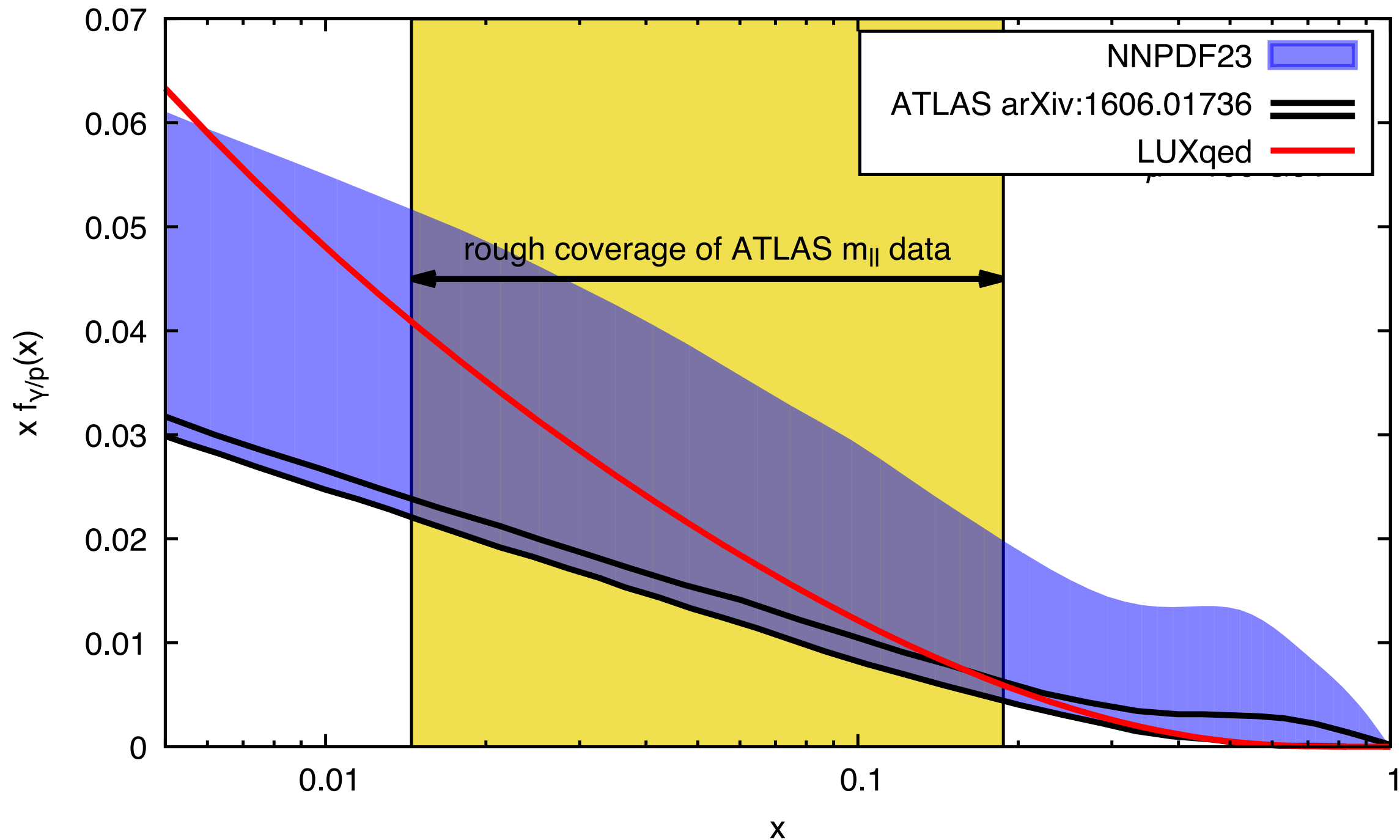


# ratio of HKR (1607.04635) to LUXqed



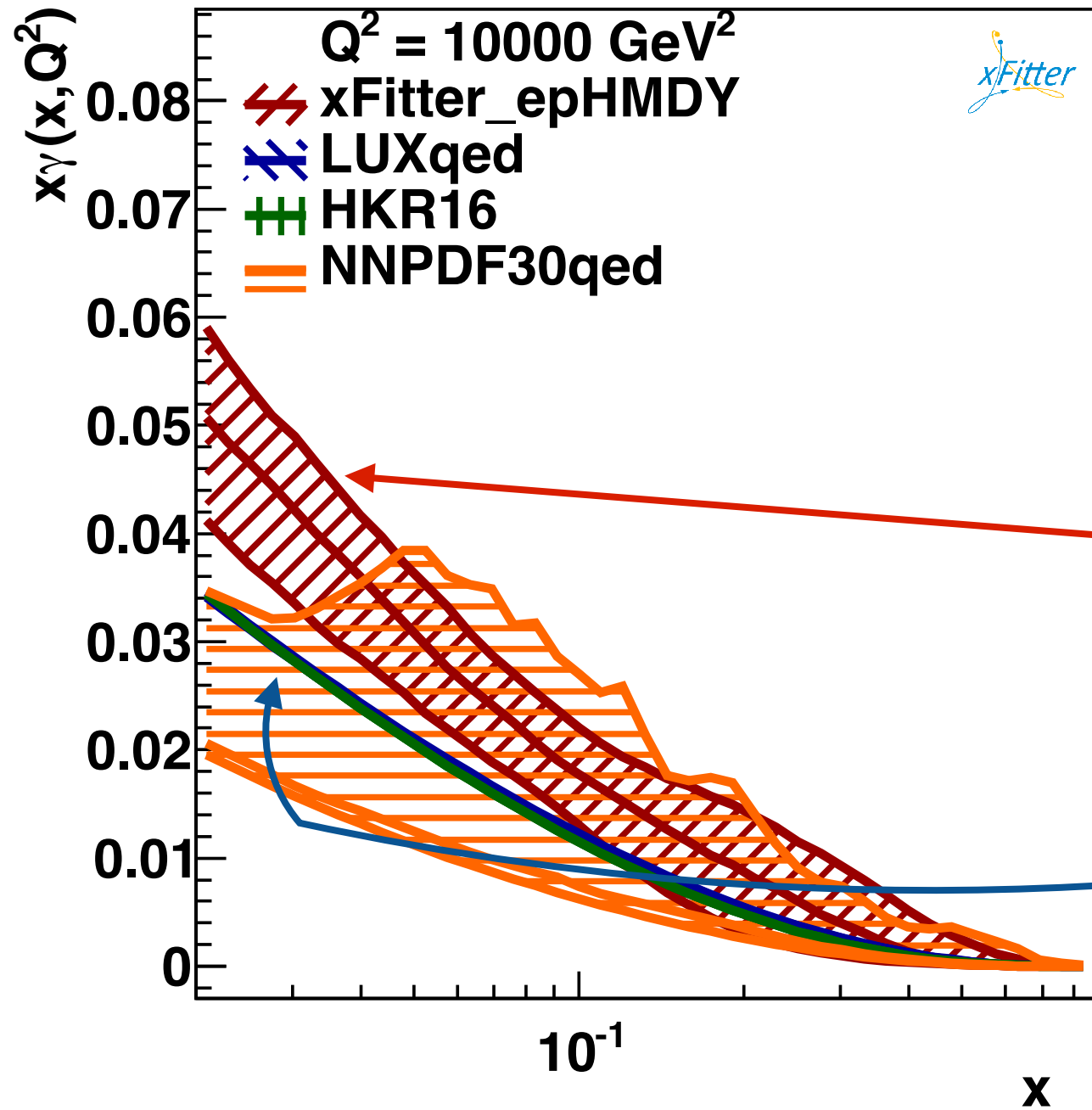
HKR based on elastic contribution (dipole approx) + model for inelastic part + evolution

# ATLAS photon (1606.01736): DY-driven reweighting of NNPDF23



ATLAS result based on reweighting of NNPDF23 with high-mass ( $M_{ll} > 116$  GeV) data

