Higgs and Beyond at Colliders Gavin P. Salam*

Rudolf Peierls Centre for Theoretical Physics All Souls College

* on leave from CERN and CNRS





European Research Council Established by the European Commission



THE ROYAL SOCIETY

Oxford Department of Physics Colloquium 16 November 2018



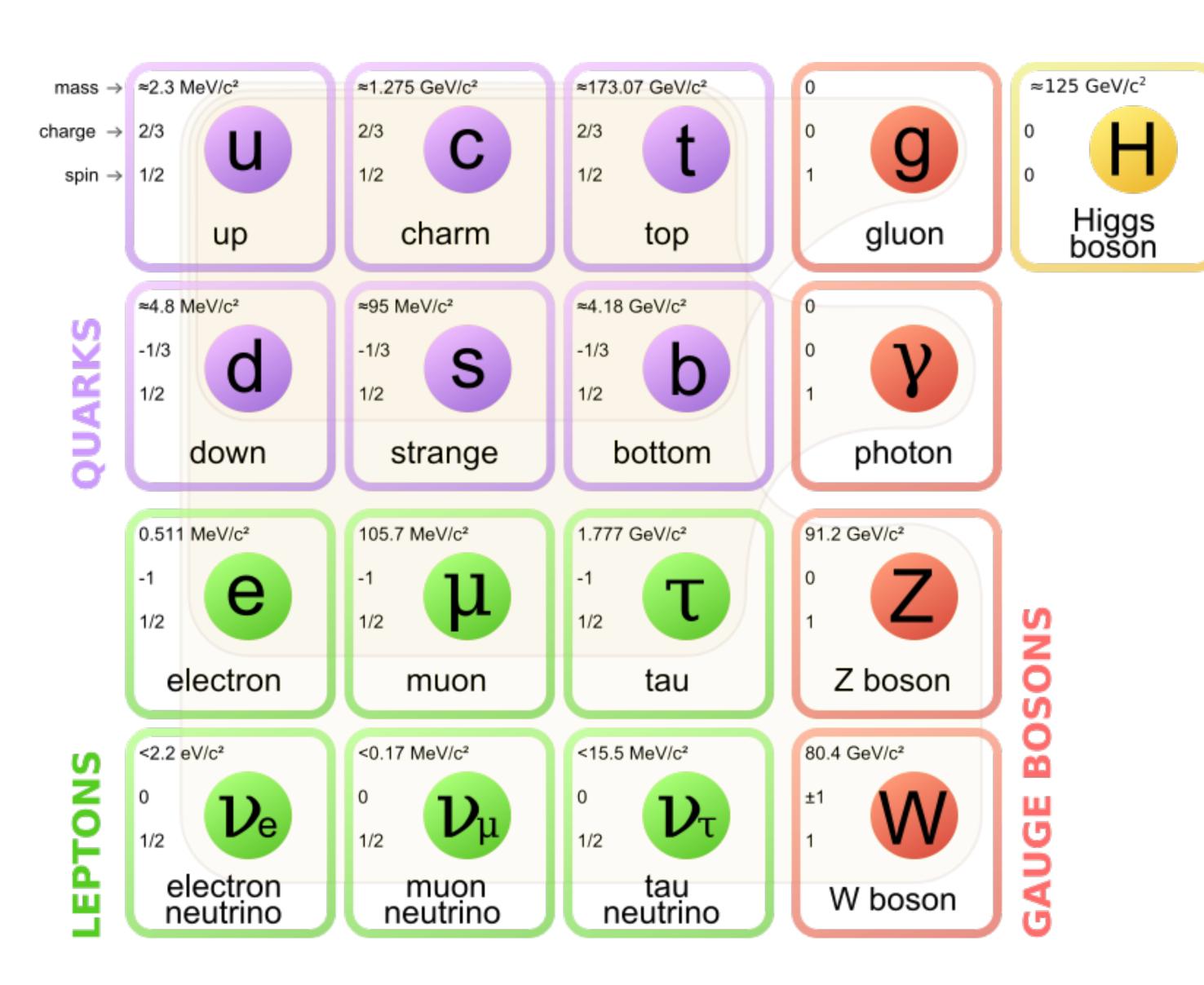
"big answerable questions" and how we go about answering them

"big unanswered questions" about fundamental particles & their interactions (dark matter, matter-antimatter asymmetry, nature of dark energy, hierarchy of scales...)

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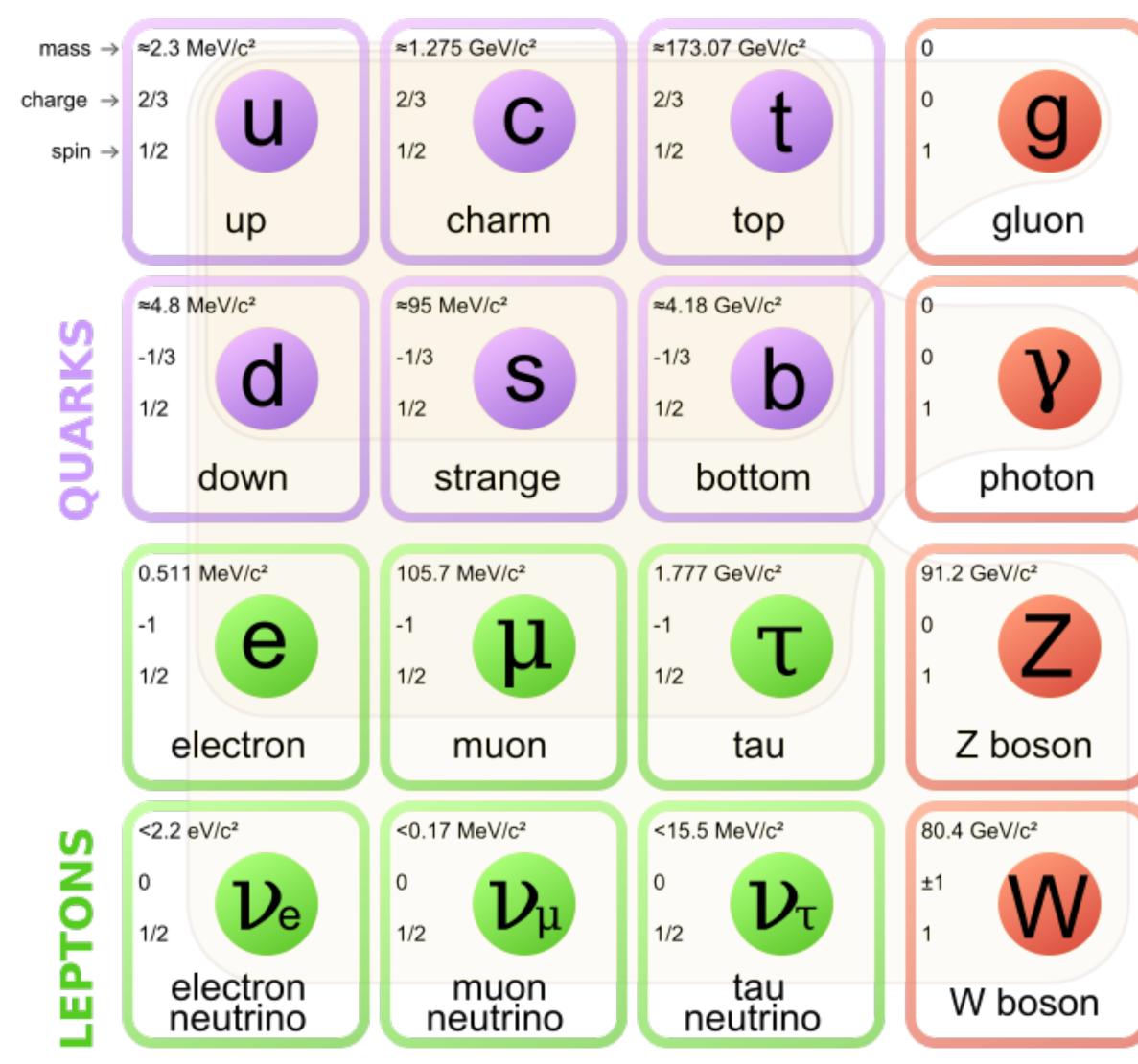










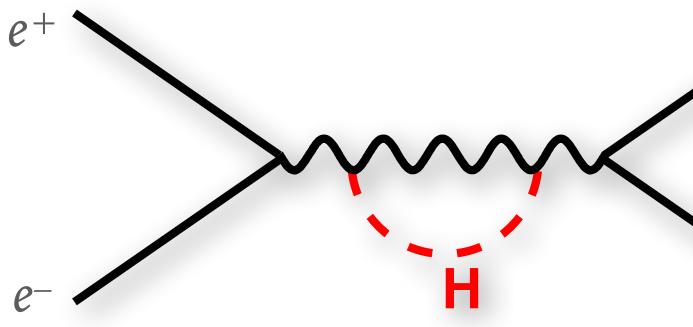






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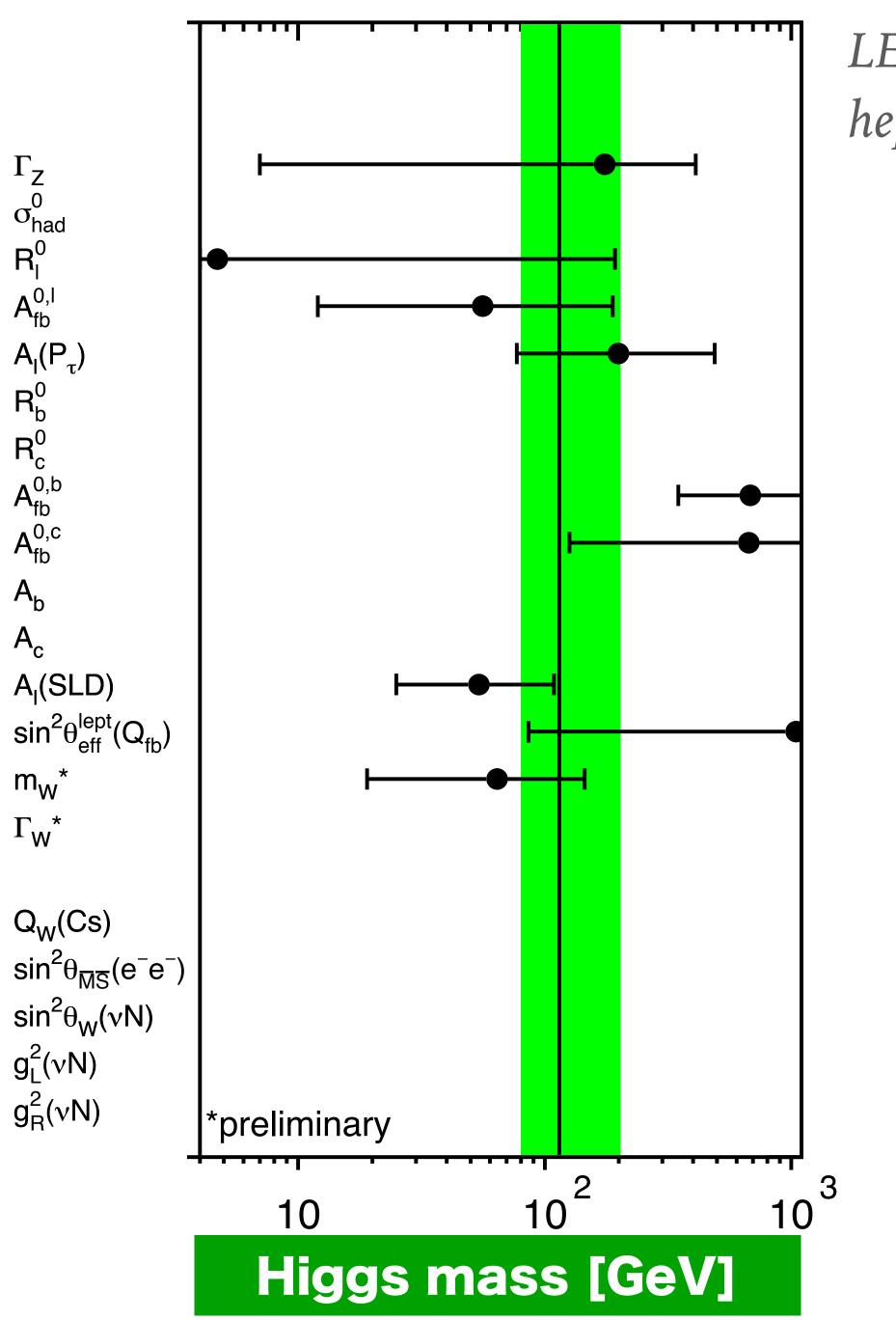


Higgs boson existence long known to be consistent with older e⁺e⁻ collider data (cf. LEP, 1989–2000 + SLD).

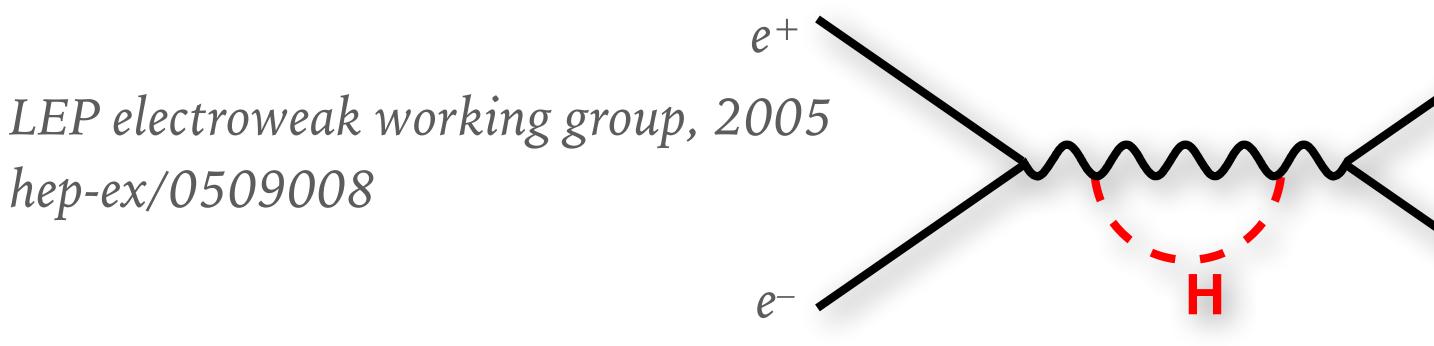
Tested through the small effect of virtual Higgs bosons on high-precision *(per-mil)* measurements.

Could be interpreted as a weak Higgs mass constraint.





hep-ex/0509008

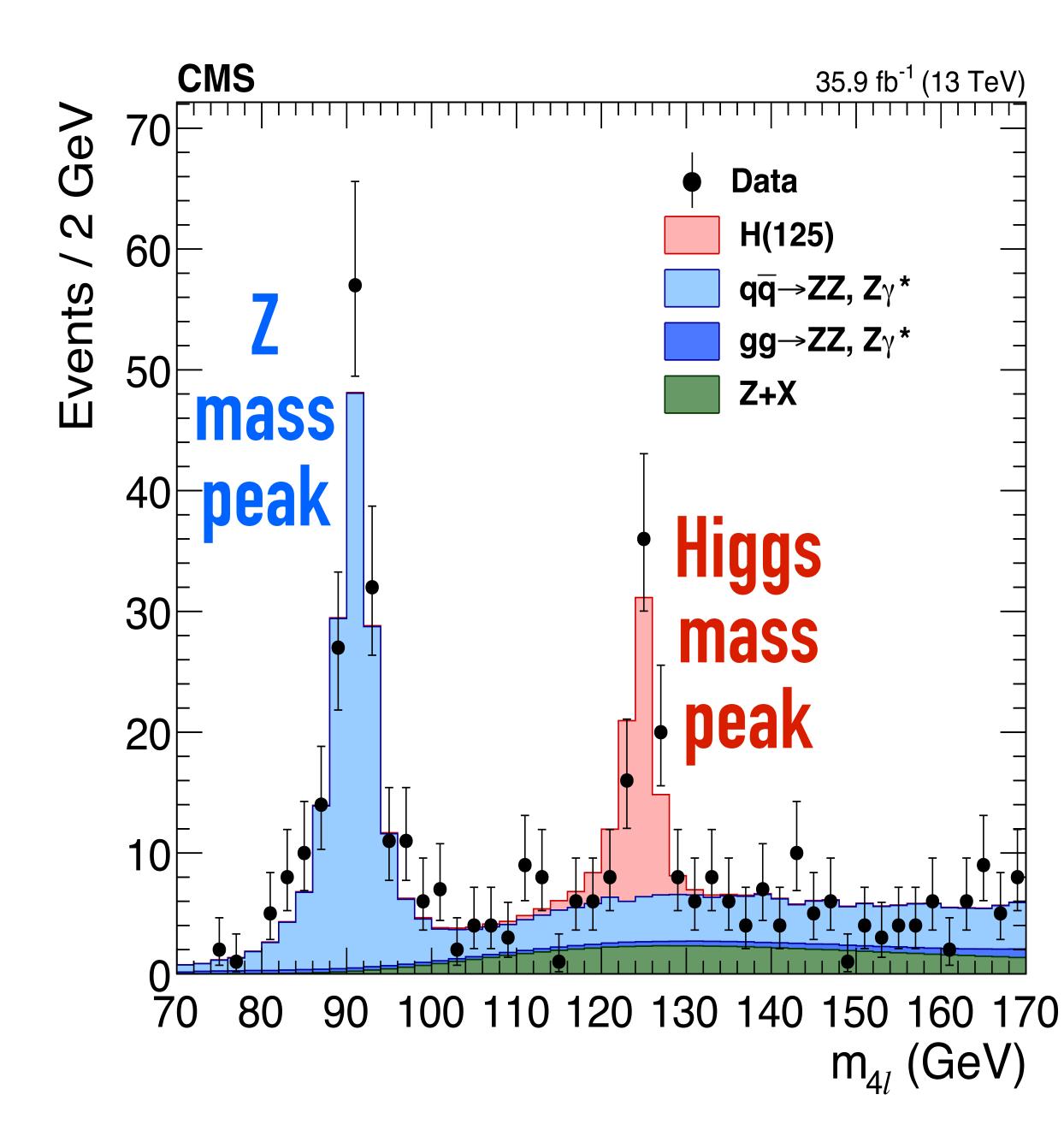


Higgs boson existence long known to be consistent with older e+e- collider data (cf. LEP, 1989-2000 + SLD).

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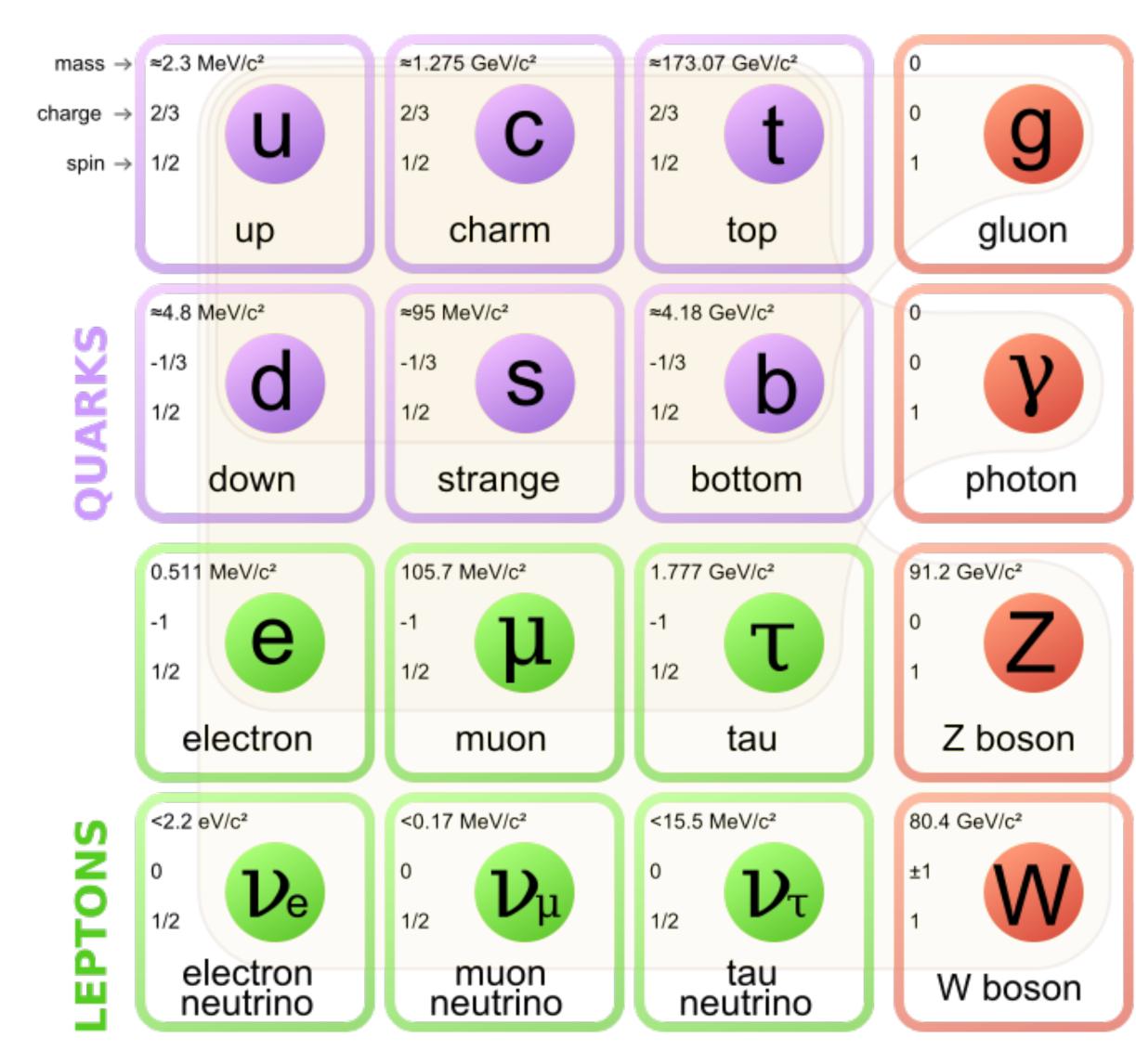
ATLAS and CMS collaborations at CERN's Large Hadron Collider (LHC):

2012 discovery of a Higgs-like boson

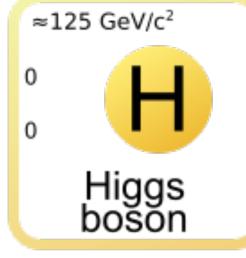
plot shows more recent data











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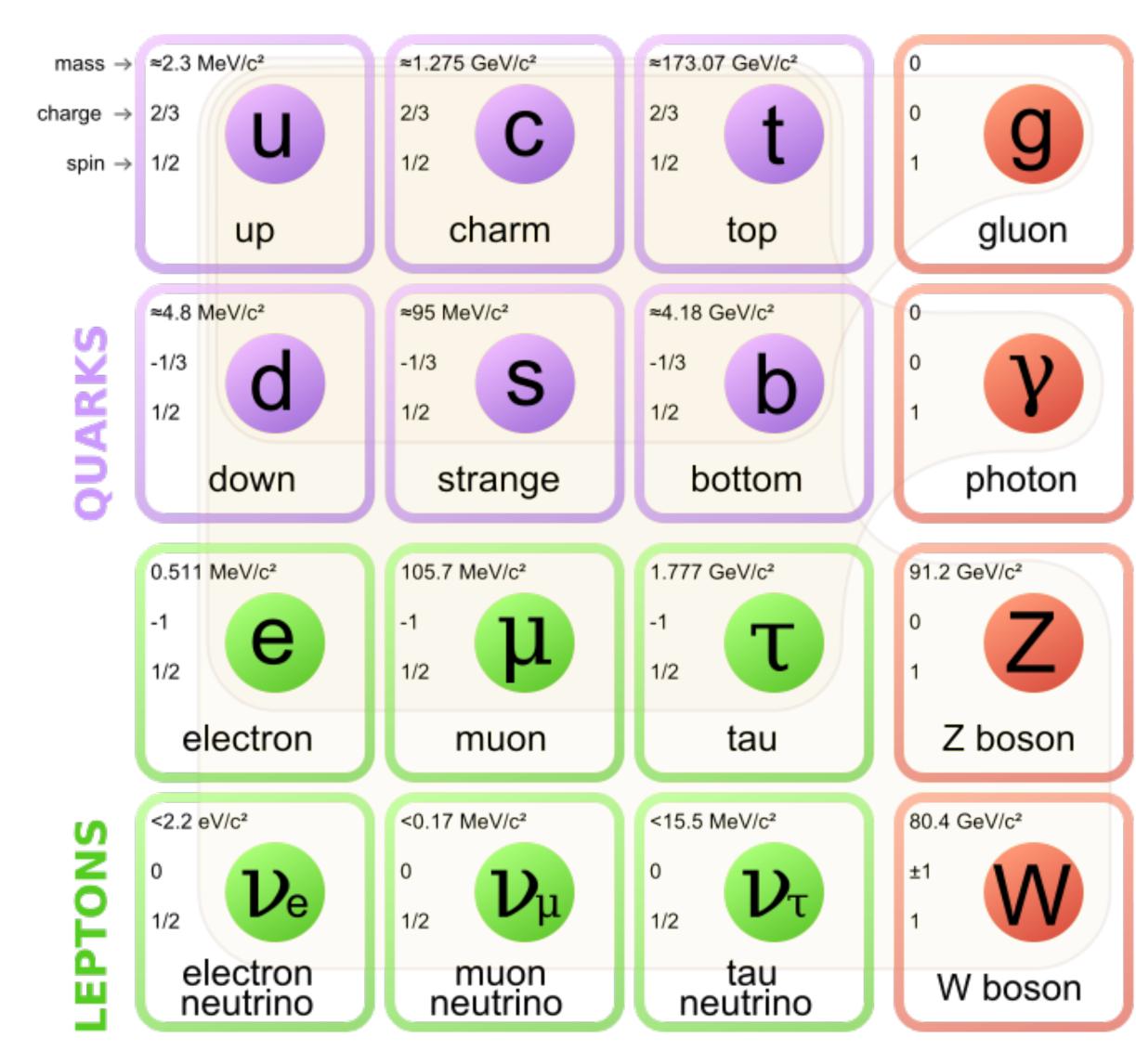
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Success!

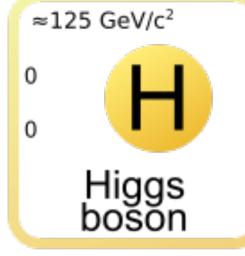
"The Standard Model is complete"











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Success!

"The Standard Model is complete"

Crisis!

No supersymmetry, no extra dimensions, there's nothing left for us to do . . .



The New York Eines

By DENNIS OVERBYE JUNE 19, 2017

|...| a cloud hanging over the physics community. [...]



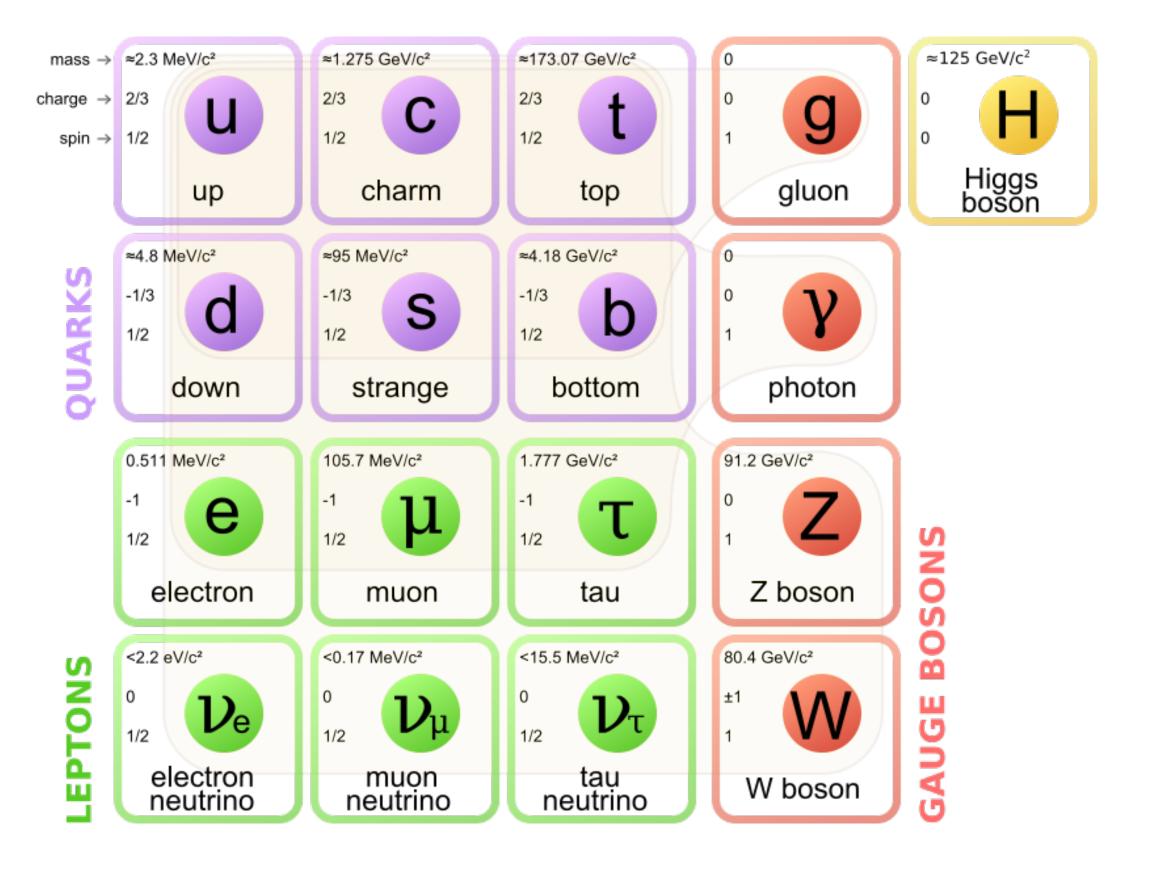
What if there is nothing new to discover? That prospect is now

https://www.nytimes.com/2017/06/19/science/cern-large-hadron-collider-higgs-physics.html





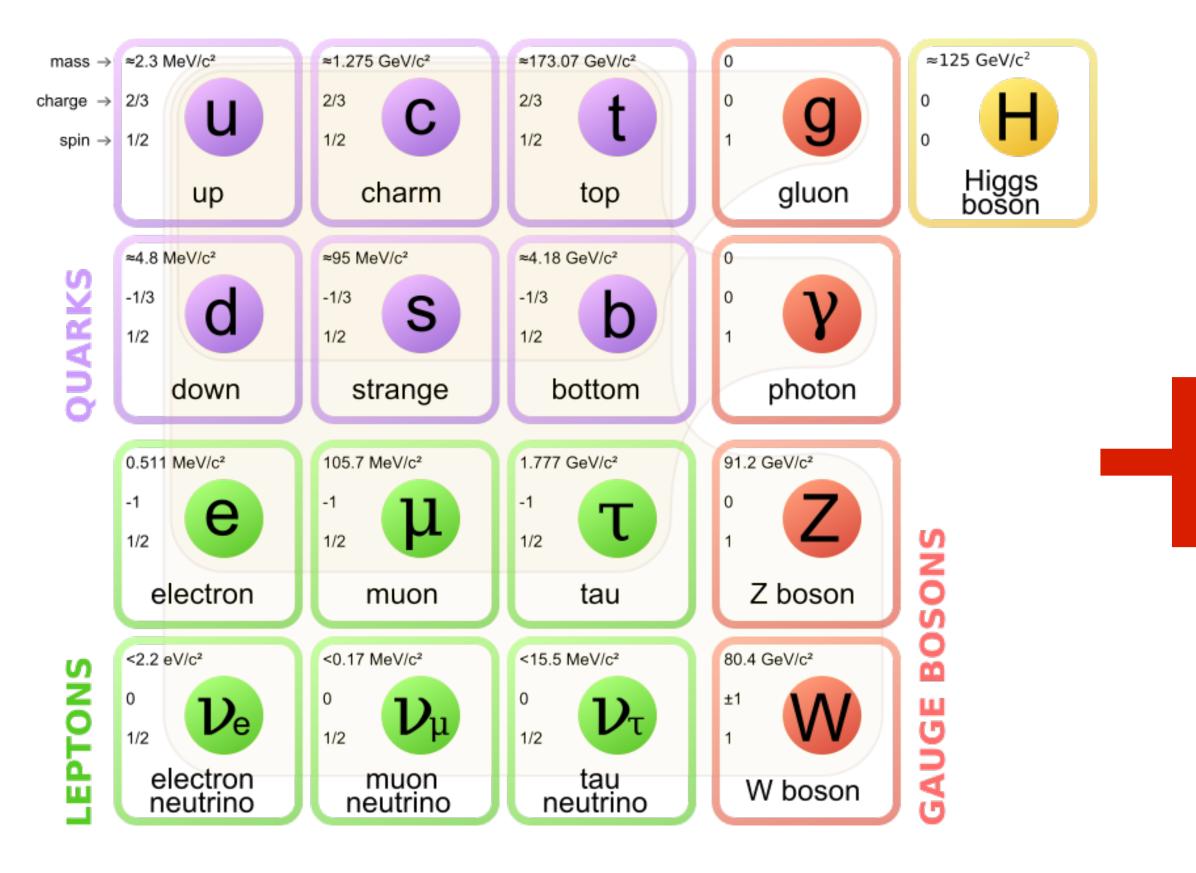
what is the Standard Model?



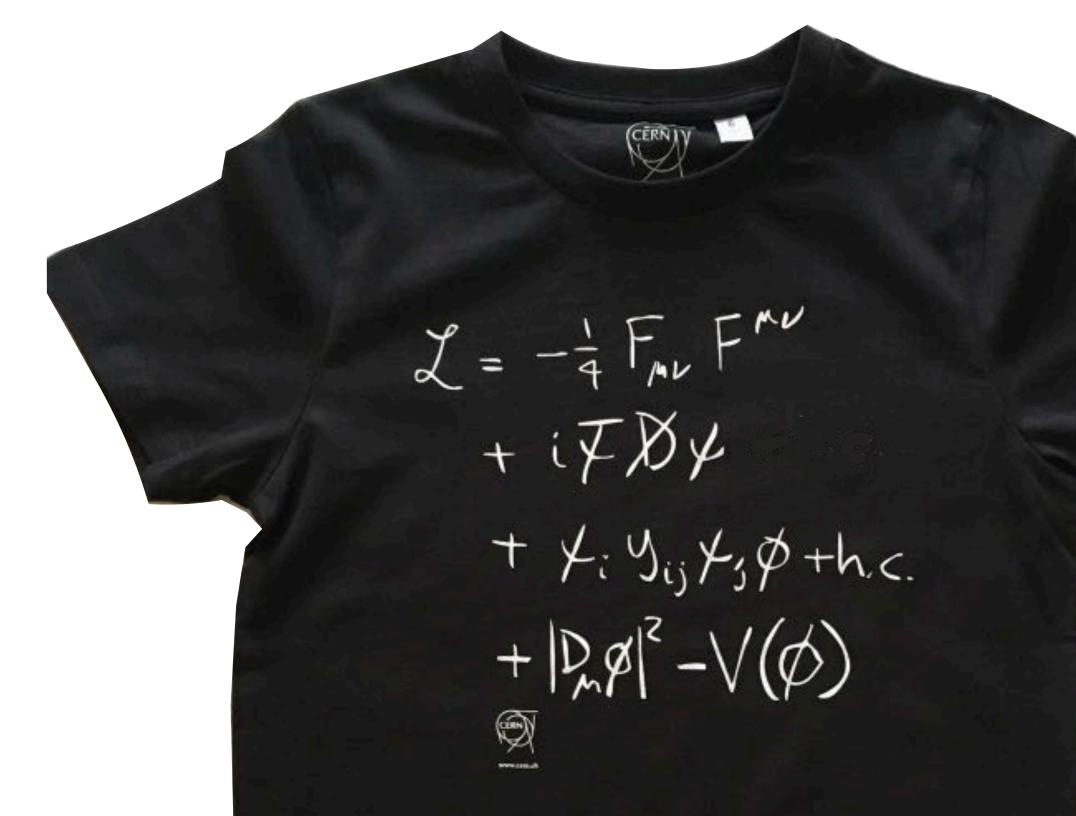
particles



what is the Standard Model?



particles



interactions





Z = - FALFALFAL + iFDY + X: Jij X; \$+h.c. + D g (-V(d))

This equation neatly sums up our current understanding of fundamental particles and forces.

STANDARD MODEL — KNOWABLE UNKNOWNS

This is what you get when you buy one of those famous CERN T-shirts







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STANDARD MODEL — KNOWABLE UNKNOWNS

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"understanding" = knowledge ? "understanding" = assumption ?



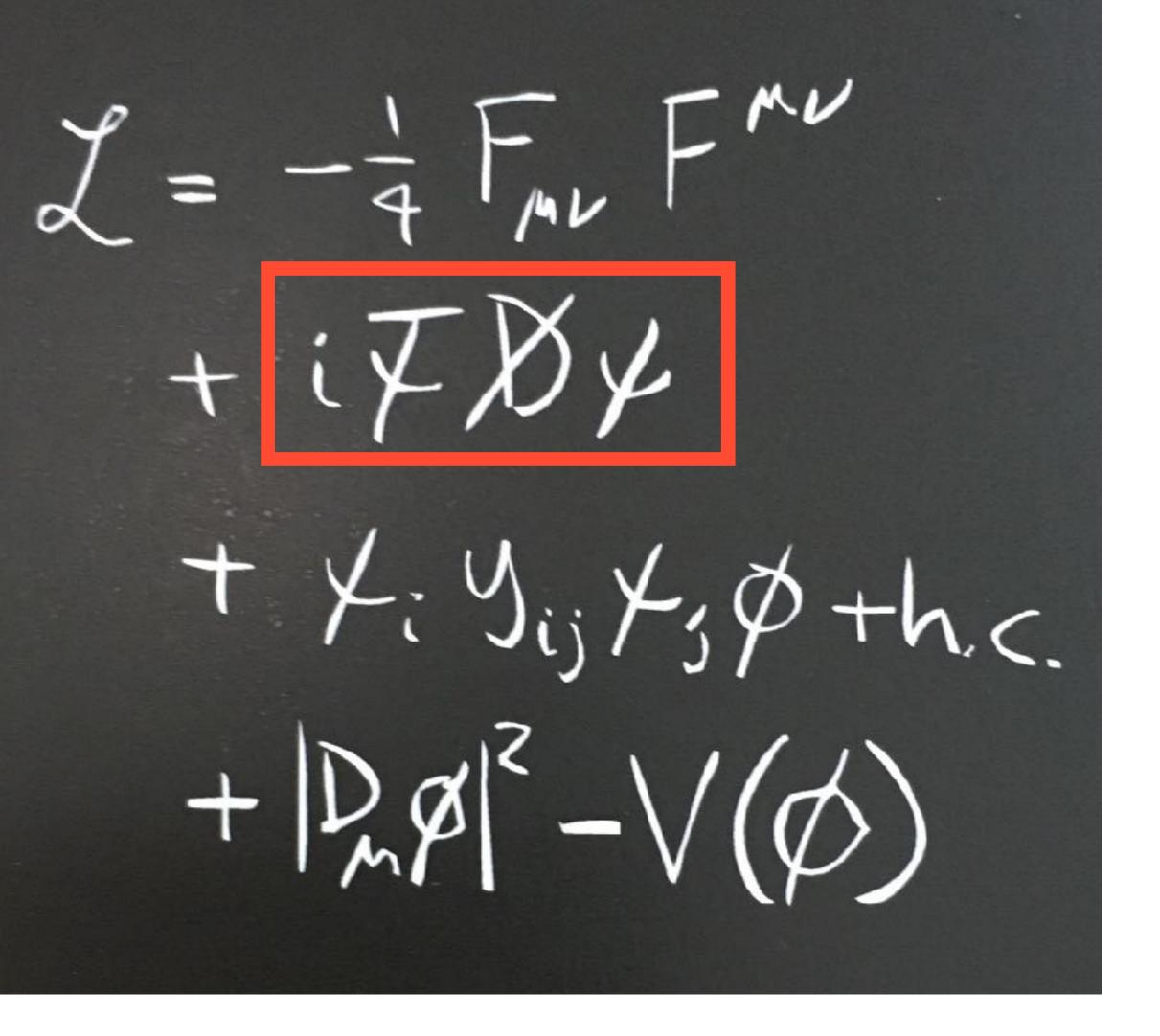












NOTATION

- A_{μ} : gauge field
 - ψ : fermion field

photons, gluons, W,Z

quarks & leptons

 ϕ : Higgs field $= \phi_0(\text{VEV}) + H(\text{Higgs})$

 $D_{\mu} = \partial_{\mu} + ieA_{\mu}$ etc. $F_{\mu\nu} \sim [D_{\mu}, D_{\nu}]$



= - + FAL F + X: Jij X; Ø+h.C. + Dg(-V(d))

e.g. $\psi D\psi \rightarrow \psi A_{\mu}\psi \rightarrow$ fermion-fermion-gauge vertex i.e. terms of \mathcal{L} map to particle interactions

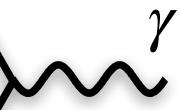
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2 = - + Fmu + i FN + 4: 5; 4; \$+h.c. $+ \left| \mathcal{D} \mathcal{P} \right|^{<} - \mathcal{V} \left(\mathcal{O} \right)$

e.g. $F_{\mu\nu}F^{\mu\nu} \to A_{\mu}A_{\nu}\partial_{\mu}A_{\nu} \to \text{triple-gauge vertex}$ i.e. terms of \mathcal{L} map to particle interactions

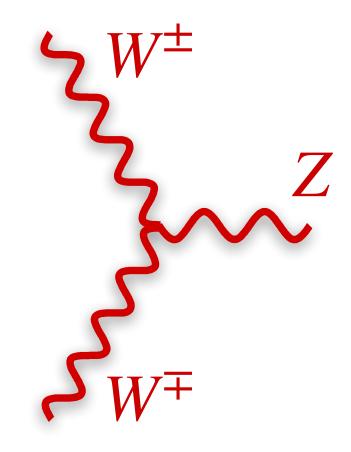
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ナ Y: Jii Y, Ø +

This equation neatly sums up our current understanding of fundamental particles and forces.

GAUGE PART

e.g. qqγ, qqZ, qqg, evW, ggg, interactions — well established in ep, e⁺e[−], pp collisions, etc. **≡ KNOWLEDGE**

(also being studied at LHC — e.g. jets, DY/Z/W, V+jets, ttbar, etc.)



 $t \chi: \mathcal{Y}_{ij} \chi_{j} \phi$ $+ |D_{\mathcal{P}}(-V(\mathcal{O}))$

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(also being studied at LHC — e.g. jets, DY/Z/W, V+jets, ttbar, etc.)

Many SM studies probe this part. In some respects dates back to 1860's, i.e. Maxwell's equations.

If you test another corner of this (as one should), don't be surprised if it works







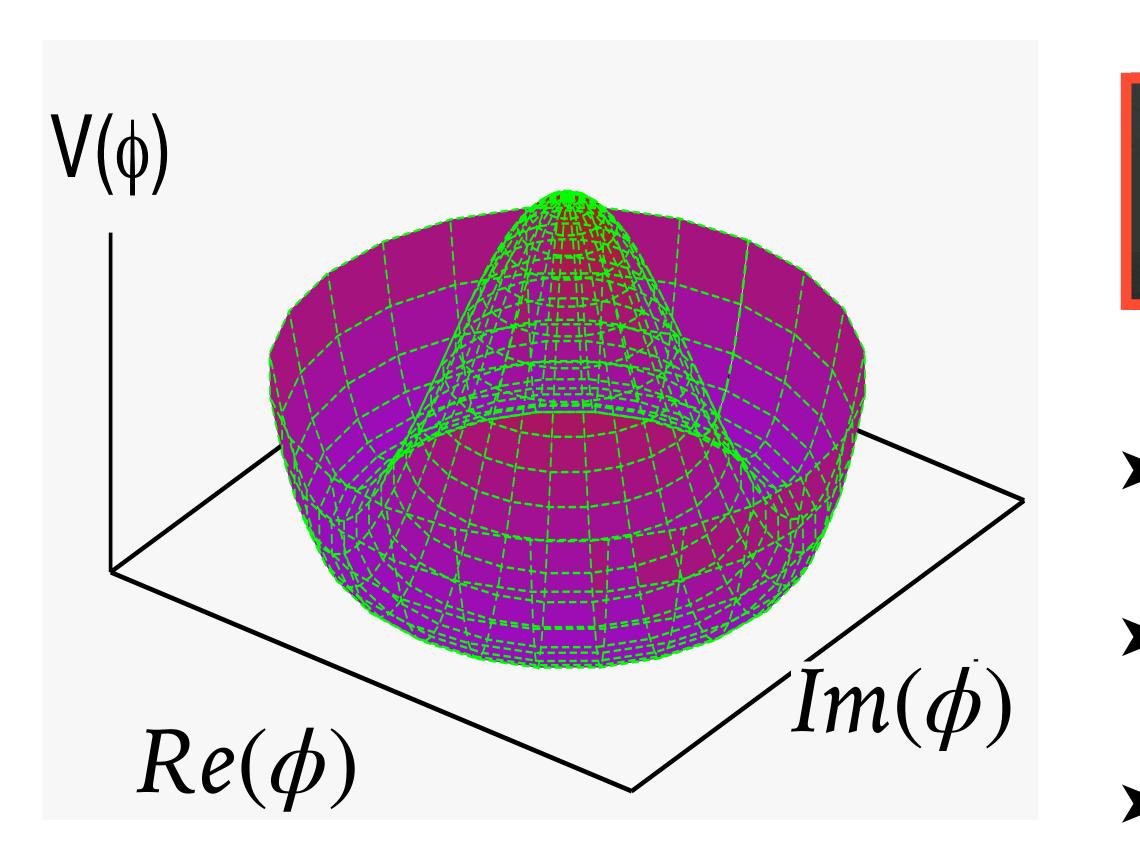


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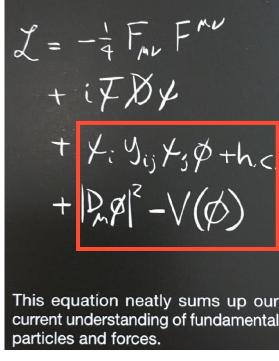
Higgs sector

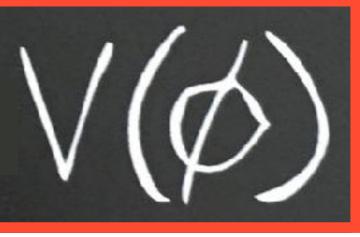
until 6 years ago none of these terms had ever been directly observed.

what terms are there in the Higgs sector? 1. Potential



Gavin Salam



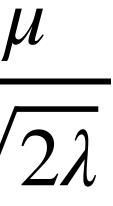


 $(\phi) = -\mu^2 \phi^2 + \lambda \phi^4$

 $\blacktriangleright \phi$ is a complex doublet

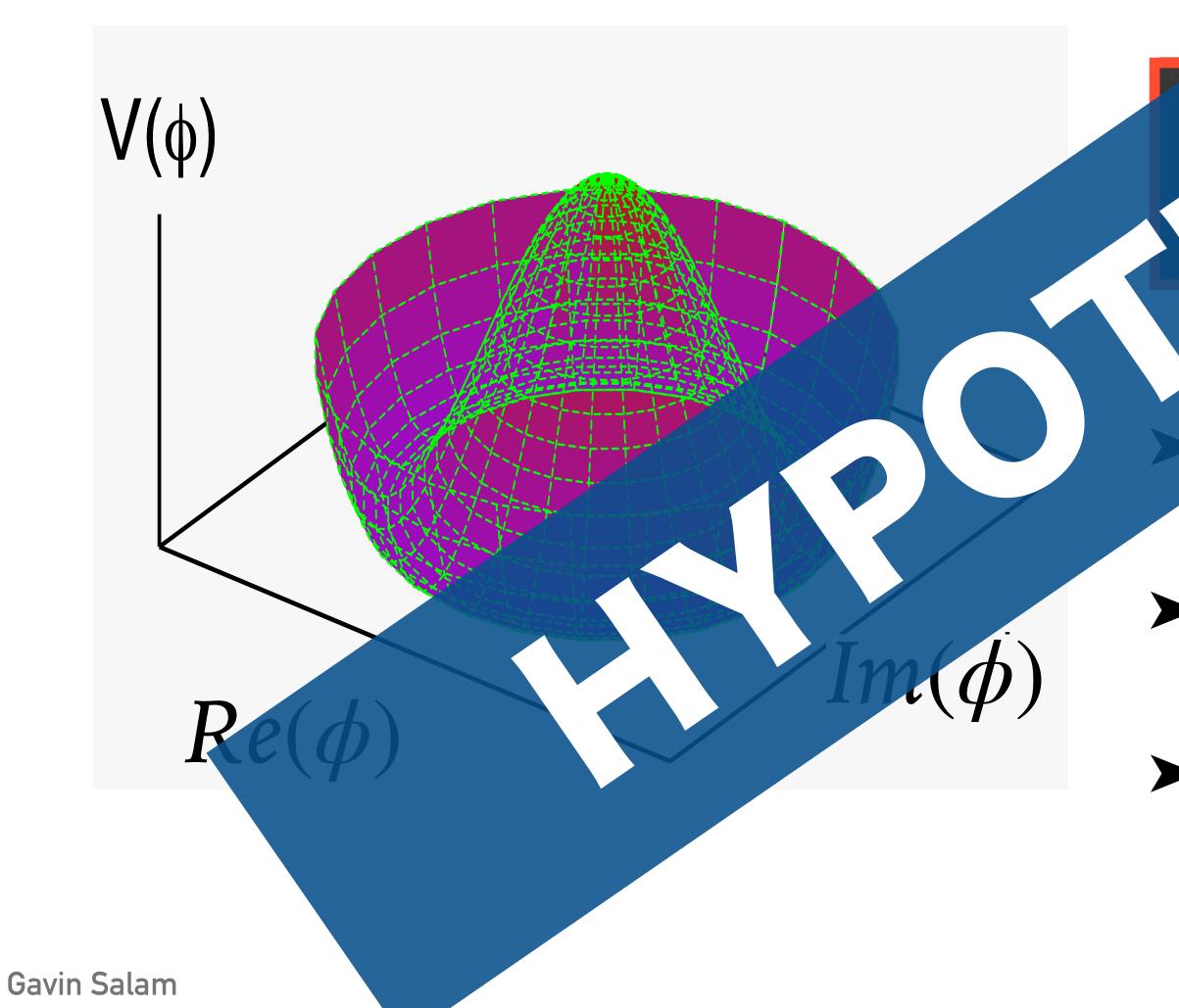
 \blacktriangleright V(ϕ) has minimum at $\phi = \phi_0 = 0$

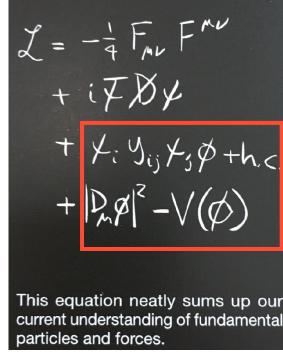
 \blacktriangleright Excitations of the φ field around φ_0 are Higgs bosons ($\varphi = \varphi_0 + H$)





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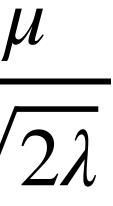


φ is a complex doublet

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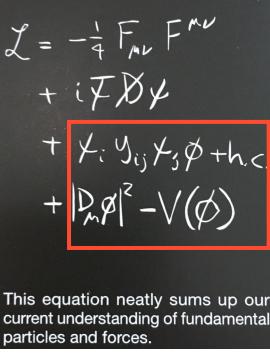
what terms are there in the Higgs sector? 2. Gauge-Higgs term



 $\to g^2 \phi_0^2 Z_{\mu} Z^{\mu} + 2g^2 \phi_0 H Z_{\mu} Z^{\mu} + \dots$

Z-boson mass term

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ZZH interaction term

 $D_{\mu} \sim \partial_{\nu} + ieZ_{\mu} + \dots$ $\phi = \phi_0 + H$

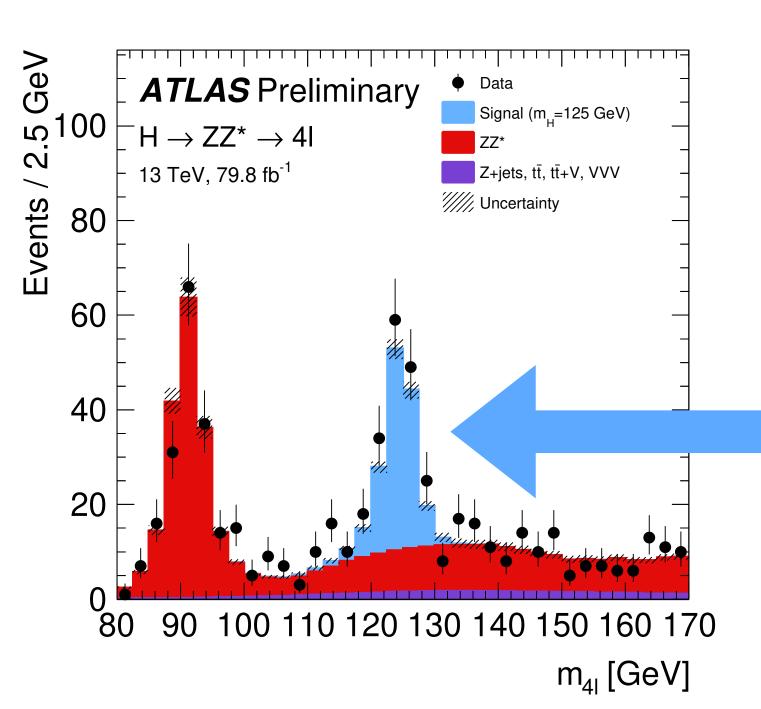


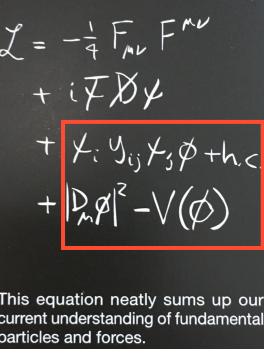
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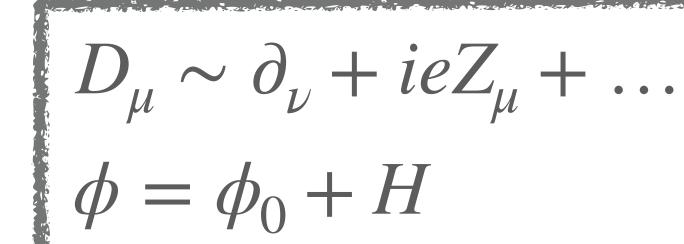
 $\rightarrow g^2 \phi_0^2 Z_\mu Z^\mu + 2g^2 \phi_0 H Z_\mu Z^\mu + \dots$

Z-boson mass term



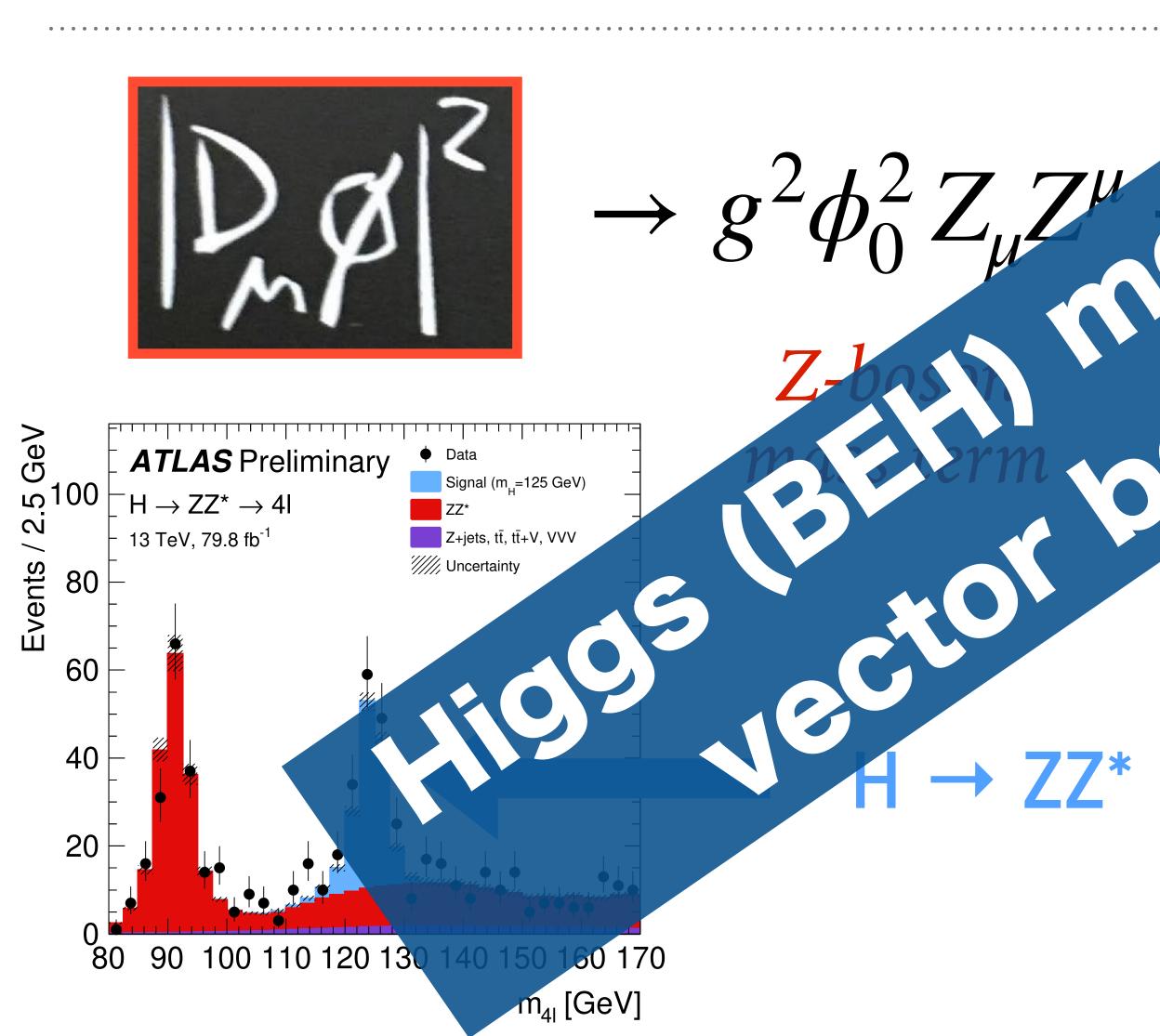


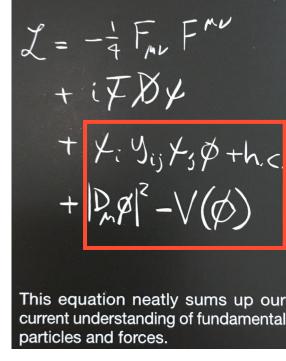
ZZH interaction term





what terms are there in the Higgs sector? 2. Gauge-Higgs term





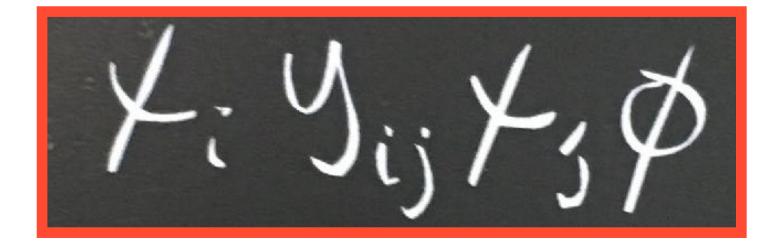
interaction

term

 $D_{\mu} \sim \partial_{\nu} + ieZ_{\mu} + \dots$ $\phi = \phi_0 + H$



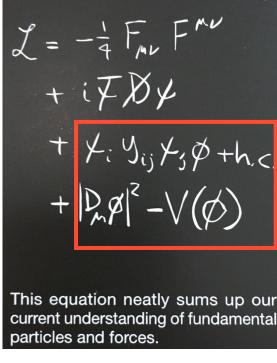
what terms are there in the Higgs sector? 3. Fermion-Higgs (Yukawa) term



fermion mass term $m_i = y_{ii}\phi_0$

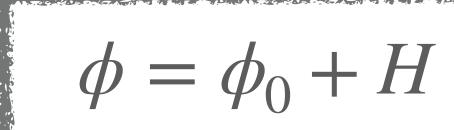
i	Уi	i	Уi
u	$2 \cdot 10^{-5}$	d	$3 \cdot 10^{-5}$
С	$8 \cdot 10^{-3}$	S	$6 \cdot 10^{-4}$
b	$3 \cdot 10^{-2}$	t	1
ν_e		е	$3 \cdot 10^{-6}$
$ u_{\mu} $	$\sim 10^{-13}$	μ	$6 \cdot 10^{-4}$
$ u_{ au} $		au	$1\cdot 10^{-4}$

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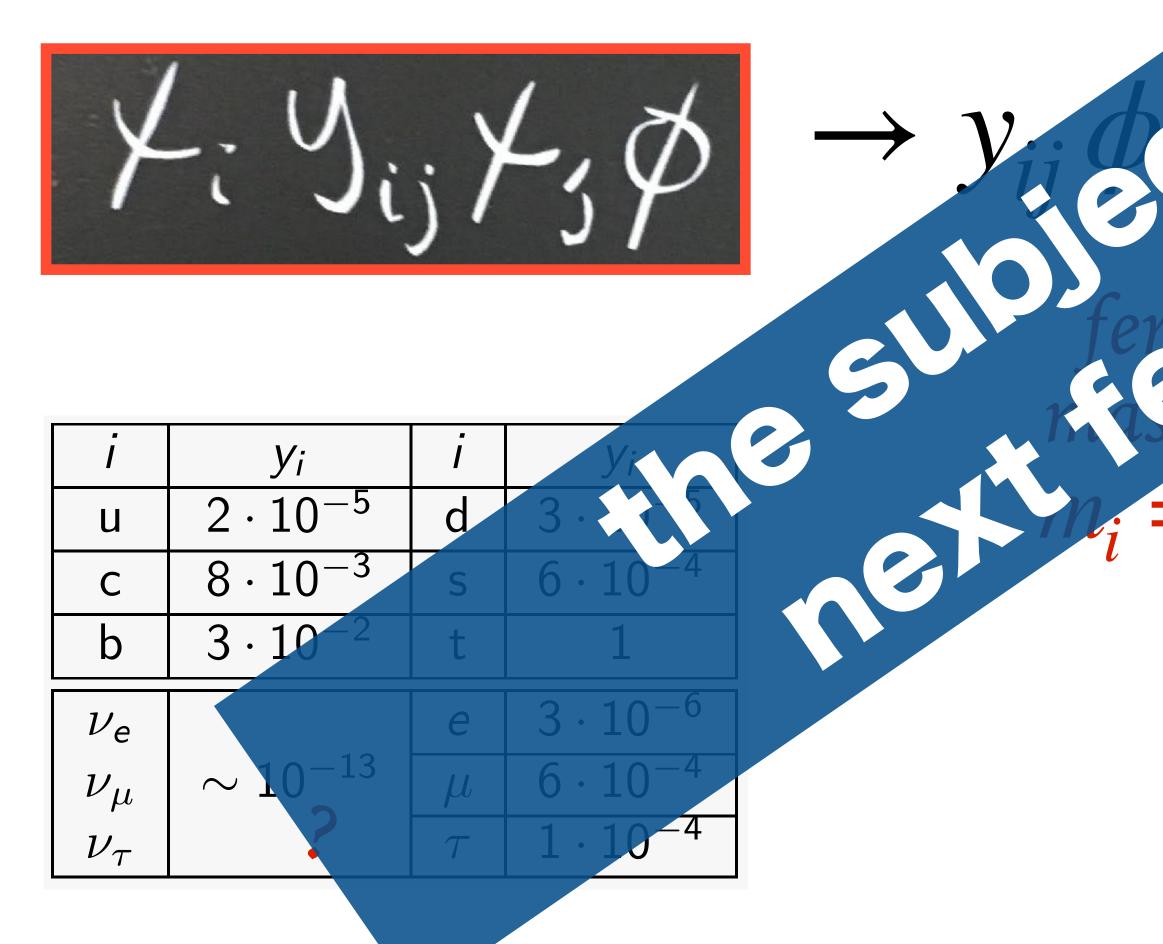


$\mathcal{Y}_{ij} \mathcal{Y}_{ij} \mathcal{Y}_{ij} \phi \rightarrow y_{ij} \phi_0 \psi_i \psi_j + y_{ij} H \psi_i \psi_i$

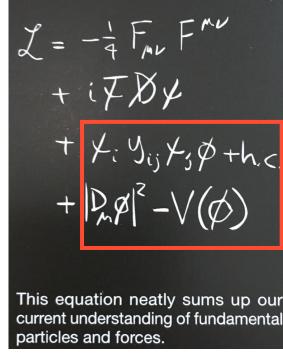
fermion-fermion-Higgs interaction term; coupling $\sim \gamma_{ii}$



what terms are there in the Higgs sector? 3. Fermion-Higgs (Yukawa) tern



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fermion-fermion-Higgs interaction term; coupling $\sim y_{ii}$

 $\psi_{ii}H\psi_i\psi_j$

 $\phi = \phi_0 + H$

concentrate on Yukawa interaction hypothesis

Yukawa couplings ~ fermion mass

first fundamental interaction that we probe at the quantum level where interaction strength is **not quantised** (i.e. no underlying unit of charge across particles)



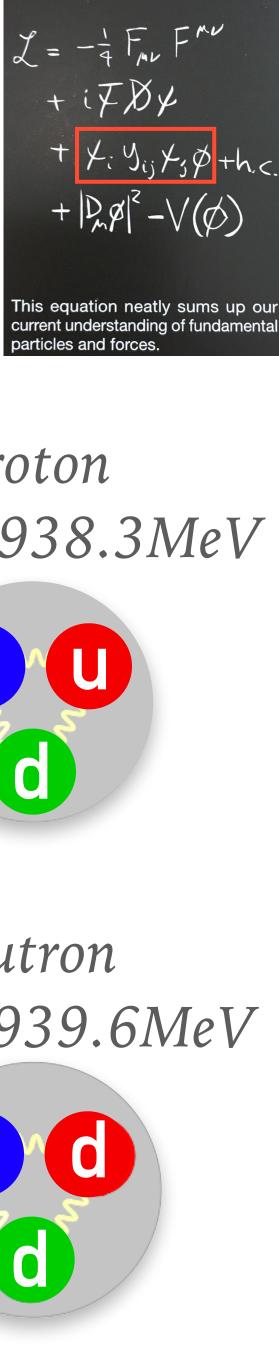
Why do Yukawa couplings matter? (1) Because, within SM conjecture, they're what give masses to all quarks

Up quarks (mass ~ 2.2 MeV) are lighter than down quarks (mass ~ 4.7 MeV)

proton **neutron** (up+down+down): 2.2 + **4.7** + 4.7 + ... = **939.6** MeV

> So protons are **lighter** than neutrons, \rightarrow protons are stable.

Which gives us the hydrogen atom, & chemistry and biology as we know it



(up+up+down): 2.2 + 2.2 + 4.7 + ... = 938.3 MeV

proton mass = 938.3 MeV

neutron mass = 939.6 MeVC

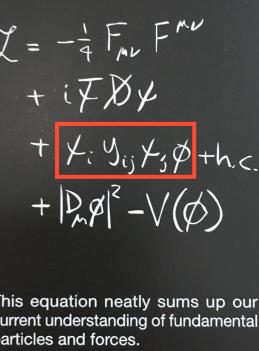
Why do Yukawa couplings matter? (2) Because, within SM conjecture, they're what give masses to all leptons

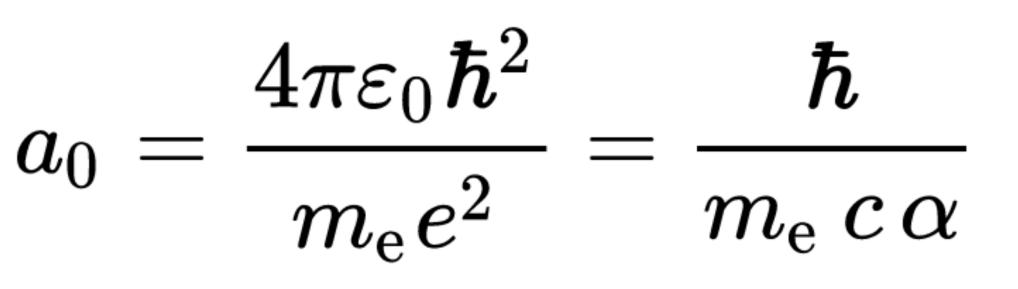
Bohr radius

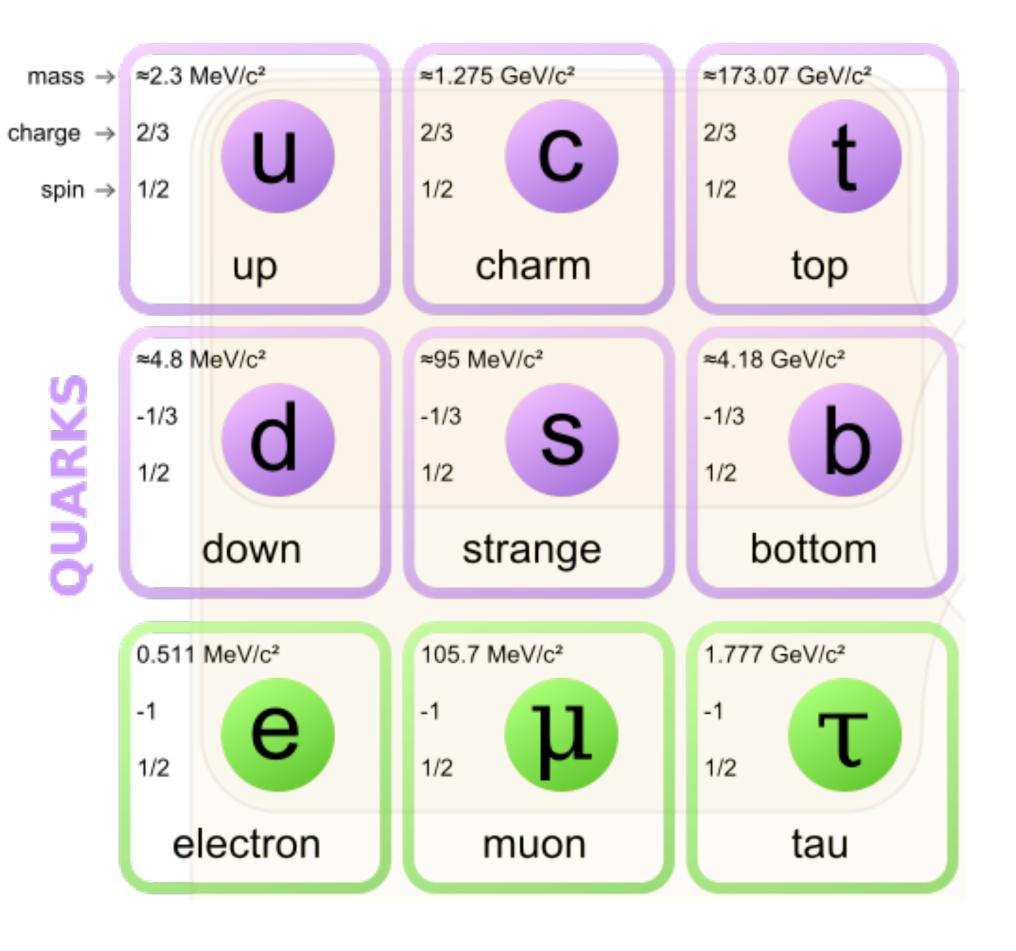
electron mass determines size of all atoms

it sets energy levels of all chemical reactions

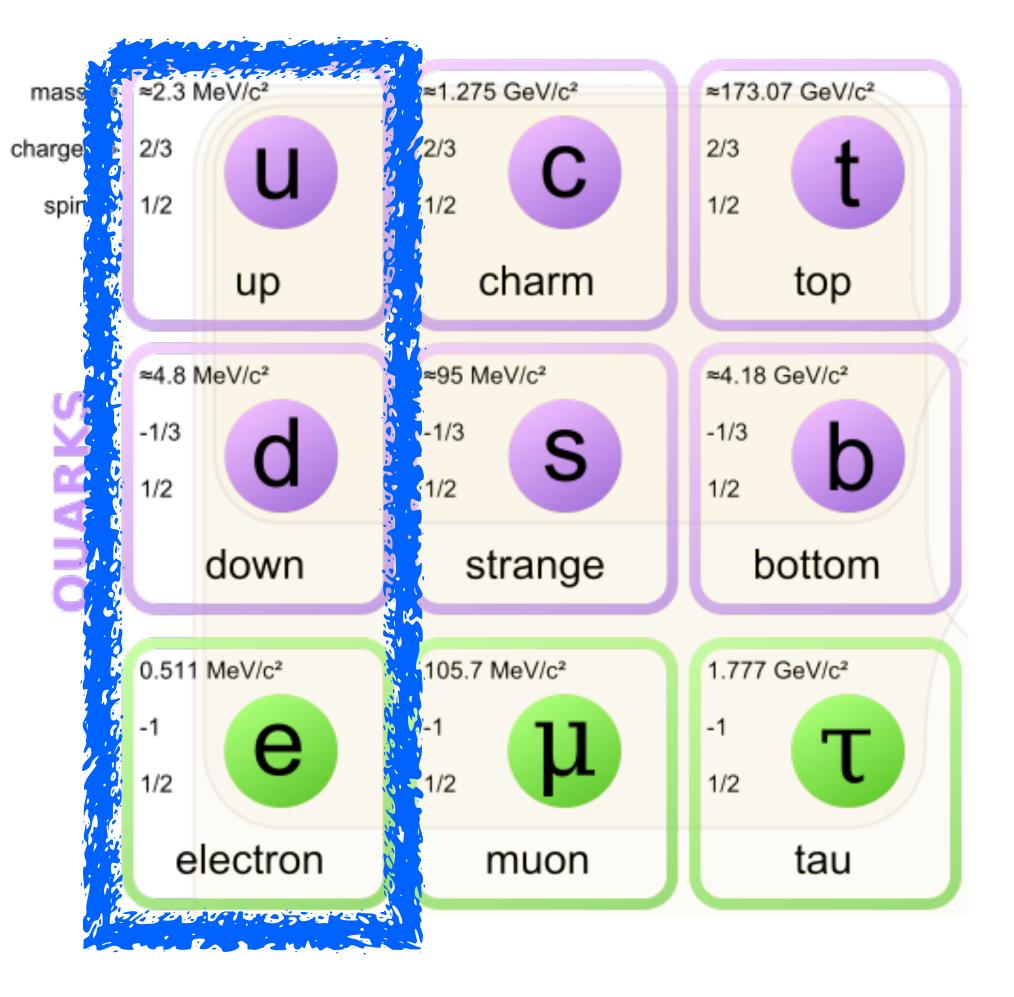
Gavin Salam





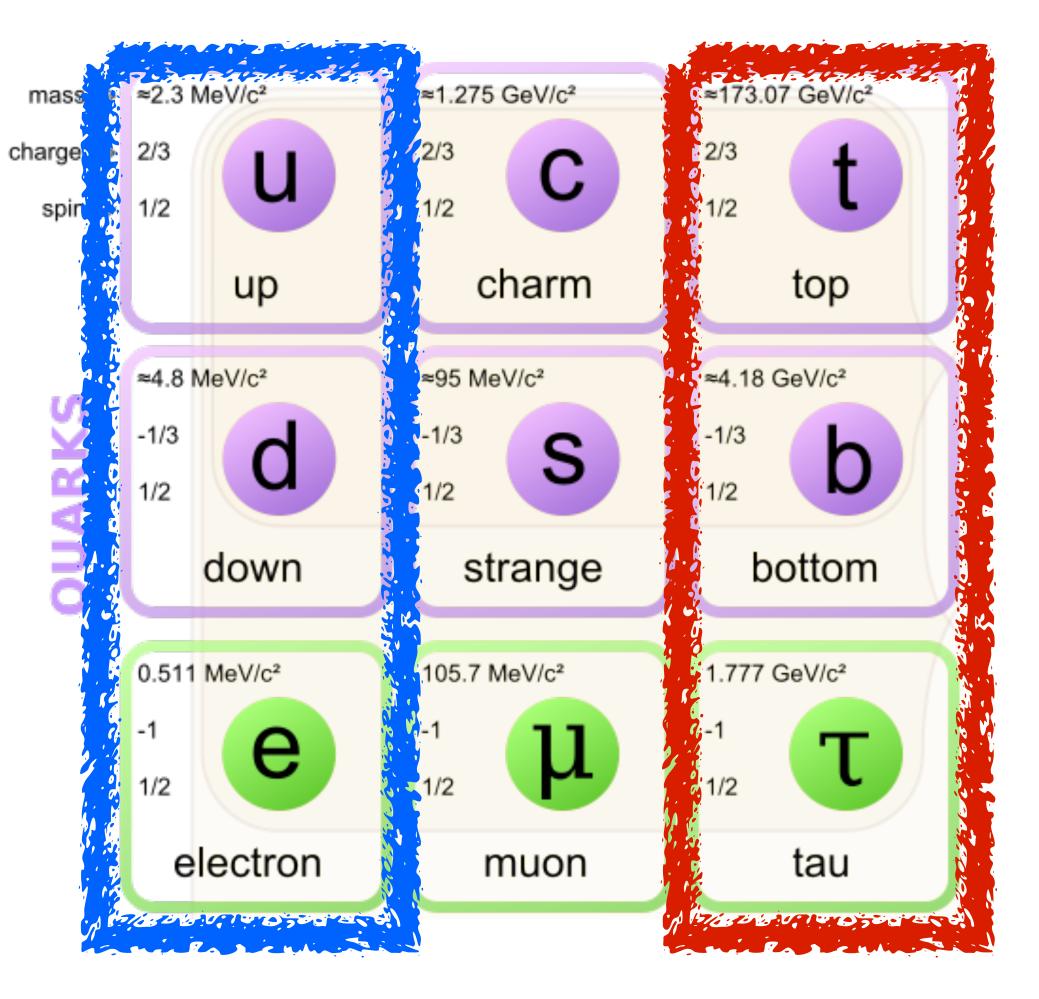






1st generation (us) has low mass because of weak interactions with Higgs field (and so with Higgs bosons): too weak to test today





1st generation (us) has low mass because of weak interactions with Higgs field (and so with Higgs bosons): too weak to test today 3rd generation (us) has high mass because of strong interactions with Higgs field (and so with Higgs bosons): can potentially be tested



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ALICE



LHC 7 TeV + 7 TeV 27 km

 \sim





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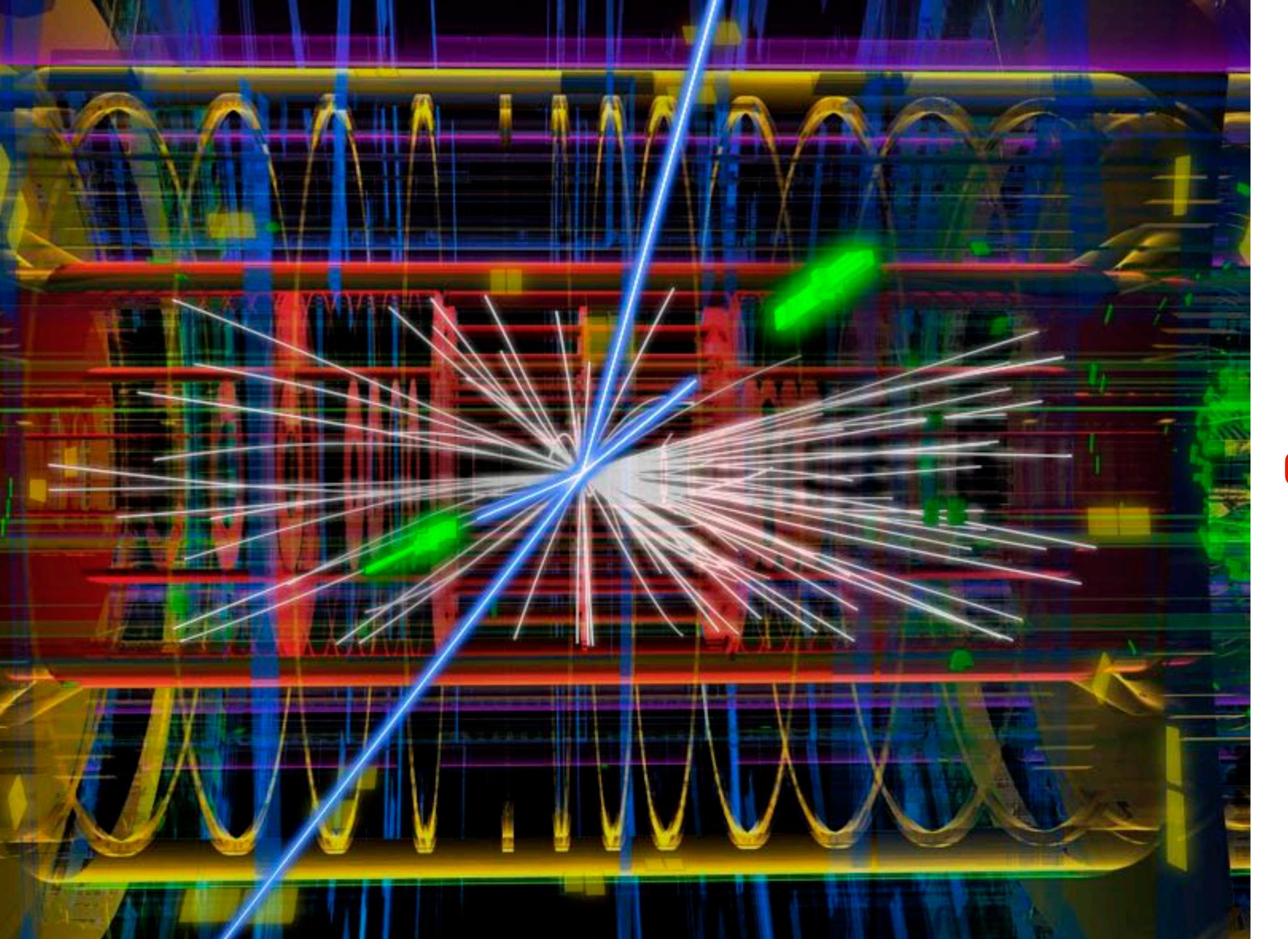


LHC 7 TeV + 7 TeV 27 km

 \sim







ATLAS & CMS **@LHC**

~up to 2 billion collisions/second

(+ lower rates at LHCb and ALICE)



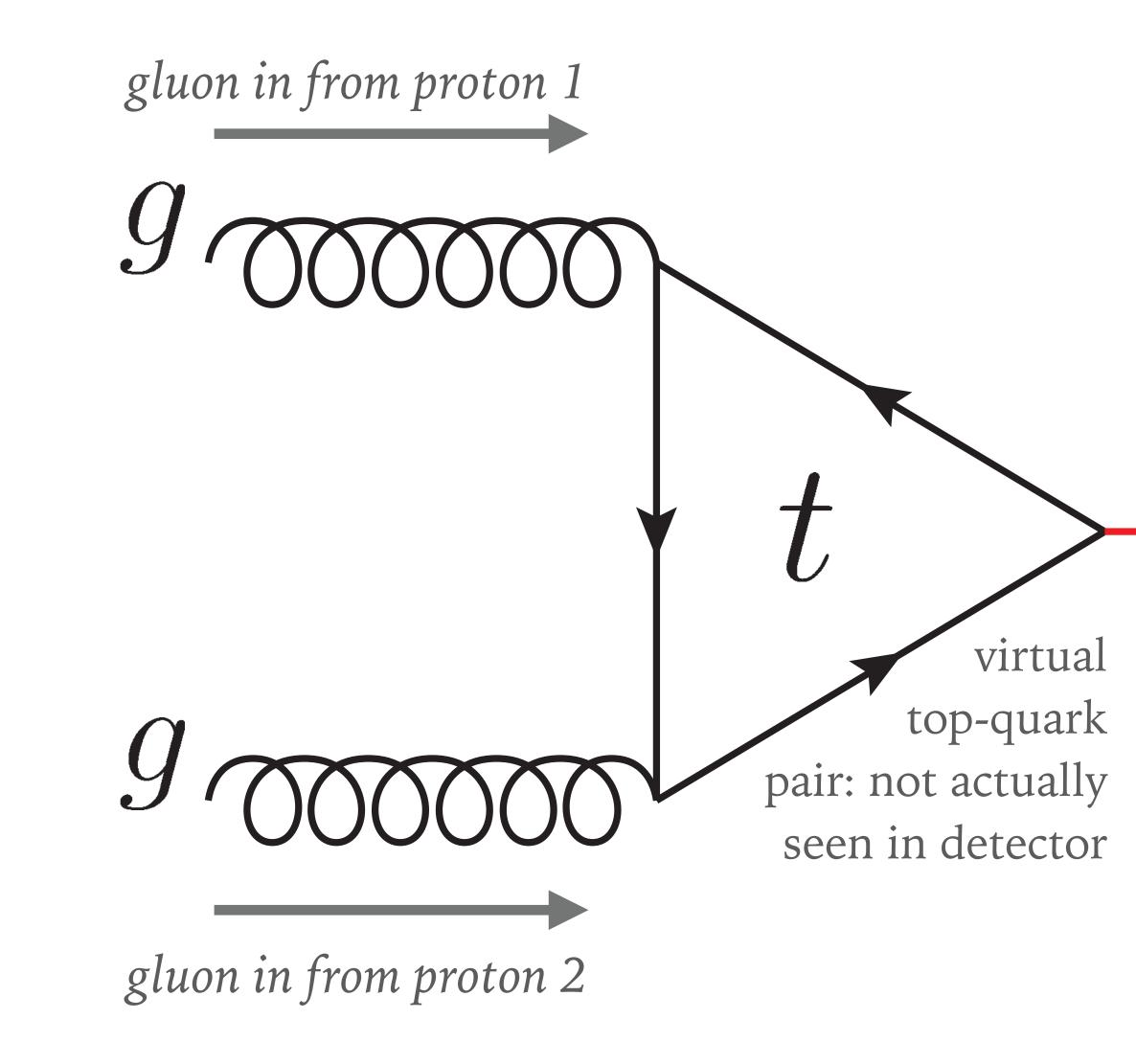


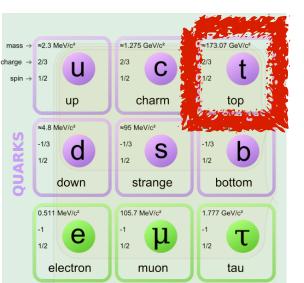




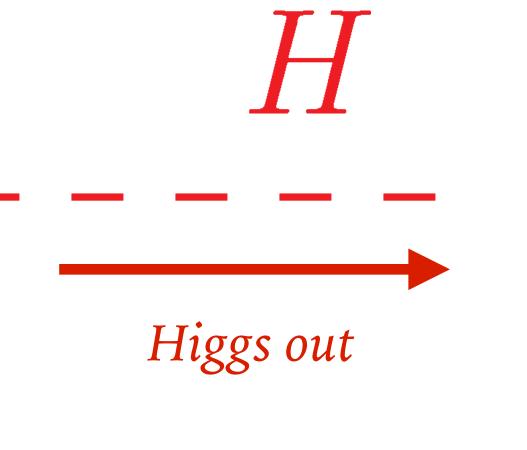
what underlying processes tell us about Yukawa interactions?