# later fit (1701.08553) to same data



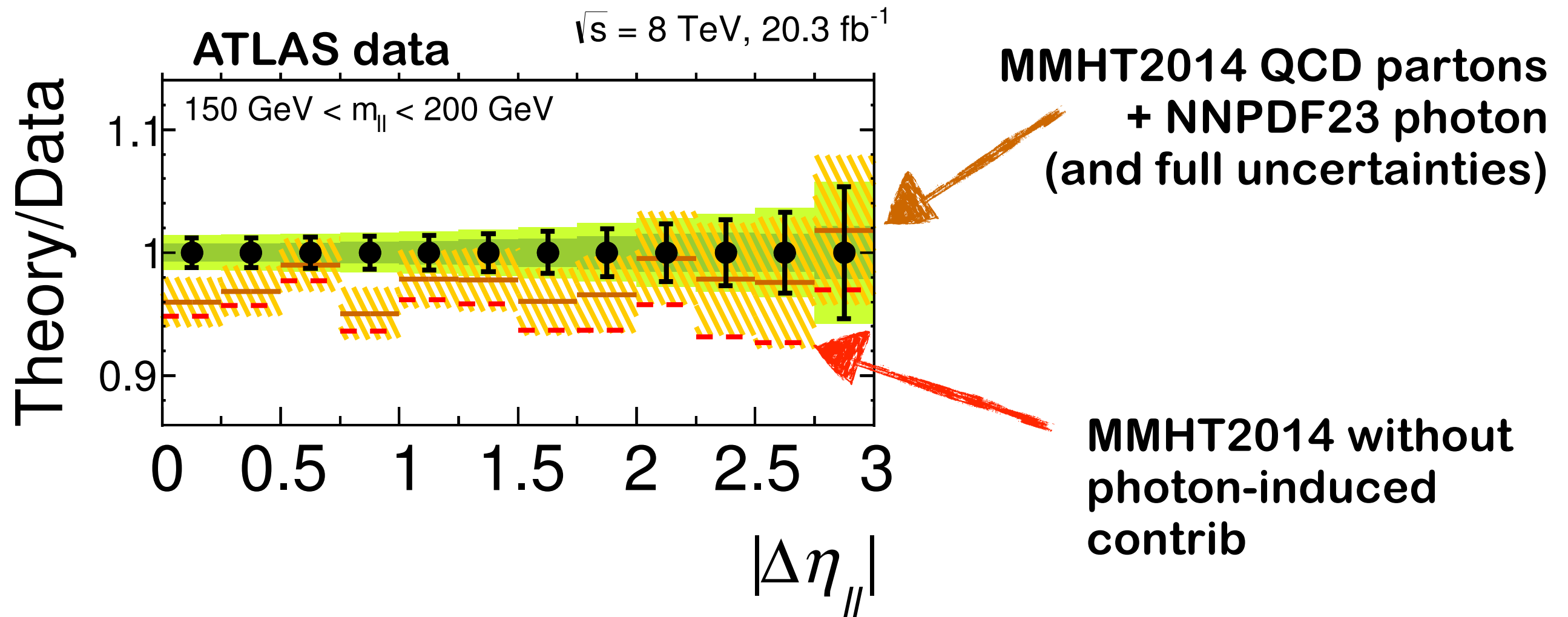
fit to

► HERA combined data

► ATLAS DY

**Fit is above LUXqed**

# ATLAS DRELL-YAN DATA (1606.01736)

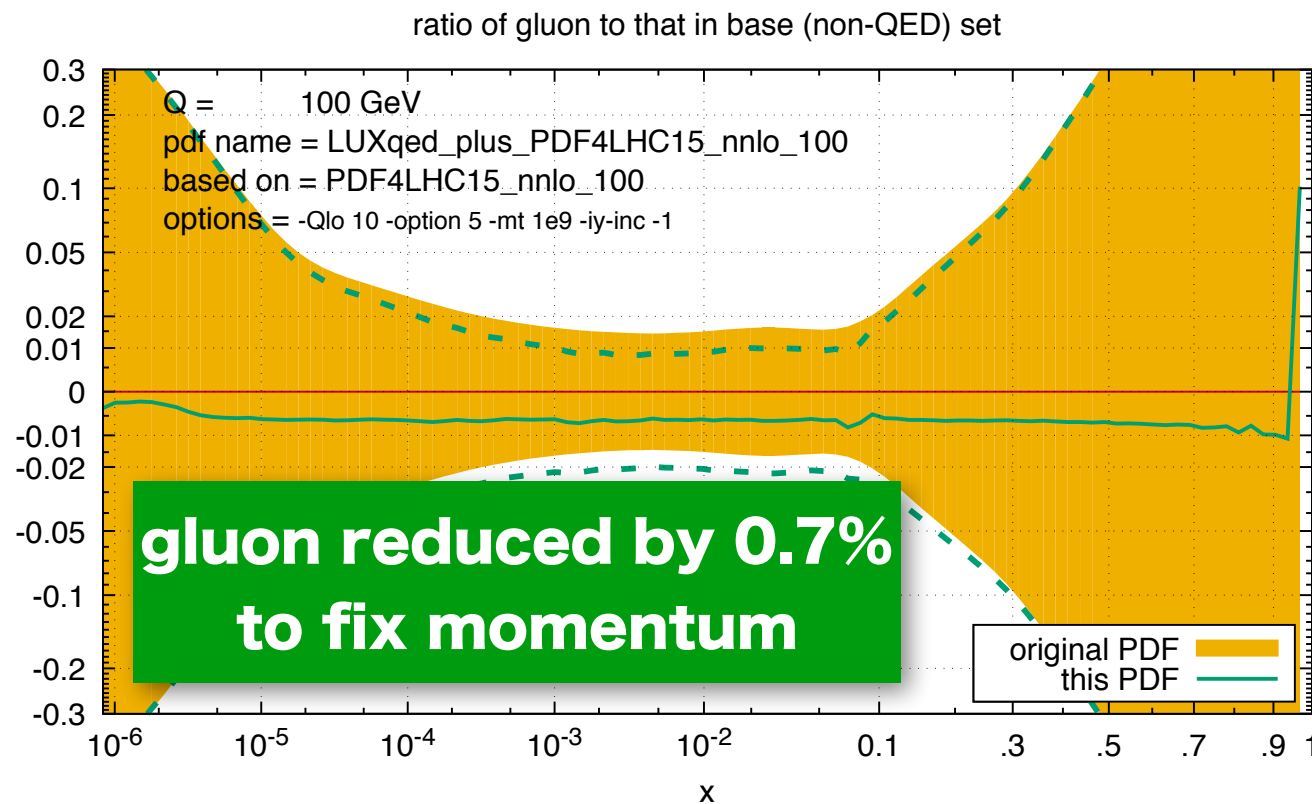


# MATCHING PROCEDURE FOR FULL SET OF PARTONS

---

- evaluate master eqn. for  $\mu = 100$  GeV (with default PDF4LHC15\_nnlo partons)
- Do  $O(\alpha_s)$  photon evolution down to  $\mu = 10$  GeV (other partons: pure QCD evln.)
- Adjust momentum sum-rule by rescaling gluon  $g(x) \rightarrow 0.993g(x)$
- Evolve back up with NNLO-QCD &  $O(\alpha_s)$  QED for all partons

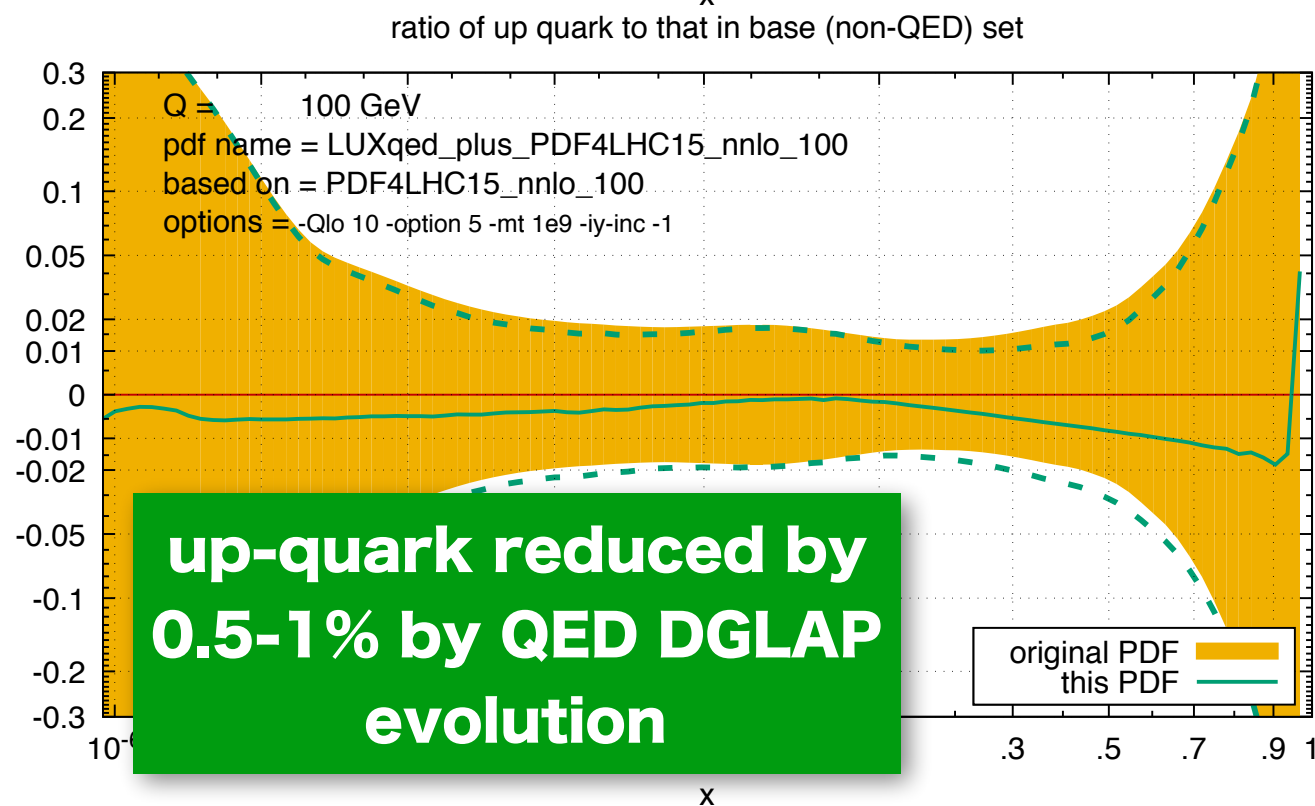
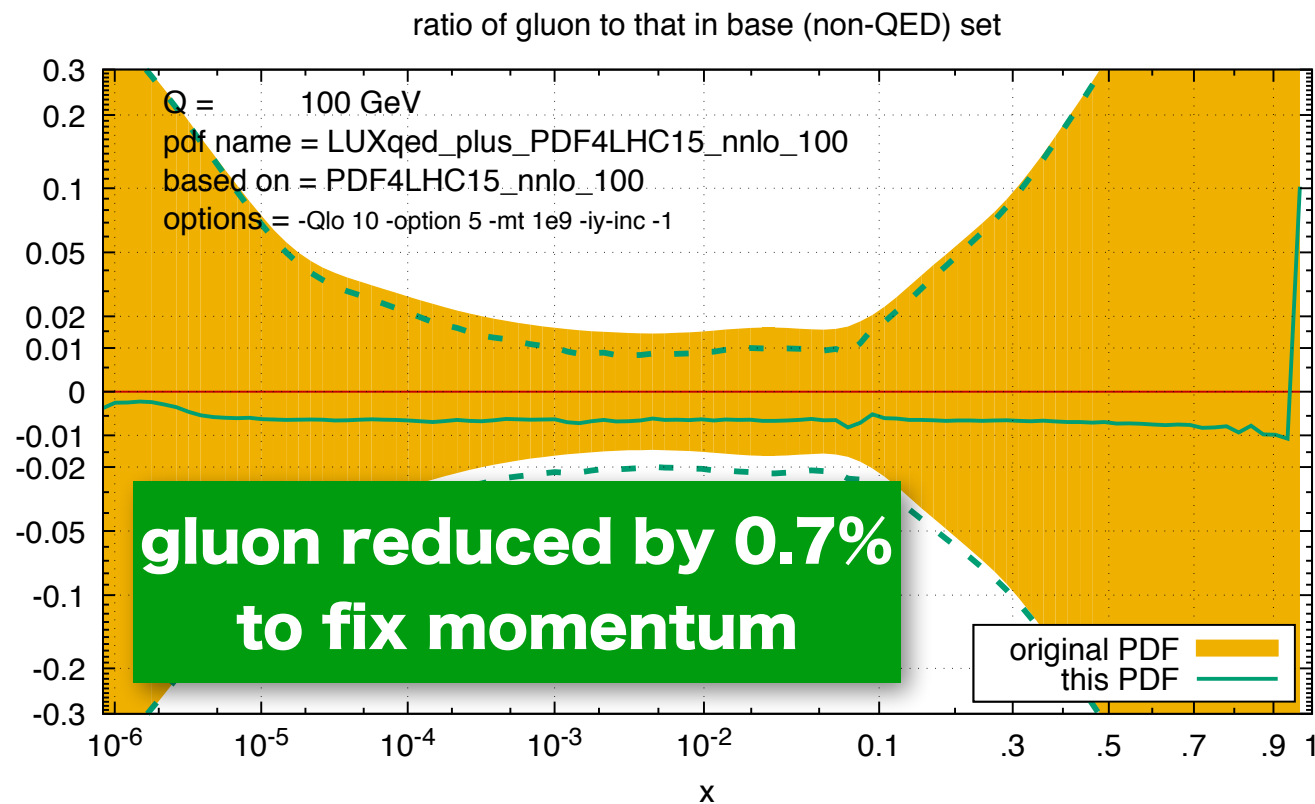
# MATCHING PROCEDURE FOR FULL SET OF PARTONS



- evaluate master eqn. for  $\mu = 100$  GeV (with default PDF4LHC15\_nnlo partons)
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 $g(x) \rightarrow 0.993g(x)$
- Evolve back up with NNLO-QCD &  $O(\alpha_s)$  QED for all partons