Higgs production: the dominant channel

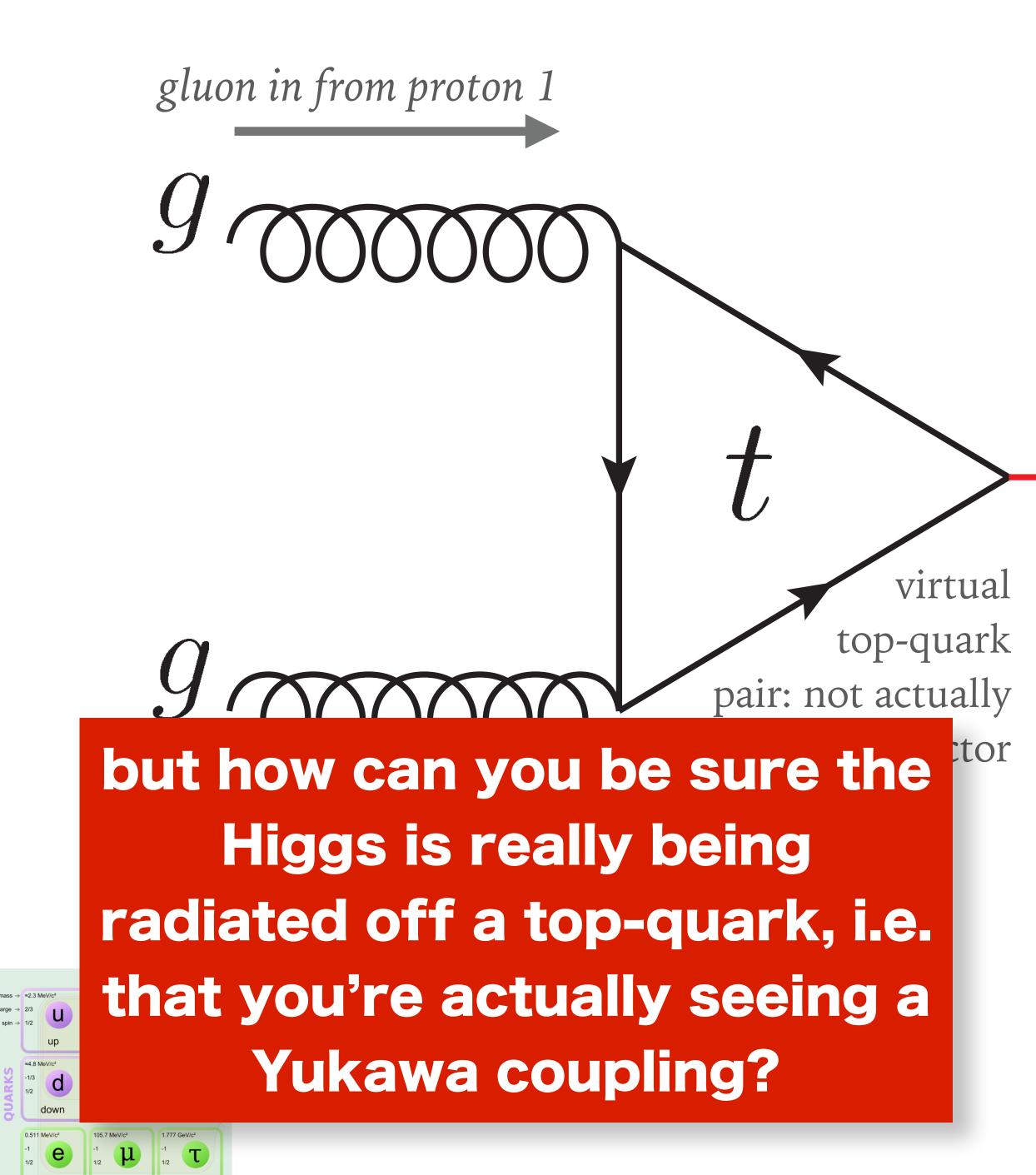


Expected to happen once for every ~2 billion inelastic proton-proton collisions

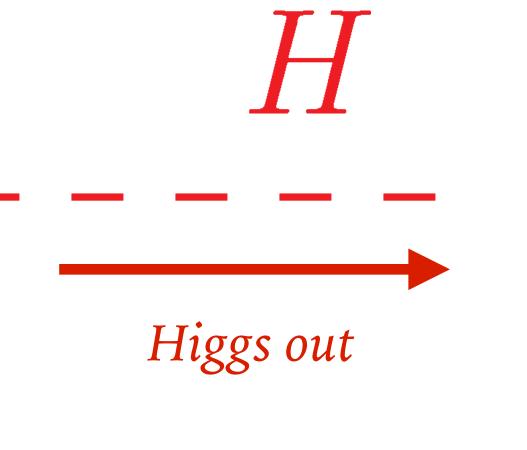
LHC data consistent with that already at discovery in 2012



24



Higgs production: the dominant channel

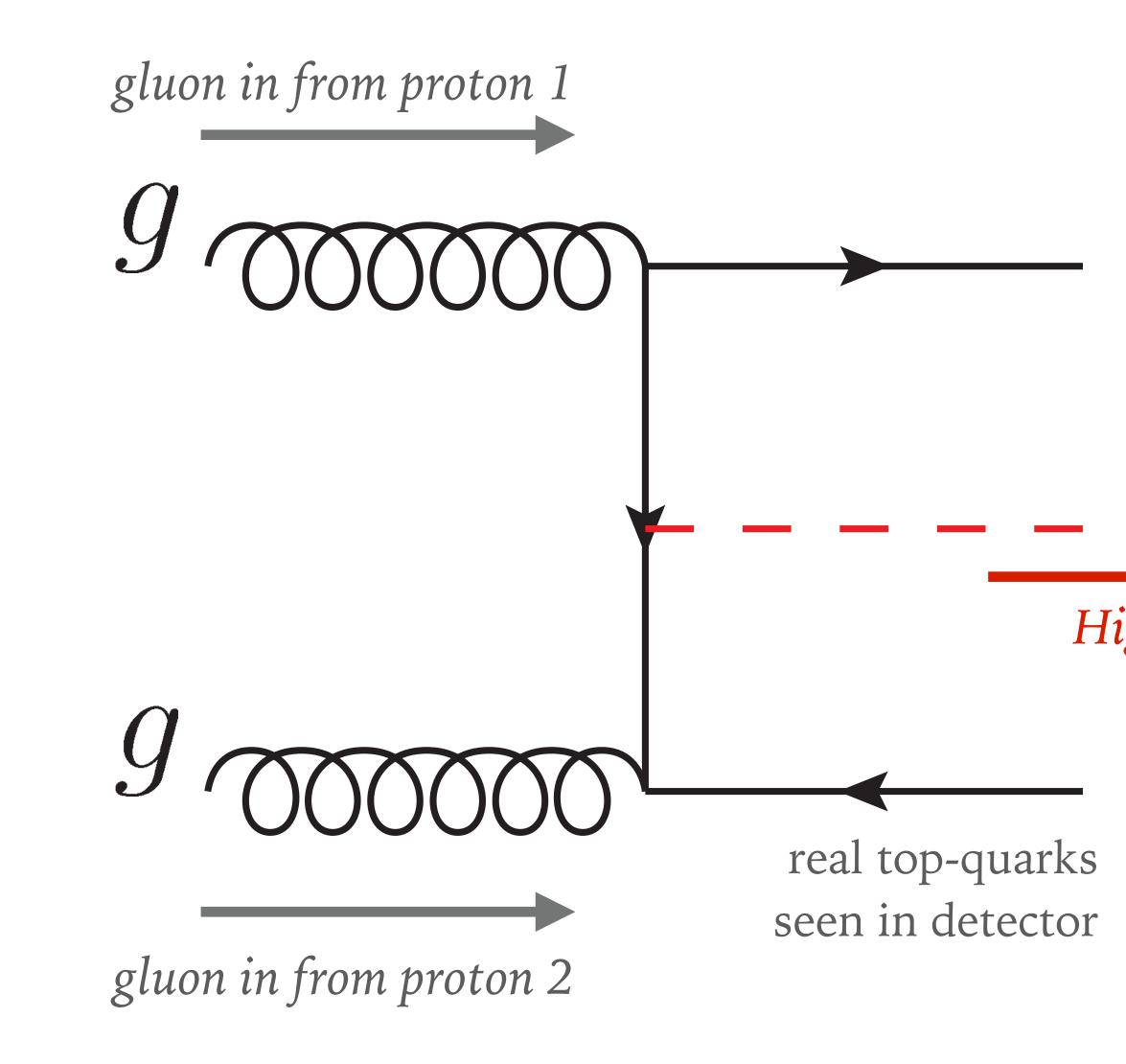


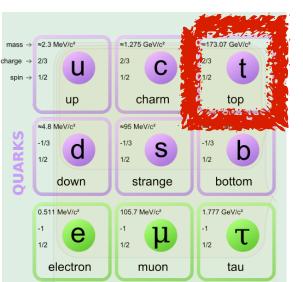
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24



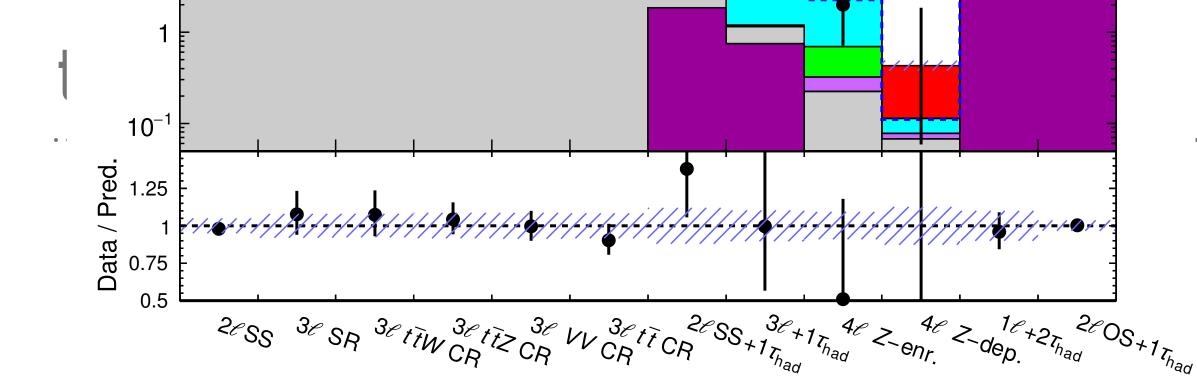


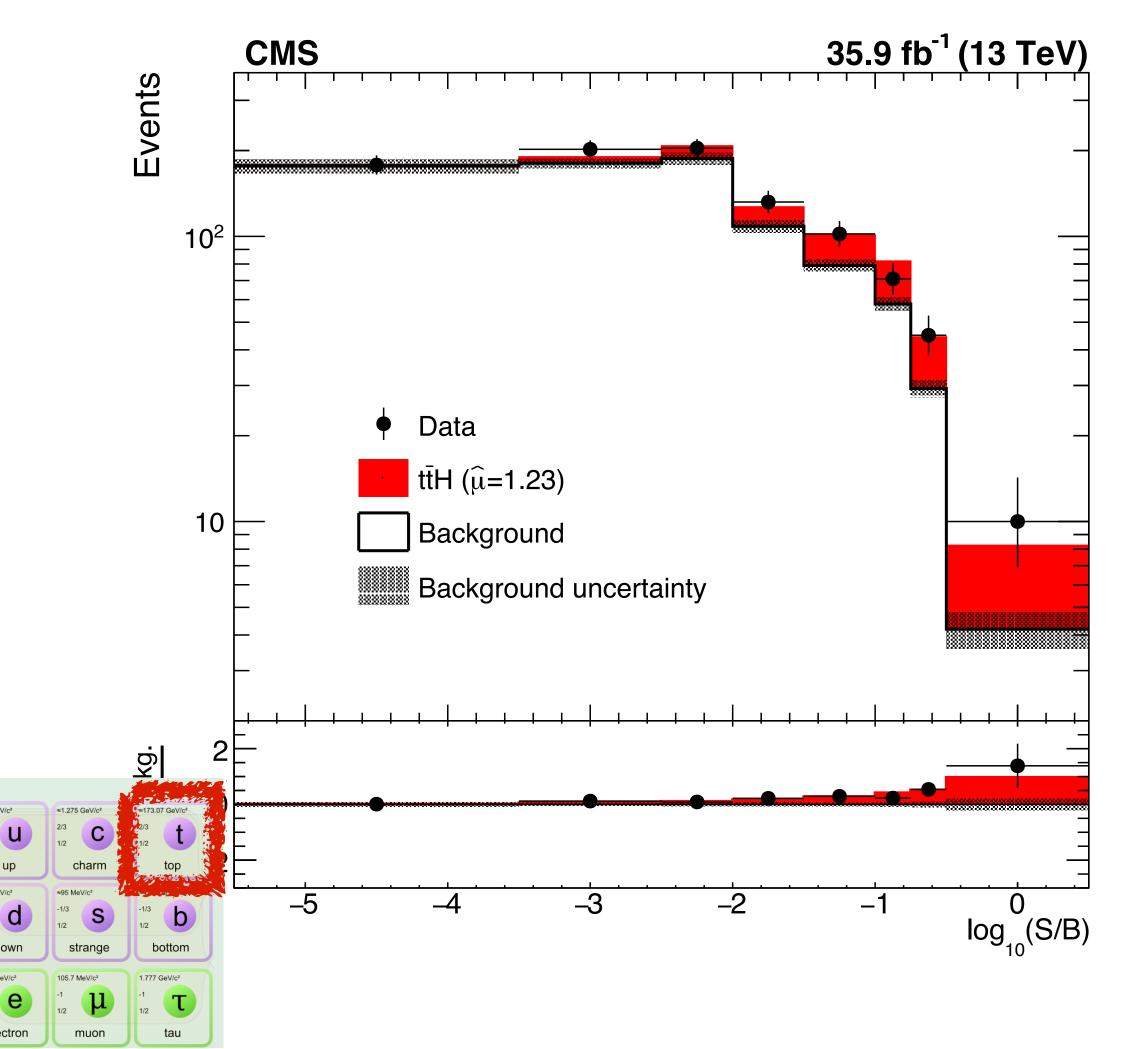
Higgs production: the ttH channel Higgs out If SM top-Yukawa hypothesis is correct, expect 1 Higgs for every 1600 top-quark pairs.

(rather than 1 Higgs for every 2 billion pp collisions)



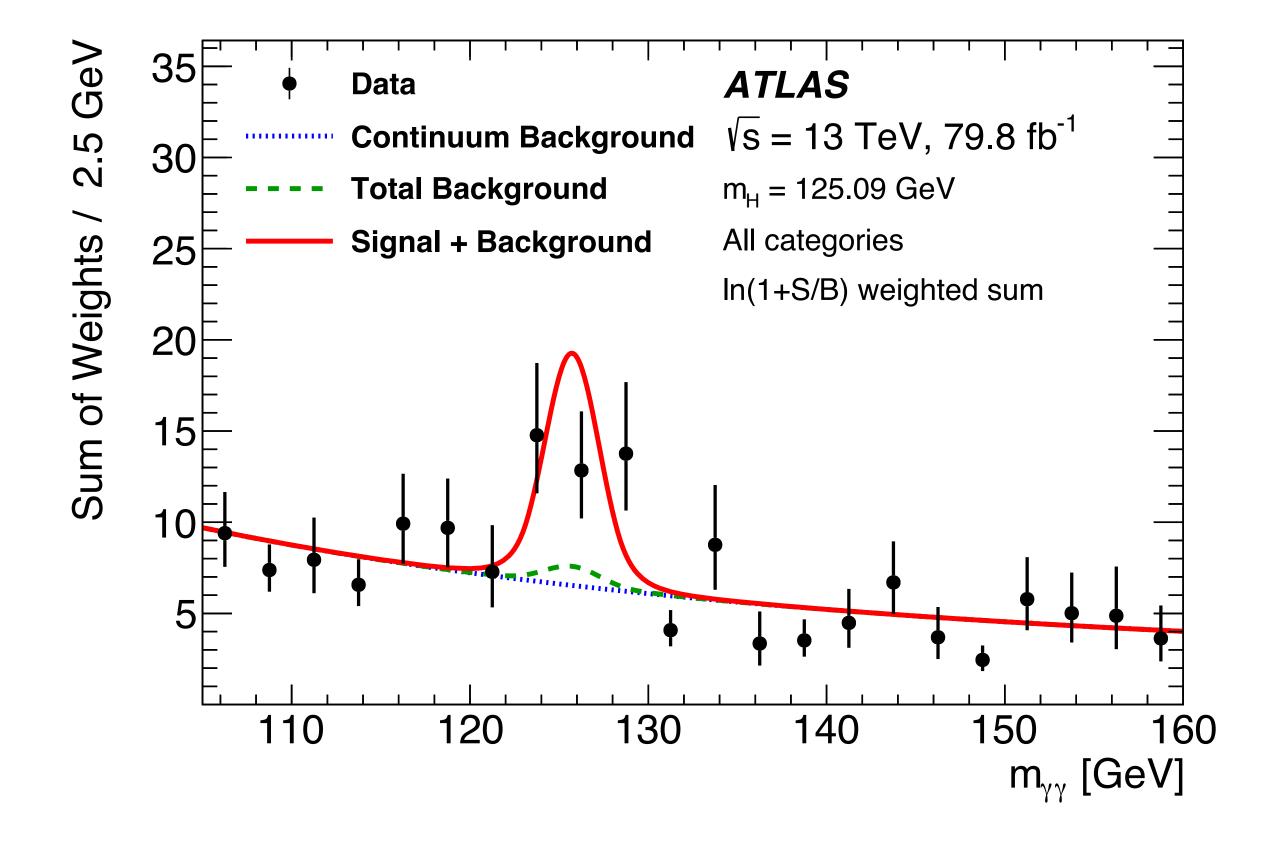




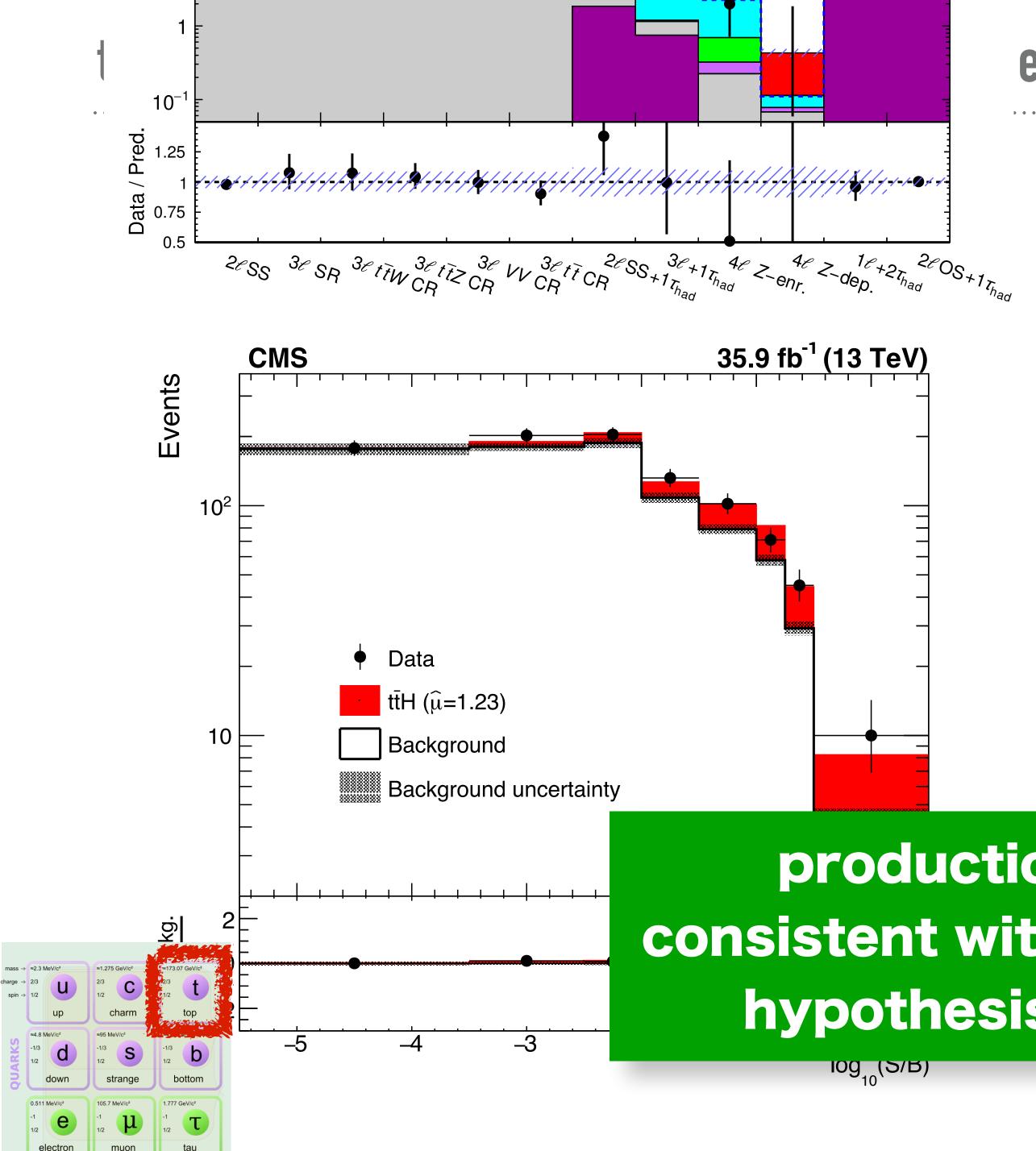


events with top-quarks & Higgs simultaneously

ATLAS > 5-sigma ttH

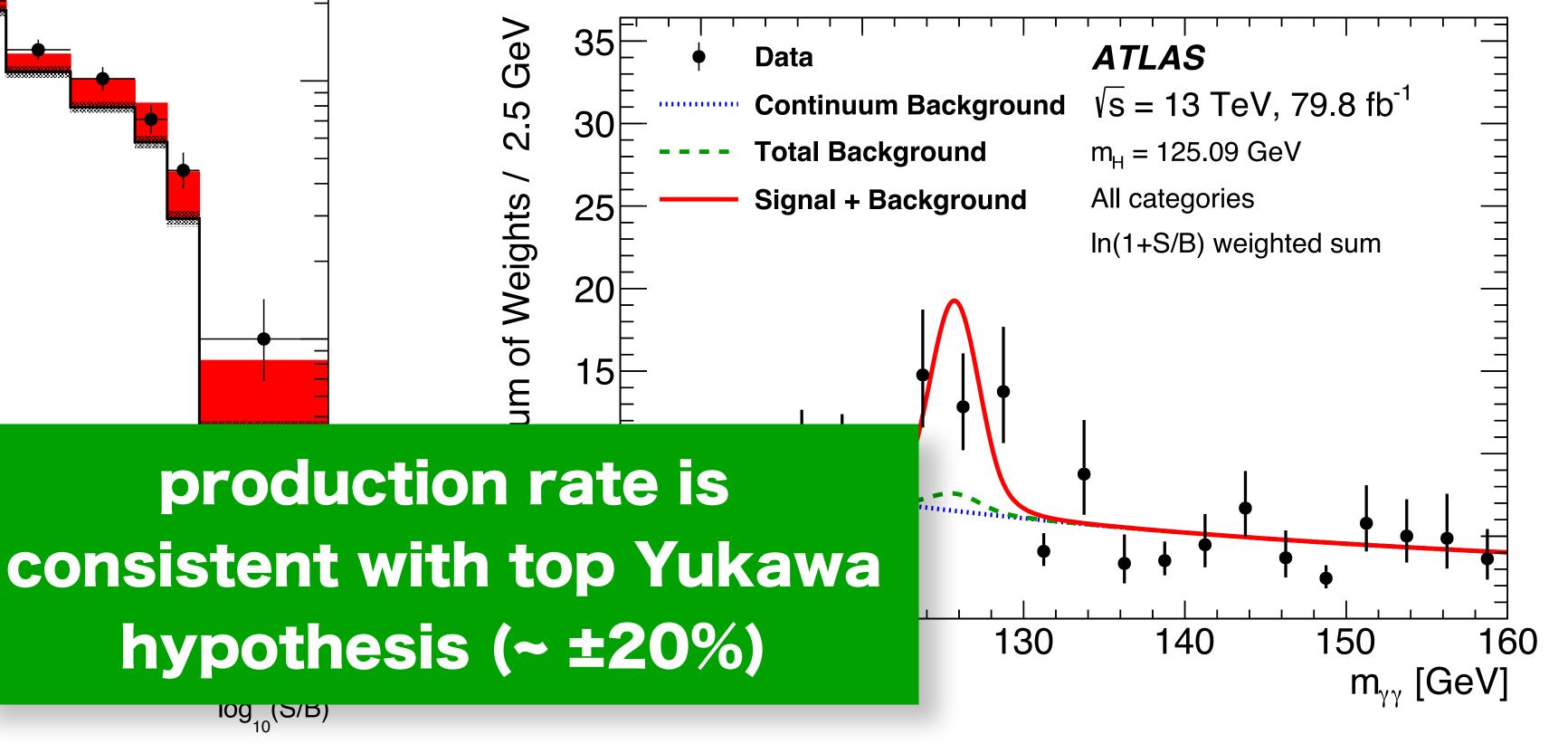




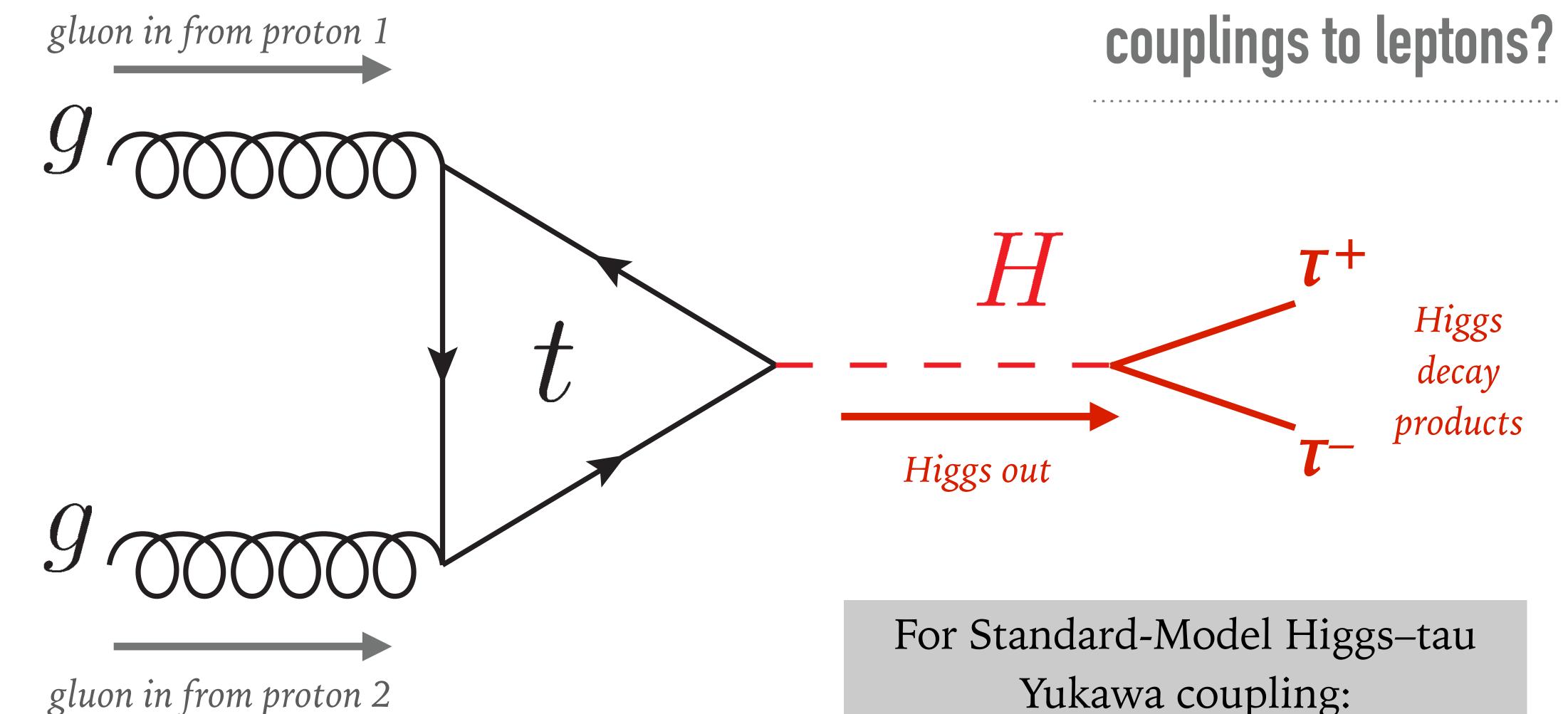


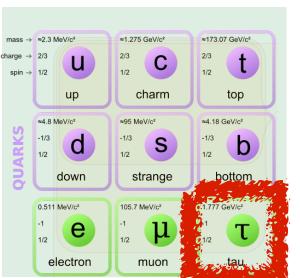
events with top-quarks & Higgs simultaneously

ATLAS > 5-sigma ttH









Yukawa coupling:

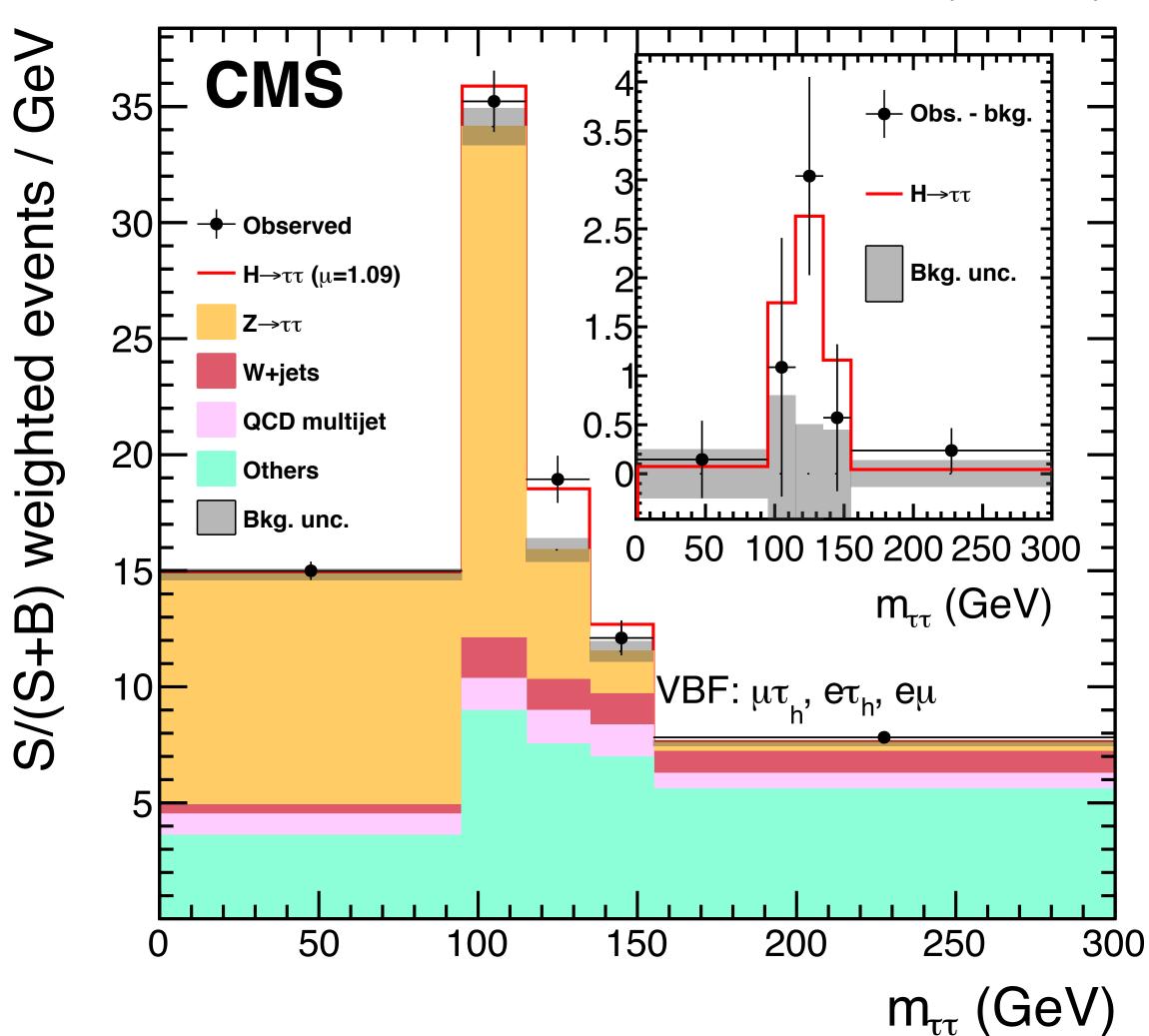
1 in every 16 Higgs bosons \sim decays to $\tau^+\tau^-$

27

observation of $H \rightarrow \tau \tau$

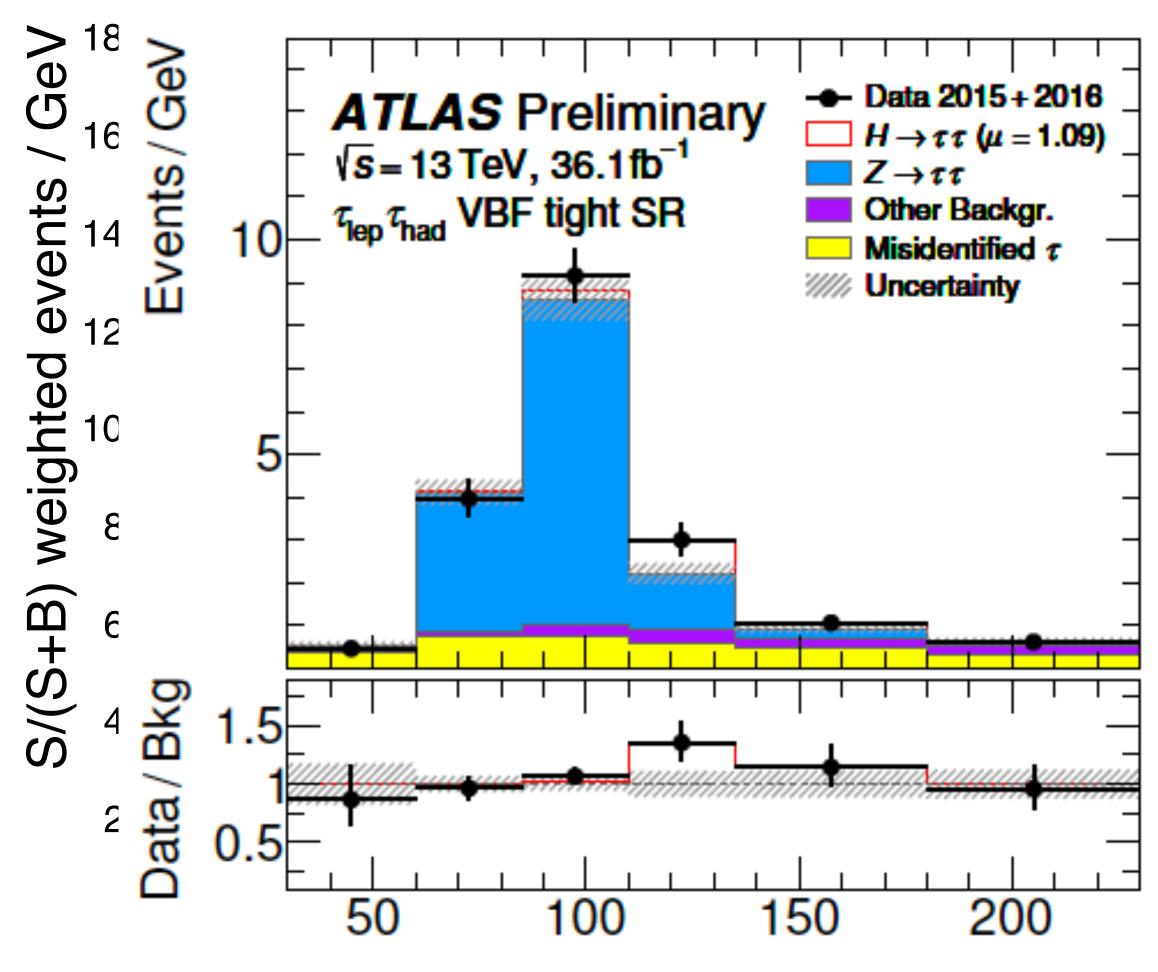
18 months ago: CMS >5-sigma H $\rightarrow \tau \tau$

35.9 fb⁻¹ (13 TeV)



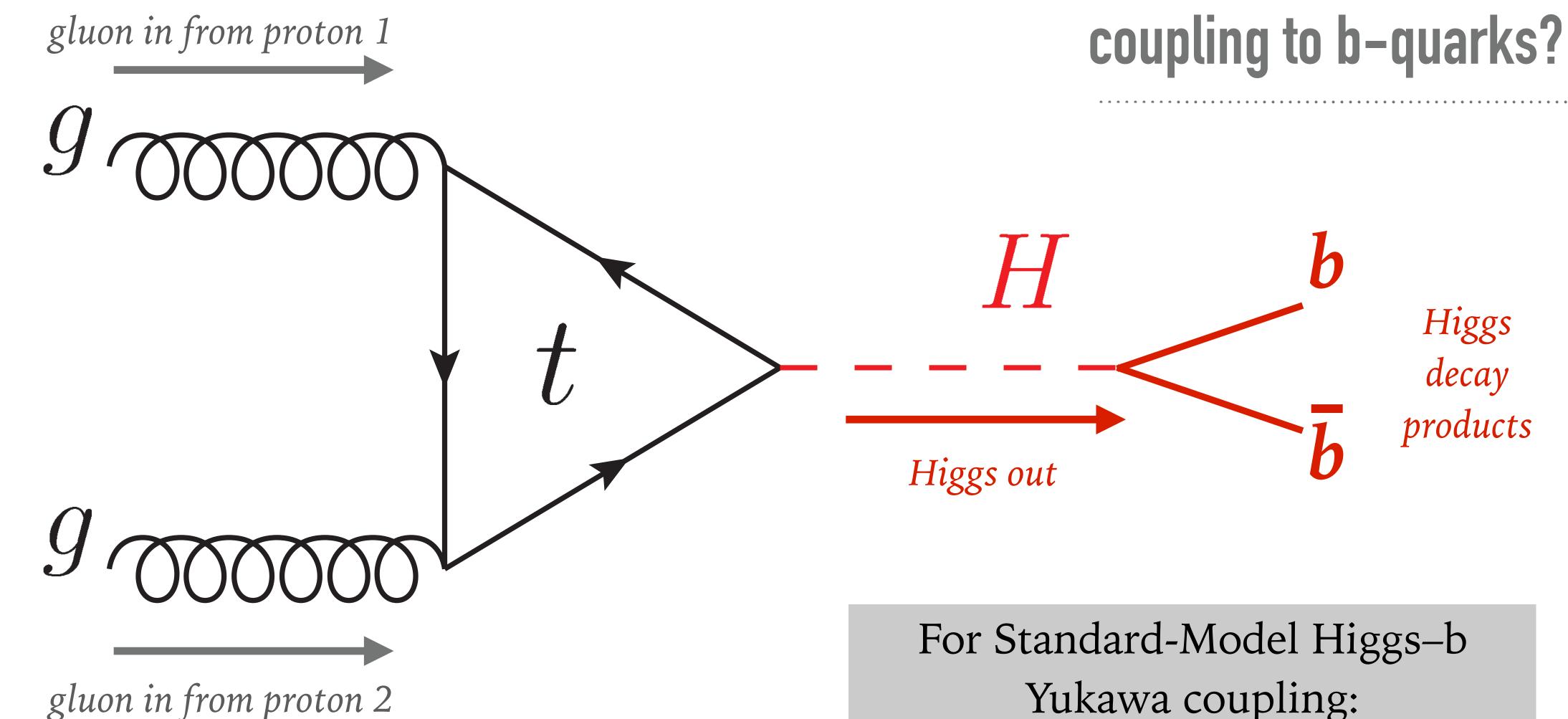
6 months ago: ATLAS >5-sigma H $\rightarrow \tau \tau$

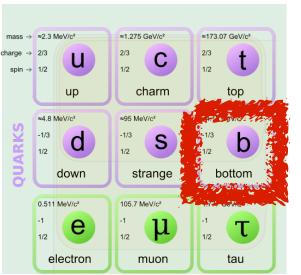
35.9 fb⁻¹ (13 TeV)



 $m_{\tau\tau}^{MMC}$ [GeV]







Yukawa coupling:

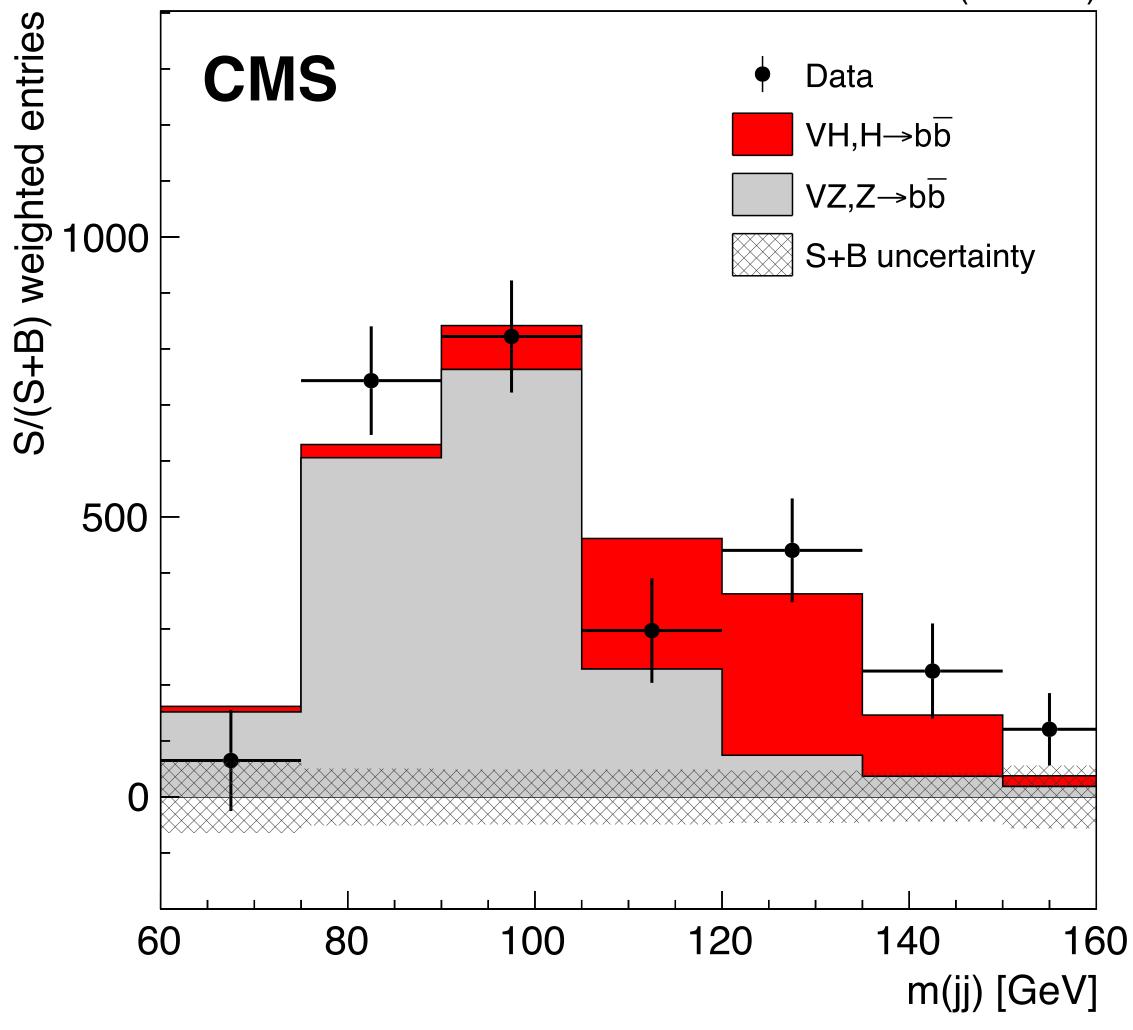
~ 58% of Higgs bosons should decay to bb

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three months ago, observation of ${\rm H} \rightarrow {\rm bb}$

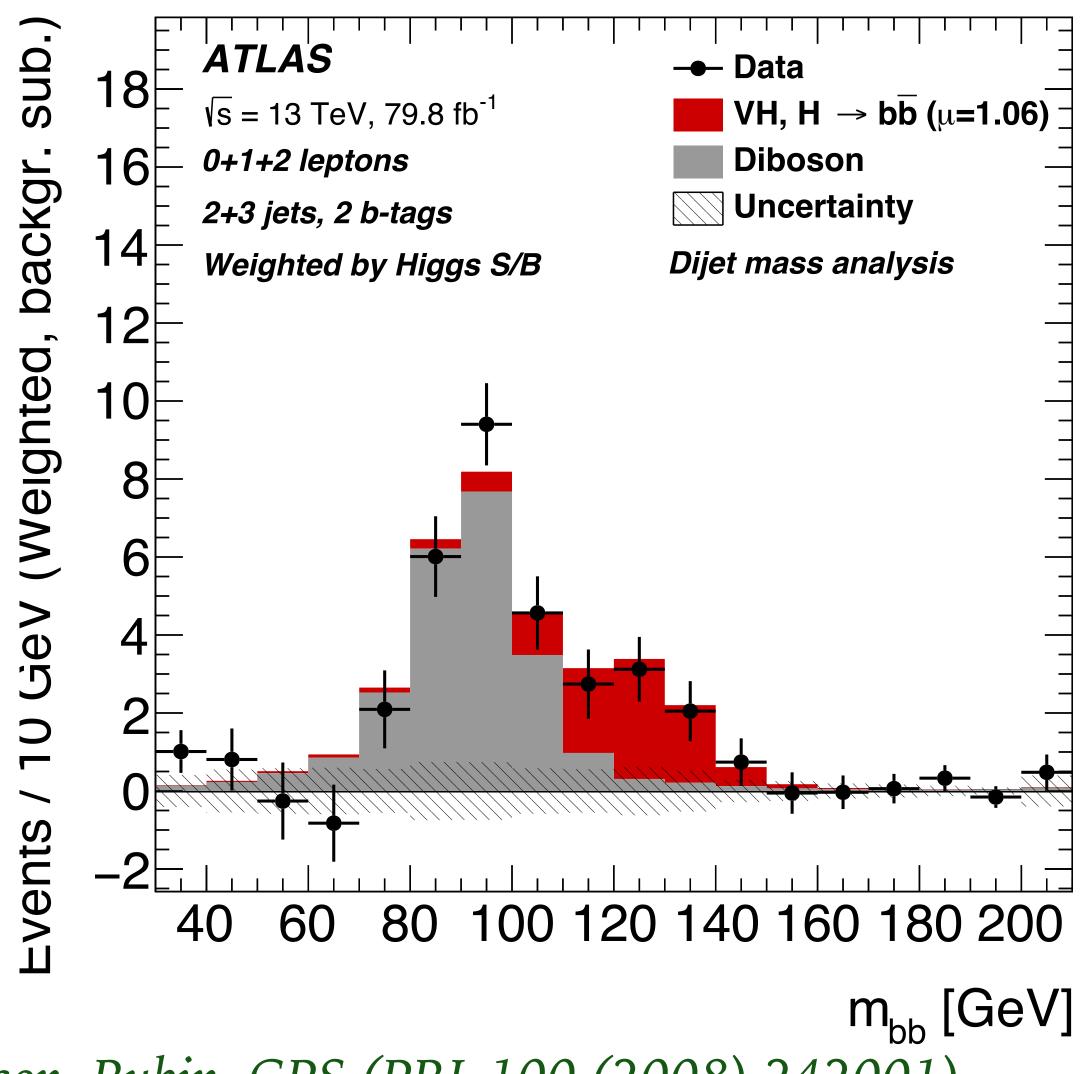
$\textbf{CMS >5-sigma H} \rightarrow \textbf{bb}$

77.2 fb⁻¹ (13 TeV)



Analysis includes key idea from Butterworth, Davison, Rubin, GPS (PRL 100 (2008) 242001)

ATLAS > 5-sigma H \rightarrow bb





The $>5\sigma$ observations of the ttH process and of H $\rightarrow \tau\tau$ and H \rightarrow bb decays, independently by ATLAS and CMS, firmly establish the existence of a new kind of fundamental interaction, Yukawa interactions.

Yukawa interactions are important not merely because they had never before been directly observed, but also because they are hypothesized to be responsible for the stability of hydrogen, and for determining the size of atoms and the energy scales of chemical reactions.

Establishing the pattern of Yukawa couplings across the full remaining set of quarks and charged leptons is one of the major challenges for particle physics today.





what could one be saying about it?