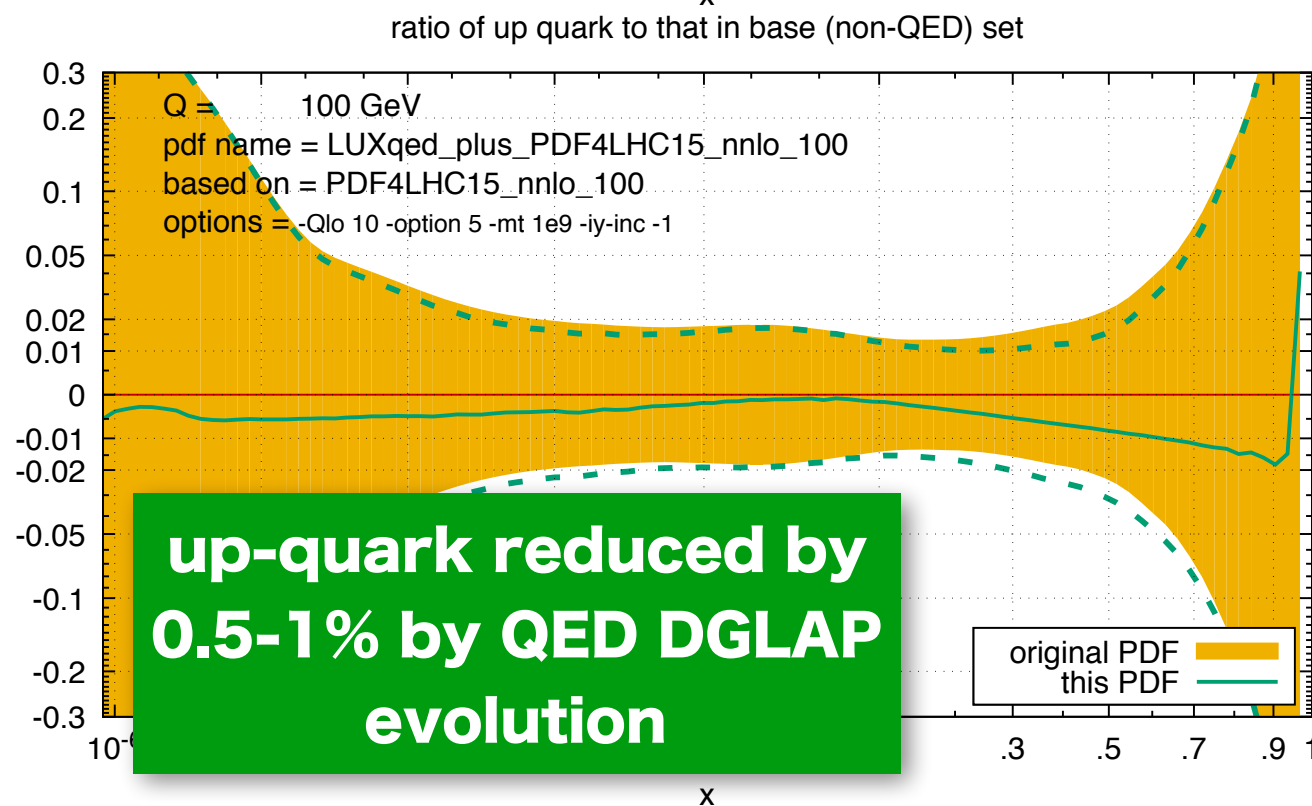
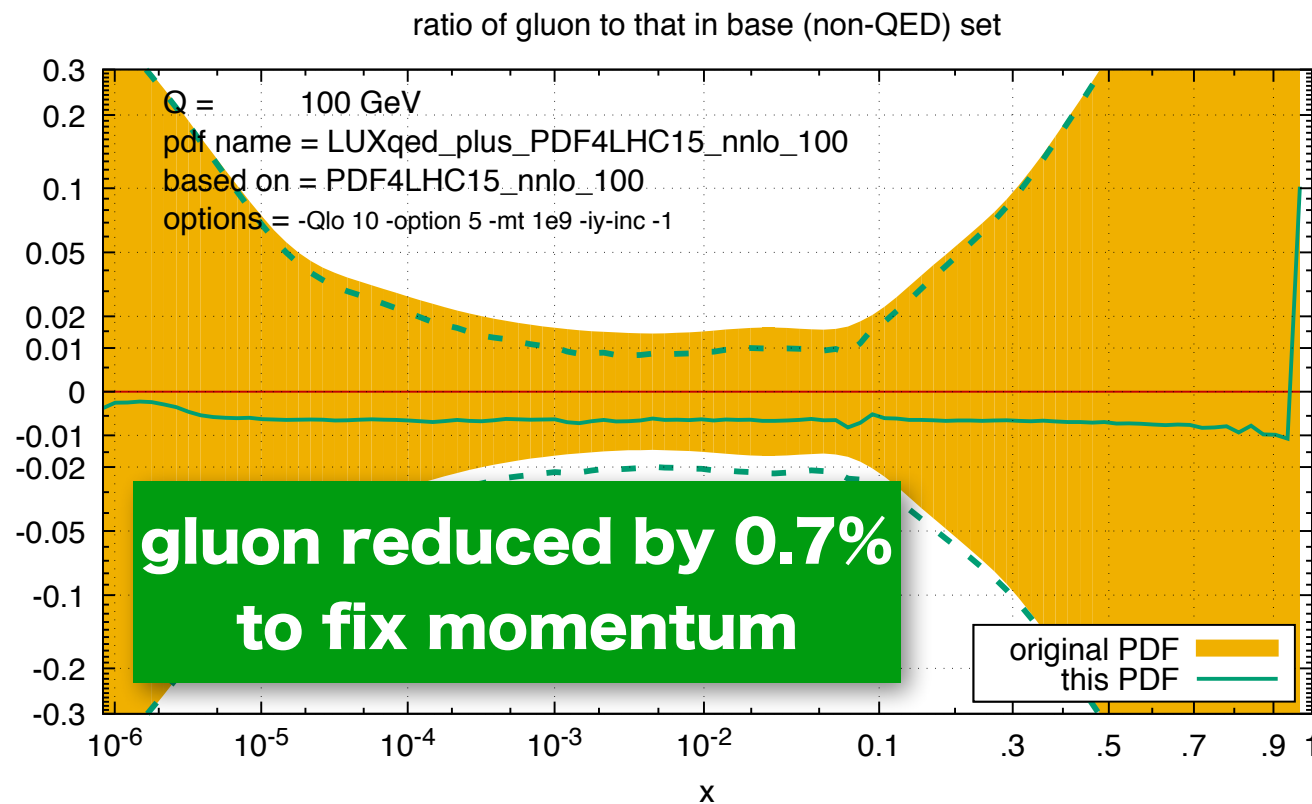


# MATCHING PROCEDURE FOR FULL SET OF PARTONS



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- Do  $O(\alpha_s)$  photon evolution down to  $\mu = 10$  GeV (other partons: pure QCD evln.)
- Adjust momentum sum-rule by rescaling gluon  $g(x) \rightarrow 0.993g(x)$
- Evolve back up with NNLO-QCD &  $O(\alpha_s)$  QED for all partons

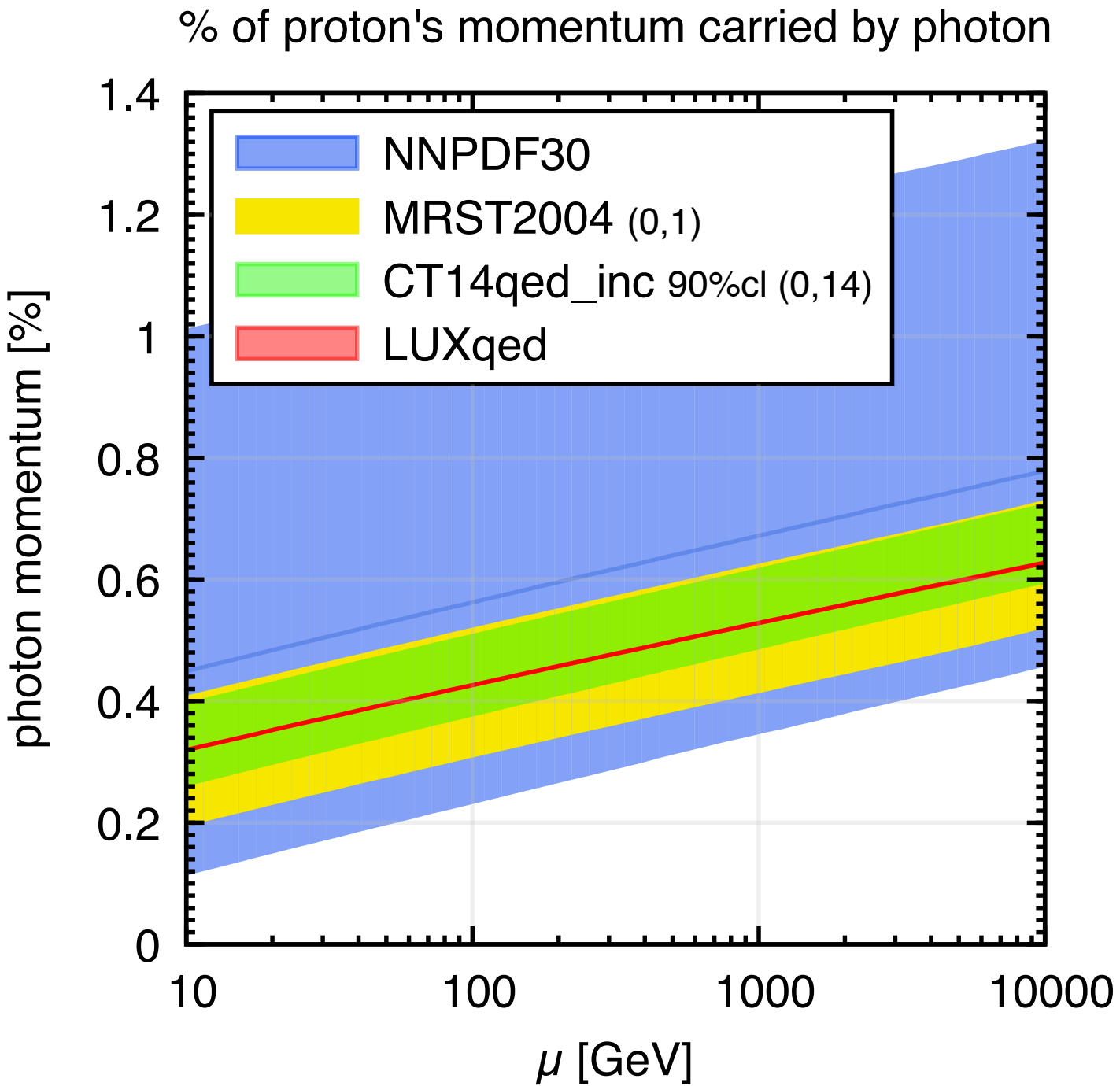
# MATCHING PROCEDURE FOR FULL SET OF PARTONS



- evaluate master eqn. for  $\mu = 100$  GeV (with default PDF4LHC15\_nnlo partons)
- Do  $O(\alpha_s)$  photon evolution down to  $\mu = 10$  GeV (other partons: pure QCD evln.)
- Adjust momentum sum-rule by rescaling gluon  $g(x) \rightarrow 0.993g(x)$
- Evolve back up with NNLO-QCD &  $O(\alpha_s)$  QED for all partons

**better approach would be full PDF re-fit for QCD partons incl. EW/QED corrections & LUXqed photon**

# MOMENTUM CARRIED BY PHOTON



momentum ( $\mu = 100$ GeV)	
gluon	$46.8 \pm 0.4\%$
up valence	$18.2 \pm 0.3\%$
down valence	$7.5 \pm 0.2\%$
light sea quarks	$20.7 \pm 0.4\%$
charm	$4.0 \pm 0.1\%$
bottom	$2.5 \pm 0.1\%$
photon	$0.426 \pm 0.003\%$

LUXqed\_plus\_PDF4LHC15\_nnlo\_100

(1+107 members, symmhessian, errors handled by LHAPDF out of the box,

**PDF valid for  $\mu > 10$  GeV (related to PDF4LHC15 issues)**

**applications**

# APPLICATION TO HIGGS PHYSICS

---

$pp \rightarrow H W^+ (\rightarrow l^+ \nu) + X$  at 13 TeV

non-photon induced contributions

$91.2 \pm 1.8$  fb

photon-induced contribs (NNPDF23)

$6.0^{+4.4}_{-2.9}$  fb

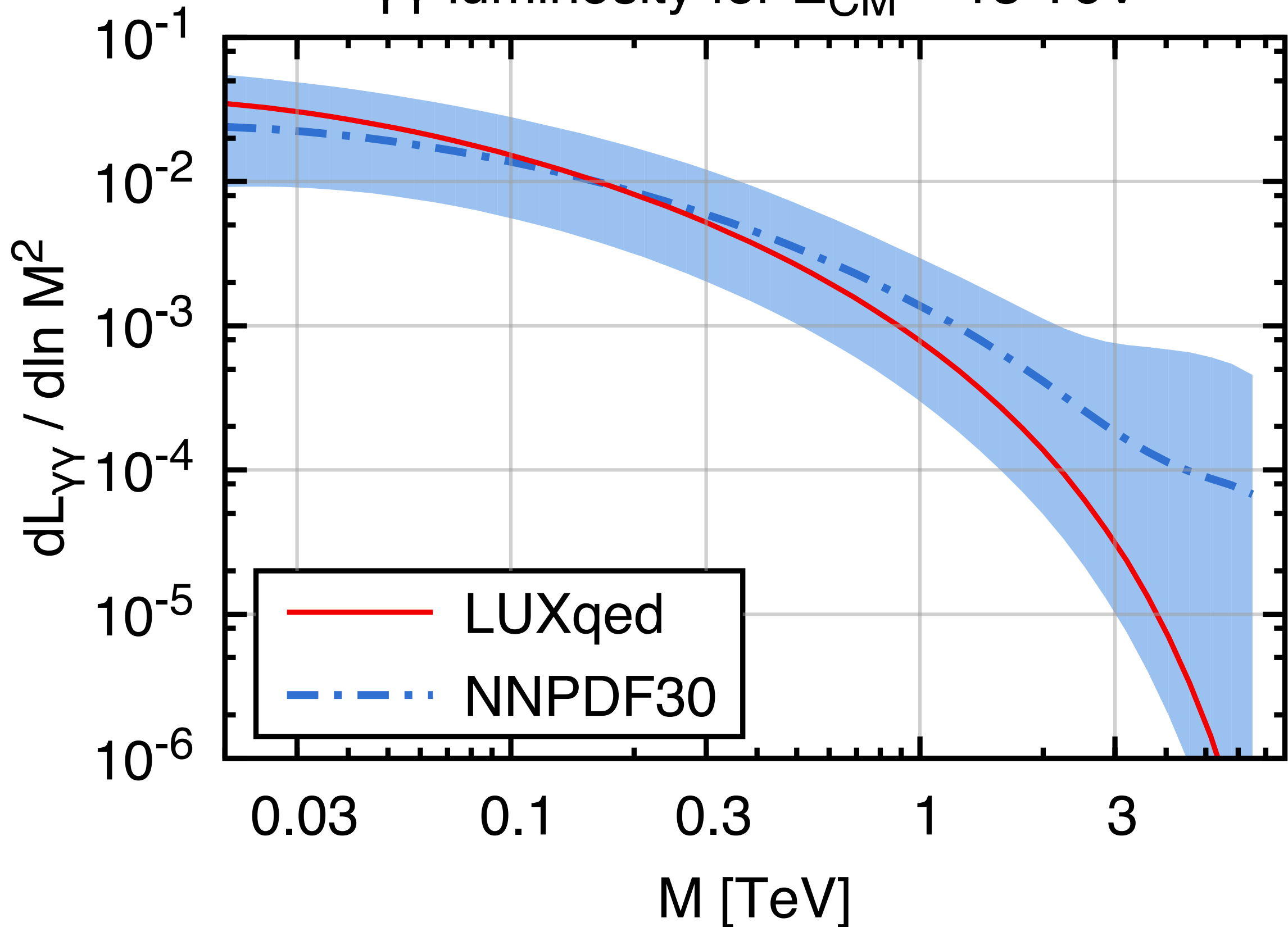
photon-induced contribs (LUXqed)