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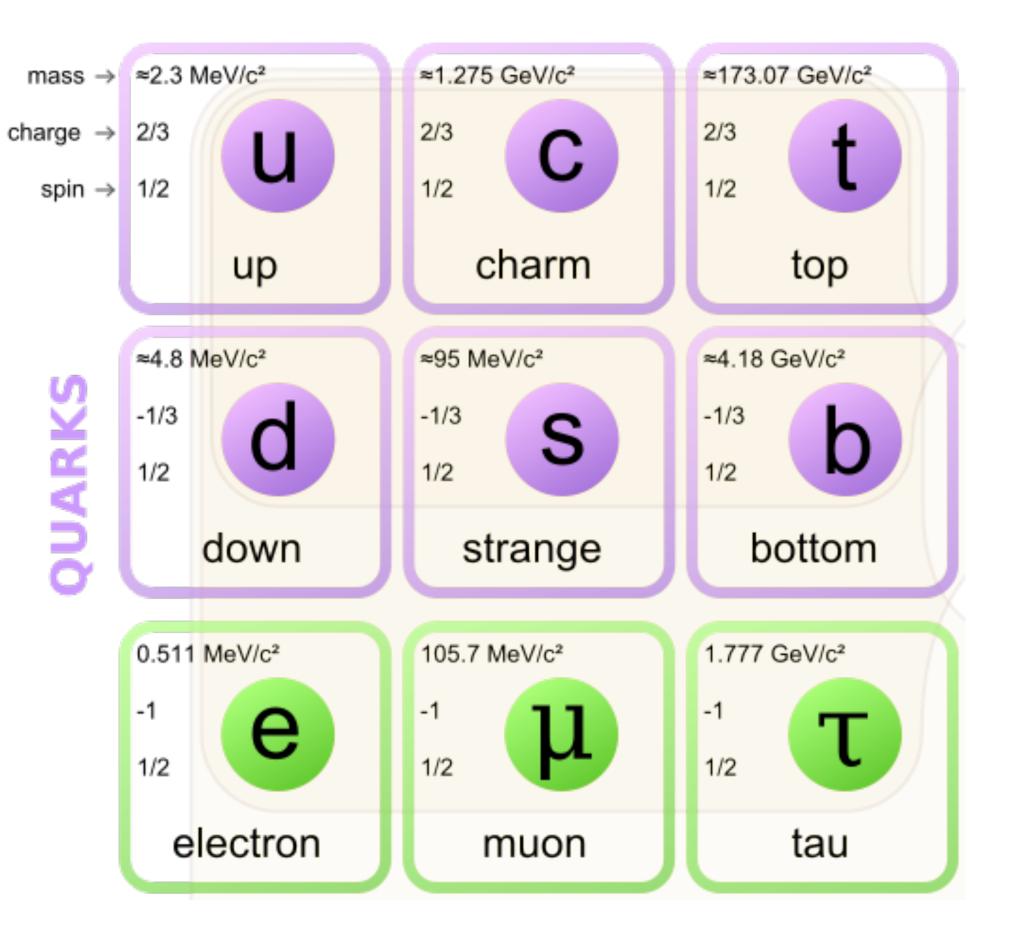
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Is this any less important than the discovery of the Higgs boson itself? My opinion: no, because fundamental interactions are as important as fundamental particles



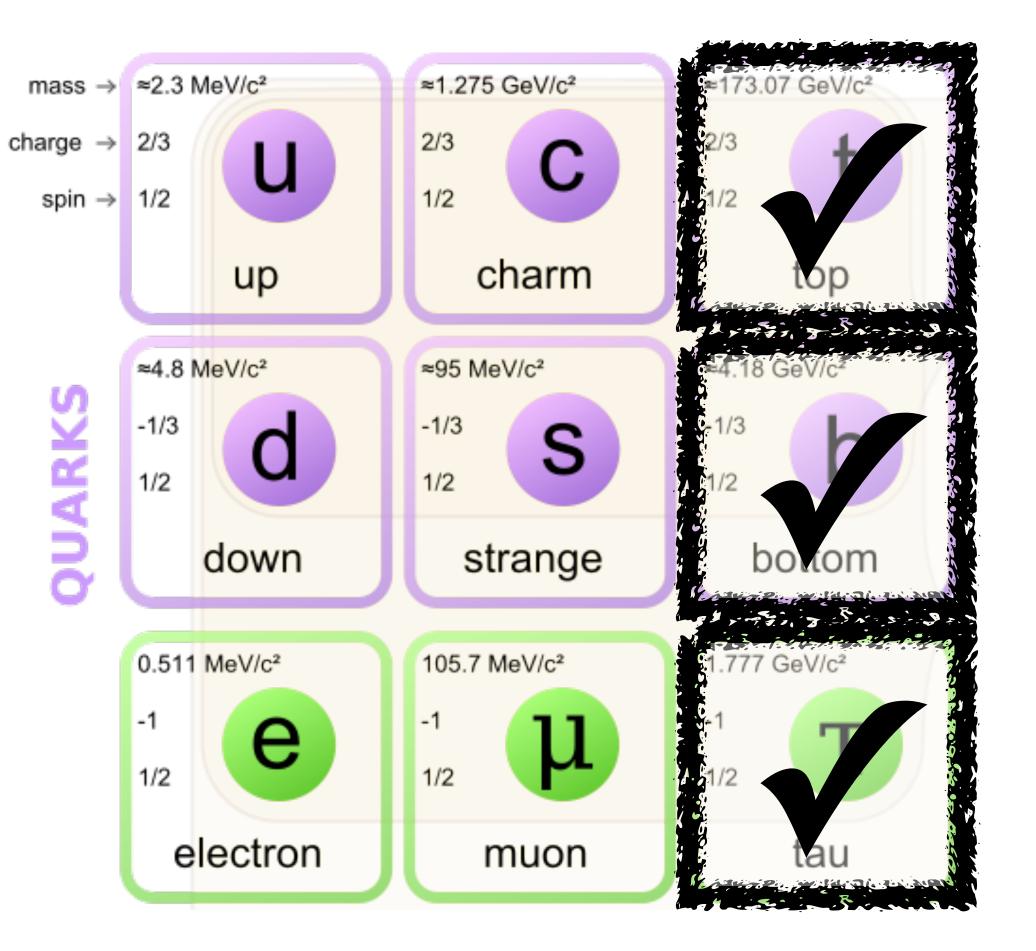


Yukawas



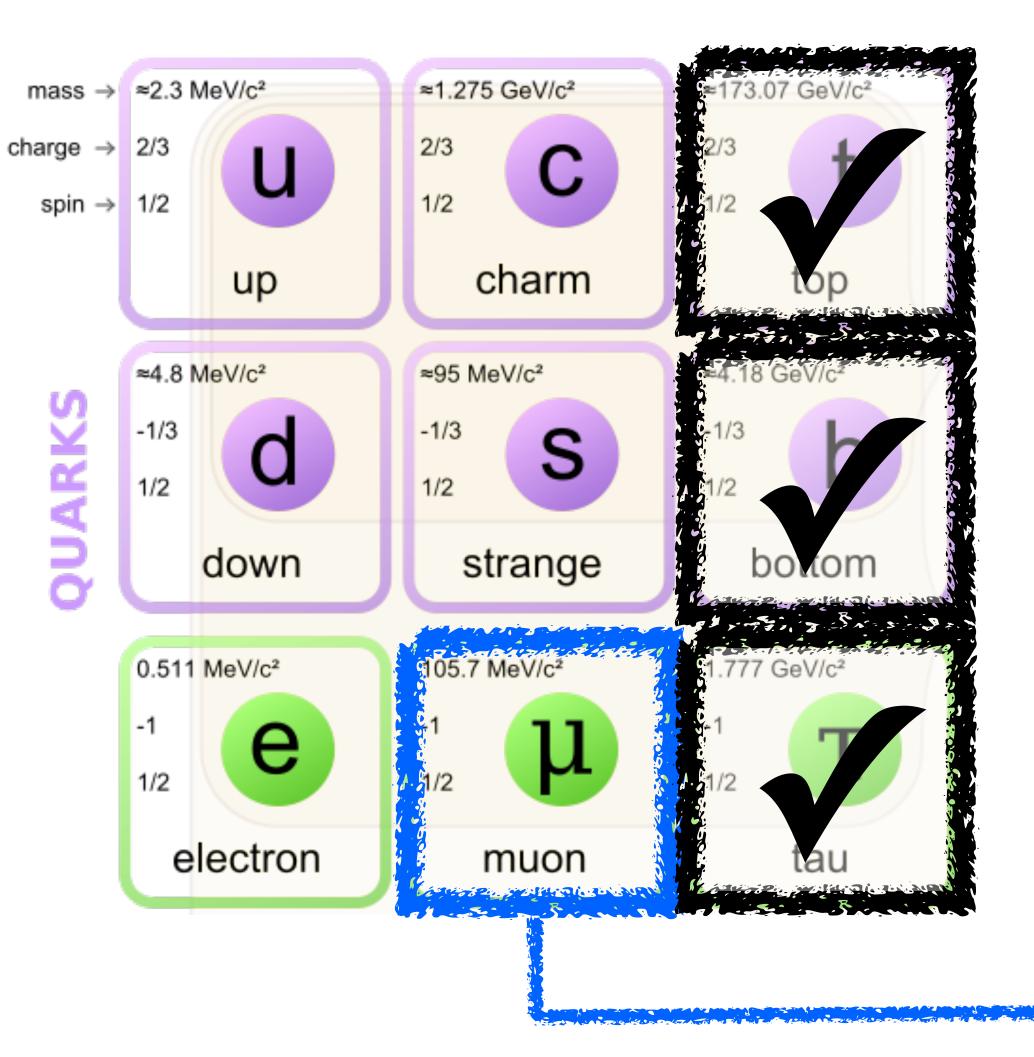


Yukawas





Yukawas



today: no evidence yet (1 in 4570 decays)

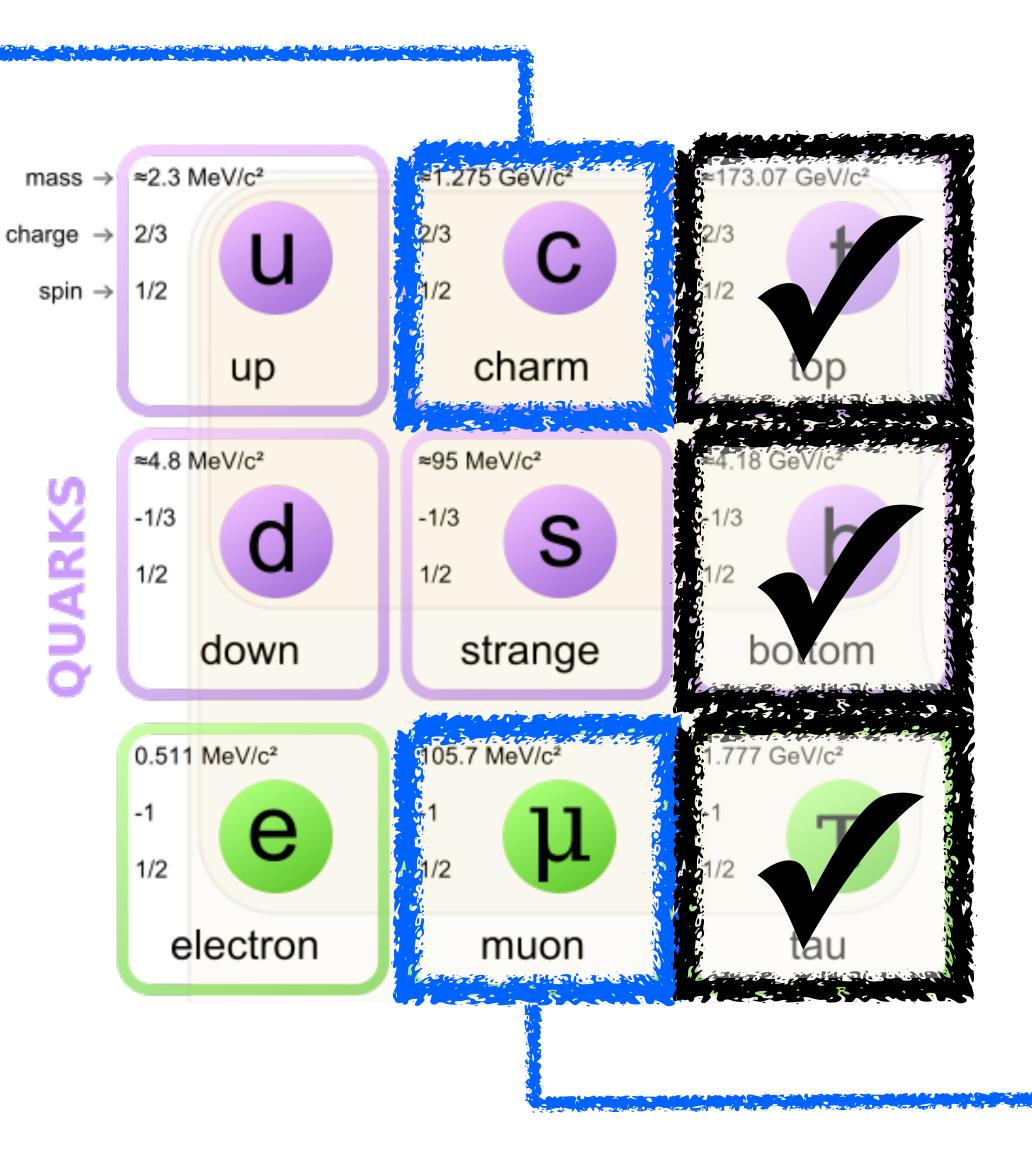
observable at the LHC within about 10 years.







today: no evidence yet (1 in 35 decays) needs a lepton collider



Yukawas

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≈2.3 MeV/c²

≈4.8 MeV/c²

0.511 MeV/c²

1/3

up

u

е

electron

down

mass -

spin

CĆ

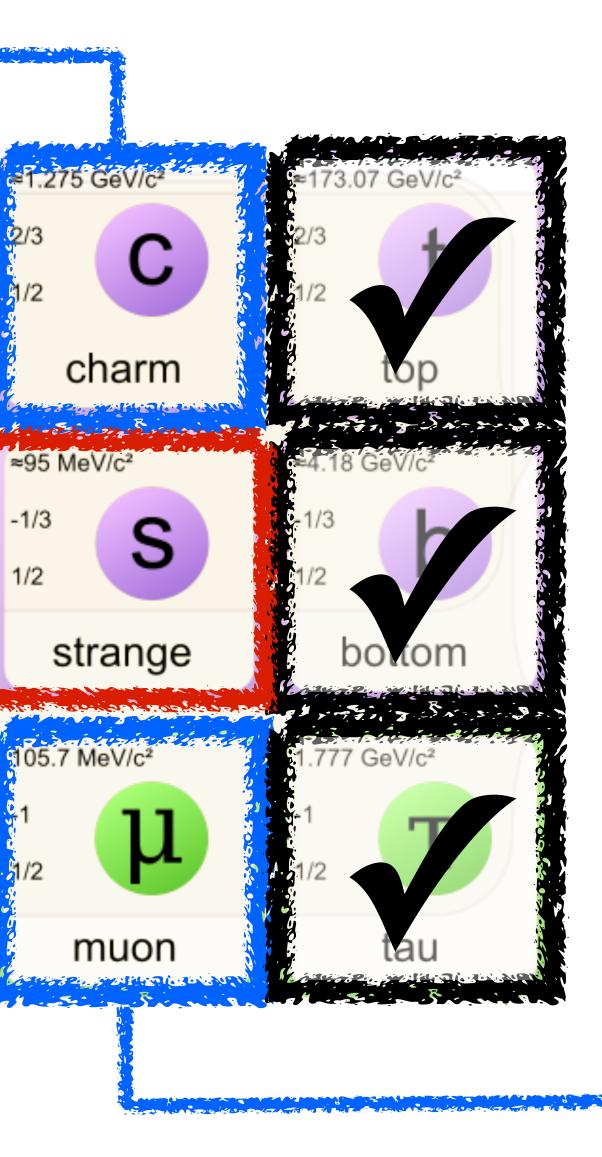
charge

today: no evidence yet (1 in 35 decays) needs a lepton collider

today: no evidence yet (1 in 4000 decays) no clear route to

establishing SM couplings at 5σ

Yukawas



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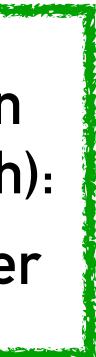
Yukawas



overall normalisation (related to Higgs width): needs a lepton collider

today: no evidence yet (1 in 4570 decays)

observable at the LHC within about 10 years.







2nd & 1st generation Yukawas

- the hierarchy of masses between generations remains a mystery (even if it's one that some people consign to the "hopeless" category)
- > Does not necessarily come from hierarchy of dimensionless Yukawa coefficients
- ► E.g. the Giudice-Lebedev mechanism (and follow-up work)

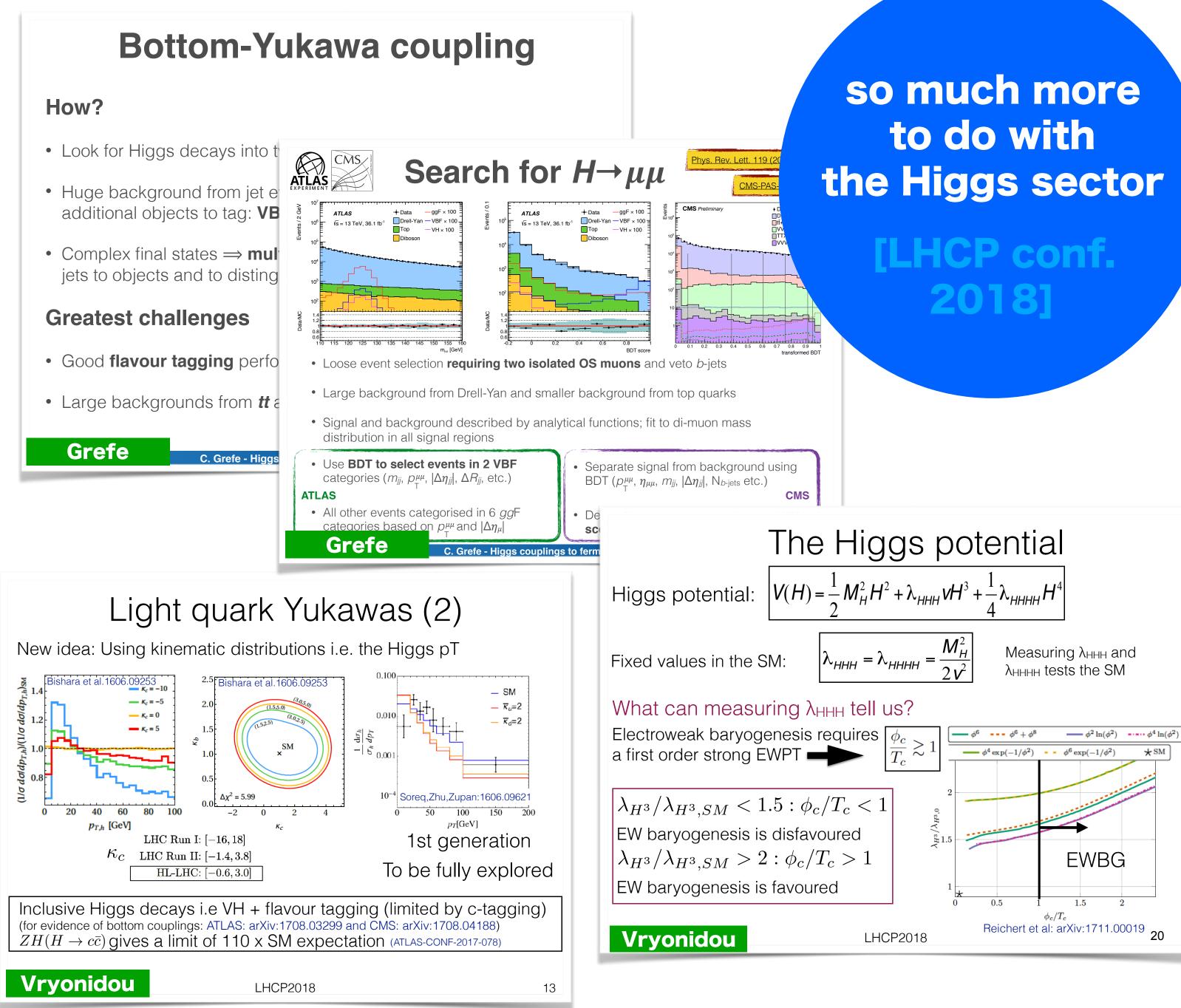
$$-\mathcal{L}_Y = Y_{ij}(\phi)\bar{\psi}_i\psi_j\phi + \text{h.c.}$$

- \blacktriangleright smallness of certain masses is consequence of vev²/M² suppression, not small c_{ij} \blacktriangleright measured Hqq interaction larger by factor (2n_{ii} + 1)
- ► cf. also various more recent discussions, e.g. by Bauer, Carena, Carmona

$$Y_{ij}(\phi) = c_{ij} \left(\frac{\phi^{\dagger}\phi}{M^2}\right)^{n_{ij}}$$

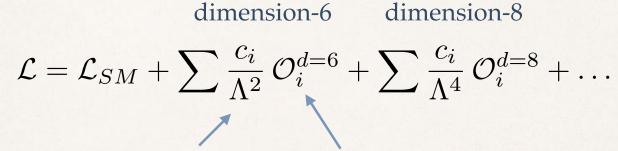
1801.00363

33

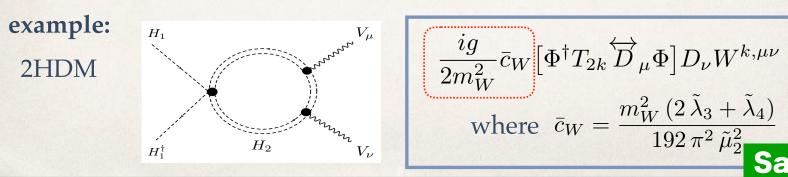


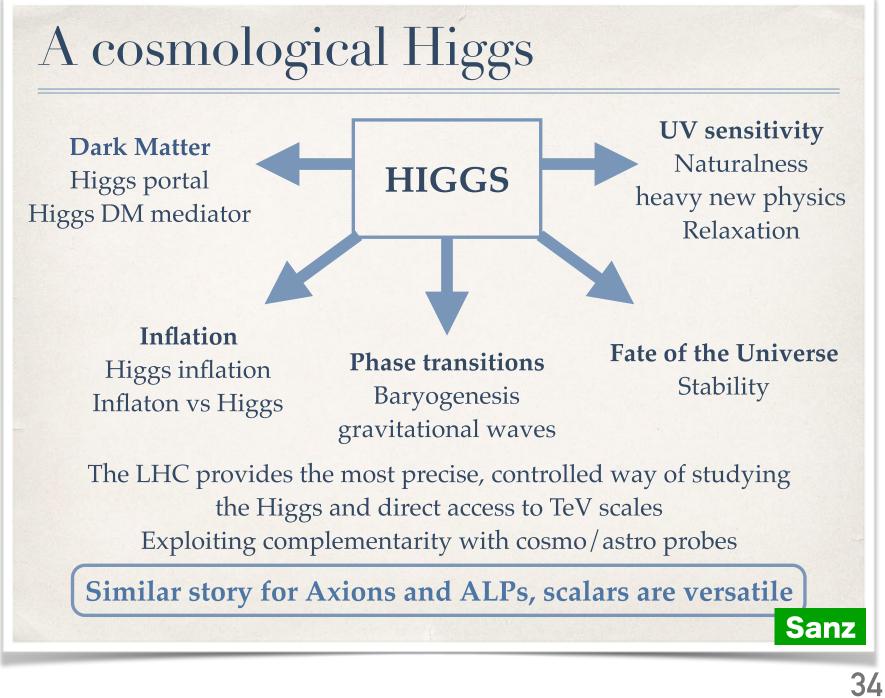
EFT approach

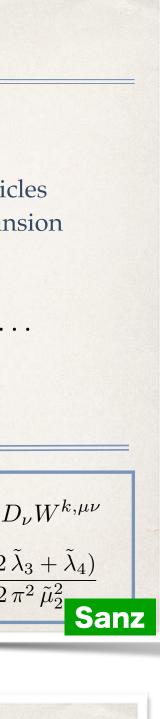
Well-defined theoretical approach Assumes New Physics states are heavy Write Effective Lagrangian with only light (SM) particles BSM effects can be incorporated as a momentum expansion



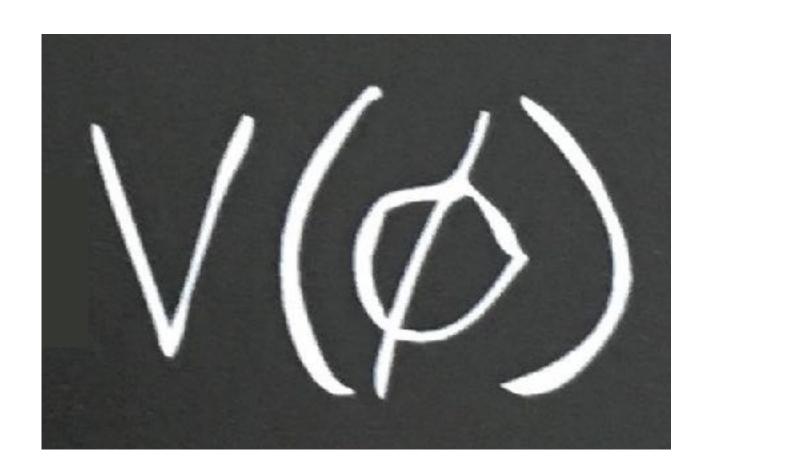
BSM effects SM particles

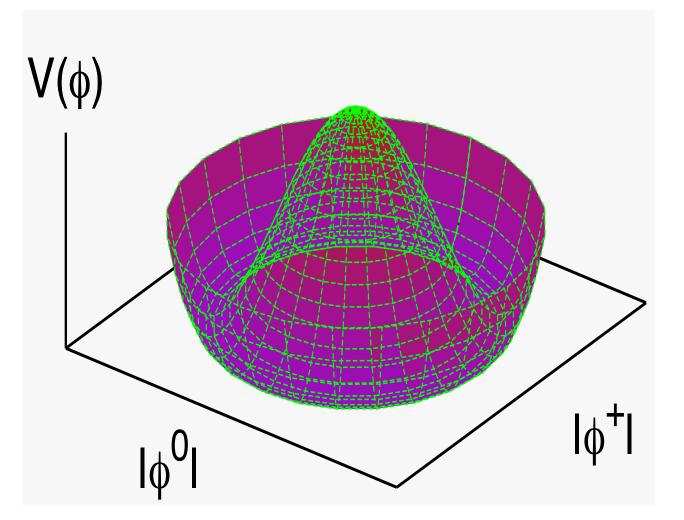






$V(\Phi) = m^2 \Phi^{\dagger} \Phi + \lambda (\Phi^{\dagger} \Phi)^2$





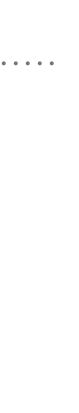
keystone of standard model

> so far φ^4 only ever seen in textbooks!

can be tested through triple-Higgs interaction

best route: a higher-energy pp collider

(at LHC, Higgs pair produced only in ~ 1 in 3 trillion collisions







for much of Higgs sector, we know what to do to get answers. What about other "big" questions

Nature of dark m Fine-tuning (e.g. sup Matter-antimatter as

- Nature of dark matter (& dark energy)
- Fine-tuning (e.g. supersymmetry and similar)
- Matter-antimatter asymmetry of the universe
 - [...]



Finding dark matter and studying it will be the biggest challenge for the Large Hadron Collider's second run

https://www.pbs.org/newshour/science/largehadron-collider-gears-find-dark-matter-newparticles-second-run

-a large LHC experiment's spokesperson [2015]



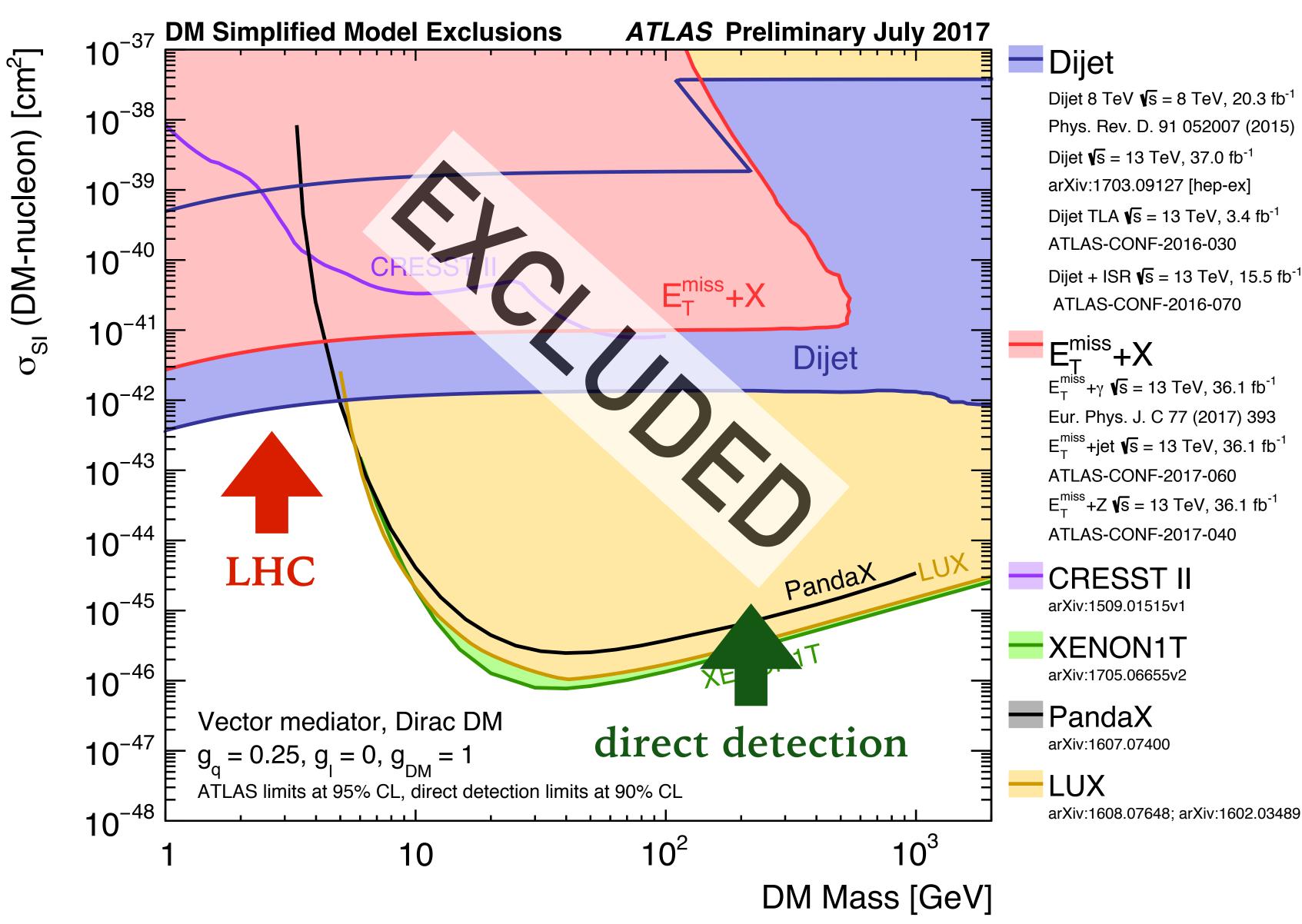


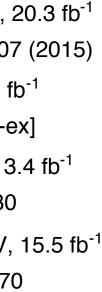
Looking beyond the SM: searches for dark matter at LHC & elsewhere

Classic dark-matter candidate: a weaklyinteracting massive particle (WIMP, e.g. from supersymmetry).

Masses ~ GeV upwards

(search interpretations strongly model dependent)







musn't be (too) disappointed at lack of dark matter signal at LHC

Evidence for dark matter exists since the 1930s.

Today we know that

- ► there are many possible models
- ► the range of parameters they span is large

We must deploy full ingenuity in searching for dark matter, including at LHC.

But must also recognise that it has remained elusive for 80–90 years, and chances of finding it in any given year are small!

Snowmass non-WIMP dark matter report, 1310.8642

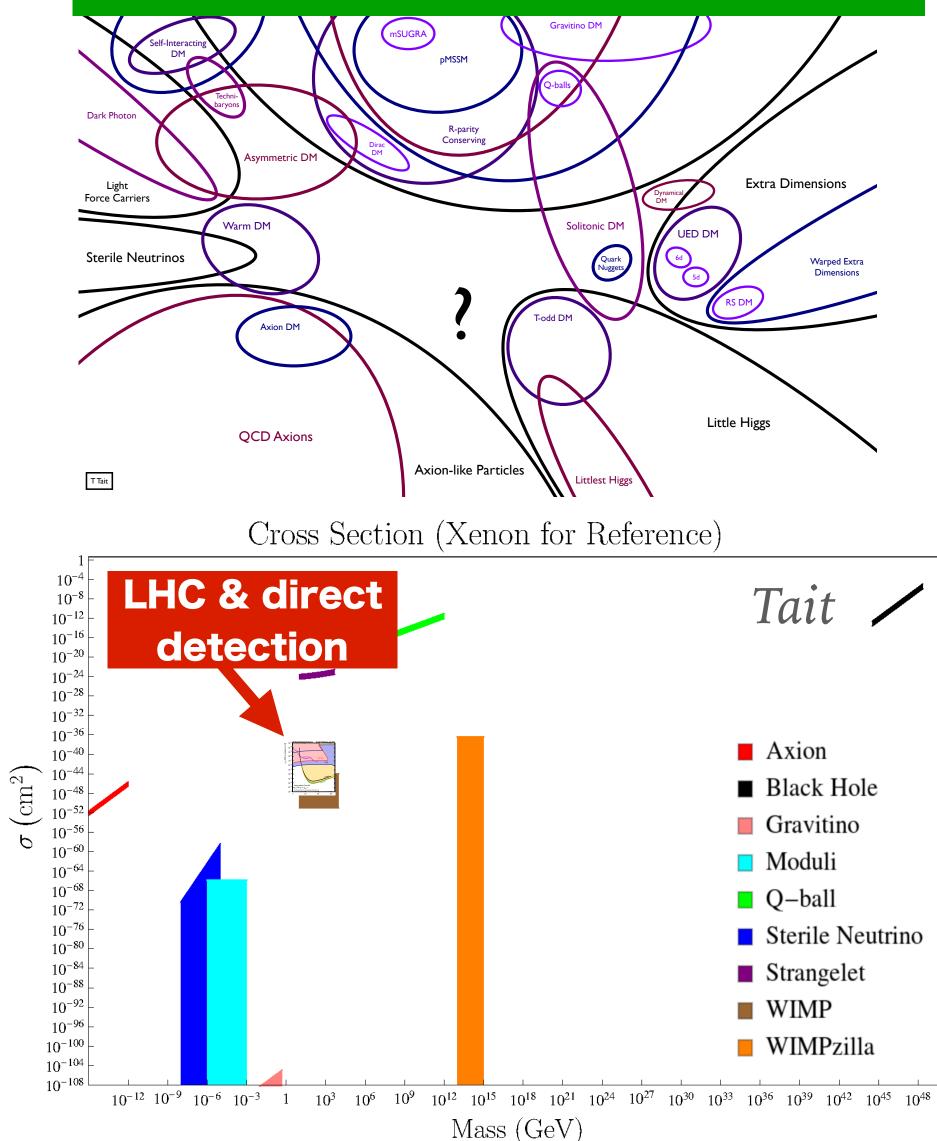


Figure 1. Graphical representation of the (incomplete) landscape of candidates. Above, the landscape of dark matter candidates due to T. Tait. Below, the range of dark matter candidates' masses and interaction cross sections with a nucleus of Xe (for illustrative purposes) compiled by L. Pearce. Dark matter candidates have an enormous range of masses and interaction cross sections.

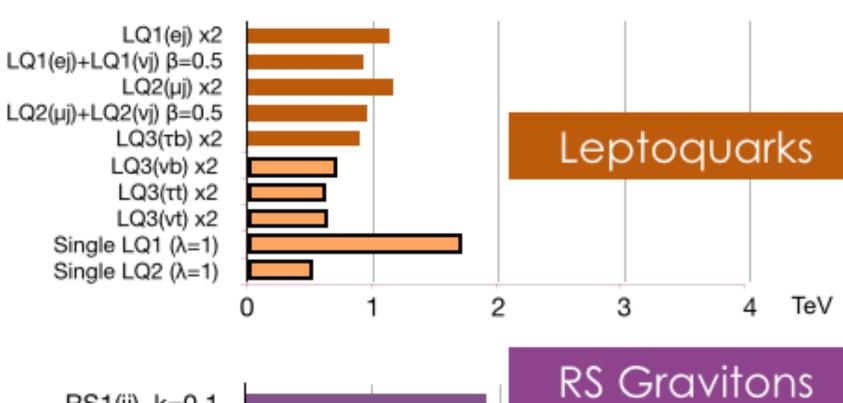


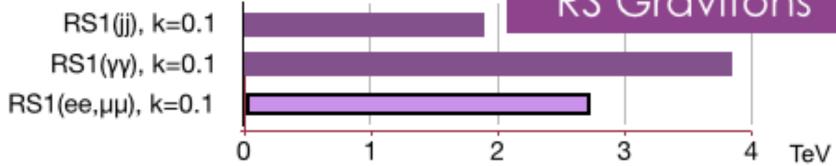
ATLAS SUSY Searches* - 95% CL Lower Limits

_	TLAS SUSY Sea lecember 2017 Model	e, μ, τ, γ					$\sqrt{s} = 7, 8 \text{ TeV}$ $\sqrt{s} = 13 \text{ TeV}$	ATLAS Preliminary $\sqrt{s} = 7, 8, 13$ TeV Reference
Inclusive Searches	$\begin{array}{l} \tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0} \\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0} \text{ (compressed)} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{1} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{\pm} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}(\ell\ell)\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qq(\ell\ell/\nu\nu)\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_{1}^{0} \\ \text{GMSB}(\tilde{\ell} \text{ NLSP}) \\ \text{GGM (bino NLSP)} \\ \text{GGM (higgsino-bino NLSP)} \\ \end{array}$	0 mono-jet 0 <i>ee</i> , μμ 3 <i>e</i> , μ 0 1-2 τ + 0-1 <i>t</i> 2 γ γ	2-6 jets 1-3 jets 2-6 jets 2 jets 4 jets 7-11 jets - 2 jets	Yes Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 36.1 14.7 36.1 36.1 36.1 36.1 36.1	710 GeV	1.57 TeV $m(\tilde{x}_{1}^{0}) < 200 \text{ GeV}, m(1^{\text{st}} \text{ gen.} \tilde{q}) = m(2^{nd} \text{ gen.} \tilde{q})$ $m(\tilde{q}) - m(\tilde{x}_{1}^{0}) < 5 \text{ GeV}$ 2.02 TeV $m(\tilde{x}_{1}^{0}) < 200 \text{ GeV}$ 2.01 TeV $m(\tilde{x}_{1}^{0}) < 200 \text{ GeV}, m(\tilde{x}^{\text{st}}) = 0.5(m(\tilde{x}_{1}^{0}) + m(\tilde{g}))$ 1.7 TeV $m(\tilde{x}_{1}^{0}) < 300 \text{ GeV},$ 1.87 TeV $m(\tilde{x}_{1}^{0}) = 0 \text{ GeV}$ 1.8 TeV $m(\tilde{x}_{1}^{0}) = 0 \text{ GeV}$ 1.8 TeV $m(\tilde{x}_{1}^{0}) = 0 \text{ GeV}$ 2.0 TeV 2.15 TeV $c\tau(\text{NLSP}) < 0.1 \text{ mm}$ 2.05 TeV $m(\tilde{x}_{1}^{0}) = 1700 \text{ GeV}, cr(\text{NLSP}) < 0.1 \text{ mm}, \mu > 0$ $m(\tilde{x}_{1}^{0}) = 1700 \text{ GeV}, cr(\text{NLSP}) < 0.1 \text{ mm}, \mu > 0$	1712.02332 1711.03301 1712.02332 1712.02332 1611.05791 1706.03731 1708.02794 1607.05979 ATLAS-CONF-2017-080 ATLAS-CONF-2017-080
3 rd gen. § med.	Gravitino LSP $\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\bar{b}\tilde{\chi}_{1}^{0}$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_{1}^{0}$	0 0-1 c, μ	mono-jet 3 b 3 b	Yes Yes Yes	20.3 36.1 36.1	^{1/2} scale 865 GeV	$\begin{split} m(\tilde{G}) > 1.8 \times 10^{-4} \ \mathrm{eV}, \ m(\tilde{g}) = m(\tilde{q}) = 1.5 \ \mathrm{TeV} \\ \hline \mathbf{1.92 \ TeV} \qquad m(\tilde{\chi}_1^0) < 600 \ \mathrm{GeV} \\ \hline \mathbf{1.97 \ TeV} \qquad m(\tilde{\chi}_1^0) < 200 \ \mathrm{GeV} \end{split}$	1502.01518 1711.01901 1711.01901
3rd gen. squarks 3 direct production	$\begin{split} \tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow b \tilde{\chi}_1^0 \\ \tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow t \tilde{\chi}_1^{\pm} \\ \tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow b \tilde{\chi}_1^{\pm} \\ \tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0 \text{ or } t \tilde{\chi}_1^0 \\ \tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow c \tilde{\chi}_1^0 \\ \tilde{t}_1 \tilde{t}_1 (natural GMSB) \end{split}$	0 2 e, μ (SS) 0-2 e, μ 0-2 e, μ (0 2 e, μ (Z) 3 e, μ (Z) 1-2 e, μ	2 b 1 b 1-2 b 0-2 jets/1-2 mono-jet 1 b 1 b 4 b		36.1 36.1 4.7/13.3 20.3/36.1 36.1 20.3 36.1 36.1 36.1	950 GeV 275-700 GeV 117-170 GeV 200-720 GeV 90-198 GeV 0.195-1.0 TeV 90-430 GeV 90-430 GeV 290-790 GeV 290-790 GeV	$\begin{split} & m(\tilde{k}_1^0) \!<\! 420 \mathrm{GeV} \\ & m(\tilde{k}_1^0) \!<\! 200 \mathrm{GeV}, m(\tilde{k}_1^\pm) \!= m(\tilde{k}_1^0) \!+\! 100 \mathrm{GeV} \\ & m(\tilde{k}_1^\pm) \!=\! 2m(\tilde{k}_1^0), m(\tilde{k}_1^0) \!\!=\!\! 55 \mathrm{GeV} \\ & m(\tilde{k}_1^0) \!\!=\!\! 1 \mathrm{GeV} \\ & m(\tilde{k}_1^0) \!\!=\!\! 5 \mathrm{GeV} \\ & m(\tilde{k}_1^0) \!\!>\!\! 150 \mathrm{GeV} \\ & m(\tilde{k}_1^0) \!\!=\!\! 0 \mathrm{GeV} \\ & m(\tilde{k}_1^0) \!\!=\!\! 0 \mathrm{GeV} \\ & m(\tilde{k}_1^0) \!\!=\!\! 0 \mathrm{GeV} \end{split}$	1708.09266 1706.03731 1209.2102, ATLAS-CONF-2016-077 1506.08616, 1709.04183, 1711.11520 1711.03301 1403.5222 1706.03986 1706.03986
EW direct	$\begin{array}{l} \tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\ell}\nu(\ell \tilde{\nu}) \\ \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{1}^{\mp}/\tilde{\chi}_{2}^{0}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\tau}\nu(\tau \tilde{\nu}), \tilde{\chi}_{2}^{0} \rightarrow \tilde{\tau}\tau(\nu \tilde{\nu}) \\ \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{L}\nu \tilde{\ell}_{L}\ell(\tilde{\nu}\nu), \ell \tilde{\nu} \tilde{\ell}_{L}\ell(\tilde{\nu}\nu) \\ \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} L \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} h \tilde{\chi}_{1}^{0}, h \rightarrow b \bar{b} / W W / \tau \tau / \gamma \gamma \\ \tilde{\chi}_{2}^{0} \tilde{\chi}_{3}^{0}, \tilde{\chi}_{2,3}^{0} \rightarrow \tilde{\ell}_{R} \ell \\ \text{GGM (wino NLSP) weak prod., } \tilde{\chi}_{1}^{0} \rightarrow \end{array}$	-	0 0 - 0 -2 jets 0-2 <i>b</i> 0 -	Yes Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 36.1 36.1 20.3 20.3 20.3 36.1	90-500 GeV 750 GeV 760 GeV 1.13 TeV 1.13 TeV 1.13 TeV 580 GeV 2.3 635 GeV 1.16 TeV	$\begin{split} & m(\tilde{\mathfrak{X}}_{1}^{0}) \!=\! 0 \\ & m(\tilde{\mathfrak{X}}_{1}^{0}) \!=\! 0, m(\tilde{\ell}, \tilde{\nu}) \!=\! 0.5(m(\tilde{\mathfrak{X}}_{1}^{\pm}) \!+\! m(\tilde{\mathfrak{X}}_{1}^{0})) \\ & m(\tilde{\mathfrak{X}}_{1}^{0}) \!=\! 0, m(\tilde{\tau}, \tilde{\nu}) \!=\! 0.5(m(\tilde{\mathfrak{X}}_{1}^{\pm}) \!+\! m(\tilde{\mathfrak{X}}_{1}^{0})) \\ & m(\tilde{\mathfrak{X}}_{1}^{\pm}) \!=\! m(\tilde{\mathfrak{X}}_{2}^{0}), m(\tilde{\mathfrak{X}}_{1}^{0}) \!=\! 0, m(\tilde{\ell}, \tilde{\nu}) \!=\! 0.5(m(\tilde{\mathfrak{X}}_{1}^{\pm}) \!+\! m(\tilde{\mathfrak{X}}_{1}^{0})) \\ & m(\tilde{\mathfrak{X}}_{1}^{\pm}) \!=\! m(\tilde{\mathfrak{X}}_{2}^{0}), m(\tilde{\mathfrak{X}}_{1}^{0}) \!=\! 0, \tilde{\ell} \text{ decoupled} \\ & m(\tilde{\mathfrak{X}}_{1}^{\pm}) \!=\! m(\tilde{\mathfrak{X}}_{2}^{0}), m(\tilde{\mathfrak{X}}_{1}^{0}) \!=\! 0, \tilde{\ell} \text{ decoupled} \\ & m(\tilde{\mathfrak{X}}_{2}^{0}) \!=\! m(\tilde{\mathfrak{X}}_{3}^{0}), m(\tilde{\mathfrak{X}}_{1}^{0}) \!=\! 0, m(\tilde{\ell}, \tilde{\nu}) \!=\! 0.5(m(\tilde{\mathfrak{X}}_{2}^{0}) \!+\! m(\tilde{\mathfrak{X}}_{1}^{0})) \\ & c\tau \!<\! 1 mm \\ c\tau \!<\!\! 1 mm \end{split}$	ATLAS-CONF-2017-039 ATLAS-CONF-2017-039 1708.07875 ATLAS-CONF-2017-039 ATLAS-CONF-2017-039 1501.07110 1405.5086 1507.05493 ATLAS-CONF-2017-080
Long-lived particles	Direct $\tilde{X}_{1}^{\dagger}\tilde{X}_{1}^{-}$ prod., long-lived \tilde{X}_{1}^{\pm} Direct $\tilde{X}_{1}^{\dagger}\tilde{X}_{1}^{-}$ prod., long-lived \tilde{X}_{1}^{\pm} Stable, stopped \tilde{g} R-hadron Stable \tilde{g} R-hadron Metastable \tilde{g} R-hadron Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow qq \tilde{X}_{1}^{0}$ GMSB, stable $\tilde{\tau}, \tilde{X}_{1}^{0} \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e, \mu)$ GMSB, $\tilde{X}_{1}^{0} \rightarrow \gamma \tilde{G}$, long-lived \tilde{X}_{1}^{0} $\tilde{g}\tilde{g}, \tilde{X}_{1}^{0} \rightarrow eev/e\mu v/\mu\mu v$	Disapp. trk dE/dx trk 0 trk dE/dx trk displ. vtx $1-2 \mu$ 2γ displ. $ee/e\mu/\mu$	- - - - - -	Yes Yes - Yes - Yes - Yes	36.1 18.4 27.9 3.2 3.2 32.8 19.1 20.3 20.3		$\begin{split} & m(\tilde{\chi}_1^{\pm})\text{-}m(\tilde{\chi}_1^{0})\text{-}160~MeV,~\tau(\tilde{\chi}_1^{\pm})\text{=}0.2~ns\\ & m(\tilde{\chi}_1^{\pm})\text{-}m(\tilde{\chi}_1^{0})\text{-}160~MeV,~\tau(\tilde{\chi}_1^{\pm})\text{<}15~ns\\ & m(\tilde{\chi}_1^{0})\text{=}100~GeV,~10~\mus\text{<}\tau(\tilde{g})\text{<}1000~s \end{split}$	1712.02118 1506.05332 1310.6584 1606.05129 1604.04520 1710.04901 1411.6795 1409.5542 1504.05162
RPV	LFV $pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e\mu/e\tau/\mu\tau$ Bilinear RPV CMSSM $\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow W\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow eev, e\mu\nu, \mu\mu\nu$ $\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow W\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow \tau\tau\nu_{e}, e\tau\nu_{\tau}$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow qqq$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow qqq$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_{1}t, \tilde{t}_{1} \rightarrow bs$ $\tilde{t}_{1}\tilde{t}_{1}, \tilde{t}_{1} \rightarrow b\ell$	1 <i>e</i> ,μ 8 1 <i>e</i> ,μ 8	- 0-3 <i>b</i> - -5 large- <i>R</i> je -10 jets/0-4 -10 jets/0-4 2 jets + 2 <i>b</i> 2 <i>b</i>	b - b -	3.2 20.3 13.3 20.3 36.1 36.1 36.1 36.7 36.1	2 1.4 1.4 1.14	1.9 TeV $\lambda'_{311}=0.11, \lambda_{132/133/233}=0.07$ 15 TeV $m(\tilde{g})=m(\tilde{g}), c\tau_{LSP}<1 \text{ mm}$ $m(\tilde{\chi}_1^0)>400 \text{GeV}, \lambda_{12k}\neq 0 \ (k=1,2)$ $m(\tilde{\chi}_1^0)>0.2\times m(\tilde{\chi}_1^{\pm}), \lambda_{133}\neq 0$ 1.875 TeV $m(\tilde{\chi}_1^0)=1075 \text{ GeV}$ 2.1 TeV $m(\tilde{\chi}_1^0)=1 \text{ TeV}, \lambda_{112}\neq 0$ 1.65 TeV $m(\tilde{t}_1)=1 \text{ TeV}, \lambda_{323}\neq 0$	1607.08079 1404.2500 ATLAS-CONF-2016-075 1405.5086 SUSY-2016-22 1704.08493 1704.08493 1710.07171 1710.05544
Othe	Scalar charm, $\tilde{c} \rightarrow c \tilde{\chi}_1^0$	0	2 <i>c</i>	Yes	20.3	510 GeV	$m(\tilde{\ell}_1^0) < 200 \text{ GeV}$	1501.01325