$4.4 \pm 0.1$  fb

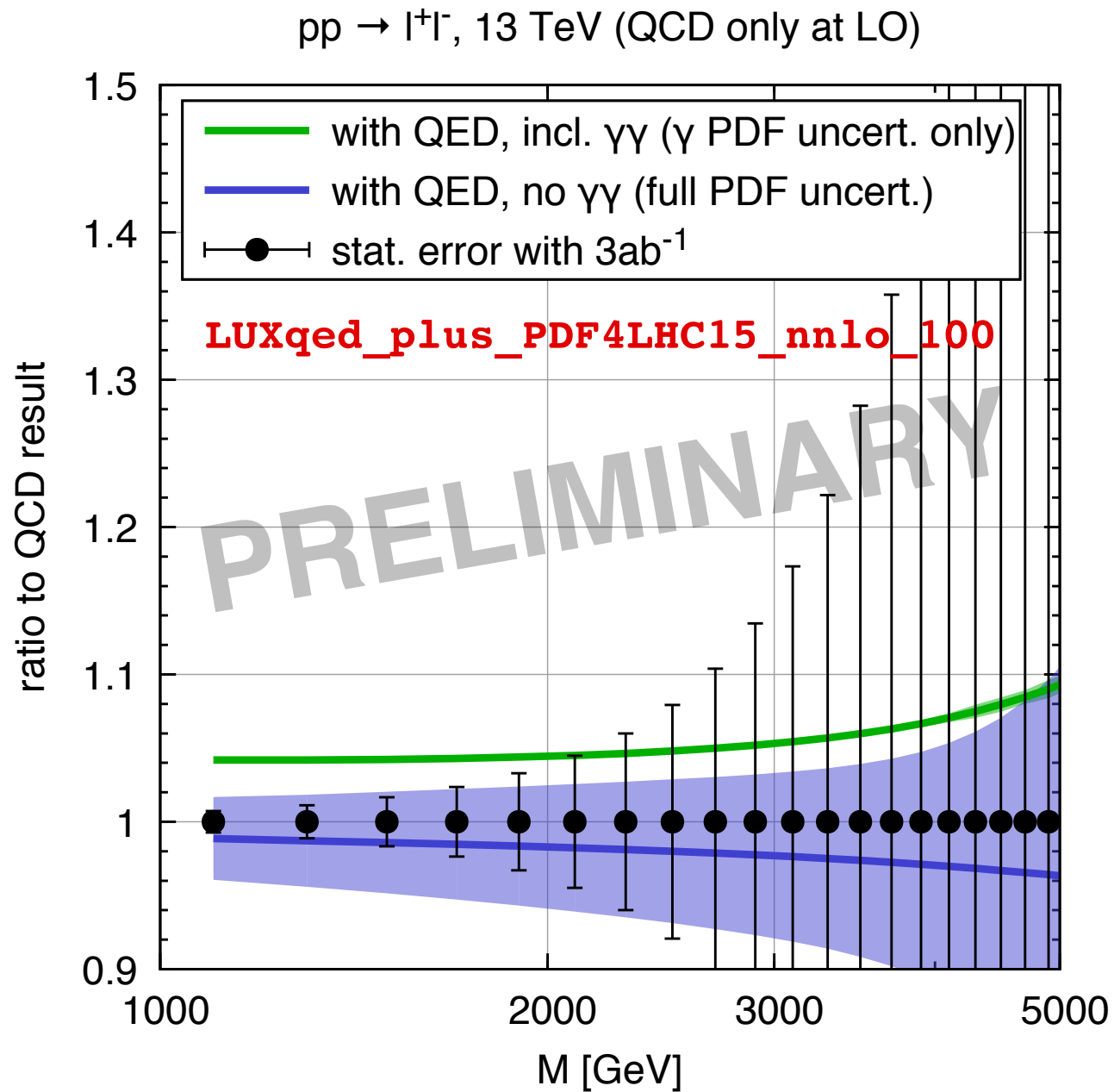
*non-photon numbers from LHCHSWG (YR4)  
including PDF uncertainties*

# $\gamma\gamma$ luminosity

$\gamma\gamma$  luminosity for  $E_{\text{CM}} = 13$  TeV



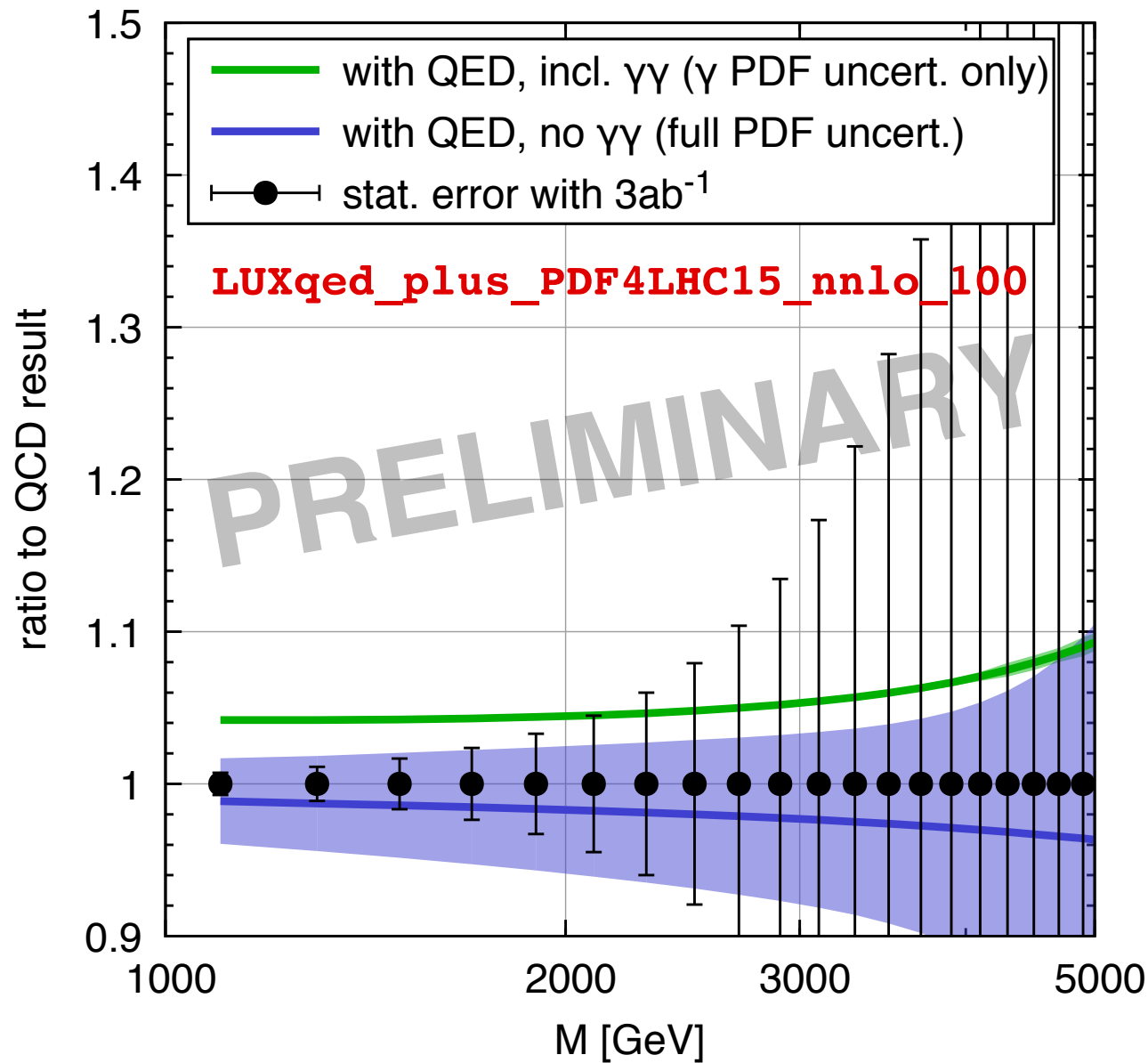
# di-lepton spectrum with $3\text{ab}^{-1}$



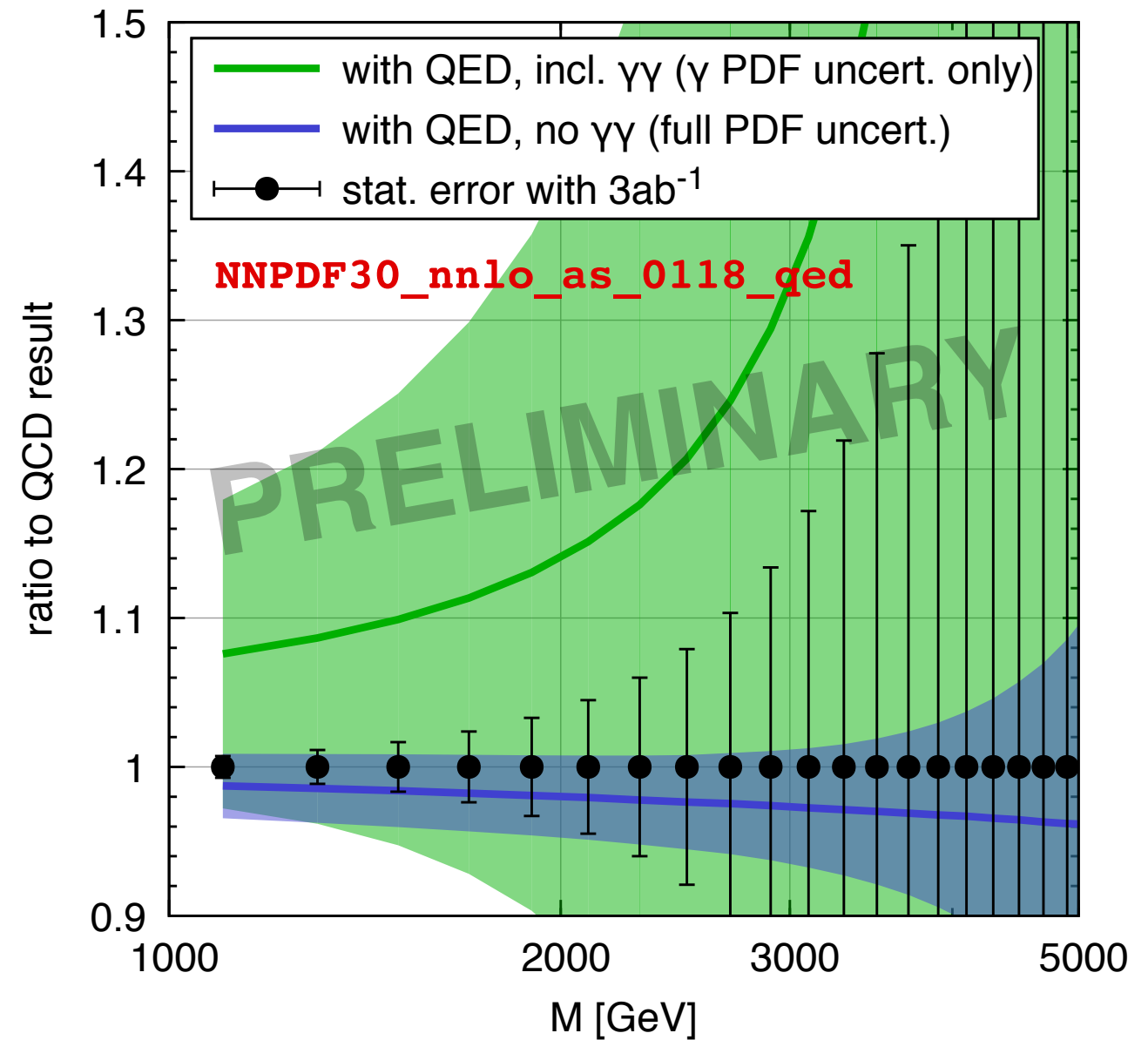
**LUXQED photon has few % effect on di-lepton spectrum and negligible uncertainties**