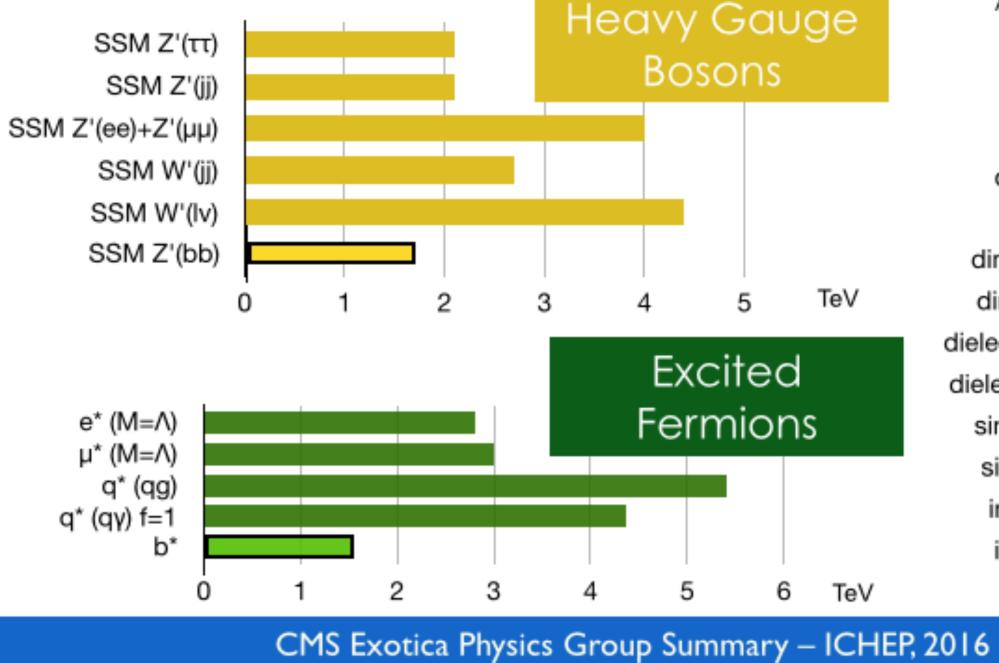
*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.



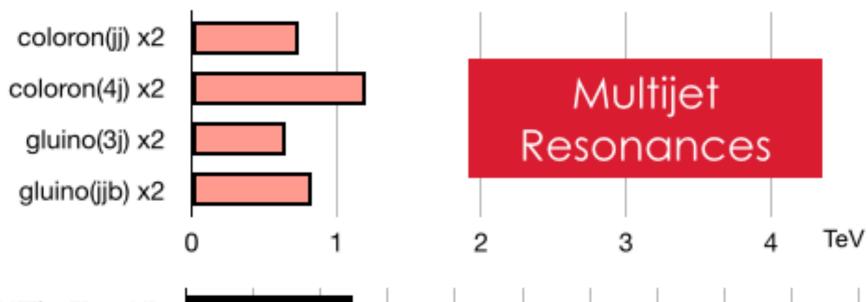


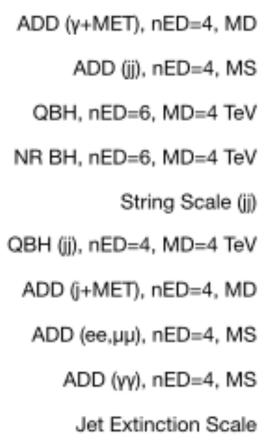


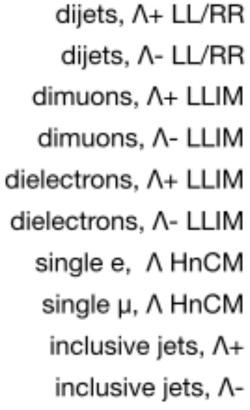
CMS Preliminary

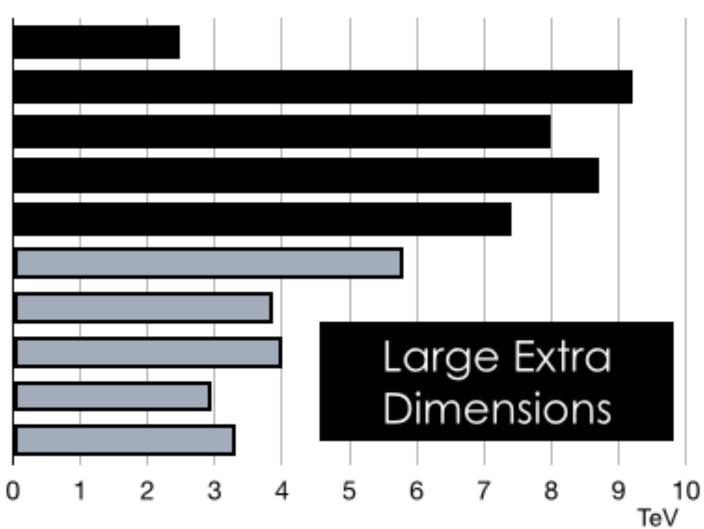


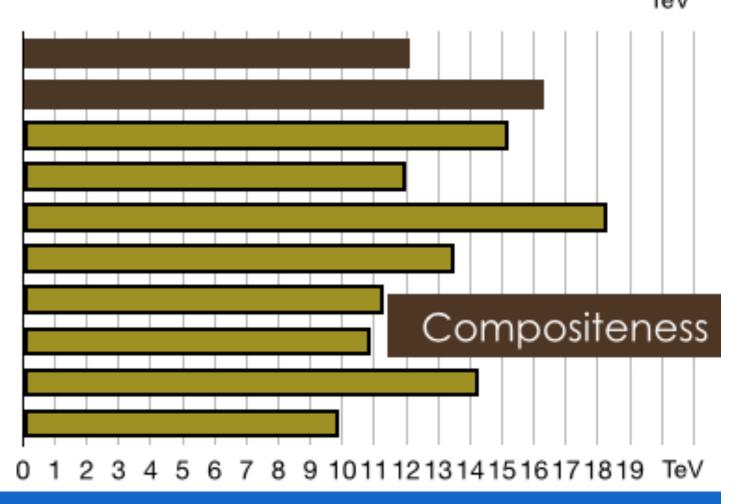
8 TeV 13 TeV













anomalies

the current place where there are hints of something happening

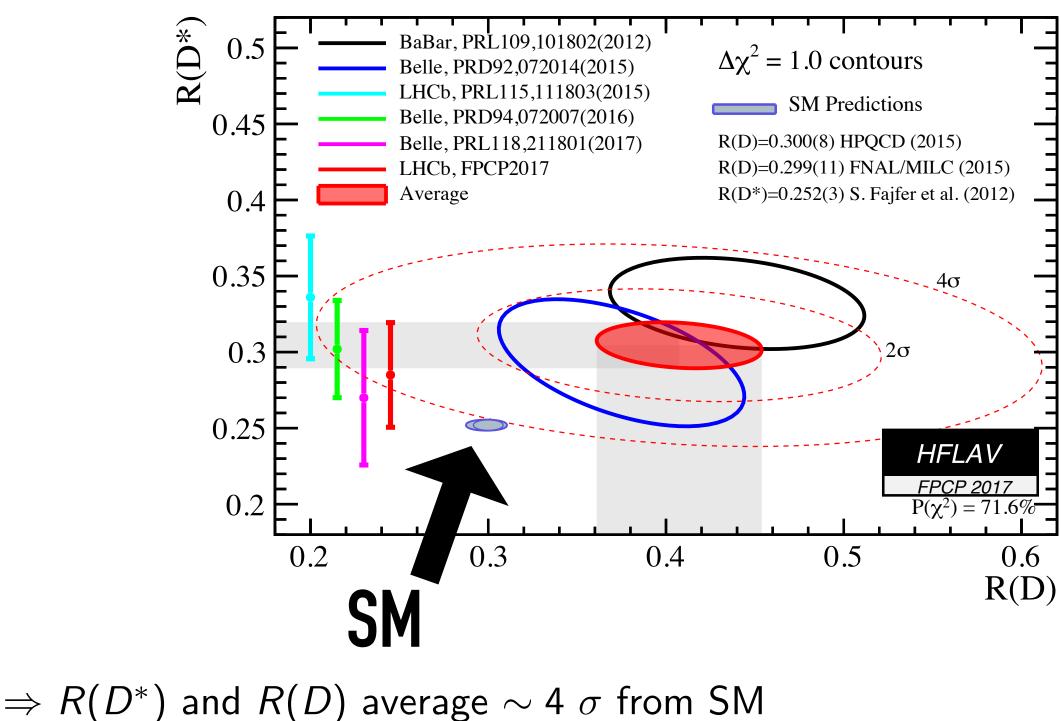




charged current

$$R(D^*) \equiv rac{\mathcal{B}(B^0
ightarrow D^{*-} au^+
u_ au)}{\mathcal{B}(B^0
ightarrow D^{*-} \mu^+
u_\mu)}$$

$R(D^*)$ and R(D) combination Combine LHCb's $R(D^*)$ results with results from B factories:



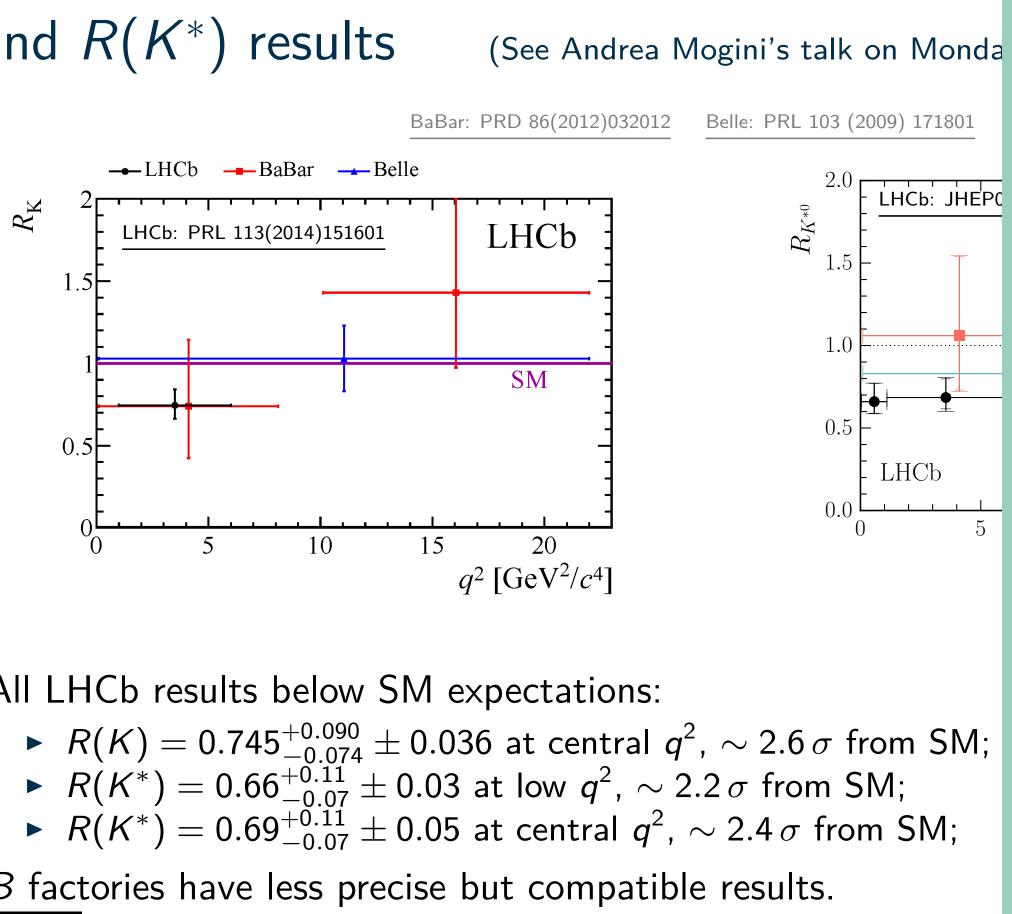
(latest SM computation: JHEP 11 (2017) 061)

Humair @ LHCP'18

neutral current

$$R(K^{(*)}) = \frac{\mathcal{B}(B \to K^{(*)}\mu^+\mu^-)}{\mathcal{B}(B \to K^{(*)}e^+e^-)}$$

R(K) and $R(K^*)$ results



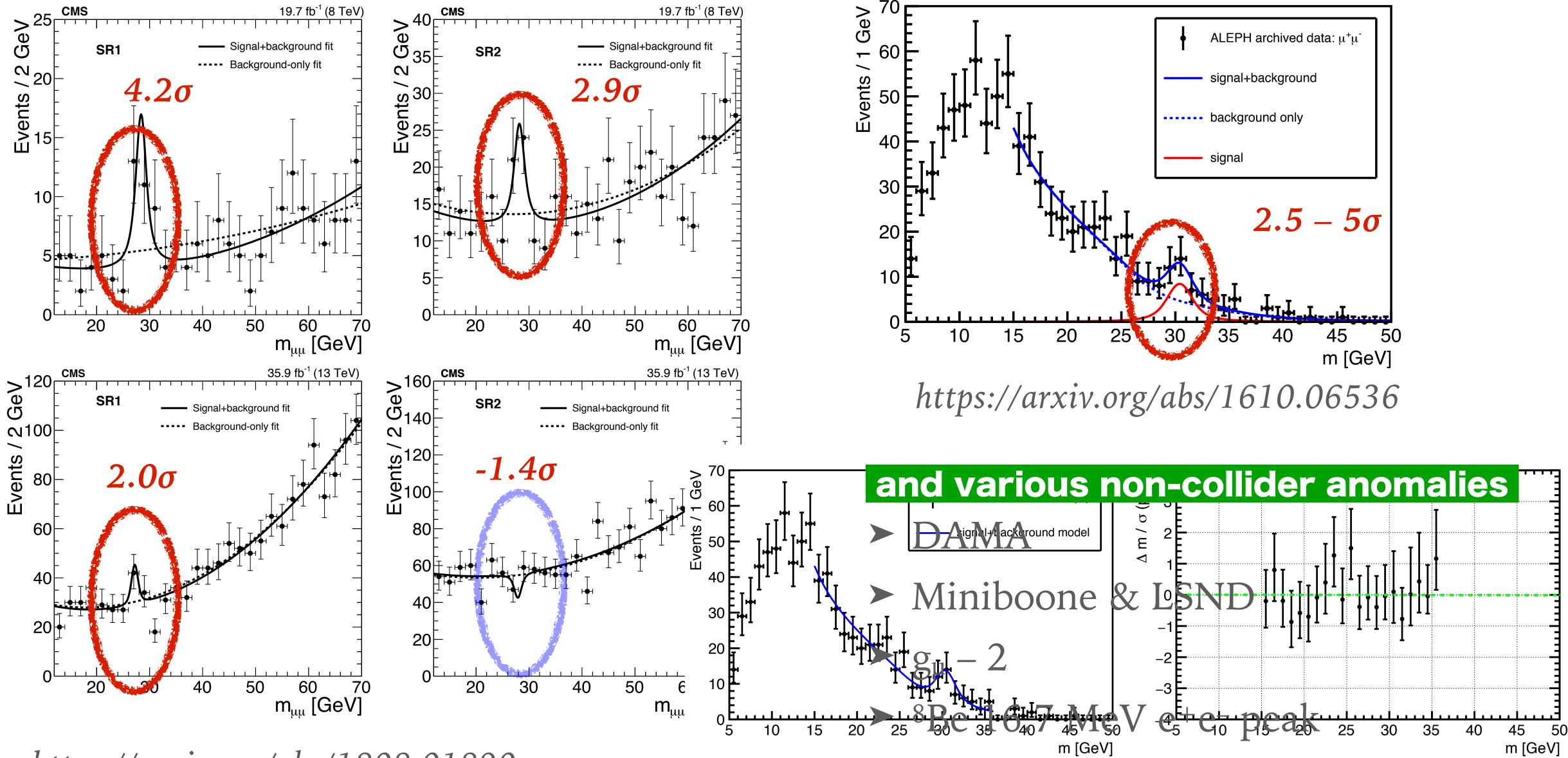
All LHCb results below SM expectations:

► *B* factories have less precise but compatible results.



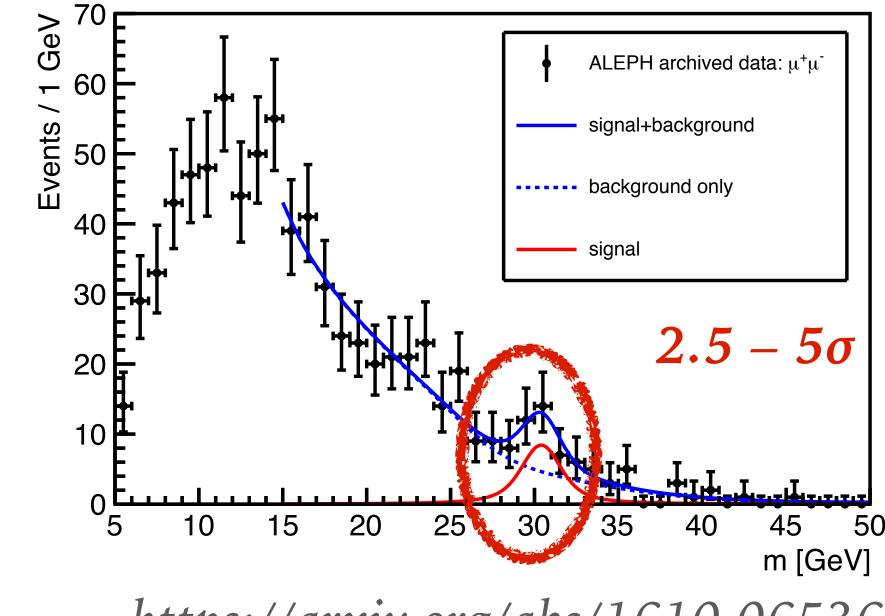


CMS pp \rightarrow b j $\mu^+\mu^-$ + X



https://arxiv.org/abs/1808.01890

ALEPH e⁺e⁻ \rightarrow b b $\mu^+\mu^-$ + X

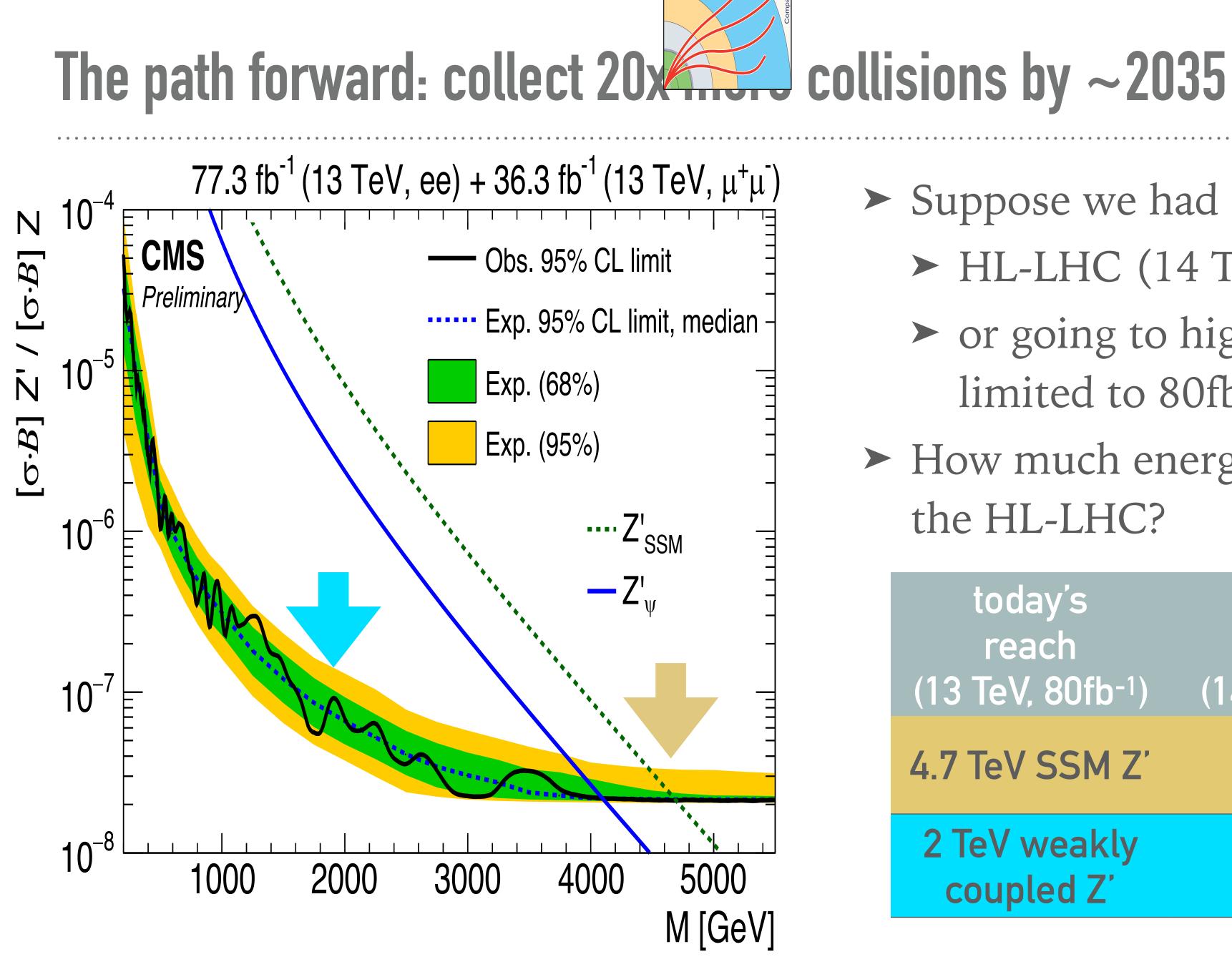




future progress?

LHC will collect ~40 times more data than used for the plots shown so far, though at mostly similar energy (13-14 TeV)





- Suppose we had a choice between ► HL-LHC (14 TeV, 3ab⁻¹)
 - or going to higher c.o.m. energy but limited to 80fb⁻¹.
- How much energy would we need to equal the HL-LHC?

today's reach (13 TeV, 80fb ⁻¹)	HL-LHC reach (14 TeV 3ab ⁻¹)	energy neede for same reac with 80fb ⁻¹
4.7 TeV SSM Z'	6.7 TeV	20 TeV
2 TeV weakly coupled Z'	3.7 TeV	37 TeV

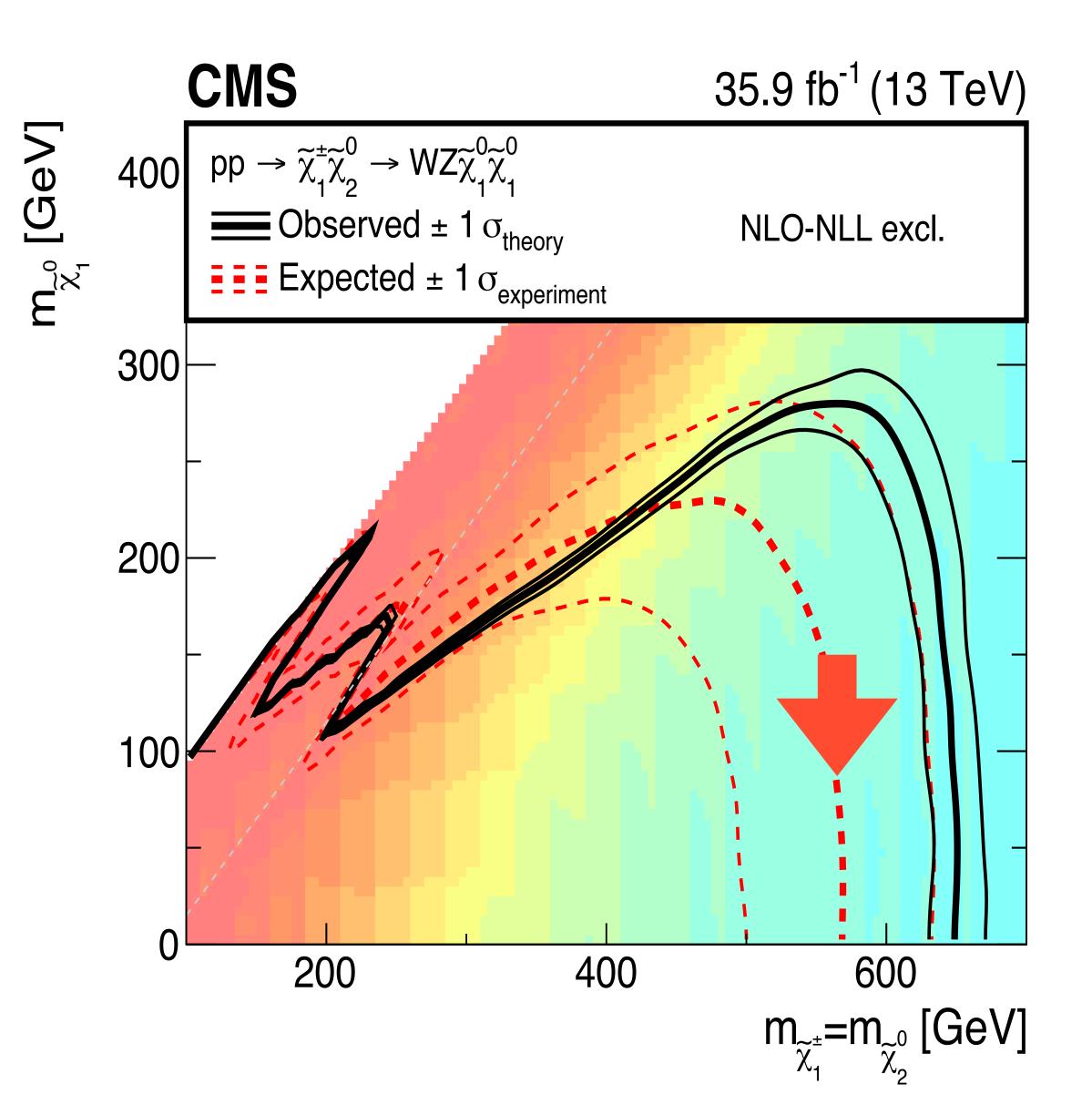
estimated with <u>http://collider-reach.cern.ch</u>, Weiler & GPS







The path forward: collect 20–30x more collisions by ~2035



- Suppose we had a choice between
 - ► HL-LHC (14 TeV, 3ab⁻¹)
 - or going to higher c.o.m. energy but limited to 80fb⁻¹.
- How much energy would we need to equal the HL-LHC?

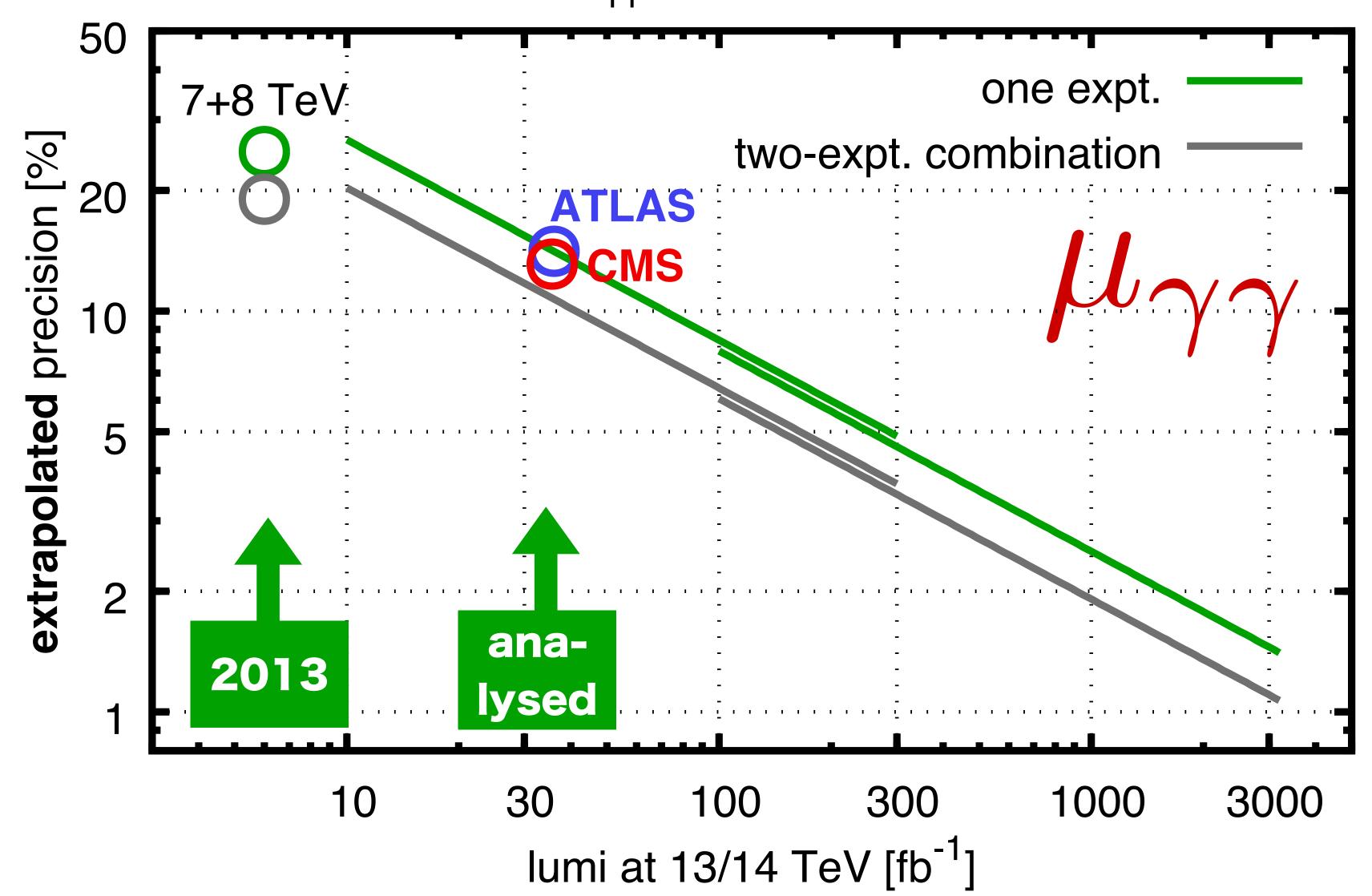
today's reach (13 TeV, 80fb ⁻¹)	HL-LHC reach (14 TeV 3ab ⁻¹)	energy neede for same reac with 80fb ⁻¹
4.7 TeV SSM Z'	6.7 TeV	20 TeV
2 TeV weakly coupled Z'	3.7 TeV	37 TeV
680 GeV chargino	1.4 TeV	54 TeV



47

Higgs precision ($H \rightarrow \gamma \gamma$) : optimistic estimate v. luminosity & time

extrapolation of μ_{vv} precision from 7+8 TeV results

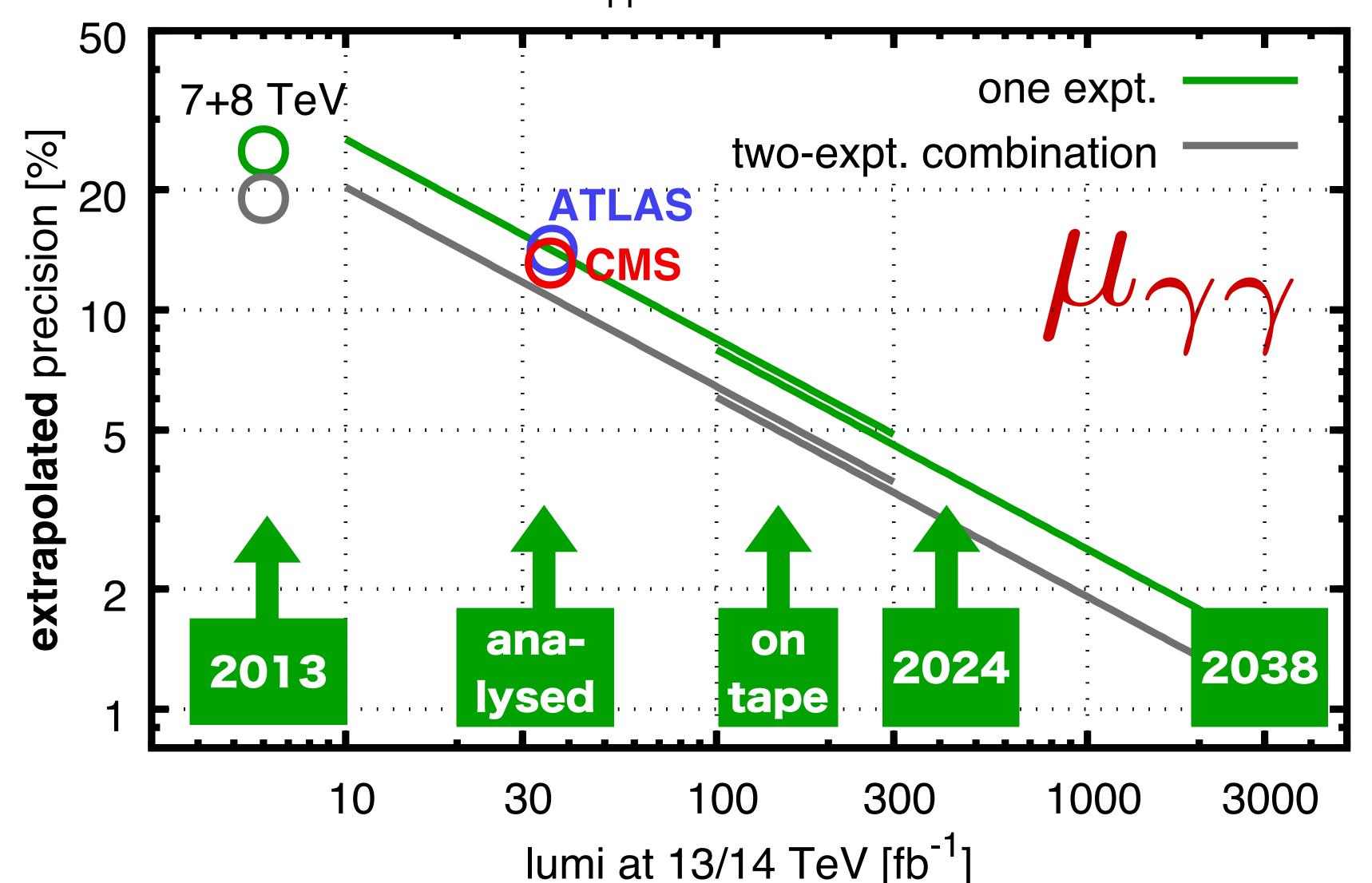


 $1 \text{ fb}^{-1} = 10^{14} \text{ collisions}$



Higgs precision (H $\rightarrow \gamma\gamma$) : optimistic estimate v. luminosity & time

extrapolation of μ_{vv} precision from 7+8 TeV results



The LHC has the statistical potential to take Higgs physics from "observation" to 1–2% precision

But only if we learn how to connect experimental observations with theory at that precision

 $1 \text{ fb}^{-1} = 10^{14} \text{ collisions}$







how is all of this made quantitative?

whether new-physics searches, Higgs physics, or other SM studies

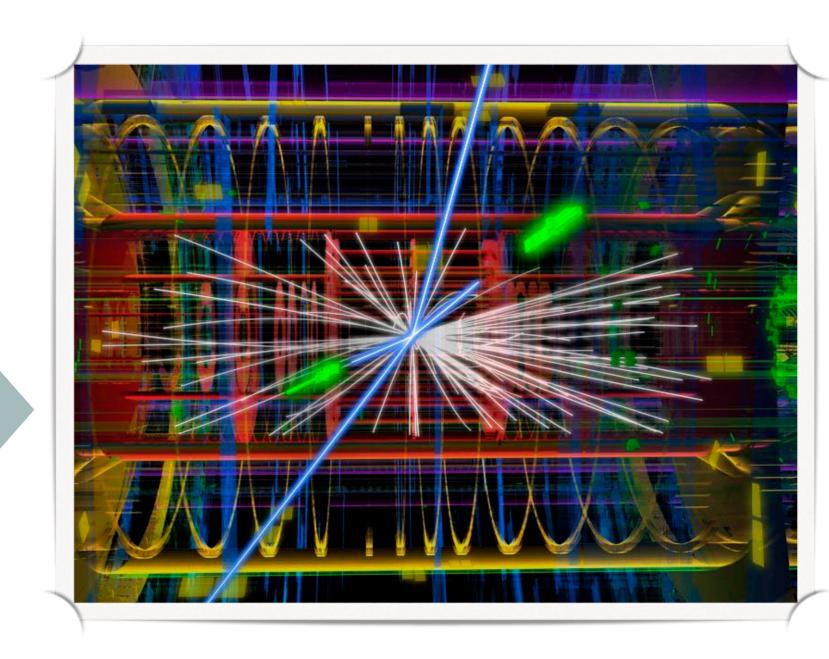


UNDERLYING THEORY

 $\begin{aligned} \mathcal{I} &= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\ &+ i F \mathcal{N} \mathcal{V} \end{aligned}$ + $\mathcal{Y}_{ij}\mathcal{Y}_{j}\phi$ +h.c + $|\mathcal{D}_{m}\phi|^{2} - V(\phi)$

EXPERIMENTAL DATA

how do you make quantitative connection?





UNDERLYING THEORY

 $\begin{aligned} \mathcal{I} &= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\ &+ i F \mathcal{B} \mathcal{F} \end{aligned}$ + $\mathcal{Y}_{ij}\mathcal{Y}_{j}\phi$ +h.c + $|\mathcal{D}_{m}\phi|^{2} - V(\phi)$

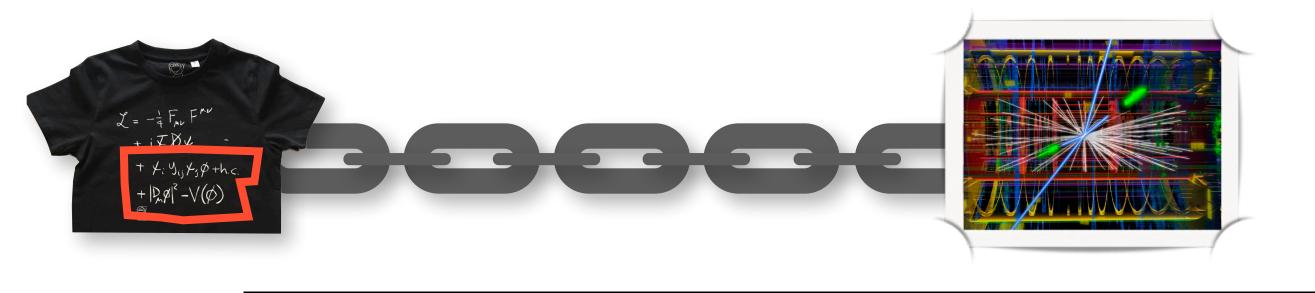
through a chain of experimental and theoretical links

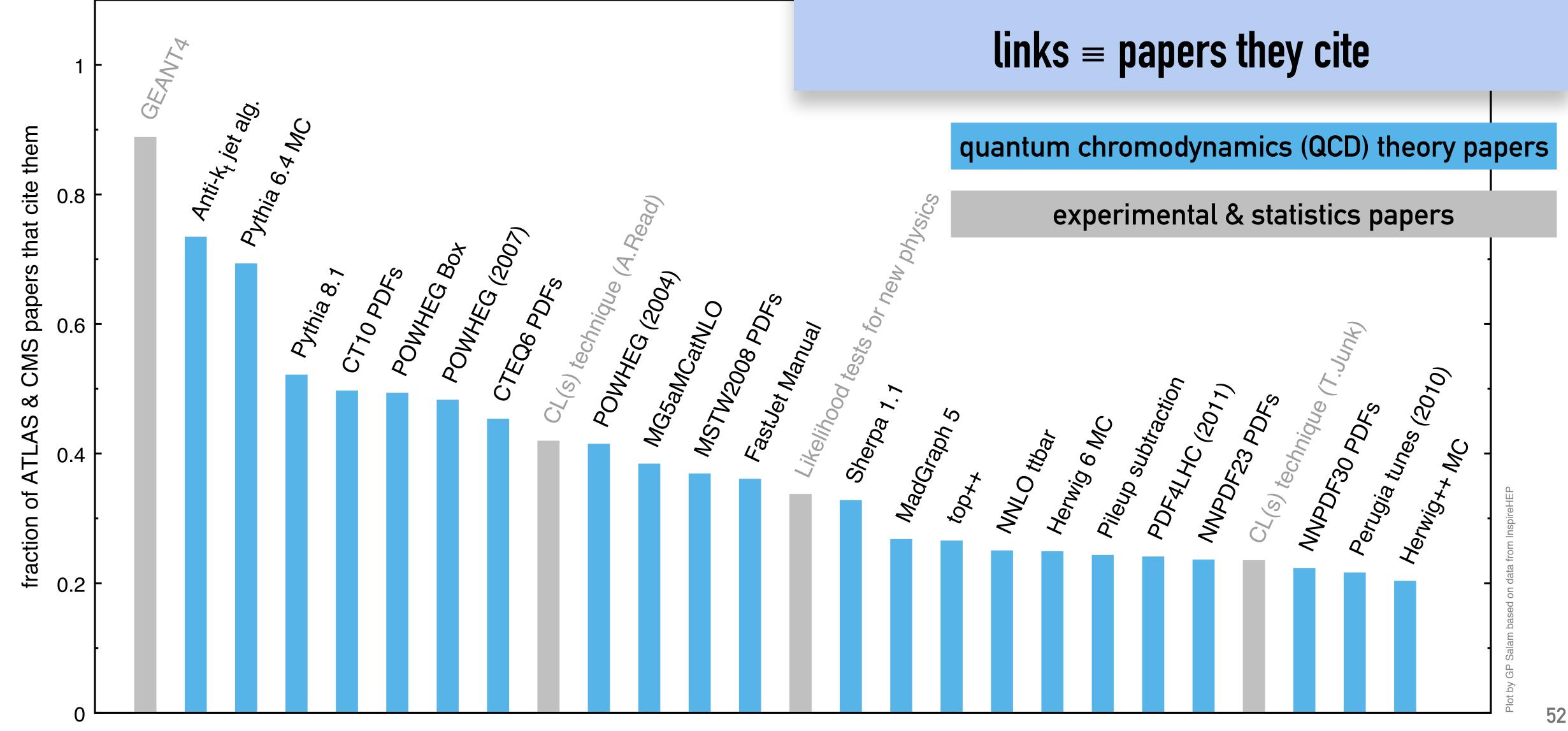
[in particular Quantum Chromodynamics (QCD)]

EXPERIMENTAL DATA

how do you make quantitative connection?







What are the links? ATLAS and CMS (big LHC expts.) have written 850 articles since 2014 links \equiv papers they cite

Pileup subtraction

Herwig 6 MC

MM Ottbar

PDF4LHC (2011)

Muddress Pors

Cl(s) technique

MNDDF30 PDFS

Likelihood tests for hew physics

Sherba 1.1

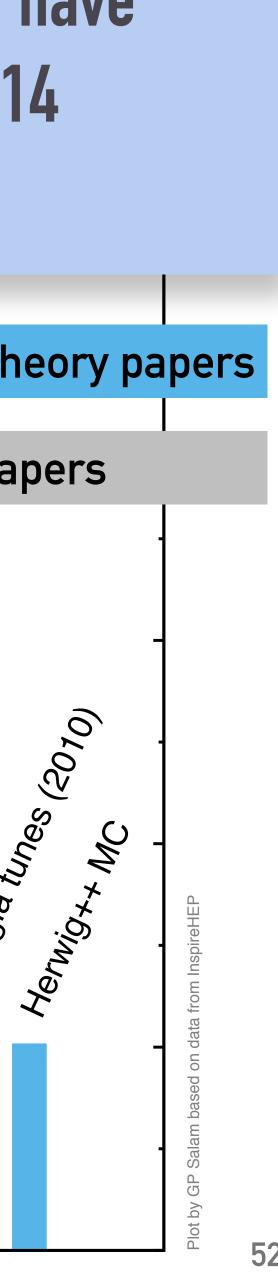
Madgraph 5

to0++

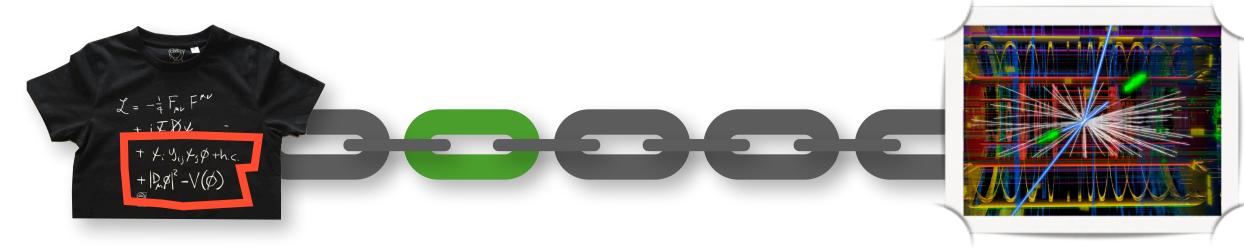
Fastuer Manual

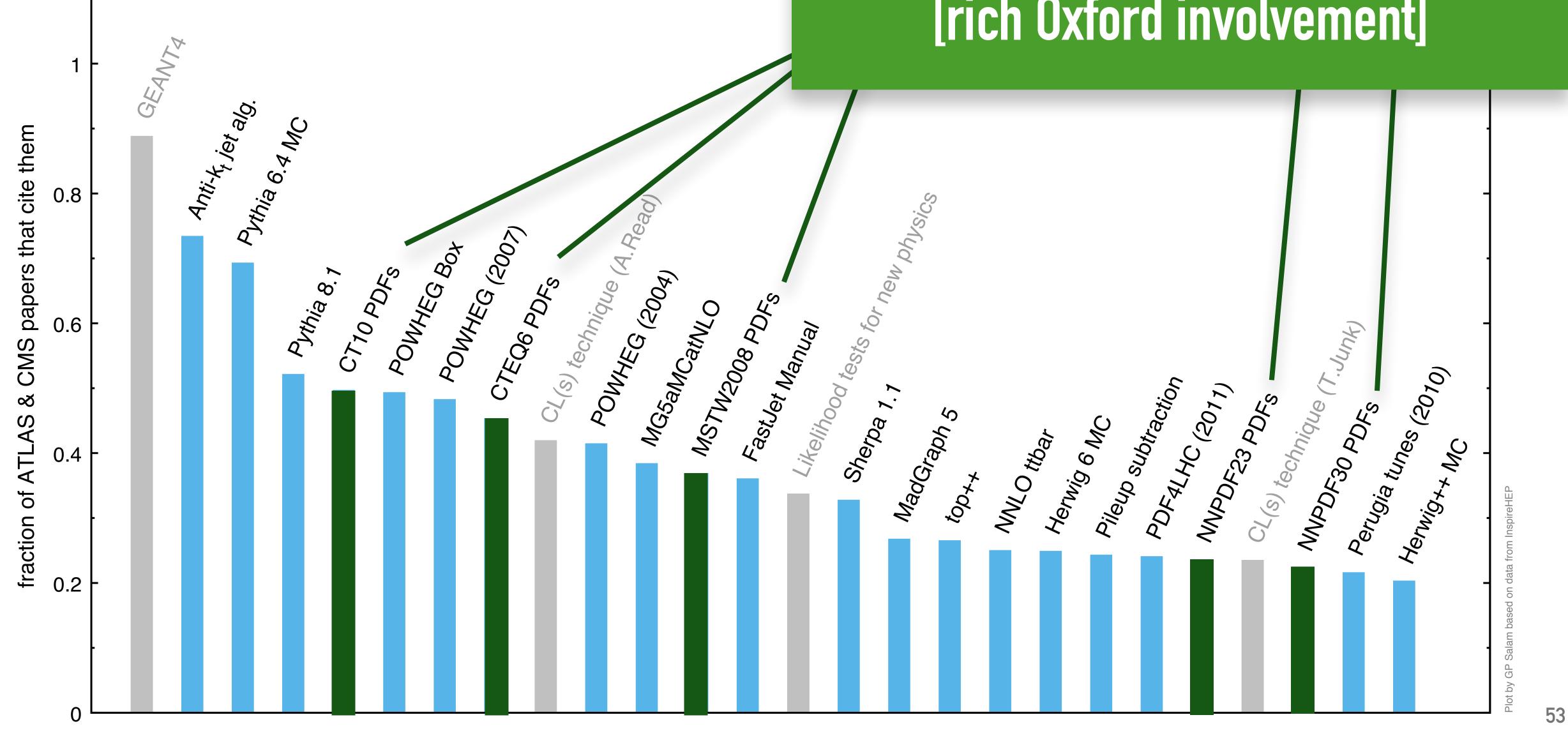
quantum chromodynamics (QCD) theory papers

experimental & statistics papers



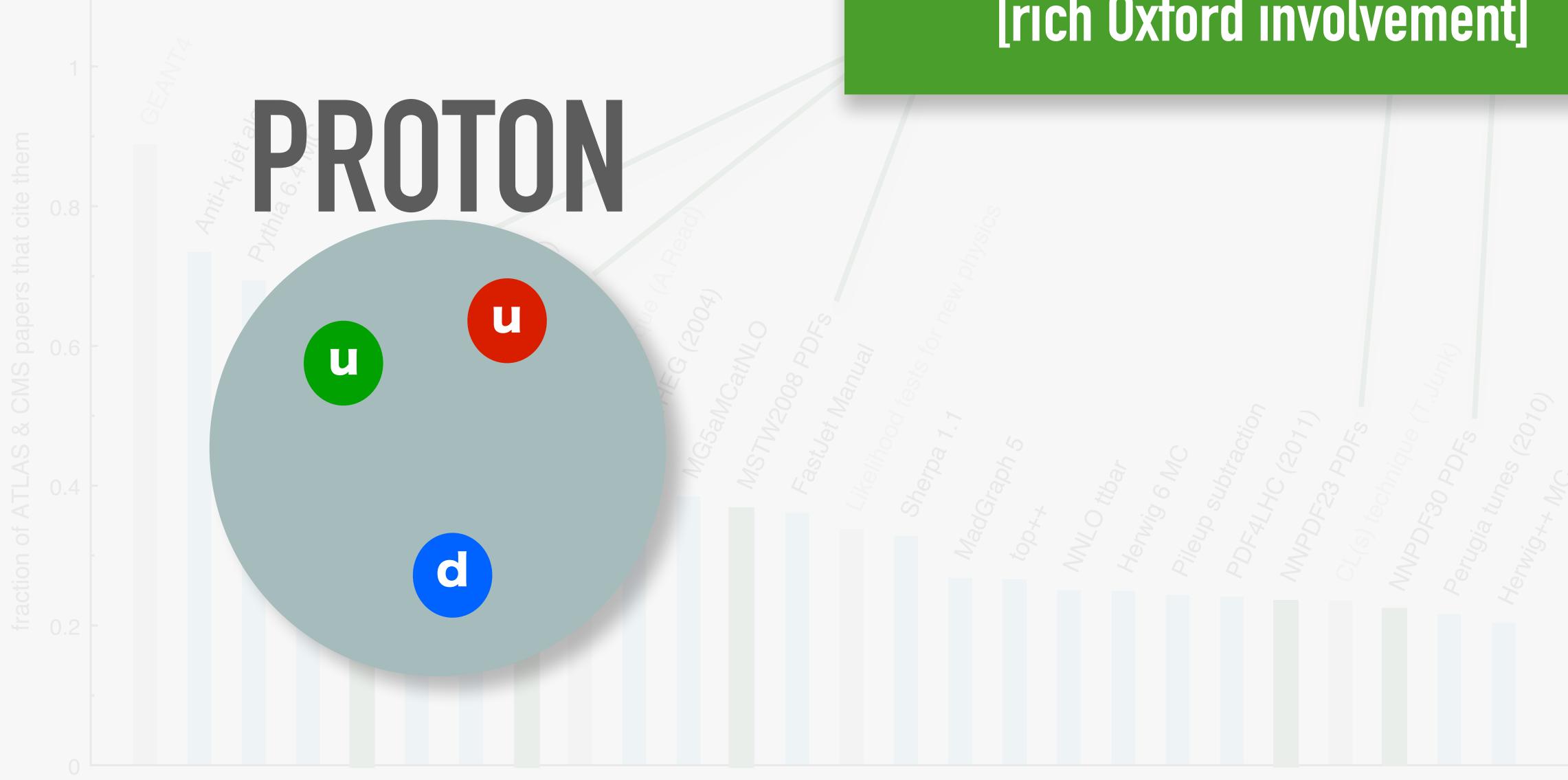
Perugia tunes (2010)





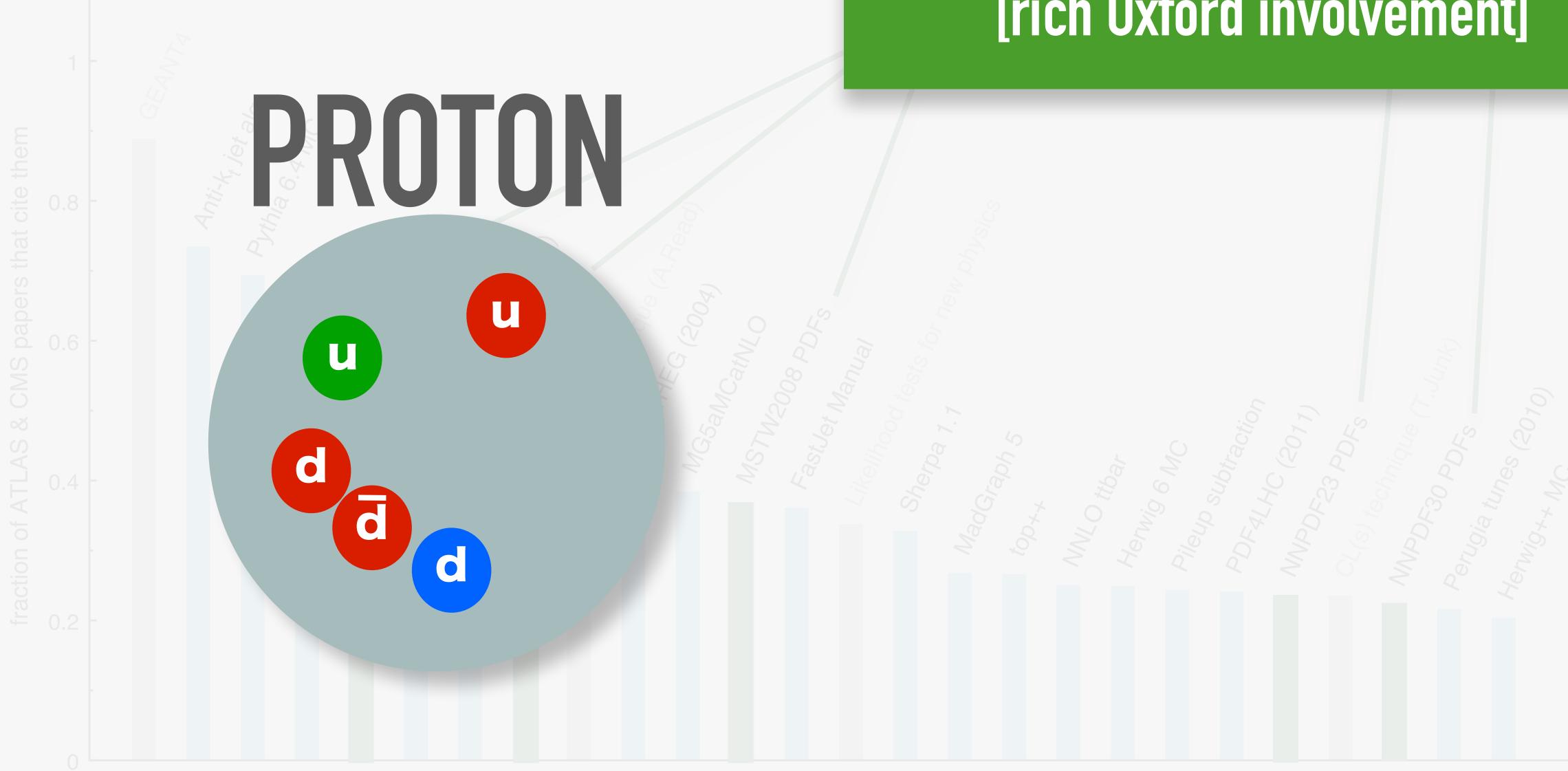






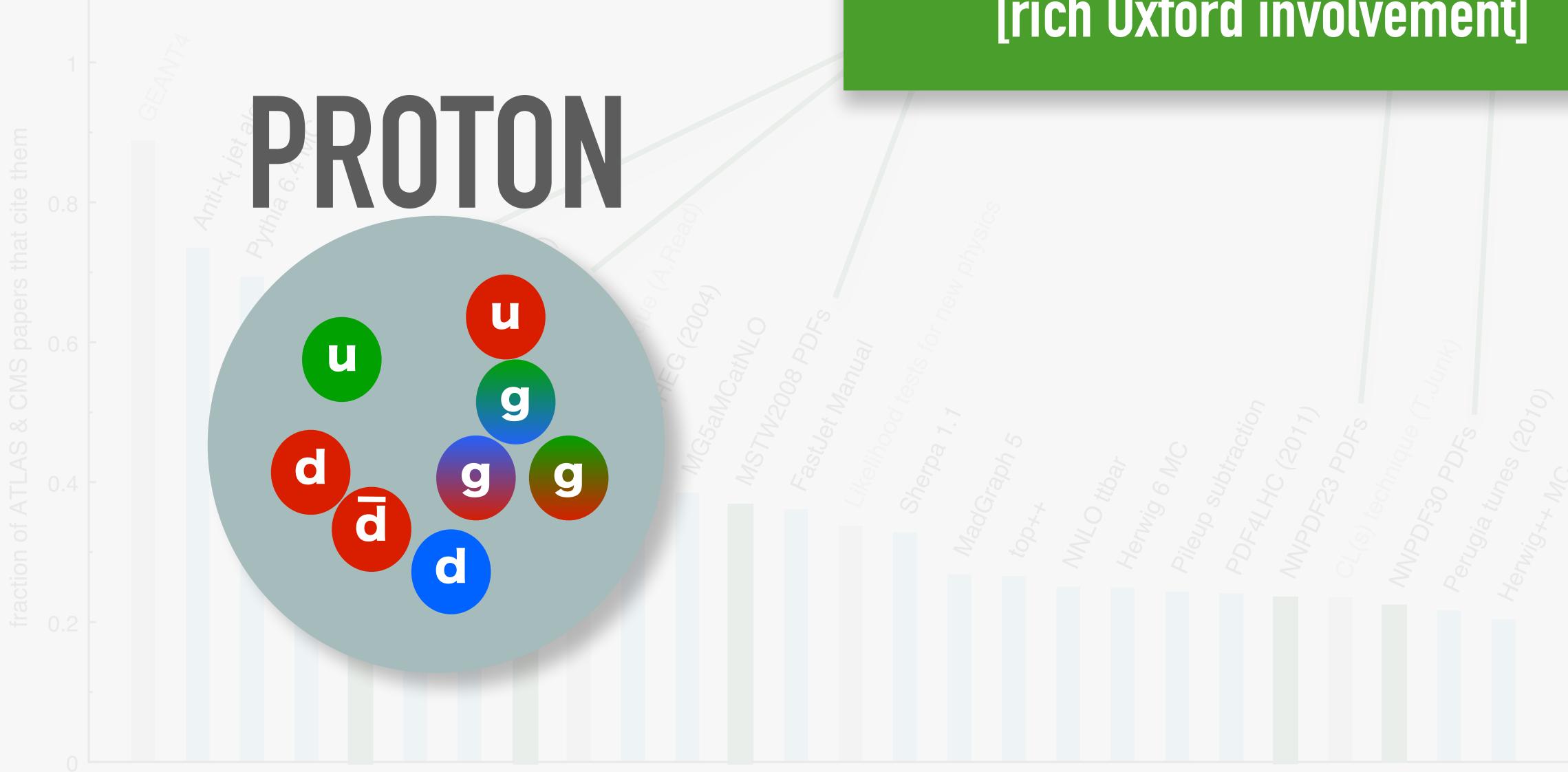






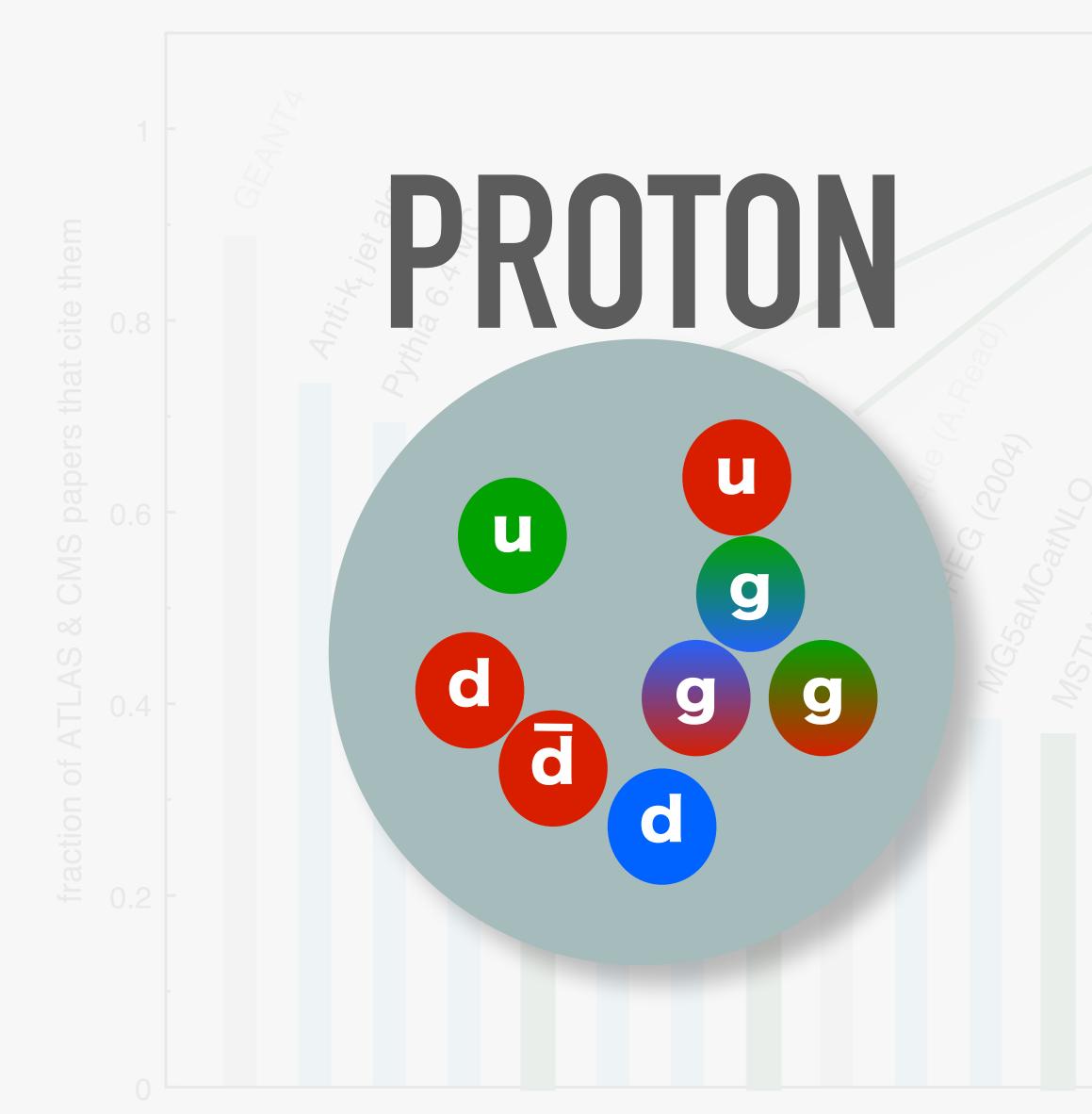












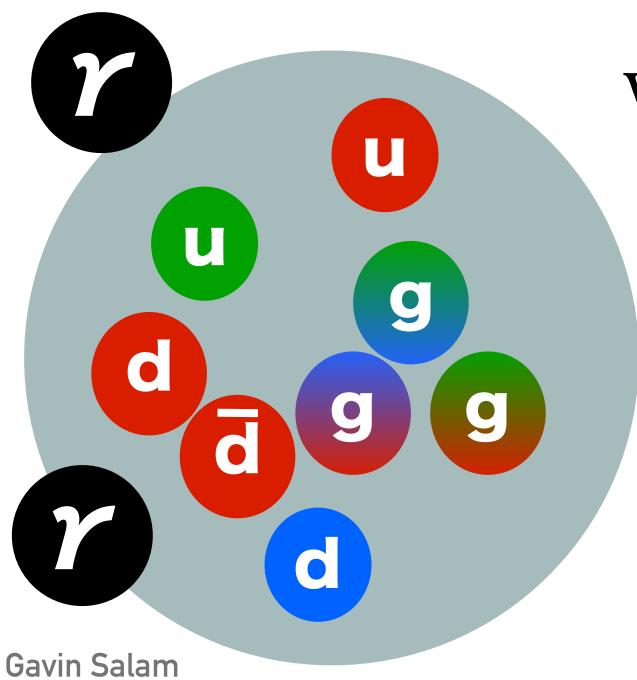
1 proton-proton collision ~ 286 ± 5 gluon-gluon collisions around the Higgs mass



how many photons accompany each proton?

A fast-moving proton comes with a cloud of photons. How many?

Number of photons accompanying an electron understood since 1934 (Fermi-Weizsäcker-Williams).



Was largest uncertainty on HW[±] production

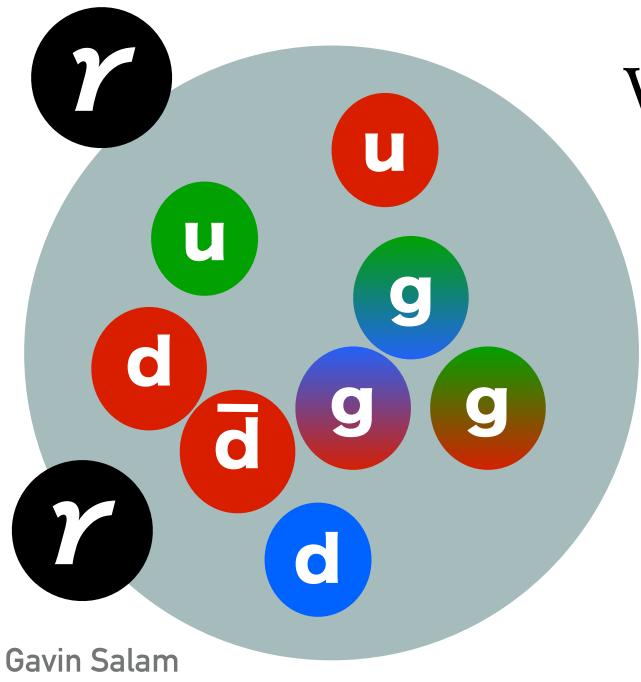
- at LHC

55

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Was largest uncertainty on HW[±] production

Solved for proton 2 years ago [Manohar, Nason, GPS & Zanderighi, Phys.Rev.Lett. 117 (2016) 242002 + JHEP 1712 (2017) 046]

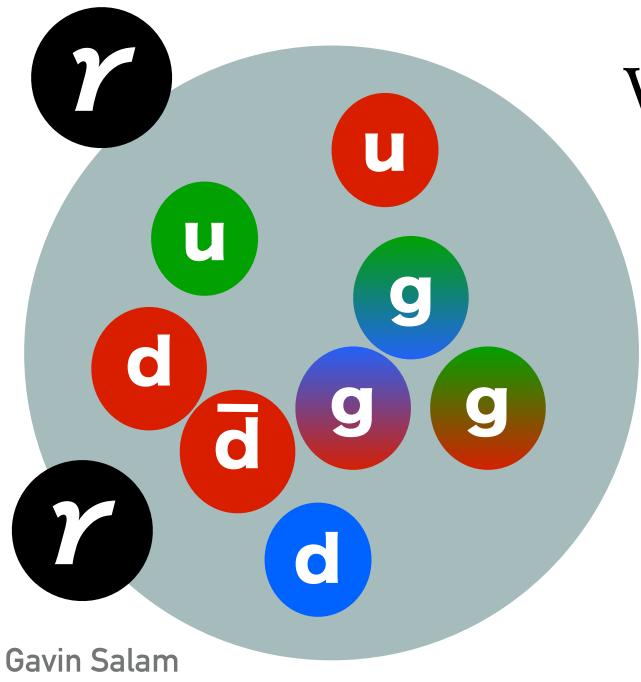
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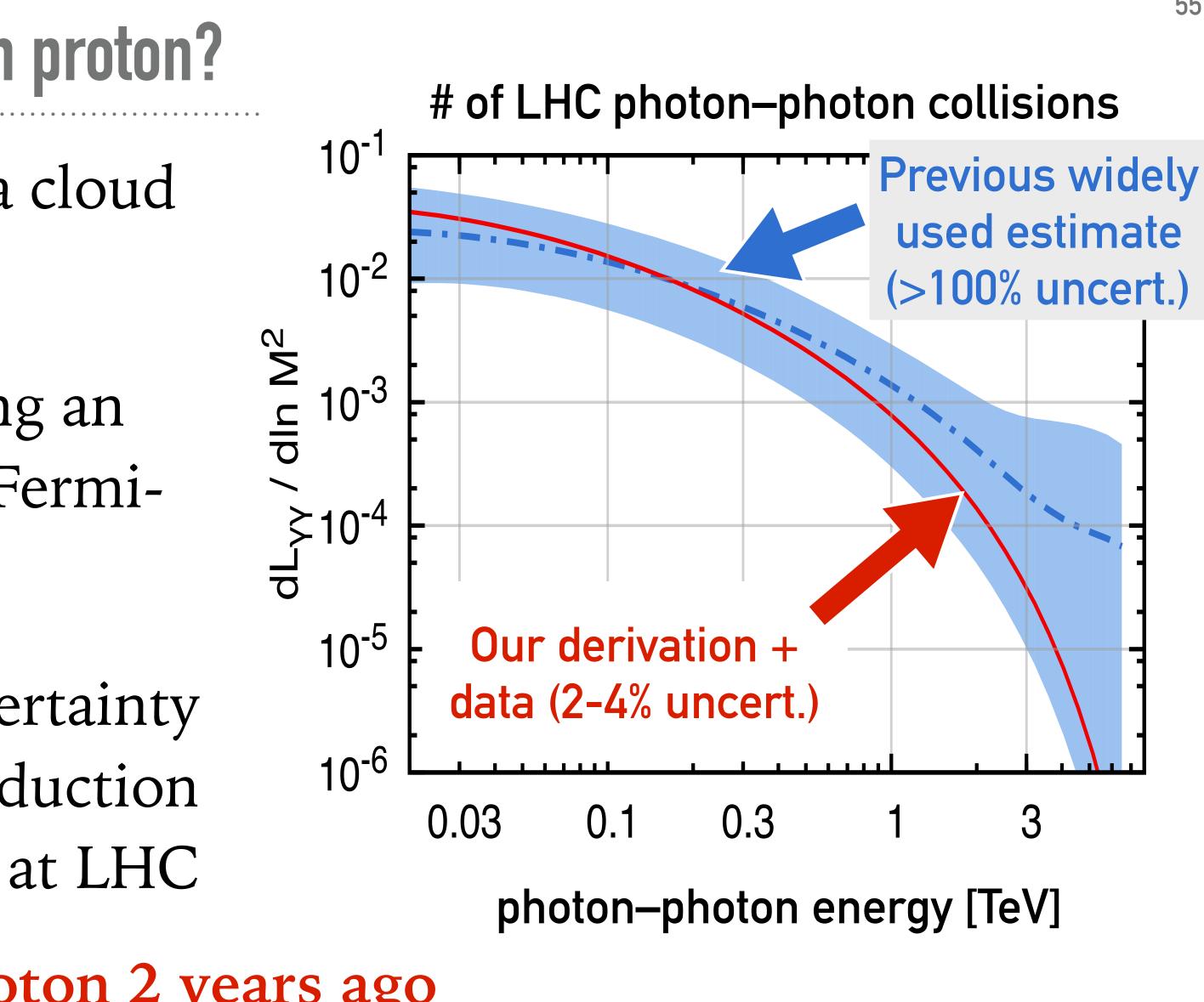
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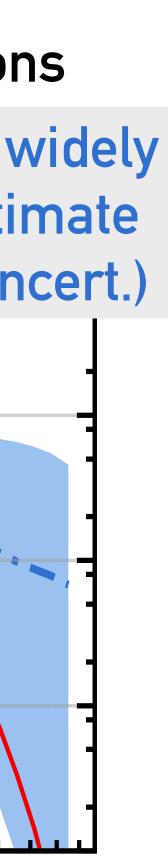
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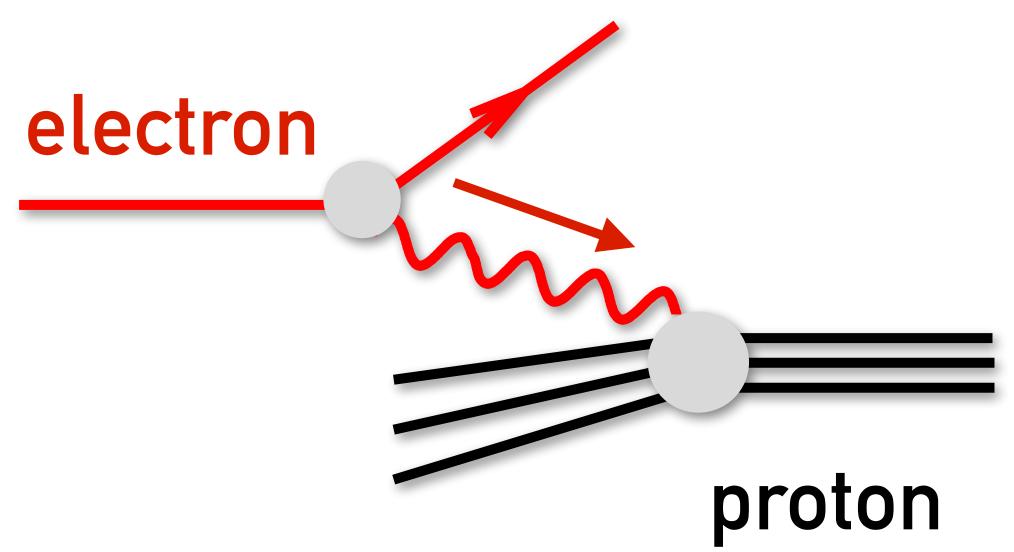


55



electron—proton scattering

- Experiments have been going on for decades
- Usually seen as photons from electron probing proton structure

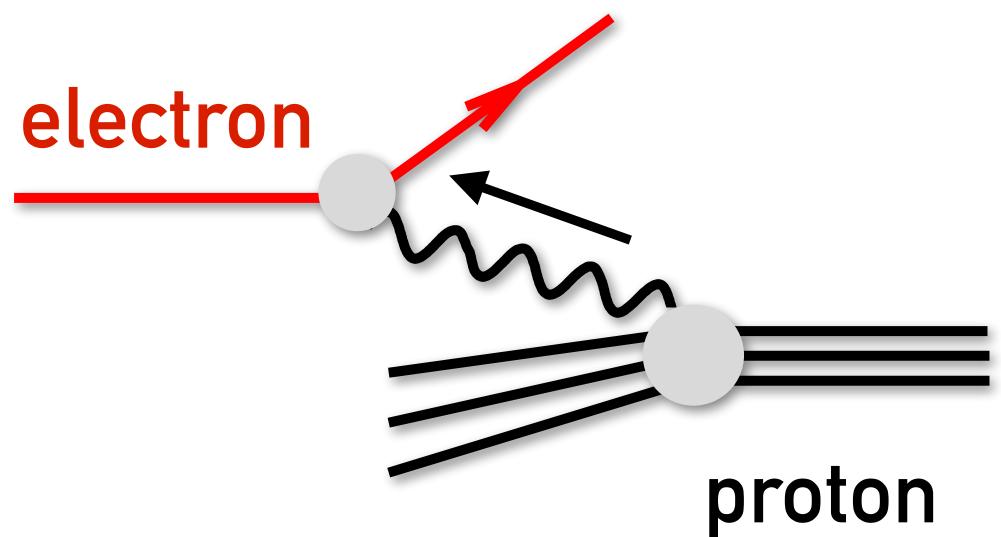


. . . .

56

electron—proton scattering

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- But can be viewed as electron probing proton's photonic field



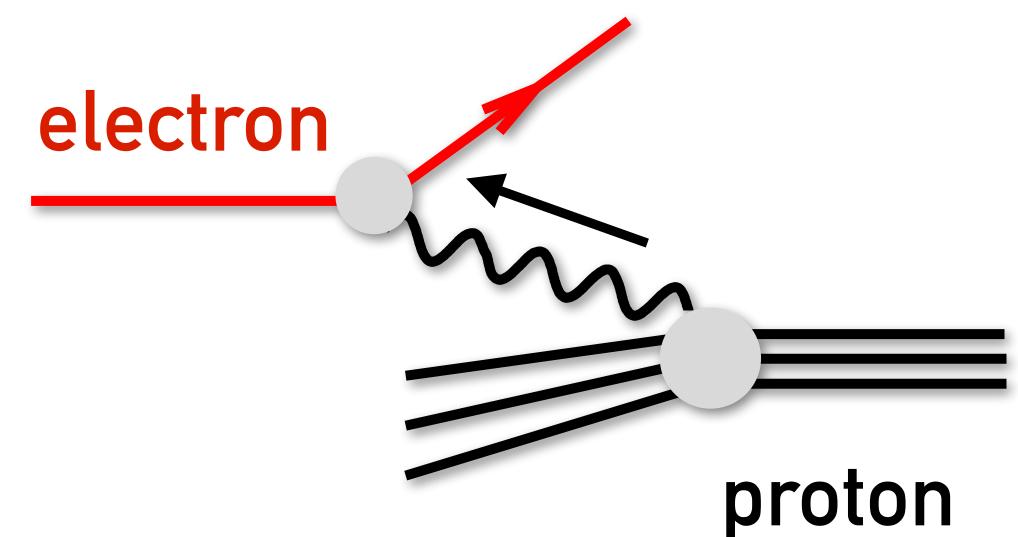
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electron-proton scattering

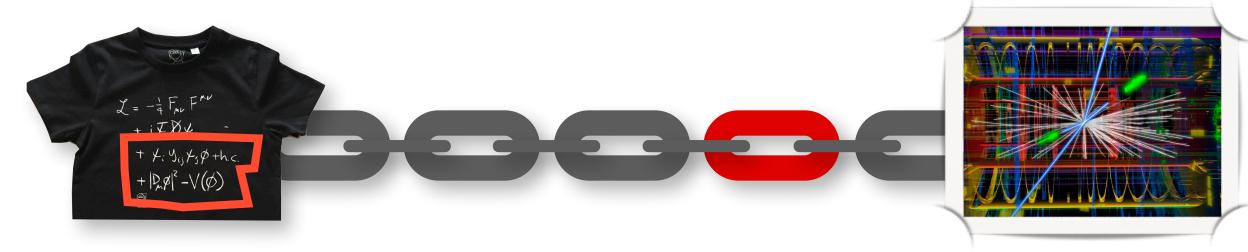
- Experiments have been going on for decades
- Usually seen as photons from electron probing proton structure
- But can be viewed as electron probing proton's photonic field
- Everything about electron-proton interaction encoded in two well measured "structure functions" $F_2(x,Q^2)$ & $F_L(x,Q^2)$

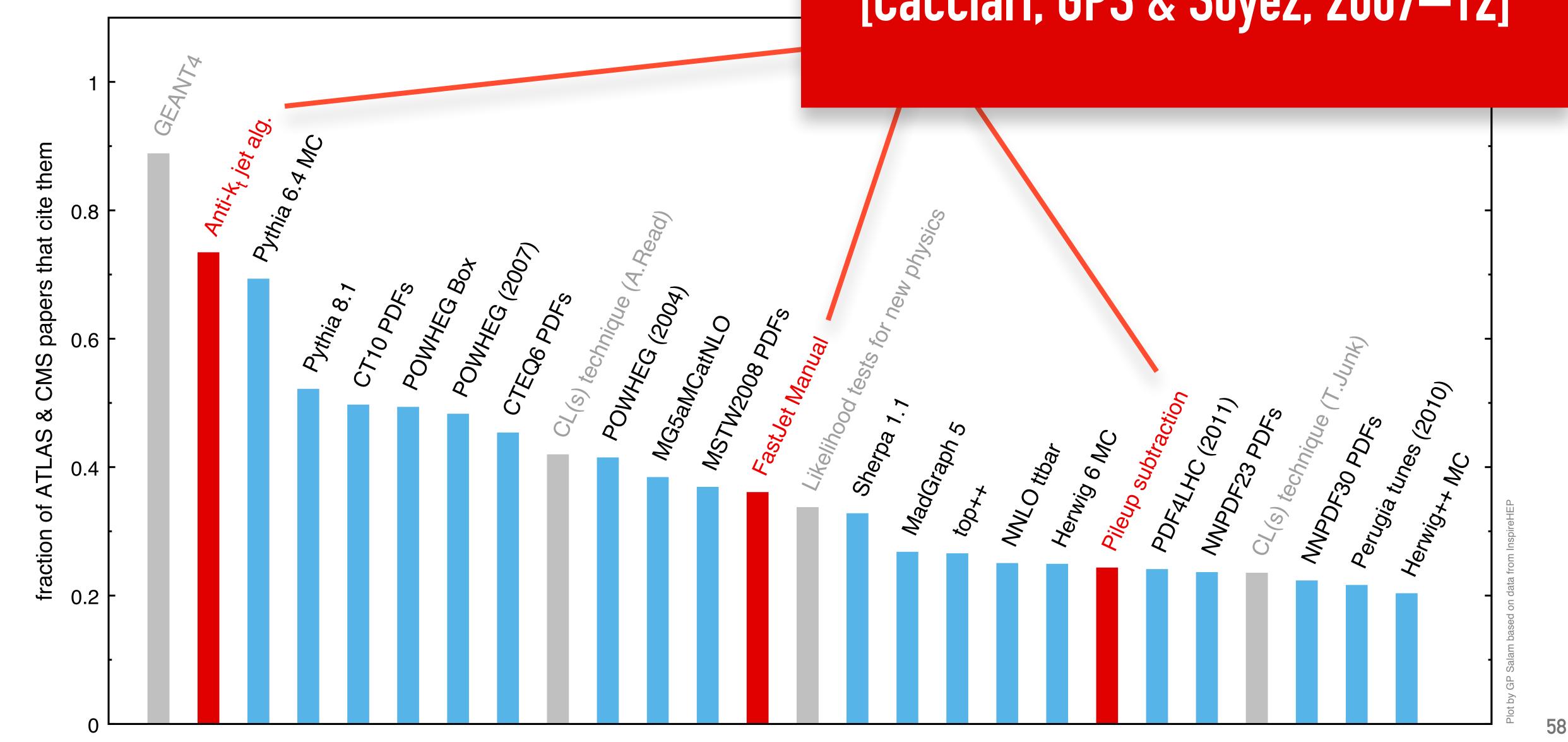
$$\frac{d\sigma}{dxdQ^2} = \frac{4\pi\alpha^2}{xQ^4} \left(\left(1 - y + \frac{y^2}{2} \left(1 + \frac{y^2}{2}\right)\right) \right) + \frac{y^2}{2} \left(1 + \frac{y^2}{2}\right) \left(1 + \frac{y^2}{2}\right) + \frac{y^2}{2} \left(1 + \frac{y^2}{2}\right)$$



 $+2x^2\frac{m_p^2}{O^2}\bigg)\bigg)F_2(x,Q^2)-\frac{y^2}{2}F_L(x,Q^2)\bigg)$







organising event information ("jets") [Cacciari, GPS & Soyez, 2007–12]

Pileup subtraction

Herwig 6 MC

MNKO ttbar

PDF4LHC (2011)