# di-lepton spectrum with $3\text{ab}^{-1}$

$pp \rightarrow l^+l^-$ , 13 TeV (QCD only at LO)



$pp \rightarrow l^+l^-$ , 13 TeV (QCD only at LO)



**LUXQED photon has few % effect on di-lepton spectrum and negligible uncertainties**



# conclusions & resources

# RESOURCES

---

- LUXqed\_plus\_PDF4LHC15\_nnlo\_100 set available from LHAPDF (for  $\mu > 10$  GeV)
- Additional plots and validation info available from <http://cern.ch/luxqed>
- Preliminary version of HOPPET DGLAP evolution code with QED (order  $\alpha$  and  $\alpha\alpha_s$ ) corrections available from hepforge:  

```
svn checkout http://hoppet.hepforge.org/svn/branches/qed hoppet-qed
```

(look at `tests/with-lhapdf/test_qed_evol_lhapdf.f90` for an example; interface may change, documentation missing; NB: APFEL code also has QED contributions in the evolution)

## CLOSING REMARKS

---

- Distribution of photons in the proton depends on the **non-perturbative QCD** physics of the proton
- But **perturbative QED** enables you to deduce the photon density from measured (non-pert.) proton structure functions
- Our public results are just to NLO (equiv. a  $\alpha_s$  in splitting functions), but higher theoretical is in the pipeline (e.g.  $\alpha^2$ , a  $\alpha_s^2$ ) — open question of whether data can follow (and whether we need it)

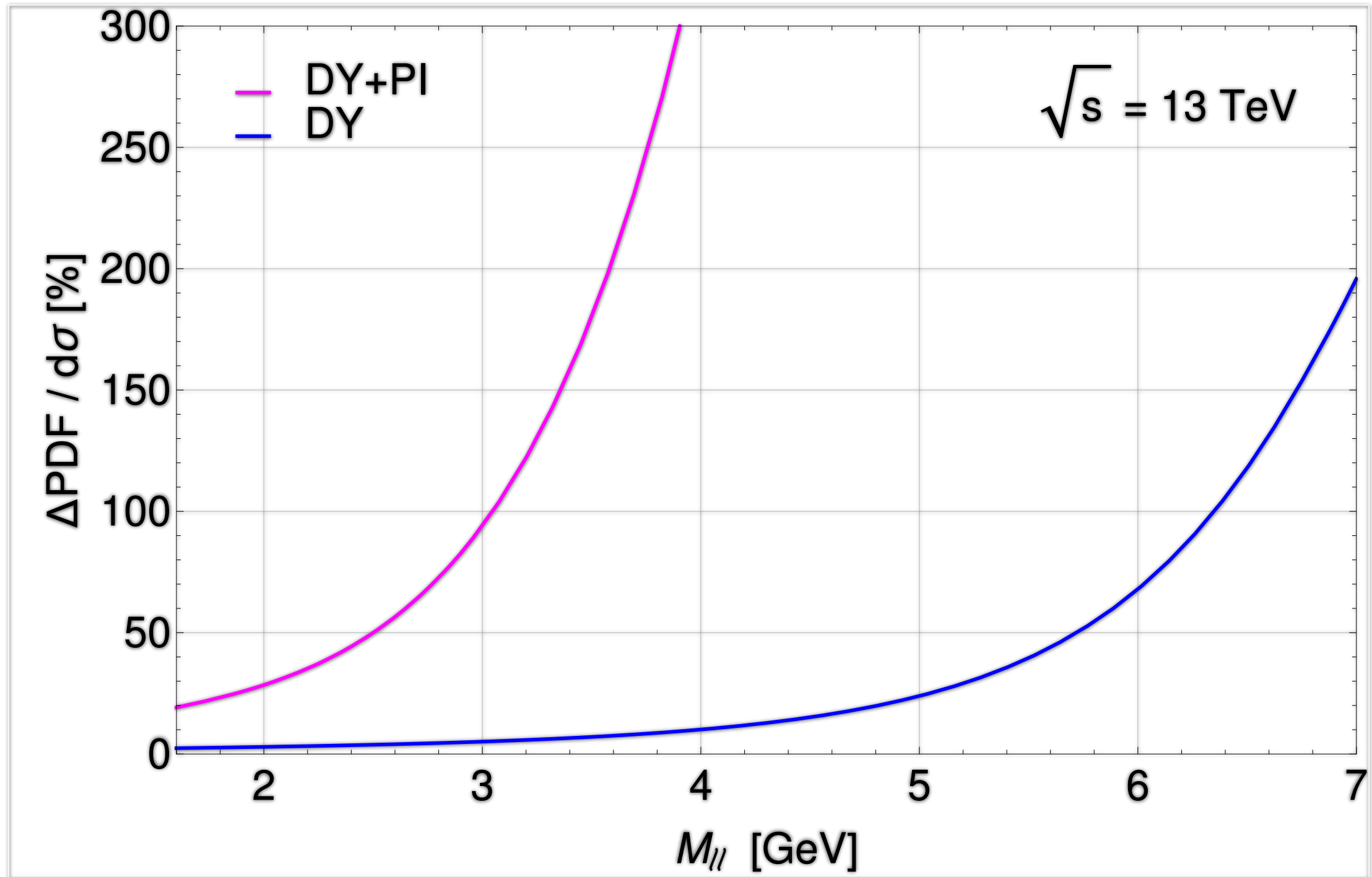
*“If you think about it, it's awesome: we are made of protons, and protons are, in some part, made of light... And now we know how much of it.”*

*[blog post](#) by Tommaso Dorigo*

**extra slides**

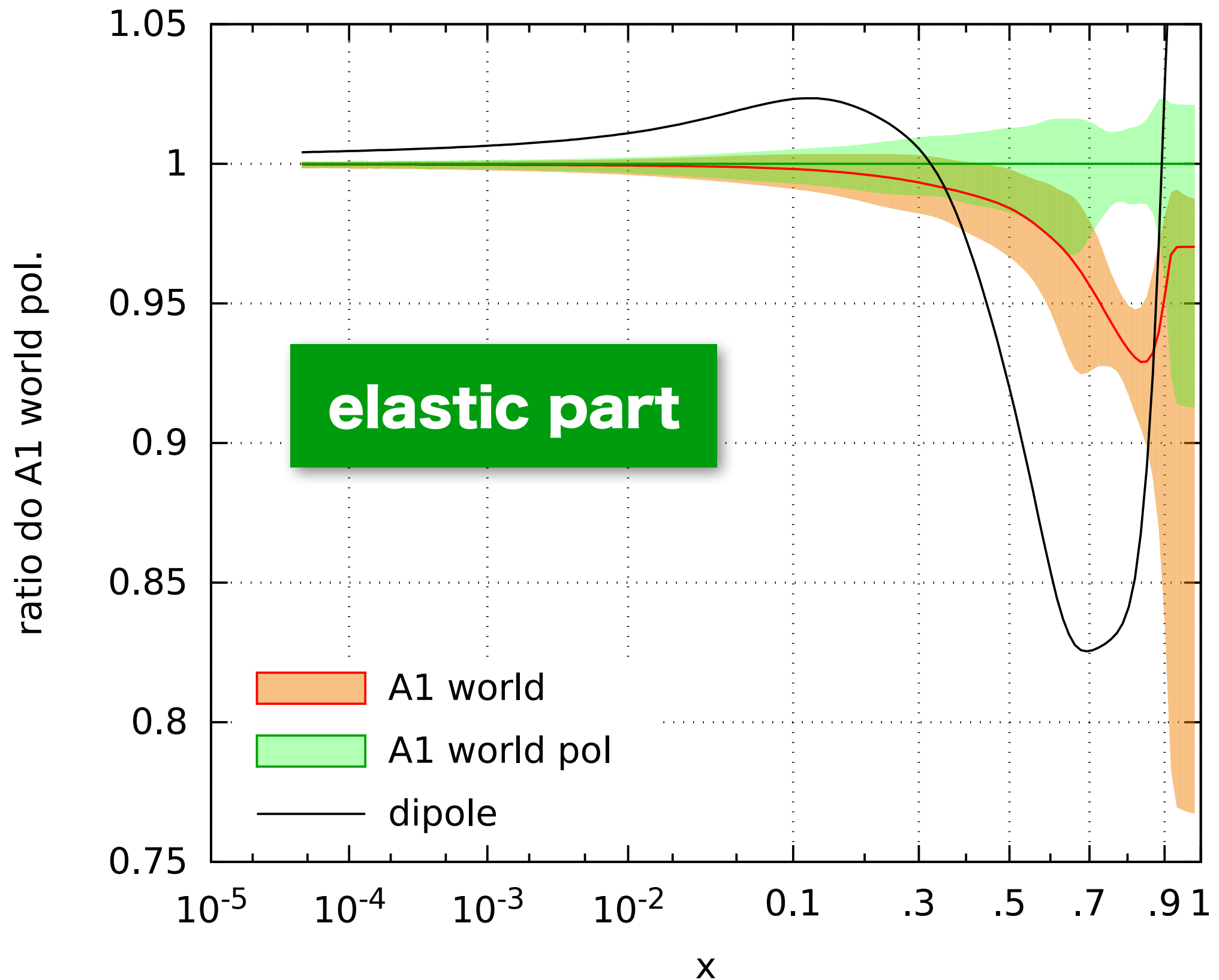
1606.06646v1.