MNDDF33 PDFS

Cl(s) technique (

MNDDF30 PDFS

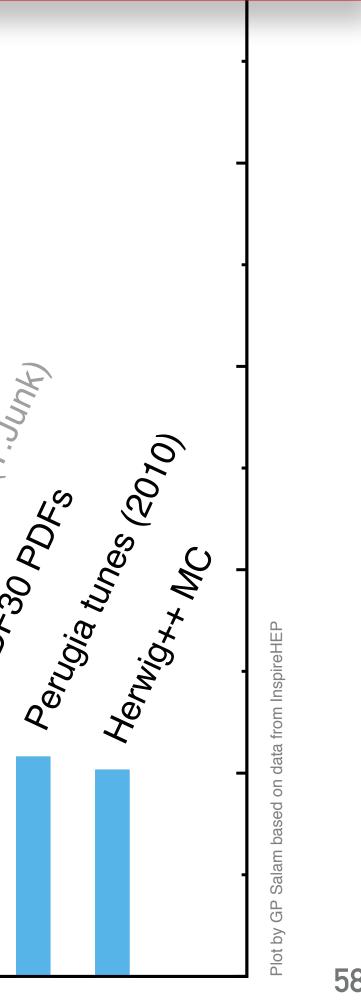
Kitelihood tests for hew physics

Sherba 1.1

MadGraph 5

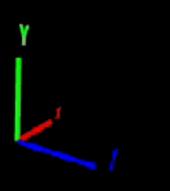
to0++

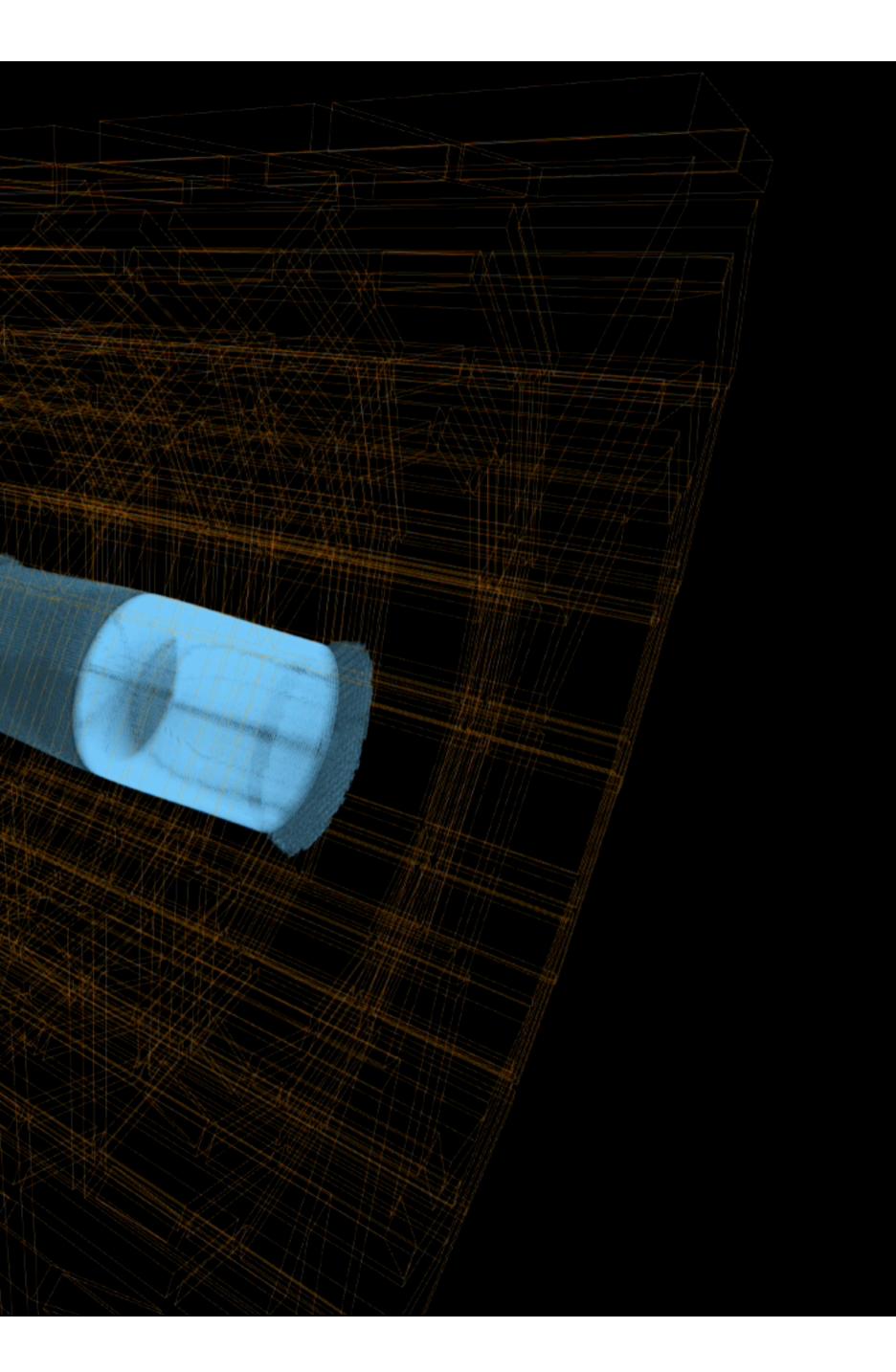
Fastuer Manual





CMS Experiment at the LHC, CERN Simulated event at 13 TeV centre-of-mass energy

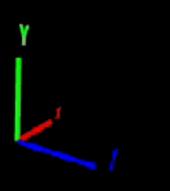


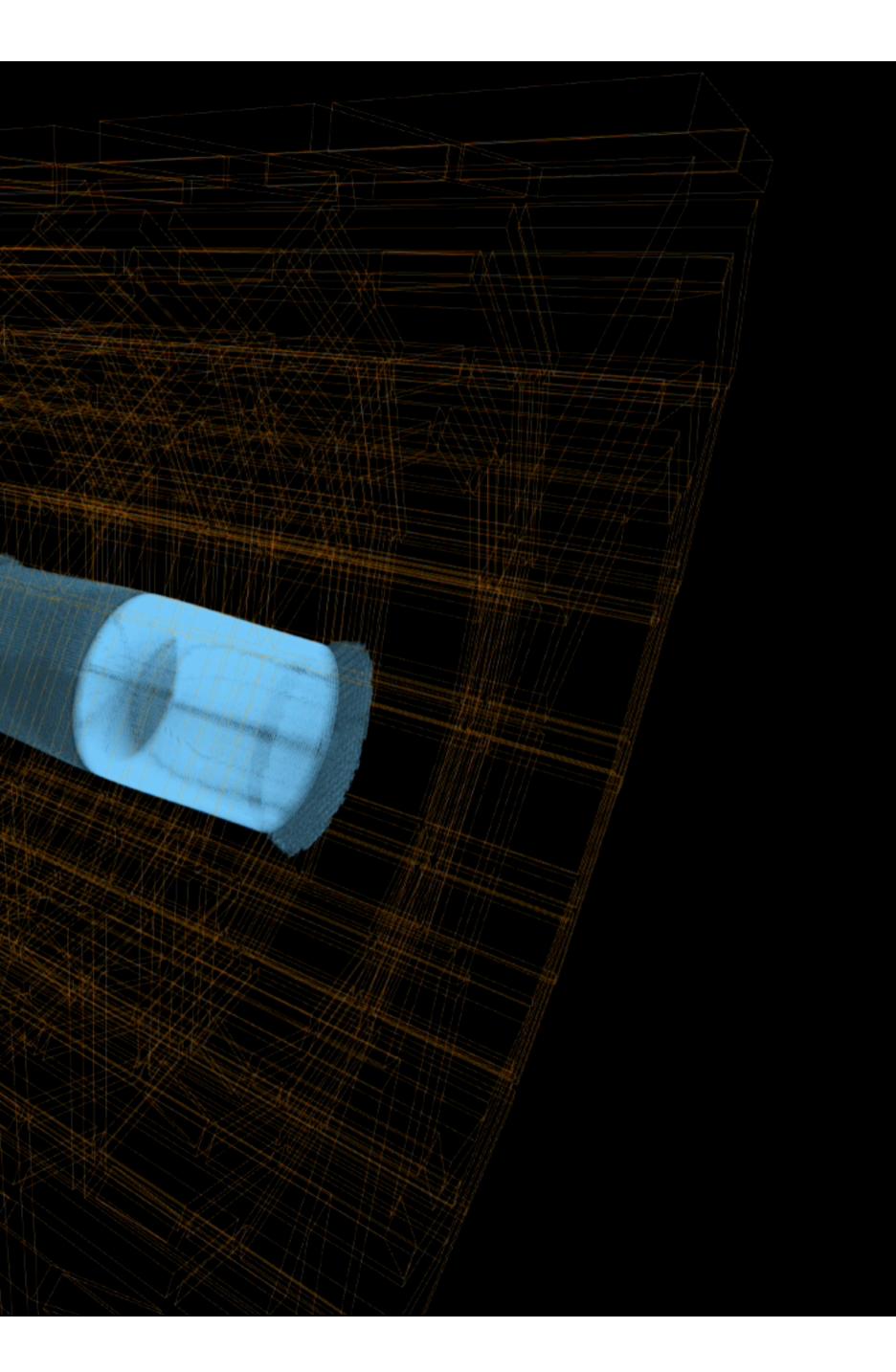


59

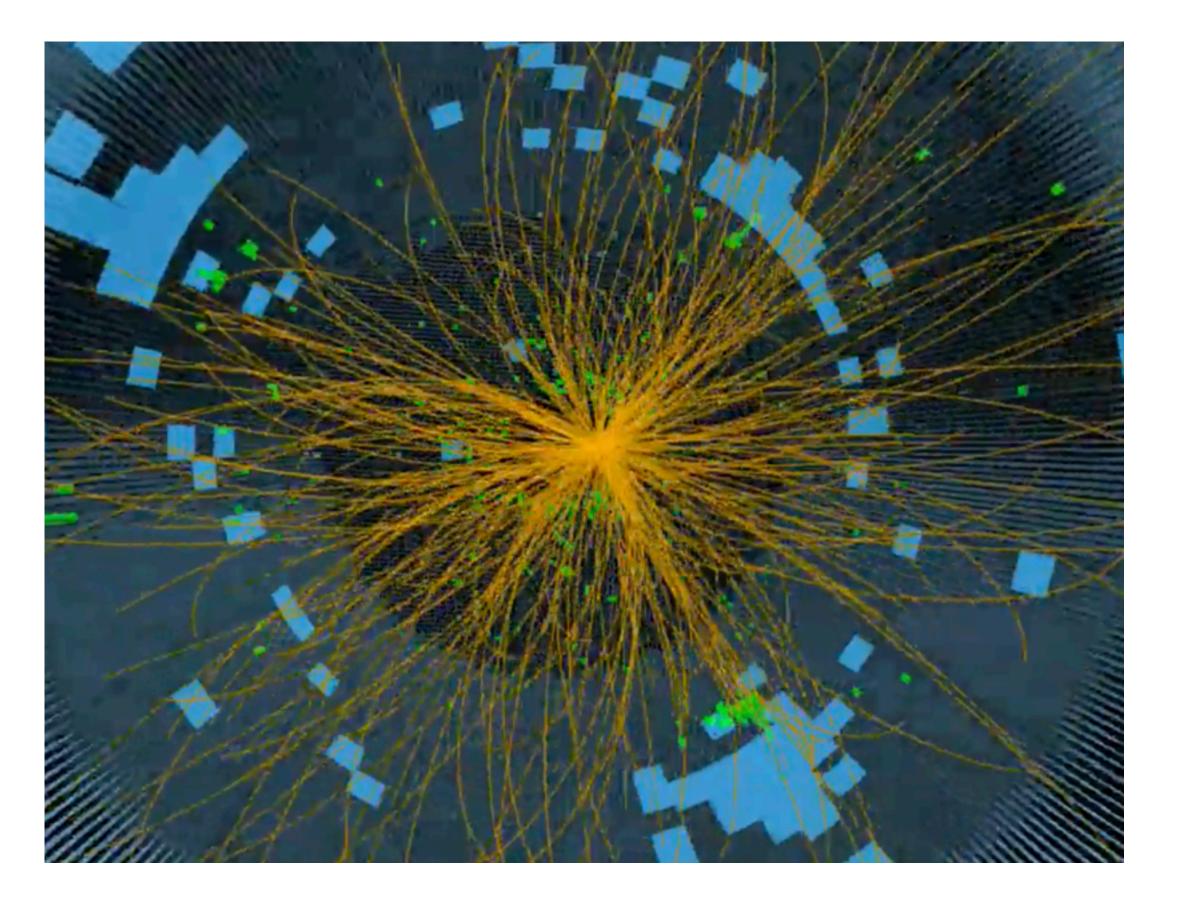


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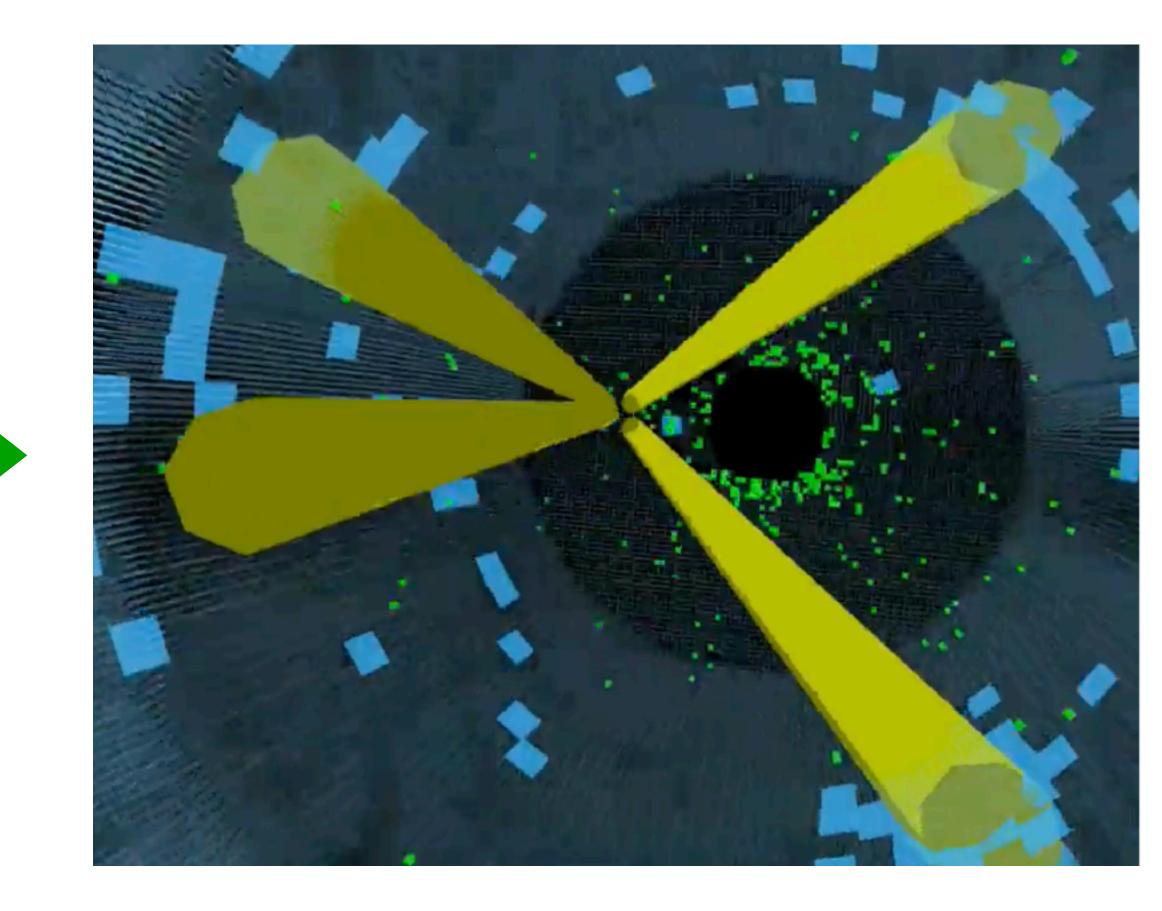


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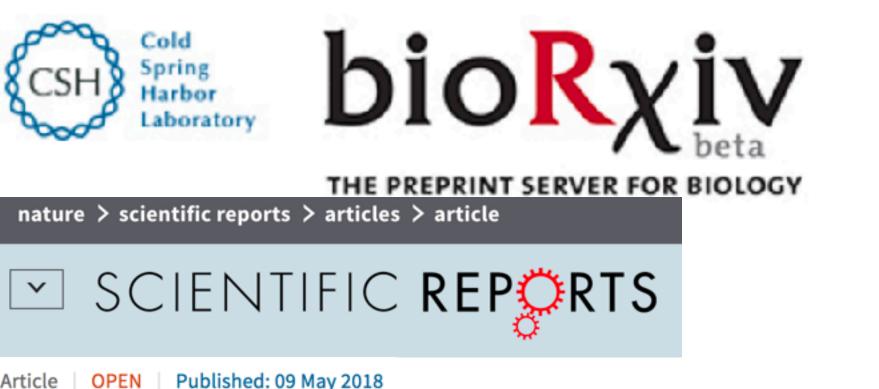


Anti-k_t algorithm, Cacciari, GPS & Soyez, 2008 (solved a 30-year old problem, first addressed by Sterman & Weinberg)

FastJet program, Cacciari, GPS & Soyez, 2005 – 18 FastJet contrib, ~ 20 contributors, 2013 – 18







For identifying spatial clusters, we have implemented both centroidlinkage hierarchical clustering using FastJet [...]

Via the qSR software, FastJet can analyze a typical super-resolution dataset within a few seconds. By storing the full tree structure, the user can quickly re-cluster data and compare the resulting clusters at varying characteristic sizes.

OPEN | Published: 09 May 2018 Article

qSR: A software for quantitative analysis of single molecule and super-resolution data

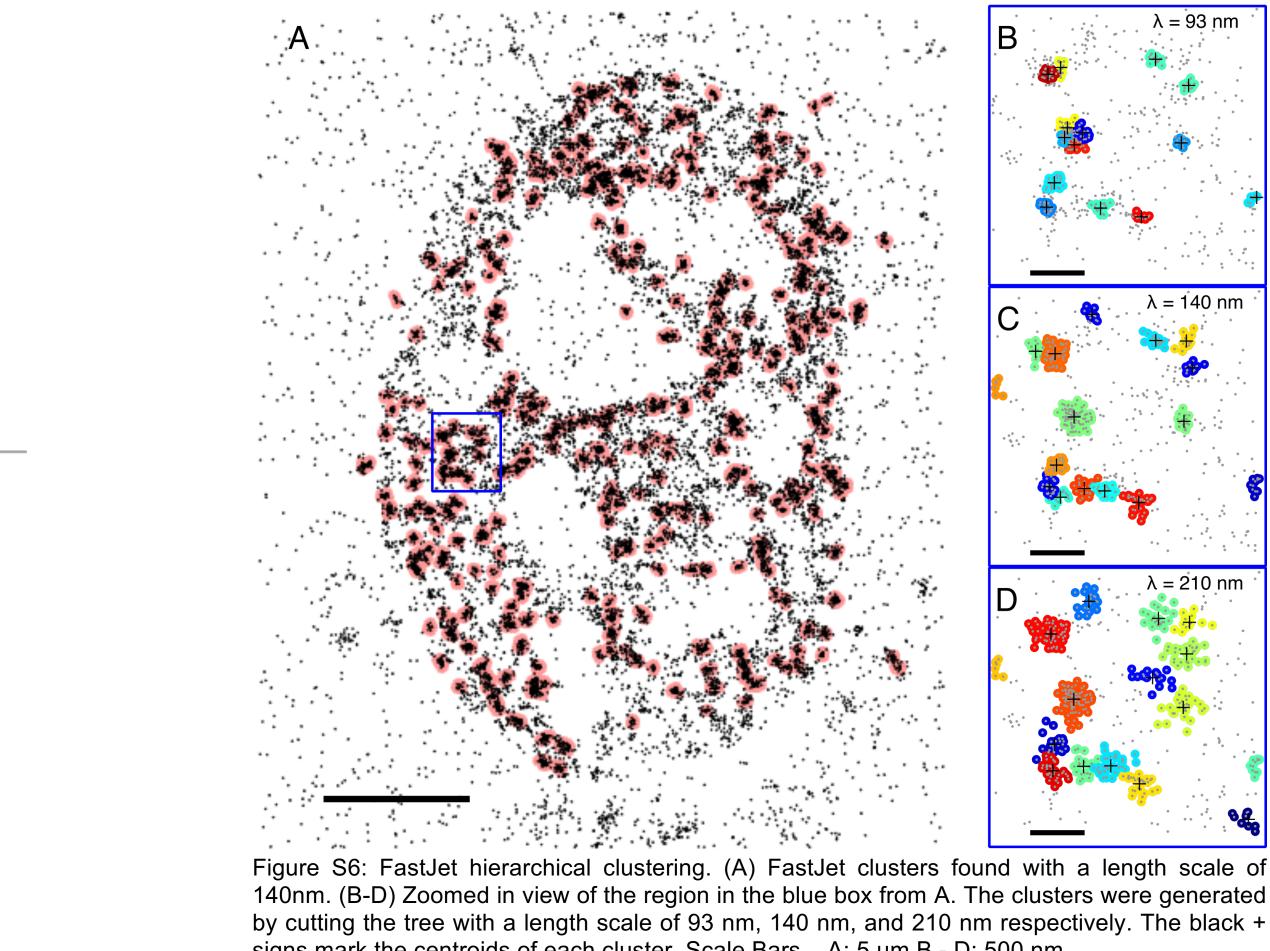
J. Owen Andrews, Arjun Narayanan, Jan-Hendrik Spille, Won-Ki Cho, Jesse D. Thaler, Ibrahim I. Cisse

doi: https://doi.org/10.1101/146241

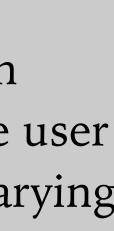
Abstract	Info/History	Metrics	Data Supplements	Preview PDF
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Abstract

We present a software for quantitative analysis of single molecule based super-resolution data. The software serves as an opensource platform integrating multiple algorithms for rigorous spatial and temporal characterizations of protein clusters in superresolution data of living, or fixed cells.

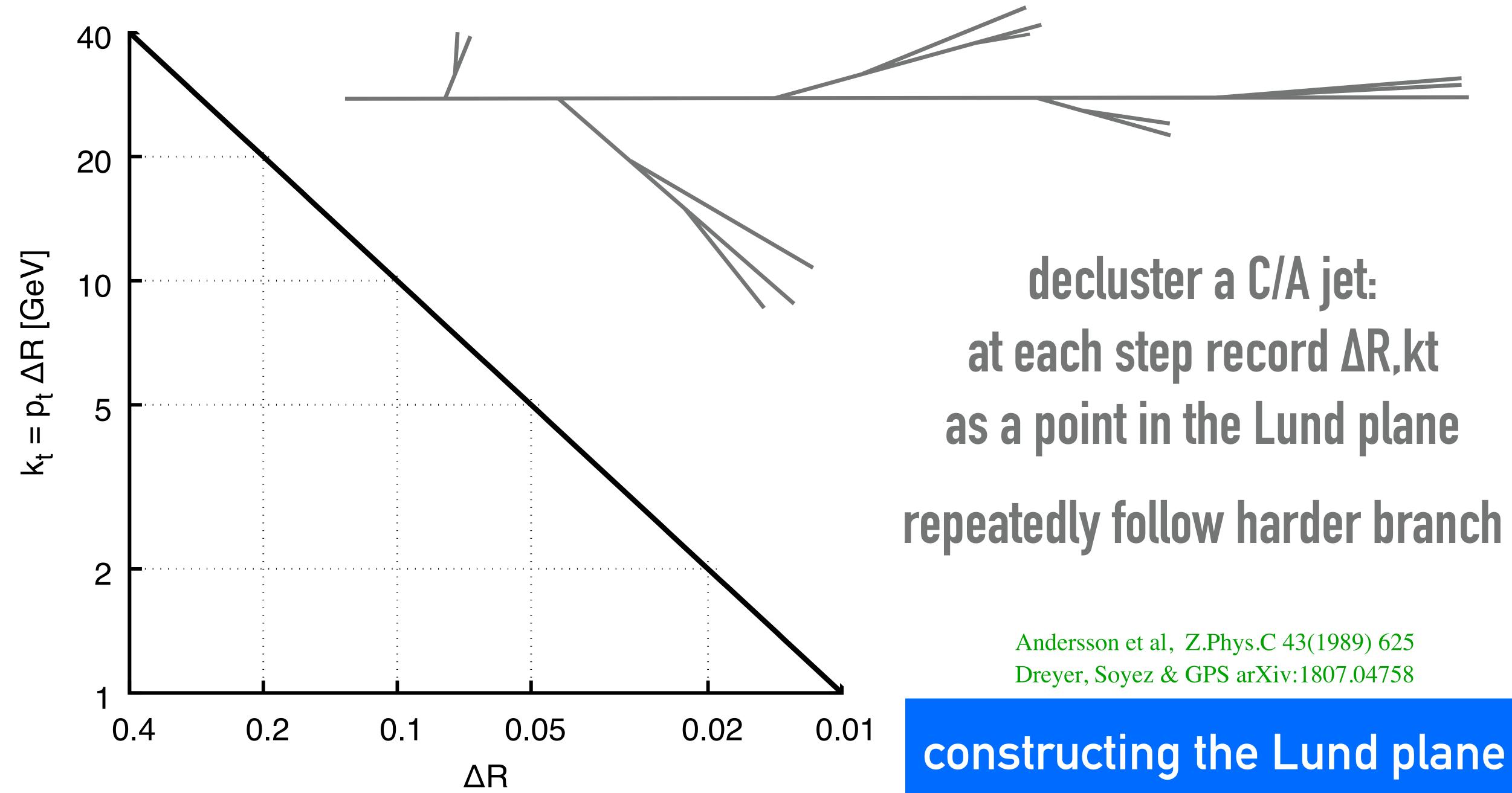


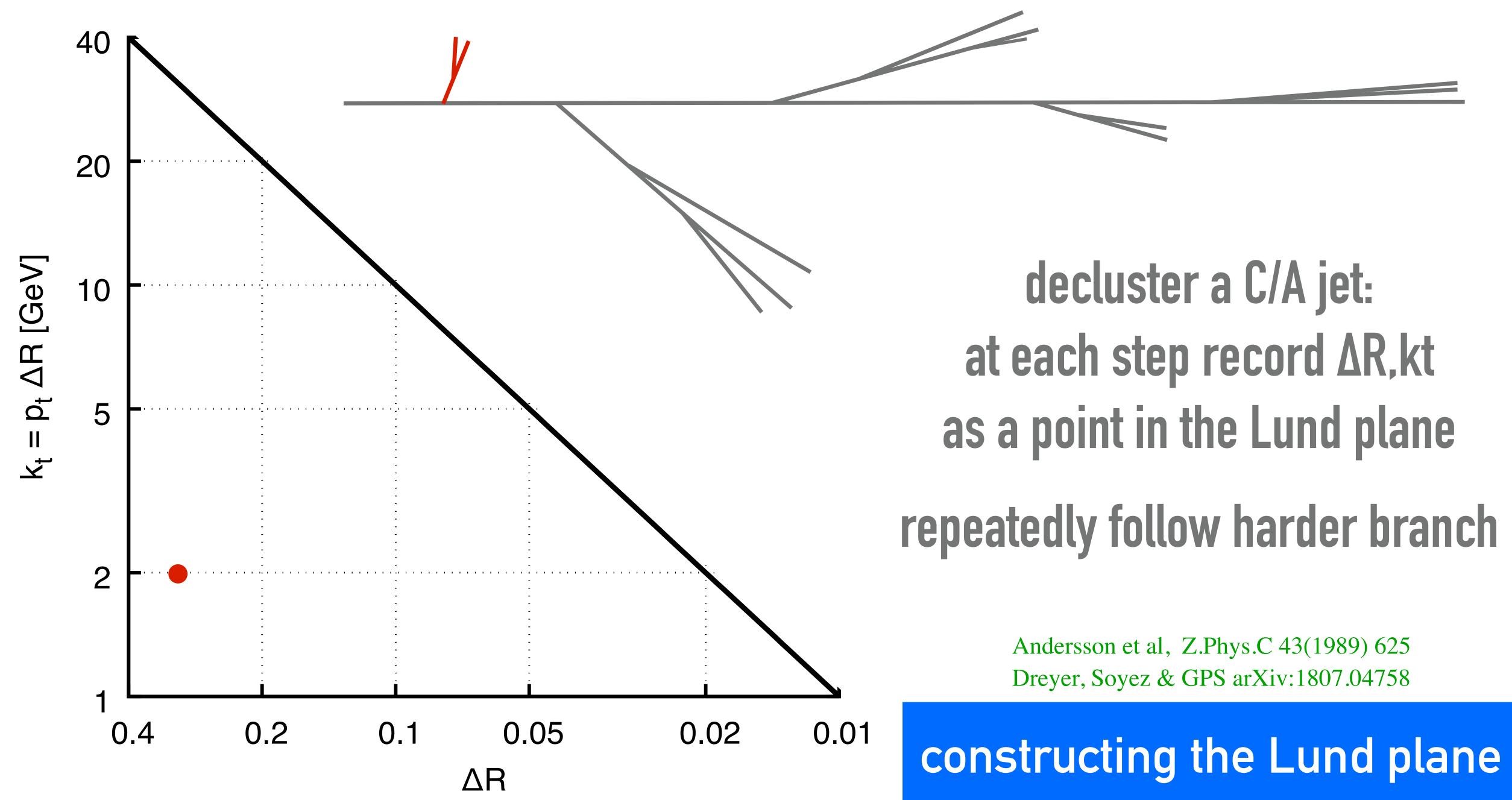
signs mark the centroids of each cluster. Scale Bars – A: 5 µm B - D: 500 nm

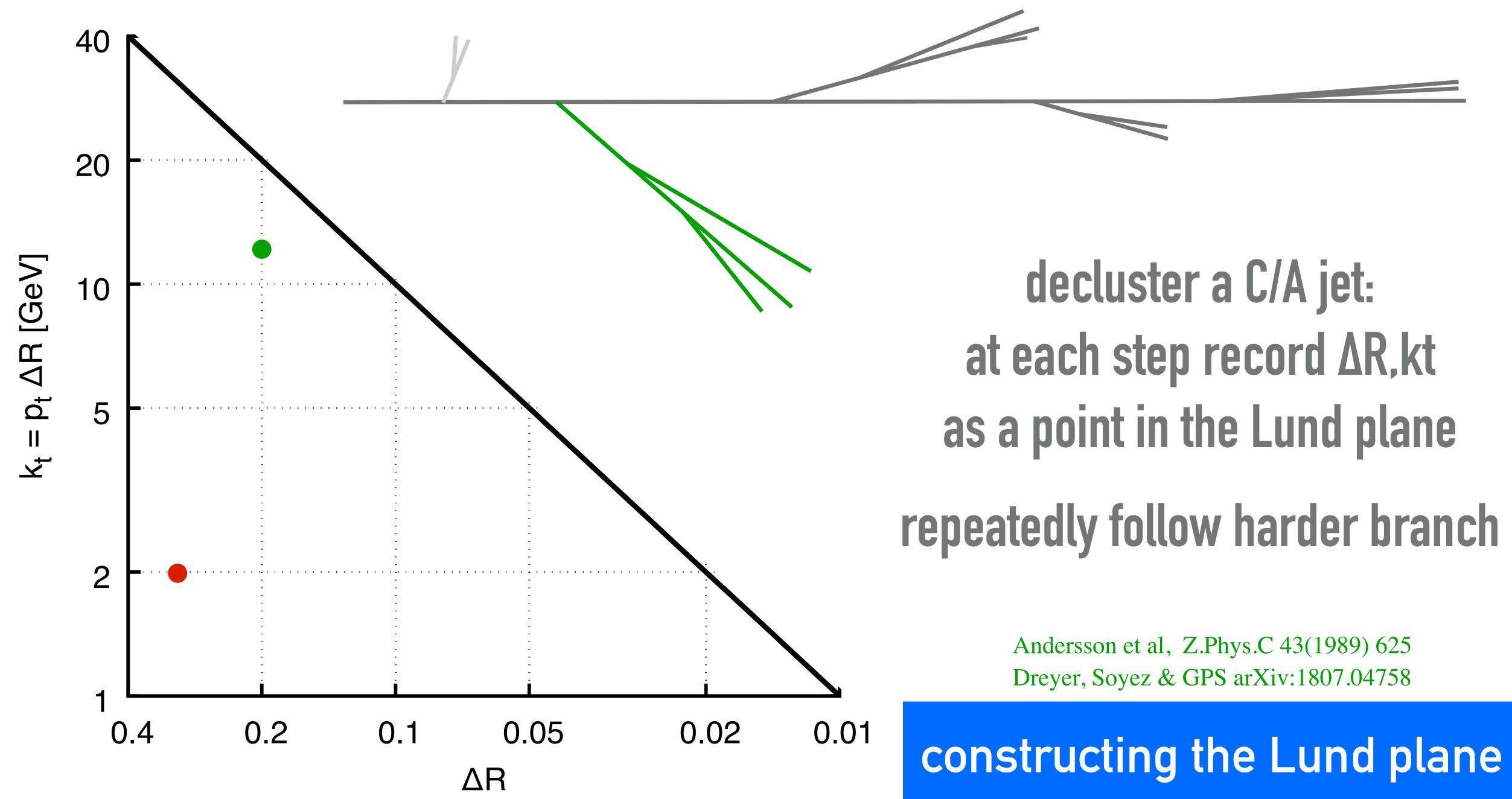


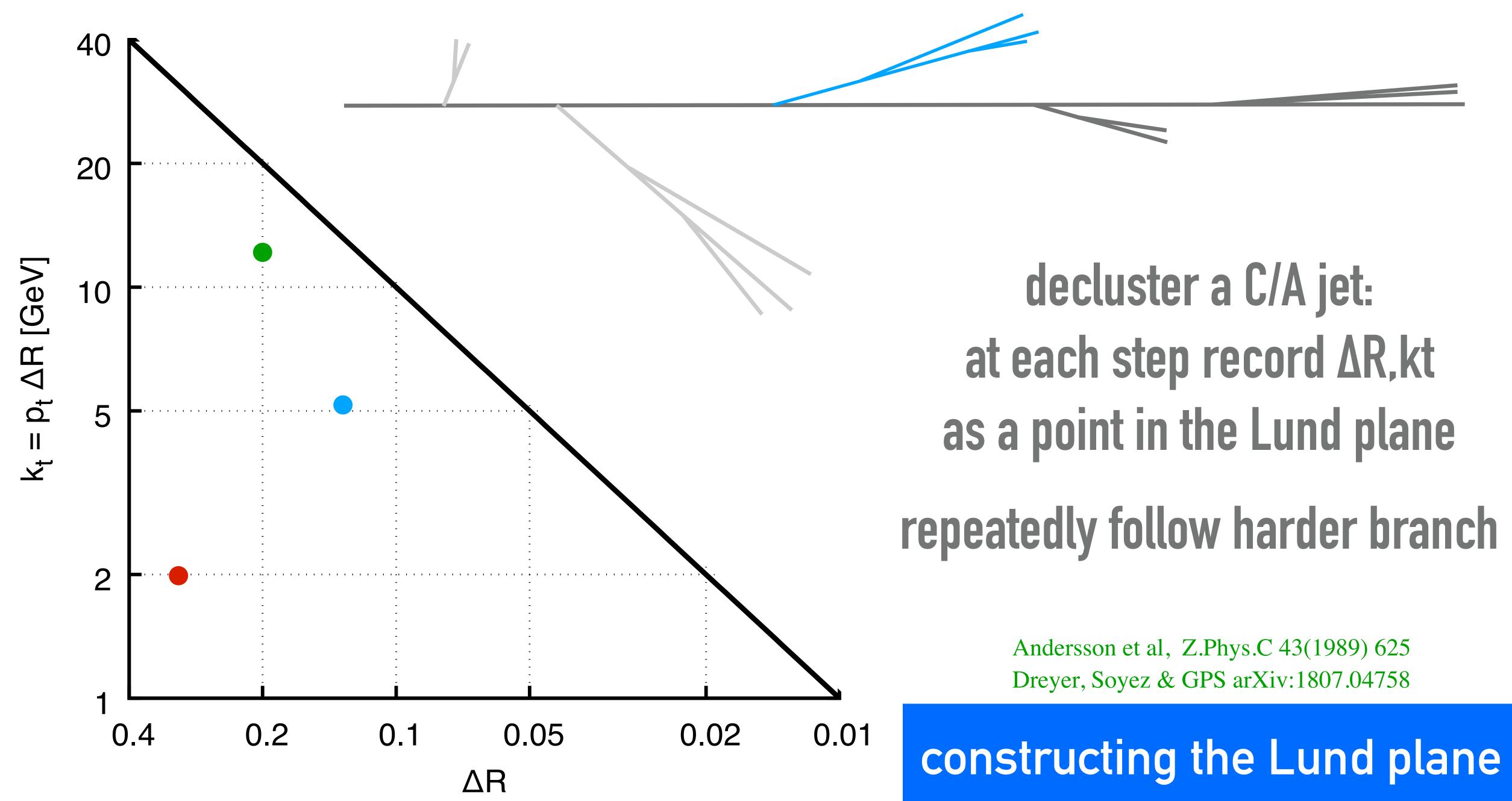


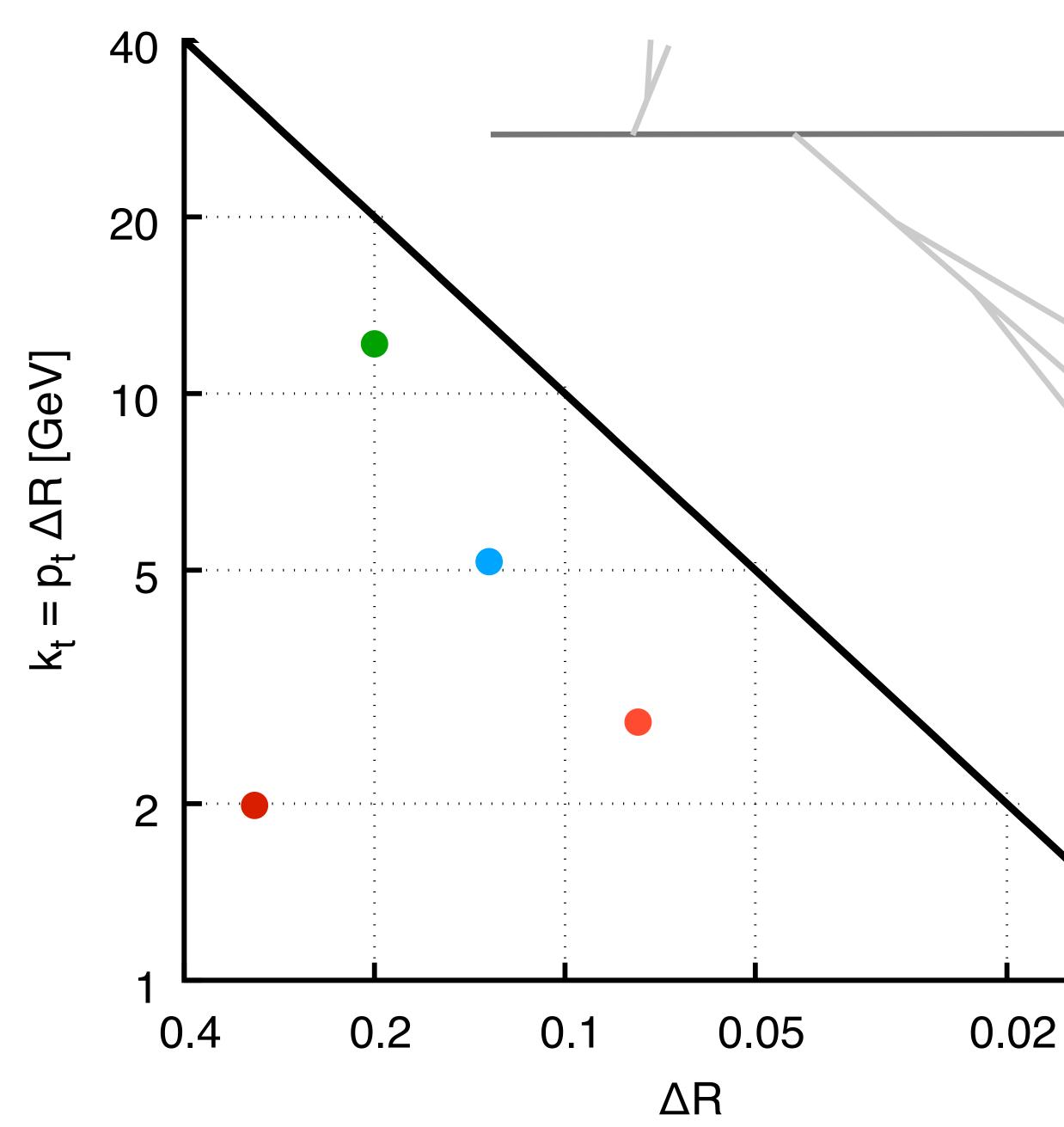
jet with R = 0.4, $p_t = 200 \text{ GeV}$









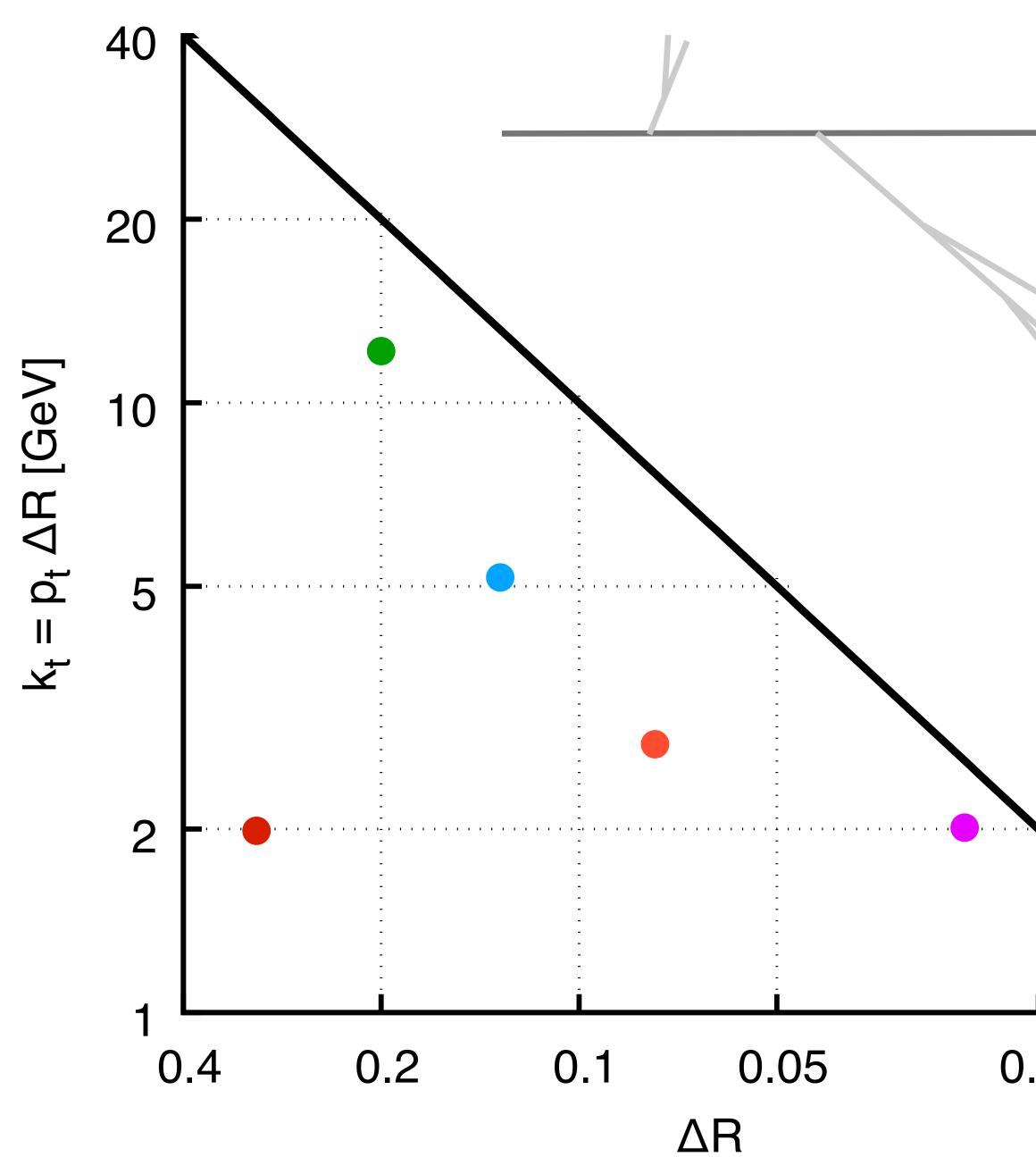


decluster a C/A jet: at each step record ΔR,kt as a point in the Lund plane repeatedly follow harder branch Andersson et al, Z.Phys.C 43(1989) 625 Dreyer, Soyez & GPS arXiv:1807.04758 0.01

constructing the Lund plane







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Andersson et al, Z.Phys.C 43(1989) 625 Dreyer, Soyez & GPS arXiv:1807.04758