Elena Accomando,<sup>1,2,\*</sup> Juri Fiaschi,<sup>1,2,†</sup> Francesco Hautmann,<sup>2,3,‡</sup>  
Stefano Moretti,<sup>1,2,§</sup> and C.H. Shepherd-Themistocleous<sup>1,2,¶</sup>



# **input data & procedures**

# ELASTIC COMPONENT & COMPARISON TO "DIPOLE" MODEL

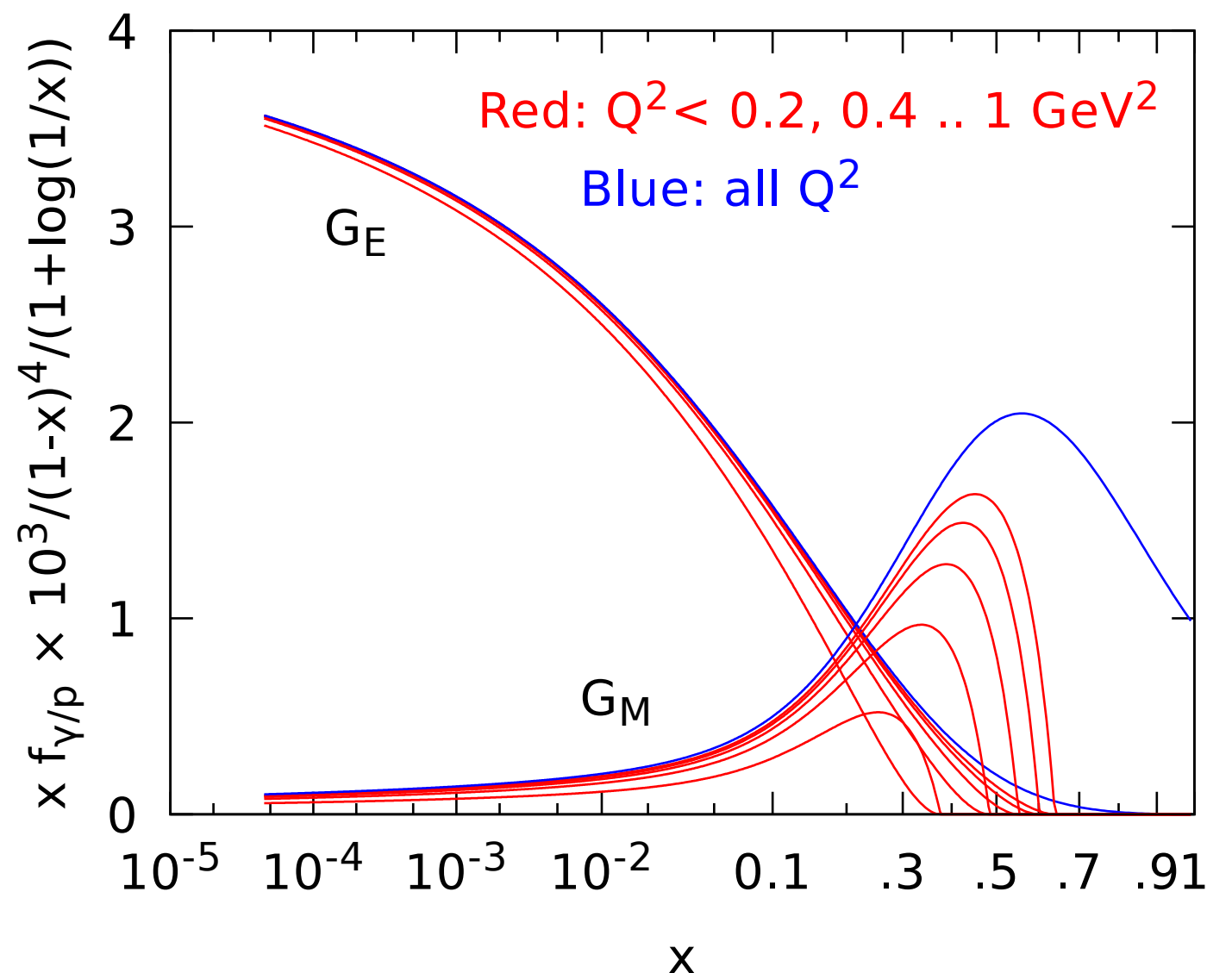


The elastic contribution to  $f_\gamma$  is

$$\begin{aligned}
 x f_\gamma^{\text{el}}(x, \mu^2) &= \frac{1}{2\pi} \int_{\frac{x^2 m_p^2}{1-x}}^{\frac{\mu^2}{1-x}} \frac{dQ^2}{Q^2} \frac{\alpha^2(Q^2)}{\alpha(\mu^2)} \left\{ \left( 1 - \frac{x^2 m_p^2}{Q^2(1-x)} \right) \frac{2(1-x) G_E^2(Q^2)}{1+\tau} \right. \\
 &+ \left. \left( 2 - 2x + x^2 + \frac{2x^2 m_p^2}{Q^2} \right) \frac{G_M^2(Q^2) \tau}{1+\tau} \right\}.
 \end{aligned}$$

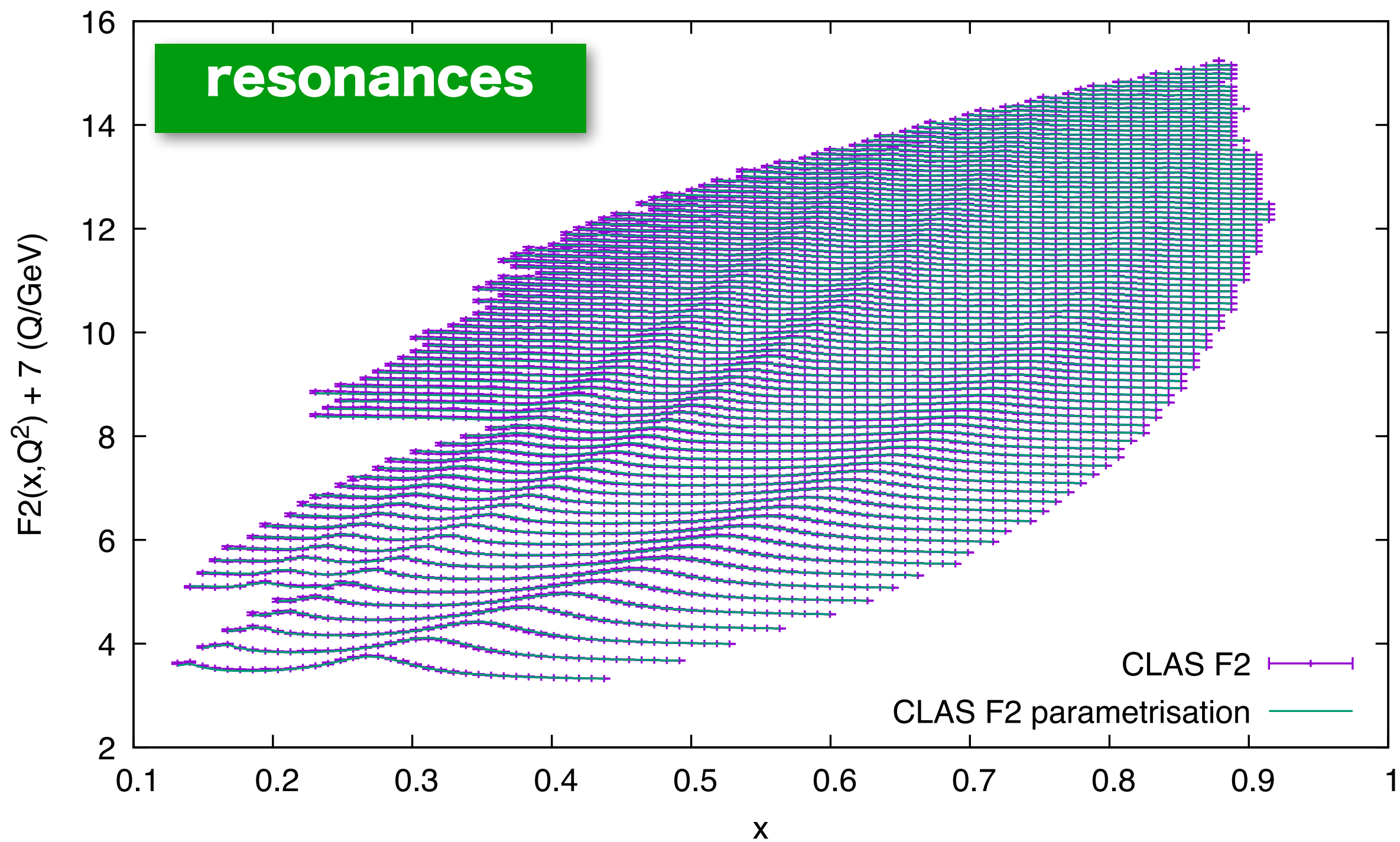
Dipole approximation,  
 $(\mu \rightarrow \infty \text{ in figure.})$

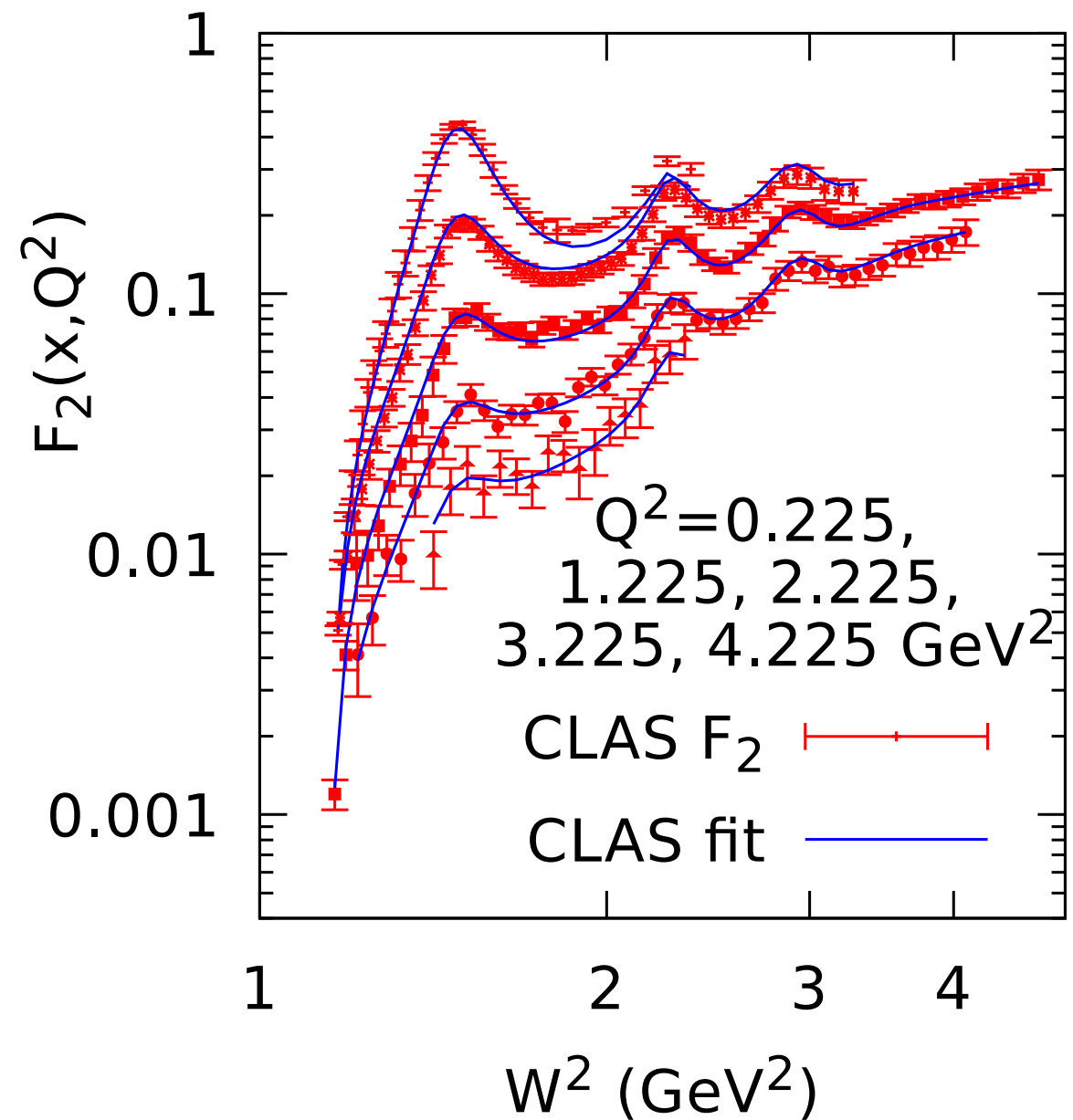
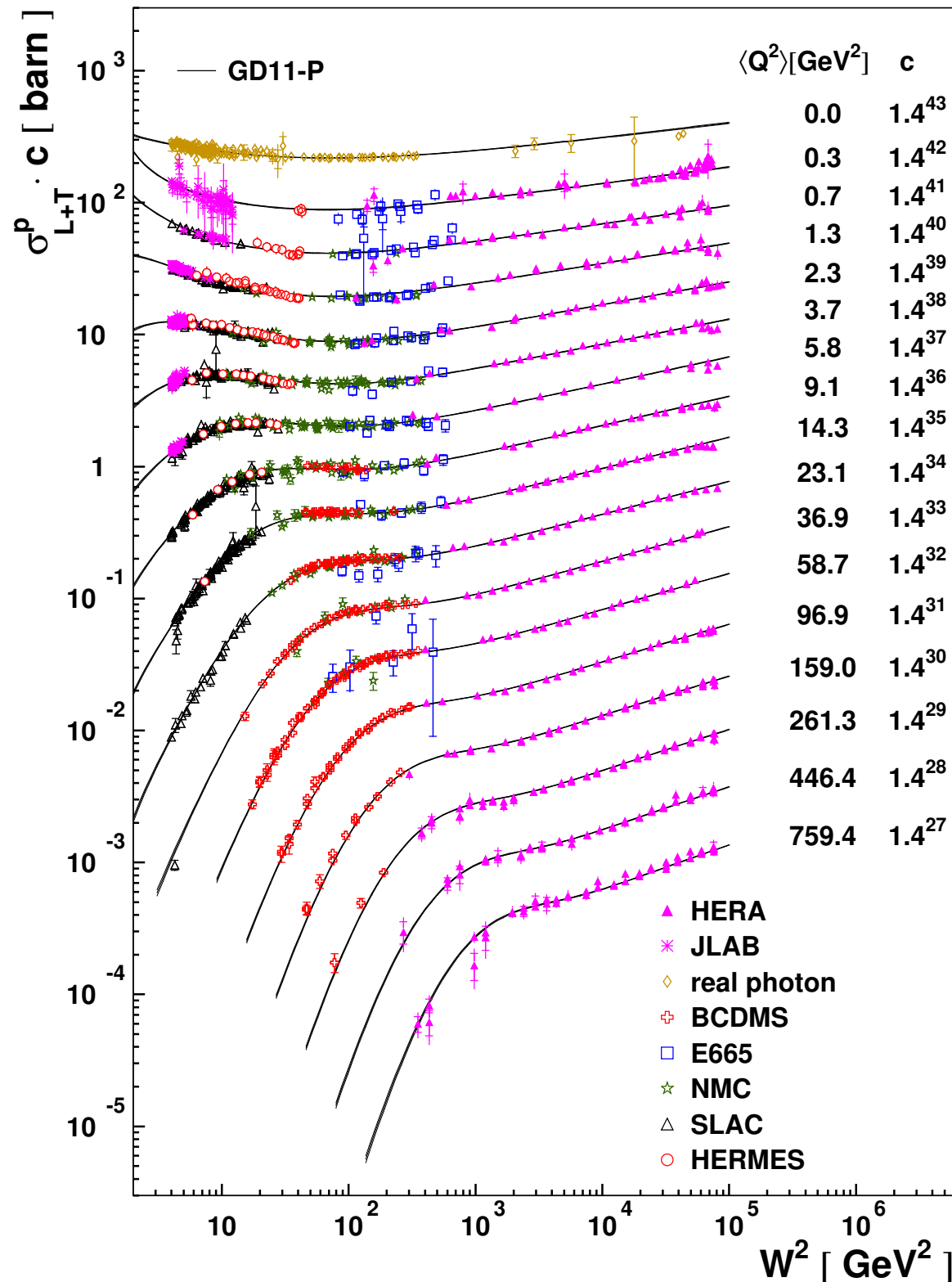
- ▶ Mostly  $G_E$  at small  $x$ .
- ▶ Mostly  $G_M$  at large  $x$ .
- ▶ Mostly from  $Q^2 < 1 \text{ GeV}$ .





# CLAS DATA





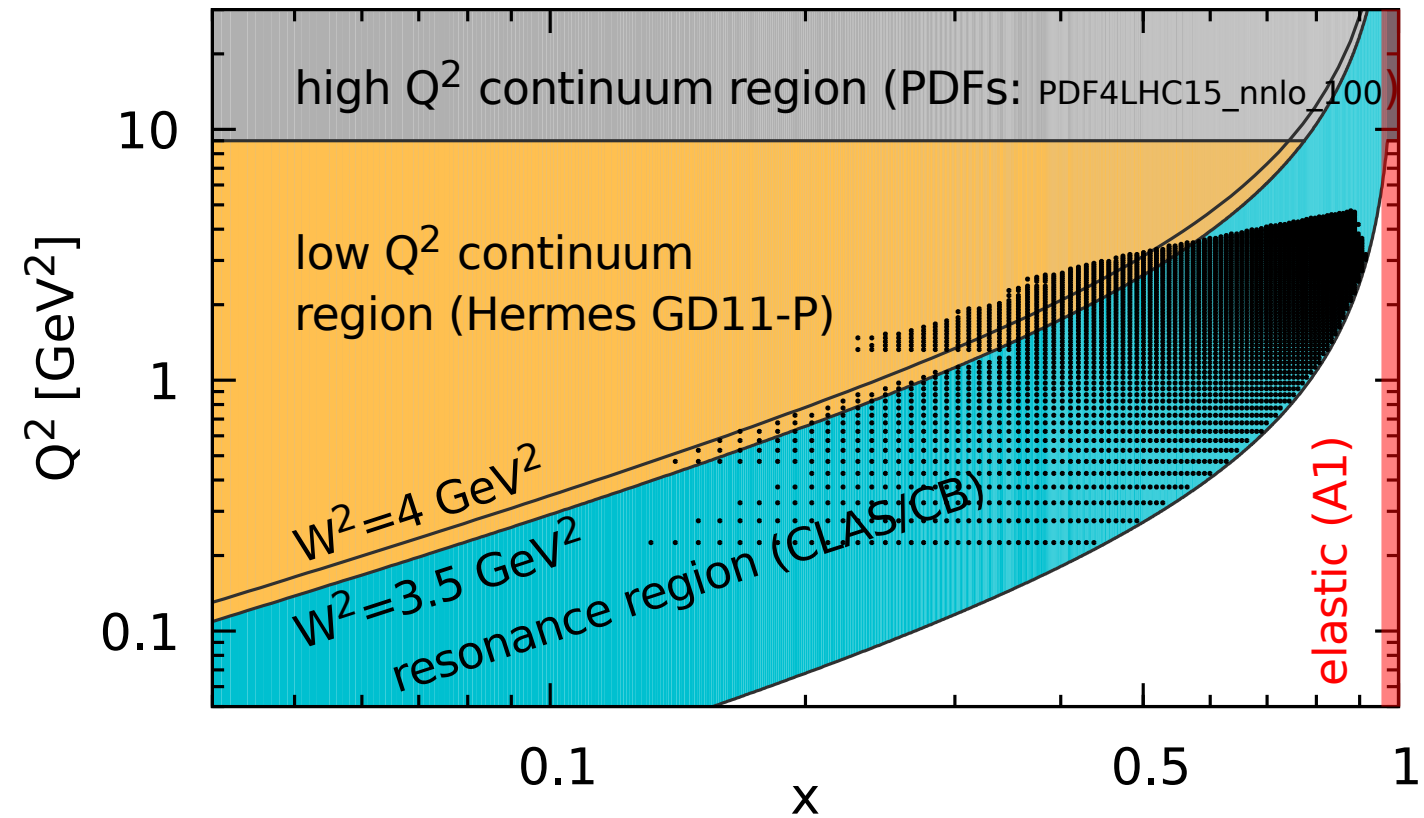
Fitted data from  $Q^2 = 0.225$  to  $4.725$  in steps of  $0.05 \text{ GeV}^2$ .

**Hermes fit:** we are interested in the region  $Q^2 < 10 \text{ GeV}^2$ .

**Continuum data region:**  $4 \text{ GeV}^2 < W^2 \lesssim 10^5 \text{ GeV}^2$  ( $x \rightarrow 10^{-4}$ ).

# Inelastic Data coverage

- ▶ Low  $Q^2$  continuum essentially covered by data.
- ▶  $F_2$  and  $F_L$  must **vanish as  $Q^2$  and  $Q^4$  at constant  $W$**  (by analyticity of  $W^{\mu\nu}$ ).



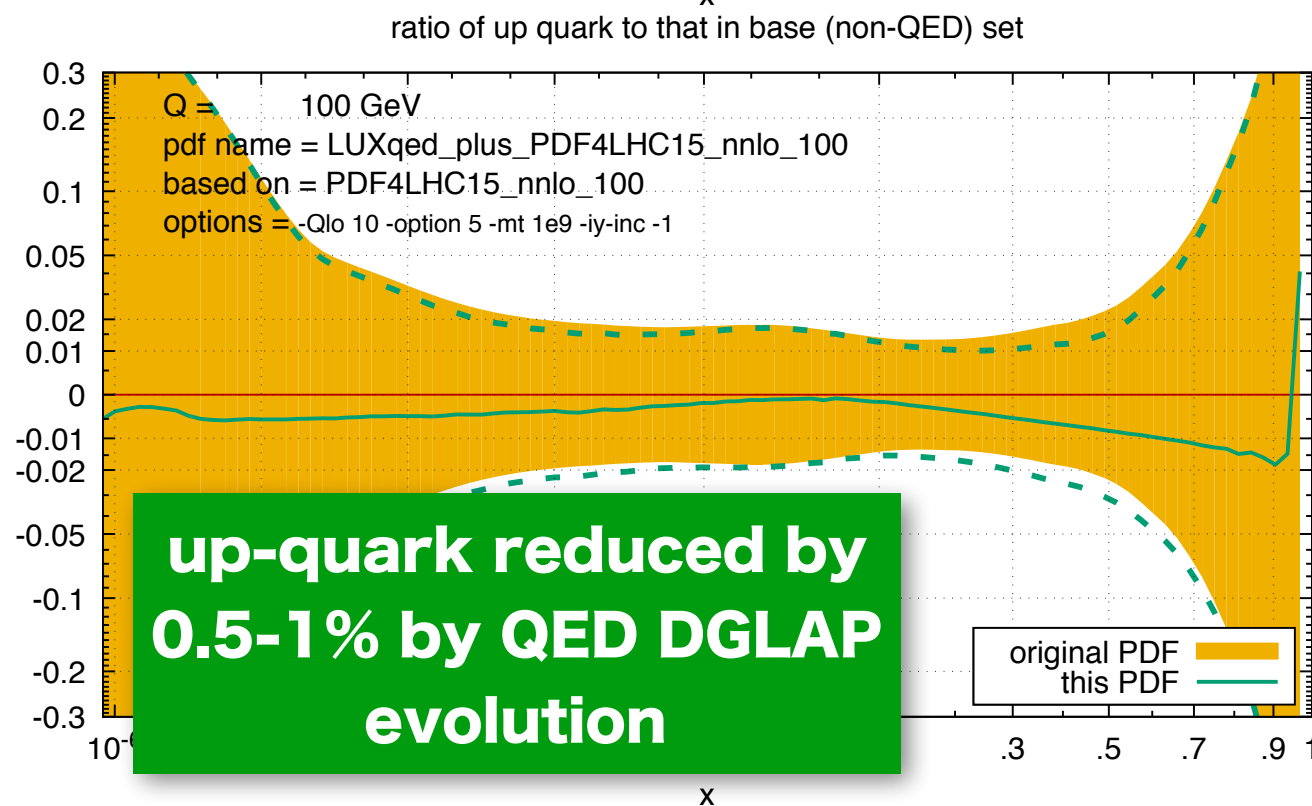
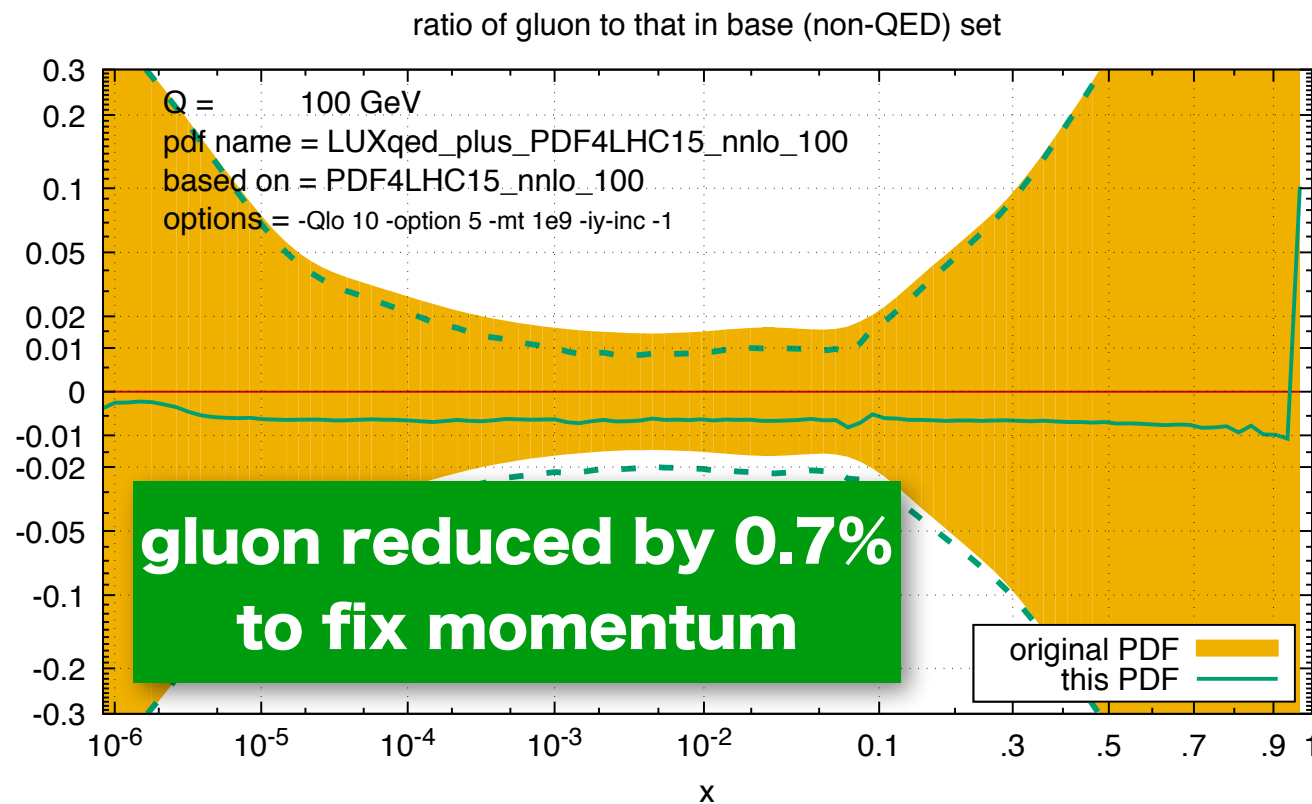
Also:

$$F_2(x, Q^2) = \frac{1}{4\pi^2\alpha} \frac{Q^2(1-x)}{1 + \frac{4x^2 m_p^2}{Q^2}} (\sigma_T(x, Q^2) + \sigma_L(x, Q^2)) \xrightarrow{Q^2 \rightarrow 0} \frac{Q^2 \sigma_{\gamma p}(W)}{4\pi^2\alpha^2}.$$

At small  $Q^2$ ,  $\sigma_T \implies \sigma_{\gamma p}(W)$ , becoming a function of  $W$  only (the  $CM$  energy in photoproduction), and  $\sigma_L$  vanishes.

**Photoproduction data included in Hermes and Christy-Bosted parametrizations.**

# MATCHING PROCEDURE FOR FULL SET OF PARTONS



- evaluate master eqn. for  $\mu = 100$  GeV (with default PDF4LHC15\_nnlo partons)
- Do  $O(\alpha_s)$  photon evolution down to  $\mu = 10$  GeV (other partons: pure QCD evln.)
- Adjust momentum sum-rule by rescaling gluon  $g(x) \rightarrow 0.993g(x)$
- Evolve back up with NNLO-QCD &  $O(\alpha_s)$  QED for all partons

**better approach would be full PDF re-fit for QCD partons incl. EW/QED corrections & LUXqed photon**

**forthcoming theory steps**

# Operator definition for unpolarized & polarized photon

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$$f_{\gamma}(x, \mu) = -\frac{1}{4\pi x p^+} \int_{-\infty}^{\infty} dw e^{-ixw p^+} \langle p | F^{n\lambda}(wn) F^n_{\lambda}(0) + F^{n\lambda}(0) F^n_{\lambda}(wn) | p \rangle_c .$$

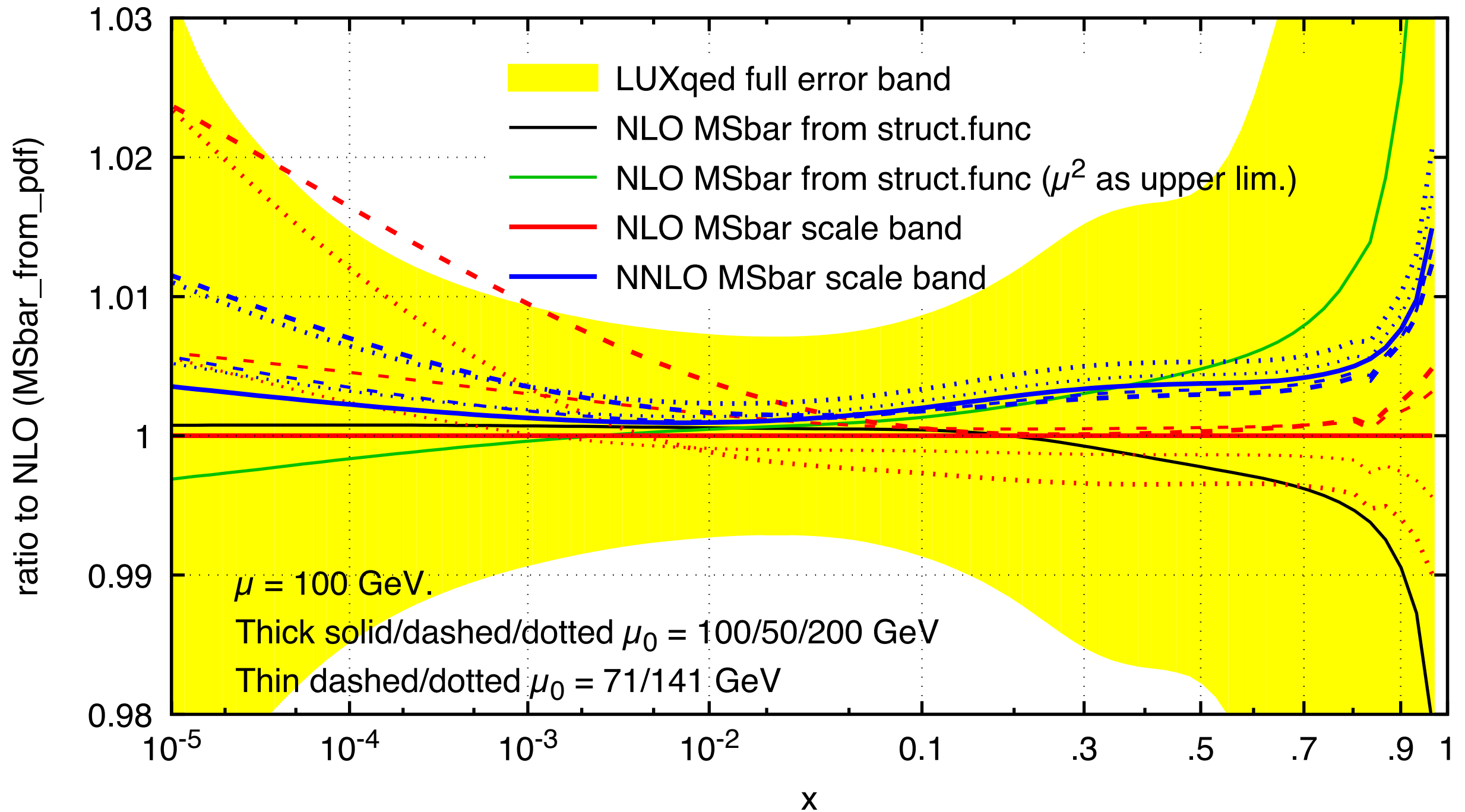
$$f_{\Delta\gamma}(x, \mu) = \frac{i}{4\pi x p^+} \int_{-\infty}^{\infty} dw e^{-ixw p^+} \langle p | F^{n\lambda}(wn) \tilde{F}^n_{\lambda}(0) - F^{n\lambda}(0) \tilde{F}^n_{\lambda}(wn) | p \rangle_c ,$$

$$\tilde{F}_{\alpha\beta} = \frac{1}{2} \epsilon_{\alpha\beta\lambda\sigma} F^{\lambda\sigma}$$

Makes it easier to go to higher orders

# higher-order contributions

NLO and NNLO (in  $\alpha_s$ ) photon extraction and scale dep.

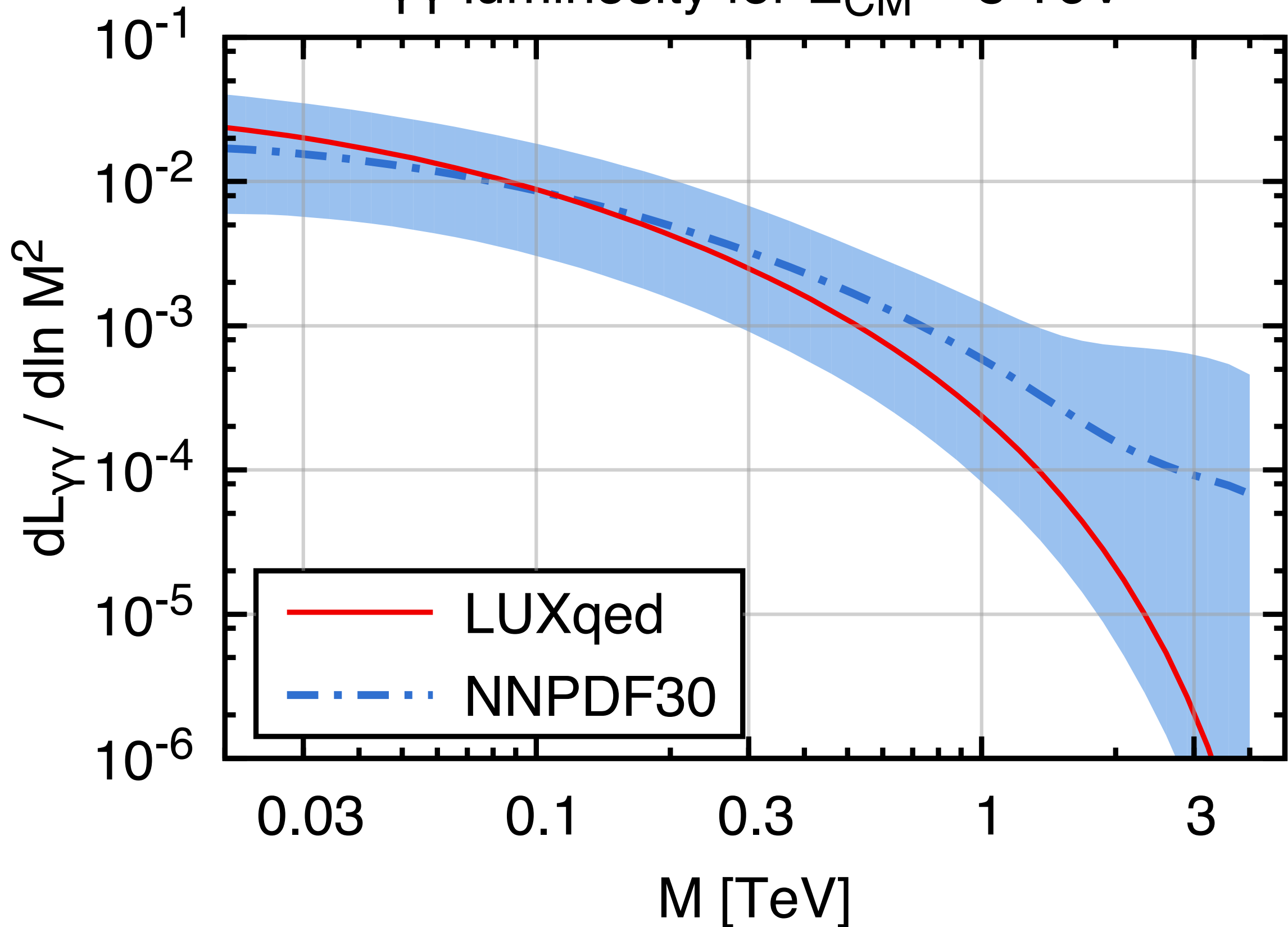


**comparisons to others**



# $\gamma\gamma$ luminosity

$\gamma\gamma$  luminosity for  $E_{\text{CM}} = 8$  TeV



# explanation does not lie with NNPDF23 v. 30 evolution differences

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