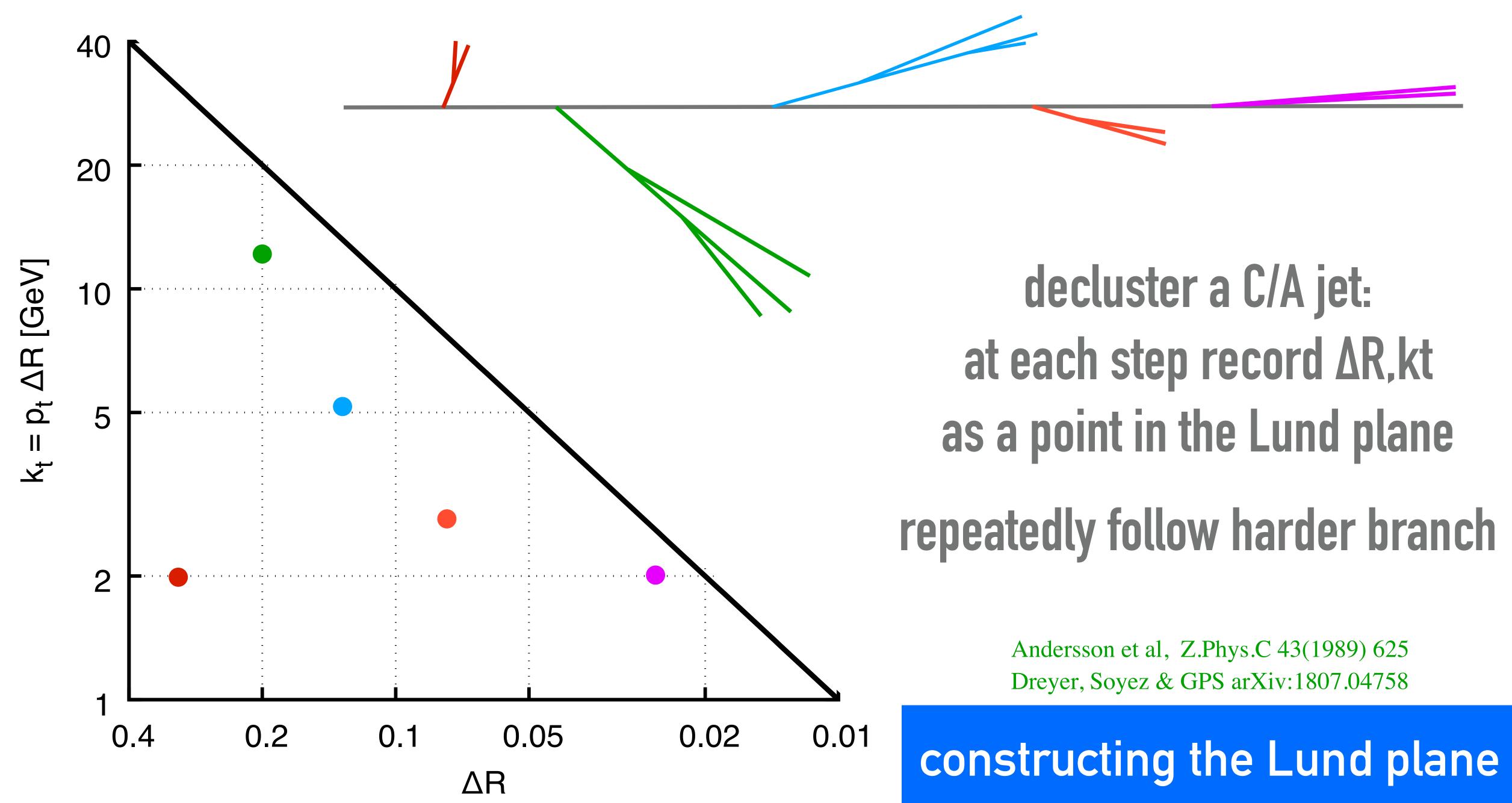
constructing the Lund plane

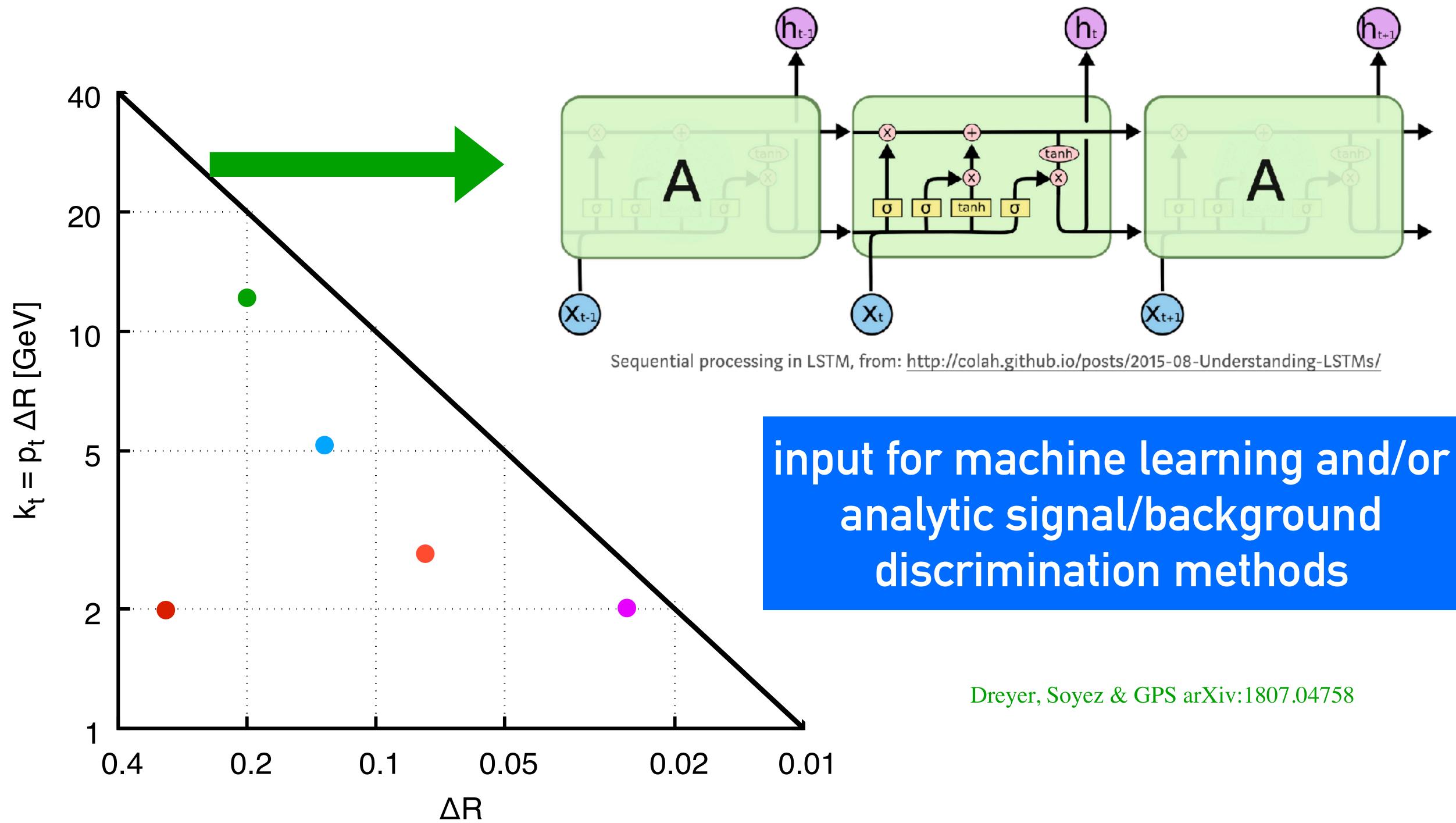
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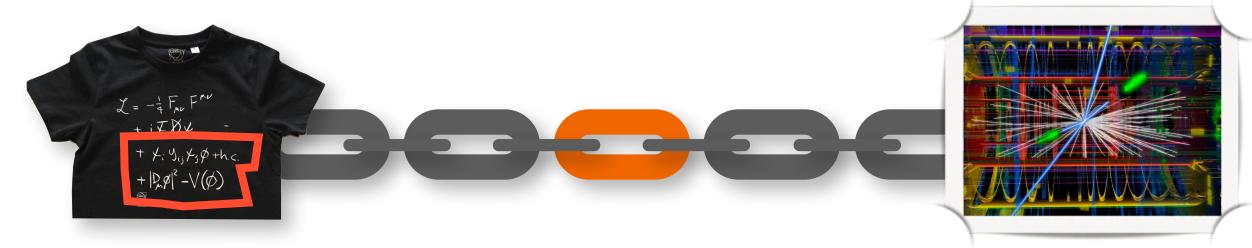
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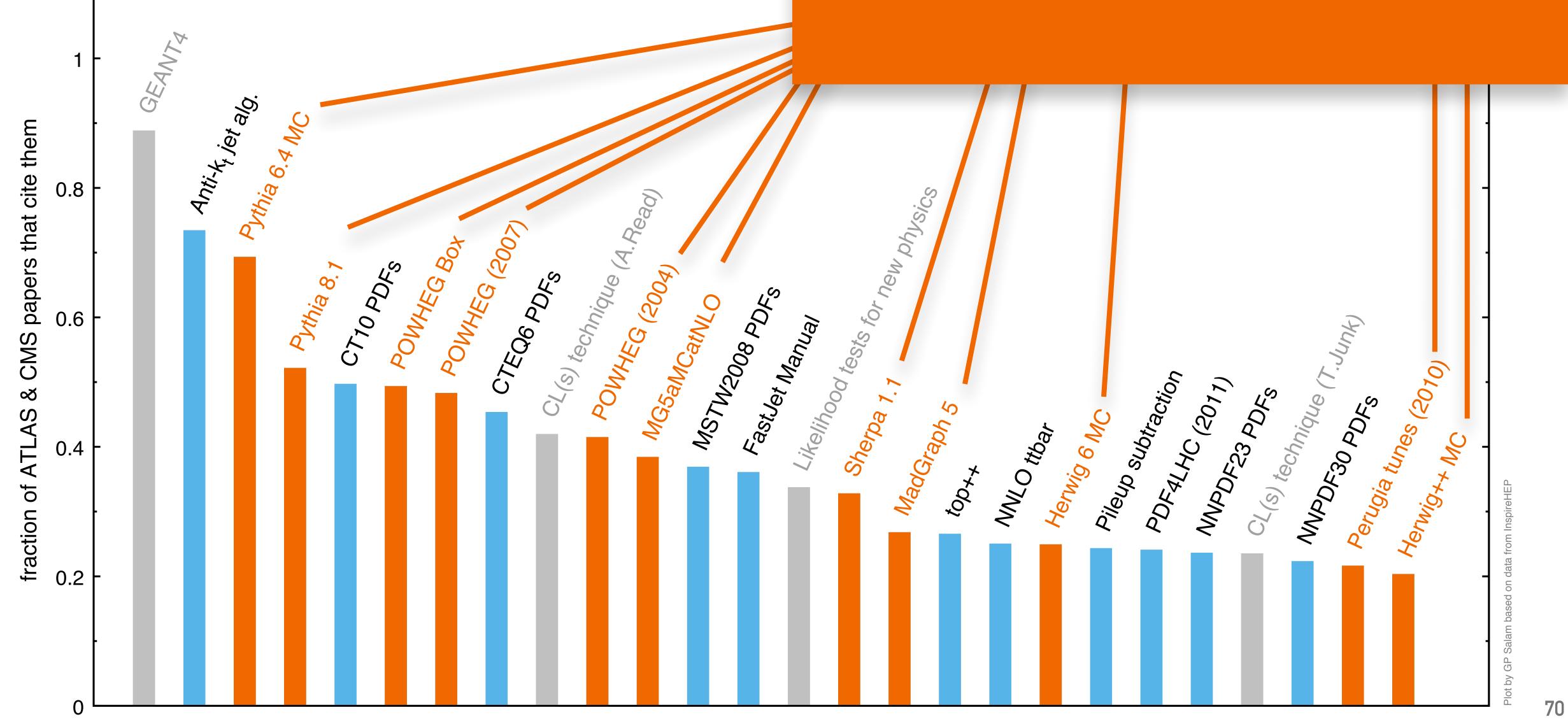








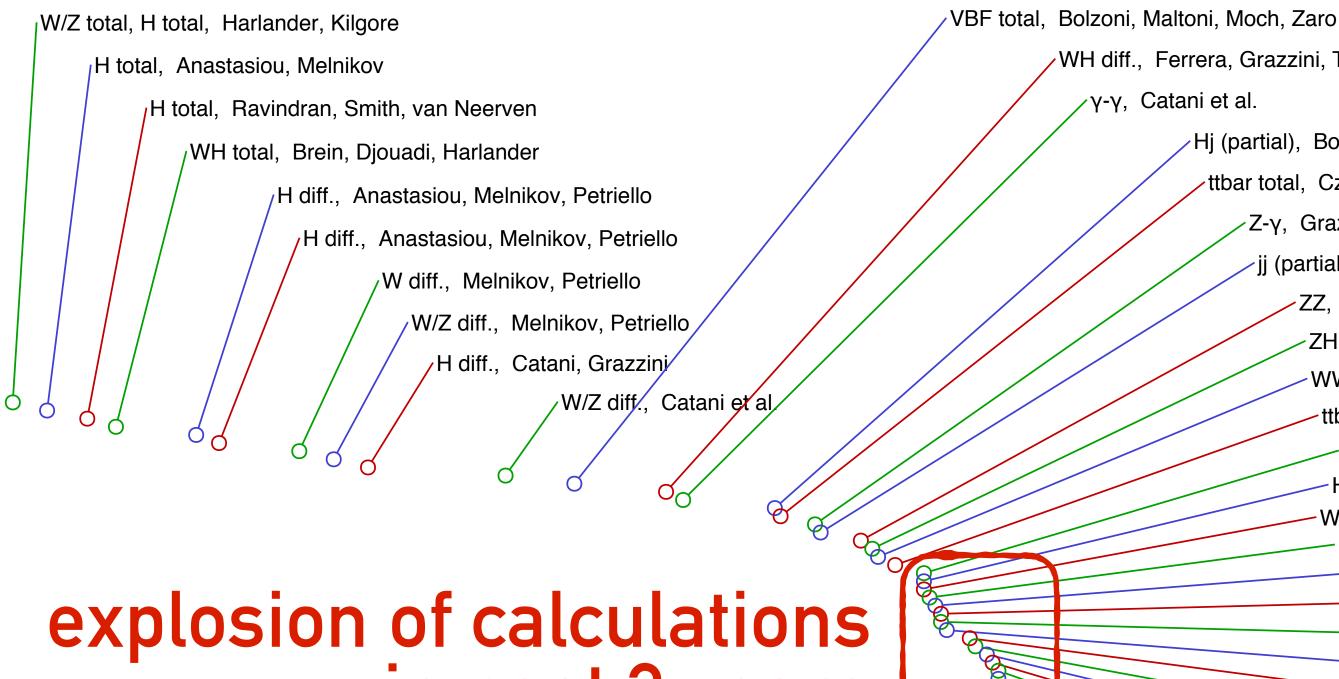




predicting full particle structure that comes out of a collision



Calculations to 3rd order (NNLO) in perturbation theory strong coupling constant (α_s)



explosion of calculations in past 3 years

2010 2006 2008 2012 2014 2002 2004 2016

as of April 2017

WH diff., Ferrera, Grazzini, Tramontano

γ-γ, Catani et al.

Hj (partial), Boughezal et al.

ttbar total, Czakon, Fiedler, Mitov

Z-γ, Grazzini, Kallweit, Rathlev, Torre

jj (partial), Currie, Gehrmann-De Ridder, Glover, Pires

ZZ, Cascioli it et al.

ZH diff., Ferrera, Grazzini, Tramontano

-WW, Gehrmann et al.

-ttbar diff., Czakon, Fiedler, Mitov

-Z-γ, W-γ, Grazzini, Kallweit, Rathlev

Hj, Boughezal et al.

Boughezal, Focke, Liu, Petriello

Hj, Boughezal et al.

VBF diff., Cacciari et al.

Zj, Gehrmann-De Ridder et al.

ZZ, Grazzini, Kallweit, Rathlev

Hj, Caola, Melnikov, Schulze

Zj, Boughezal et al.

WH diff., ZH diff., Campbell, Ellis, Williams

γ-γ, Campbell, Ellis, Li, Williams

WZ, Grazzini, Kallweit, Rathlev, Wiesemann

WW, Grazzini et al.

MCFM at NNLO, Boughezal et al.

p_{tZ}, Gehrmann-De Ridder et al.

MCFM at NNLO, Berger, Gao, C.-Yuan, Zhu

MCFM at NNLO, de Florian et al.

ptH, MCFM at NNLO, Chen et al.

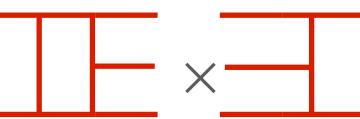
p_{tZ}, Gehrmann-De Ridder et al.

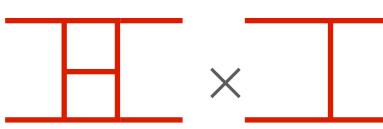
jj, Currie, Glover, Pires

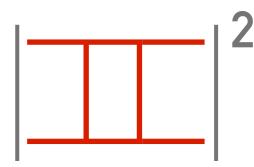
YX, Campbell, Ellis, Williams

γj, Campbell, Ellis, Williams



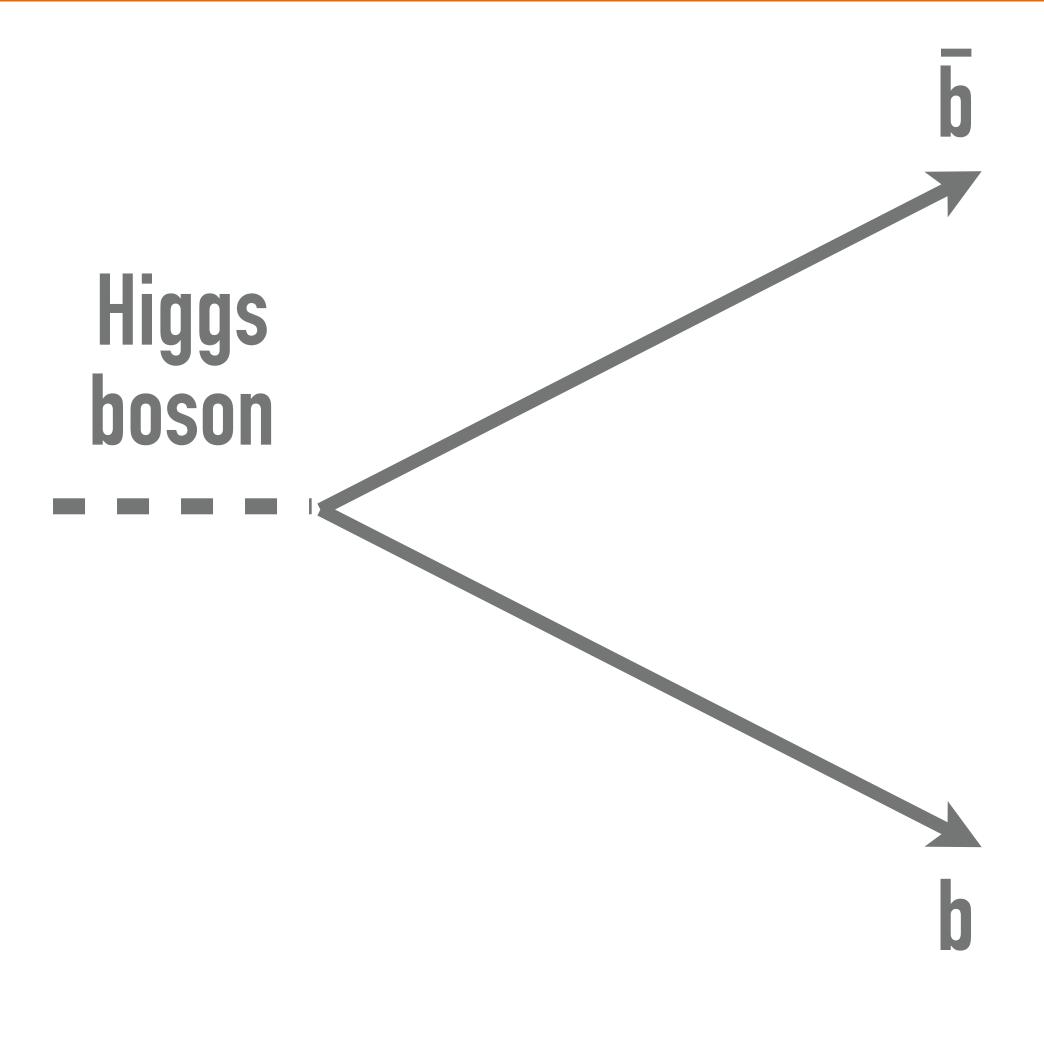


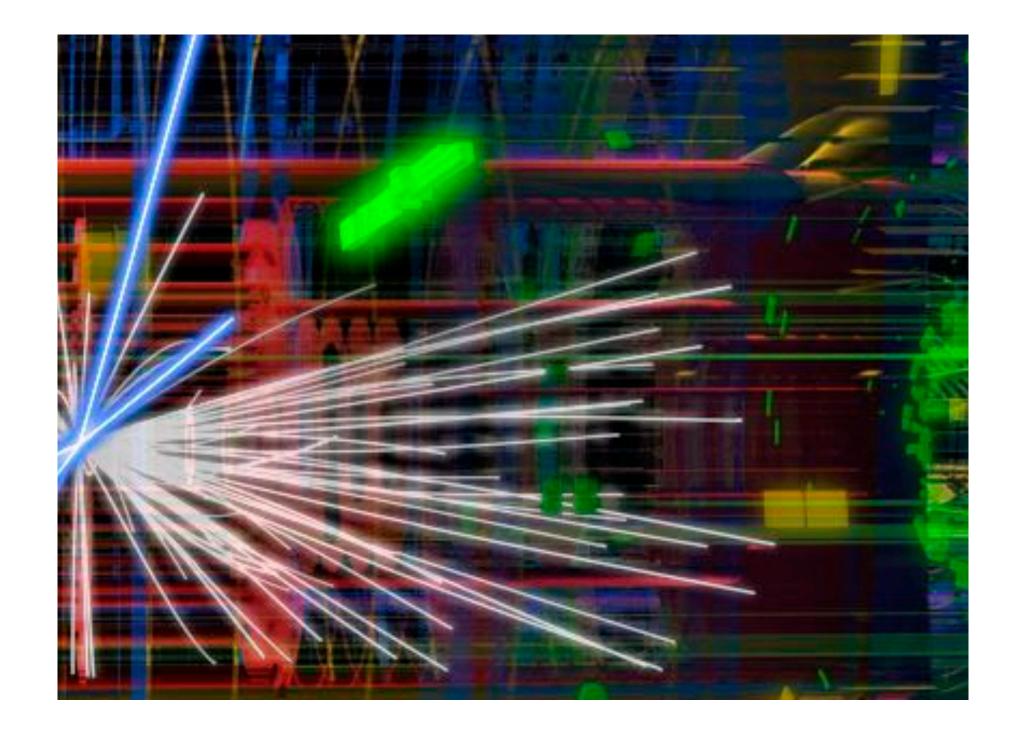






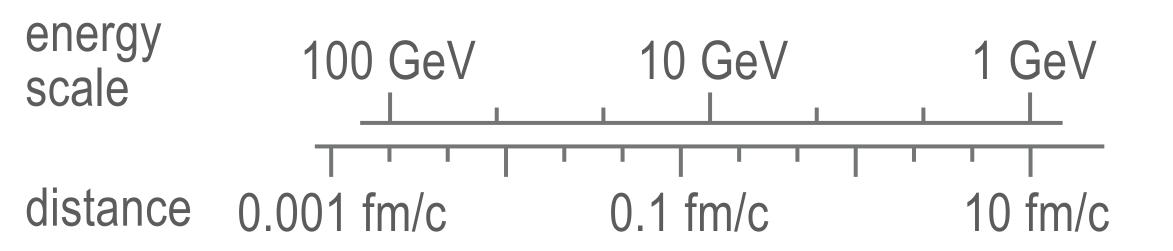
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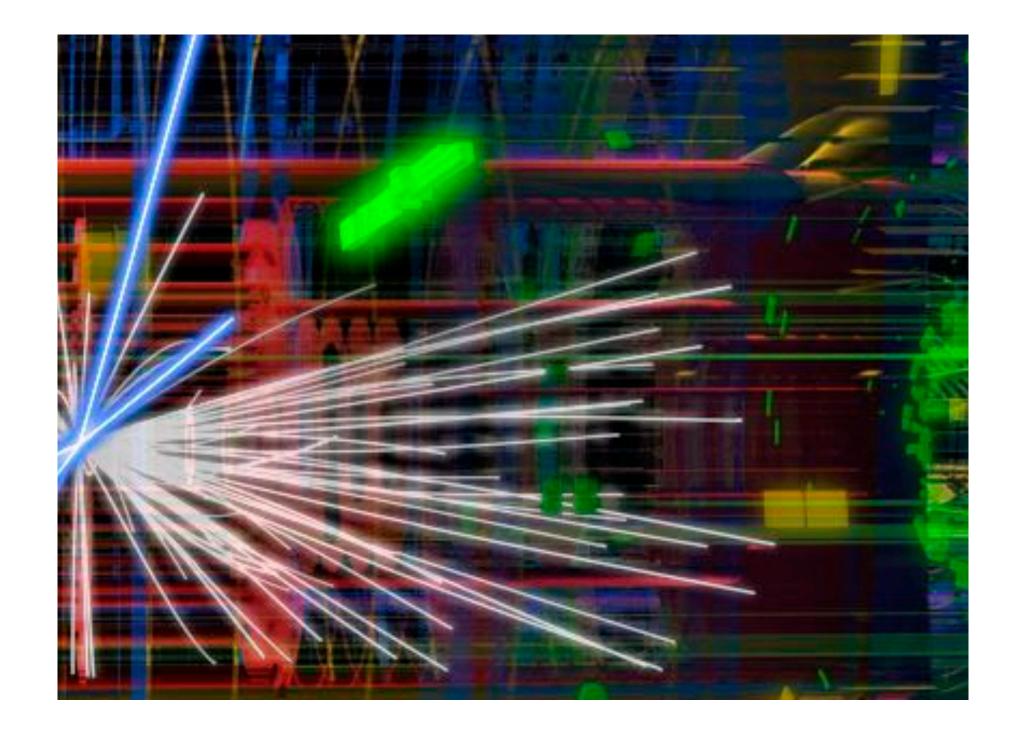




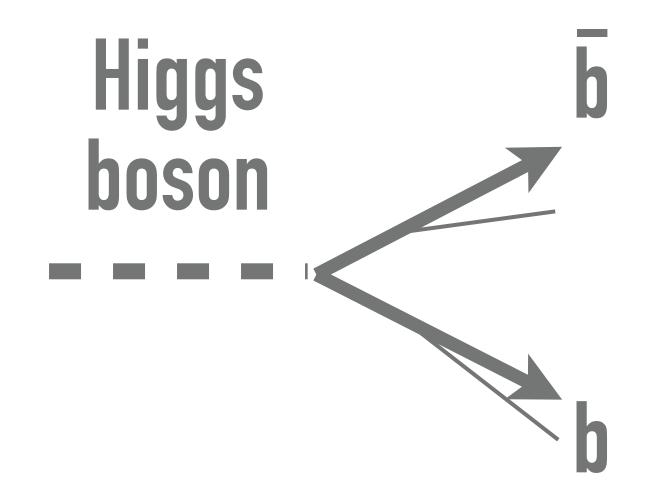


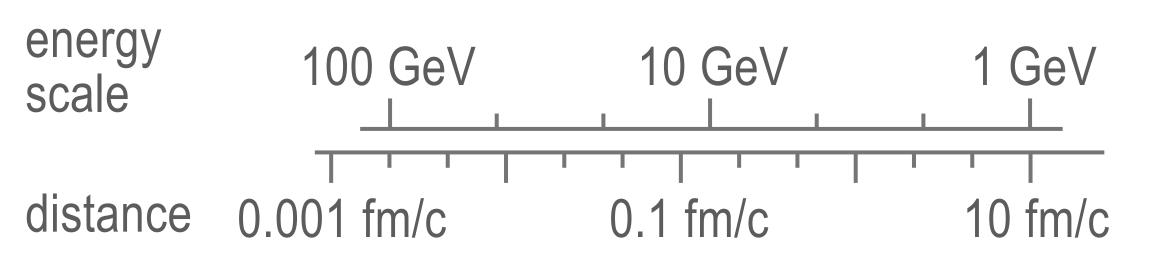
Higgs boson

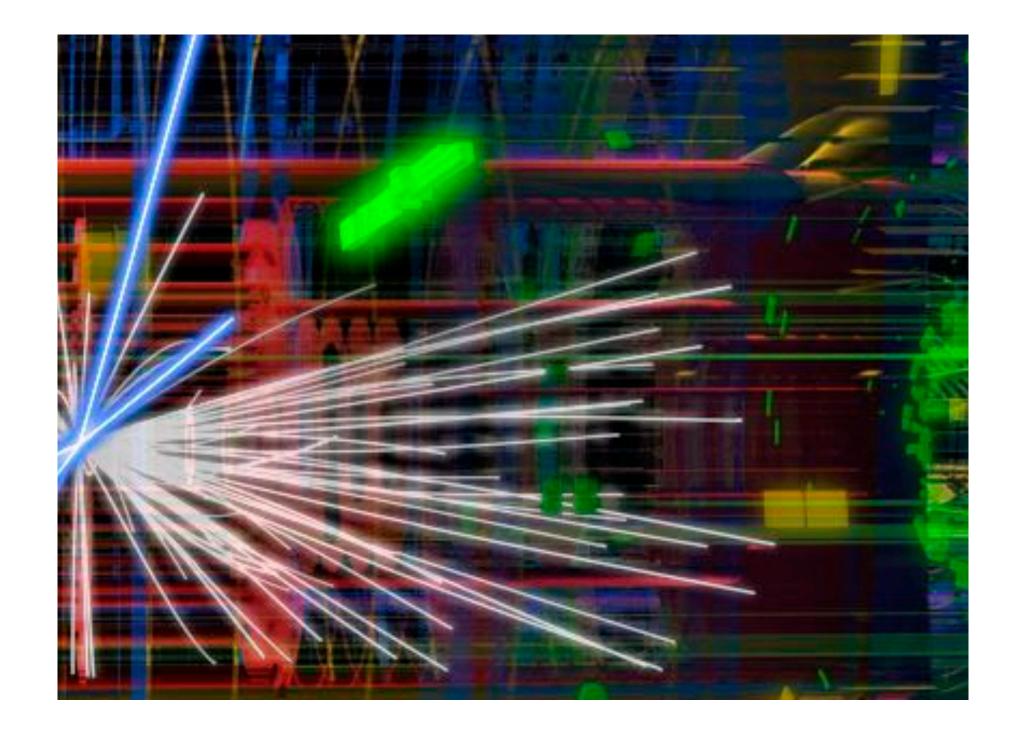




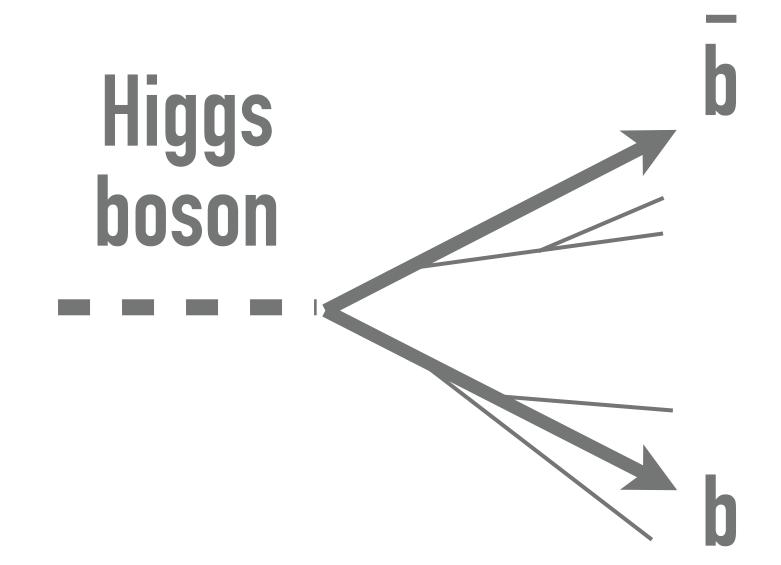


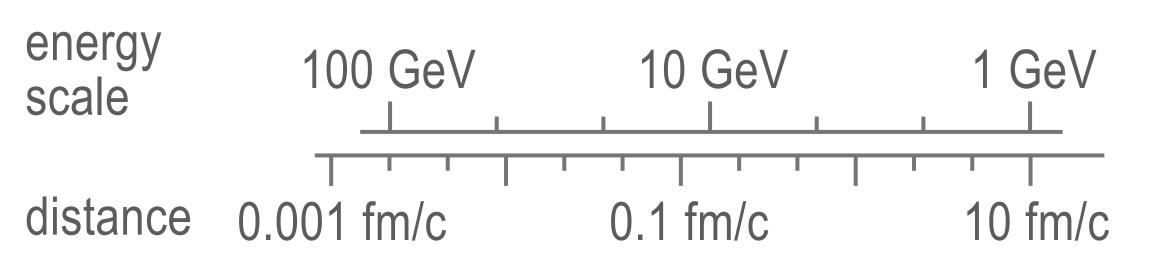


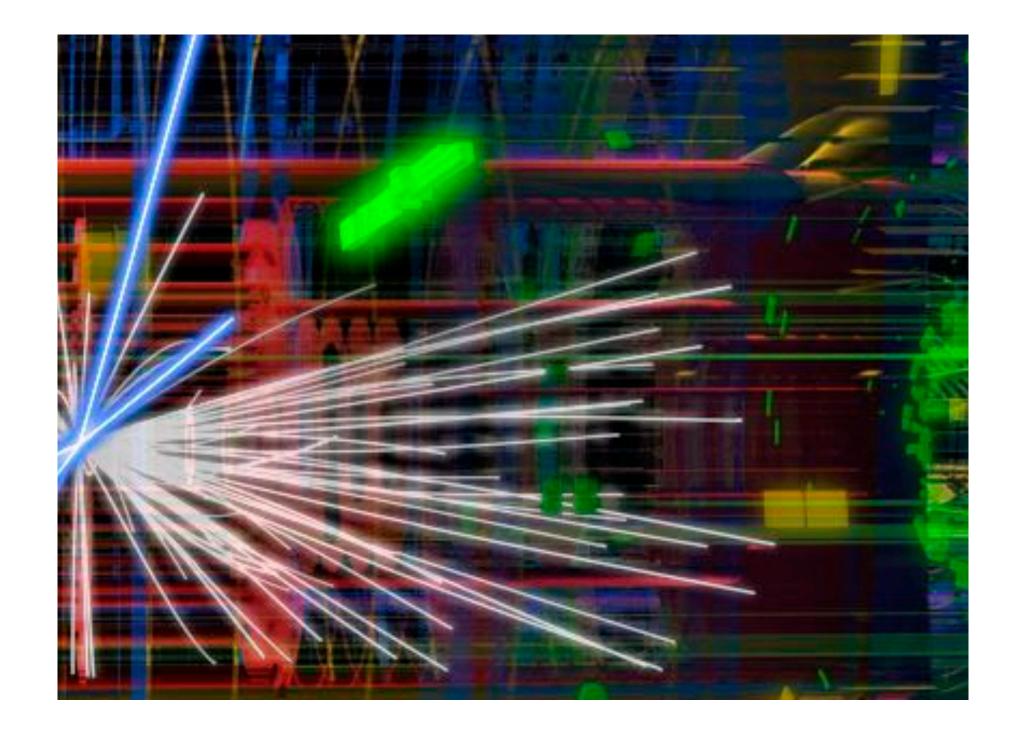




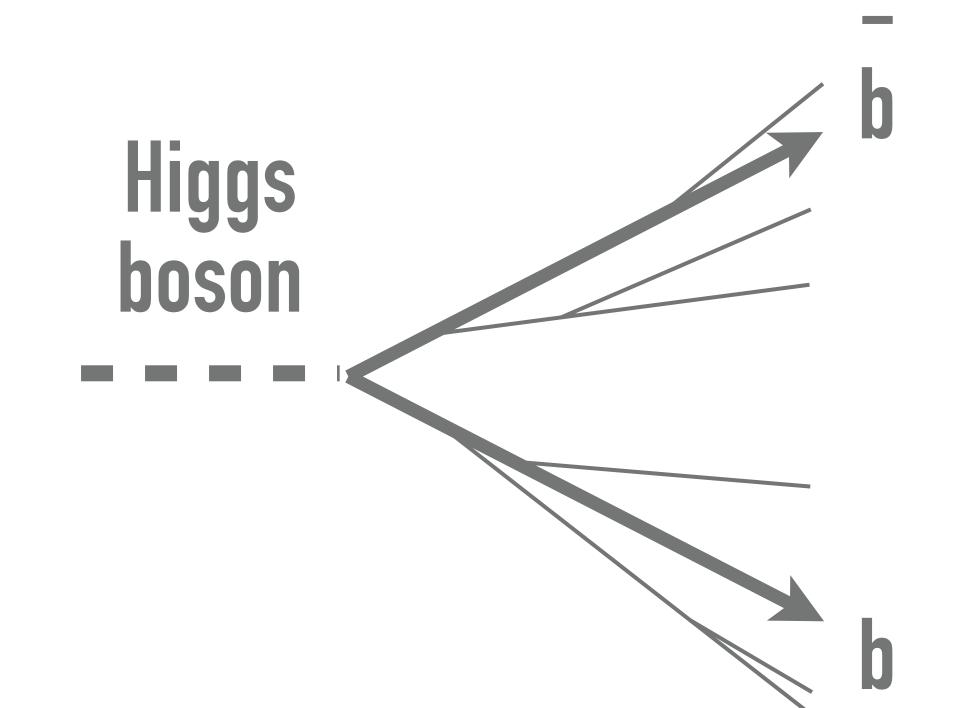


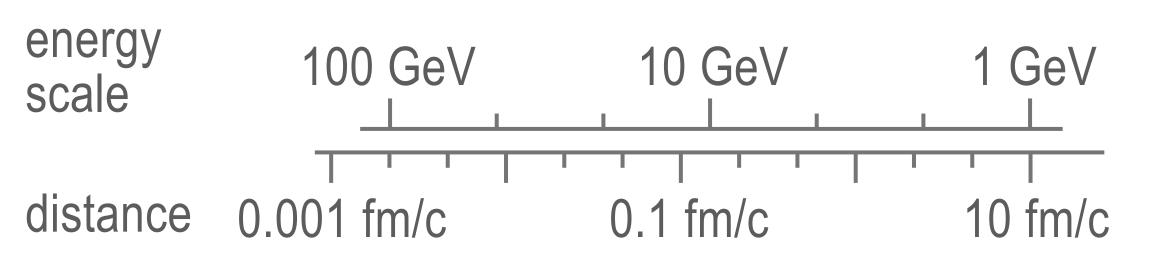


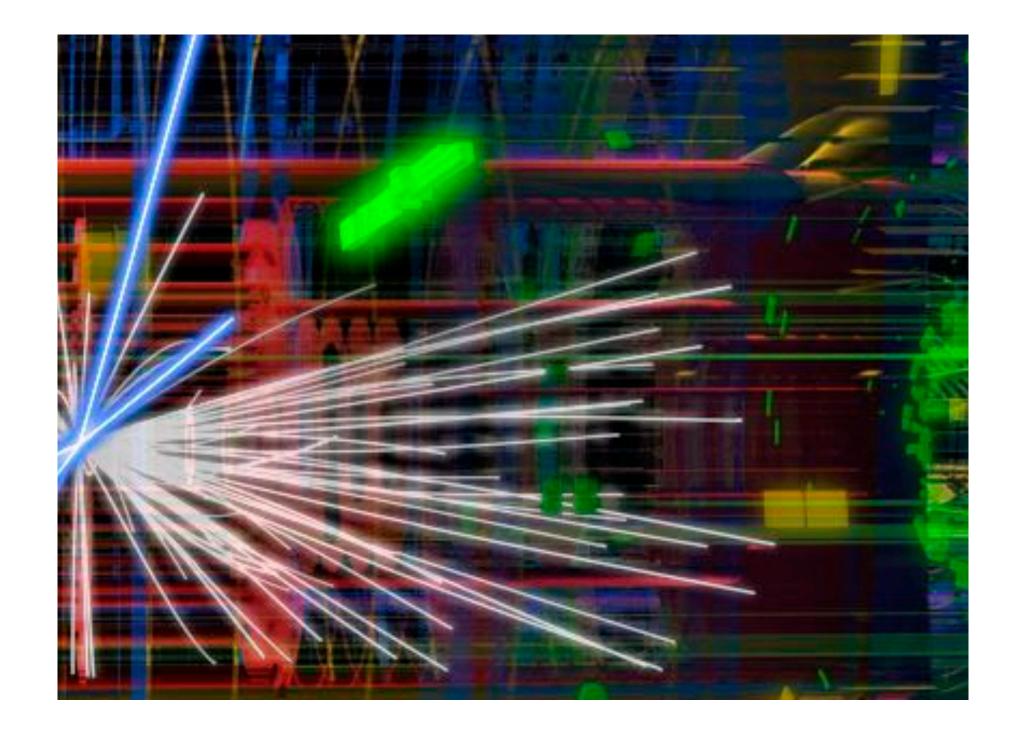




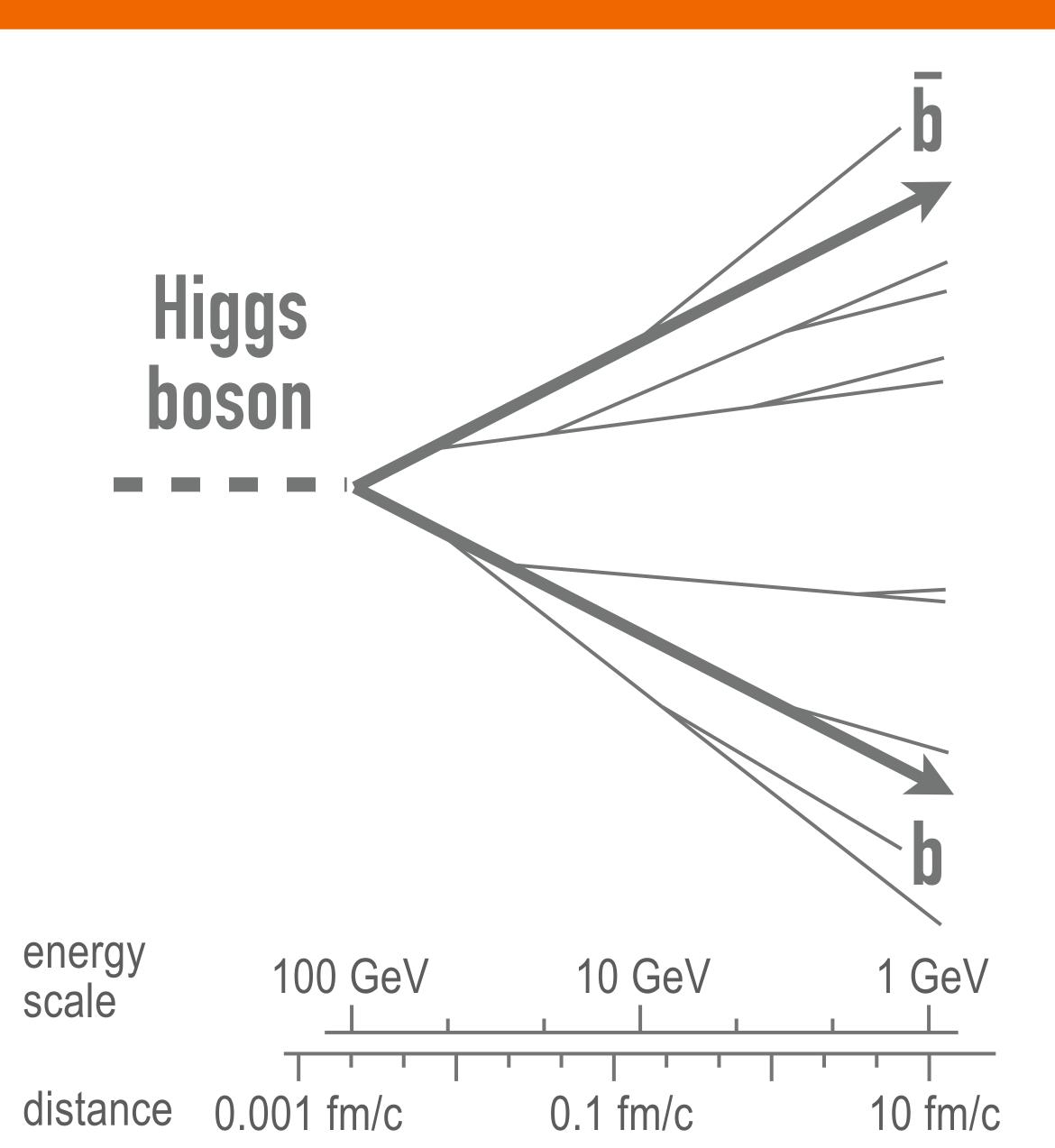


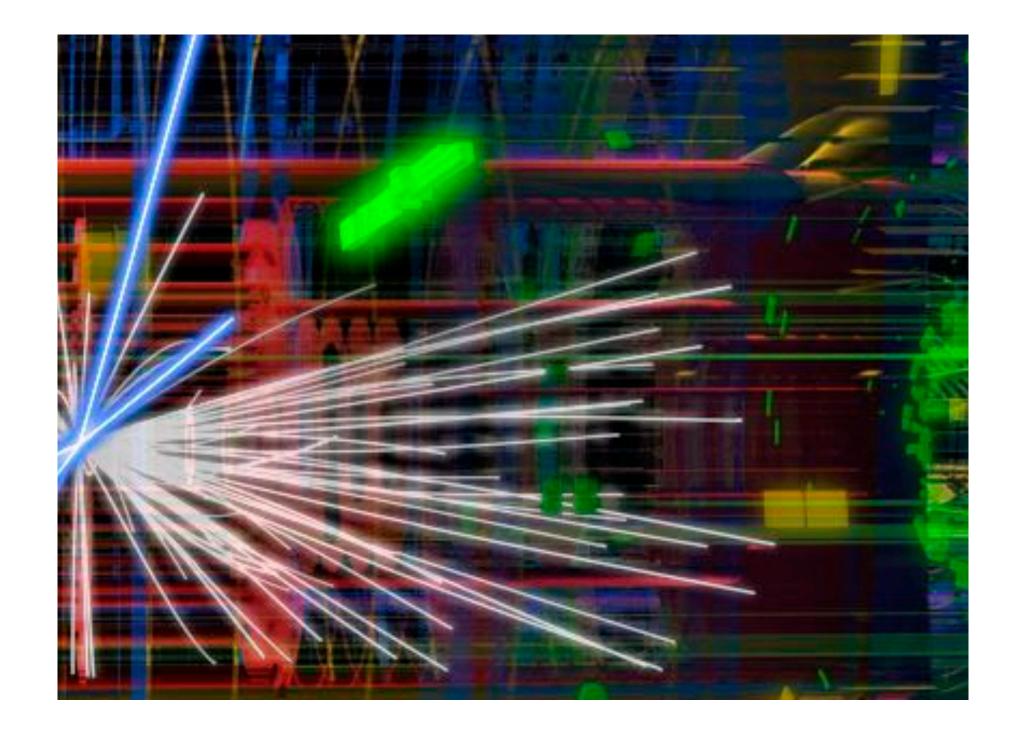




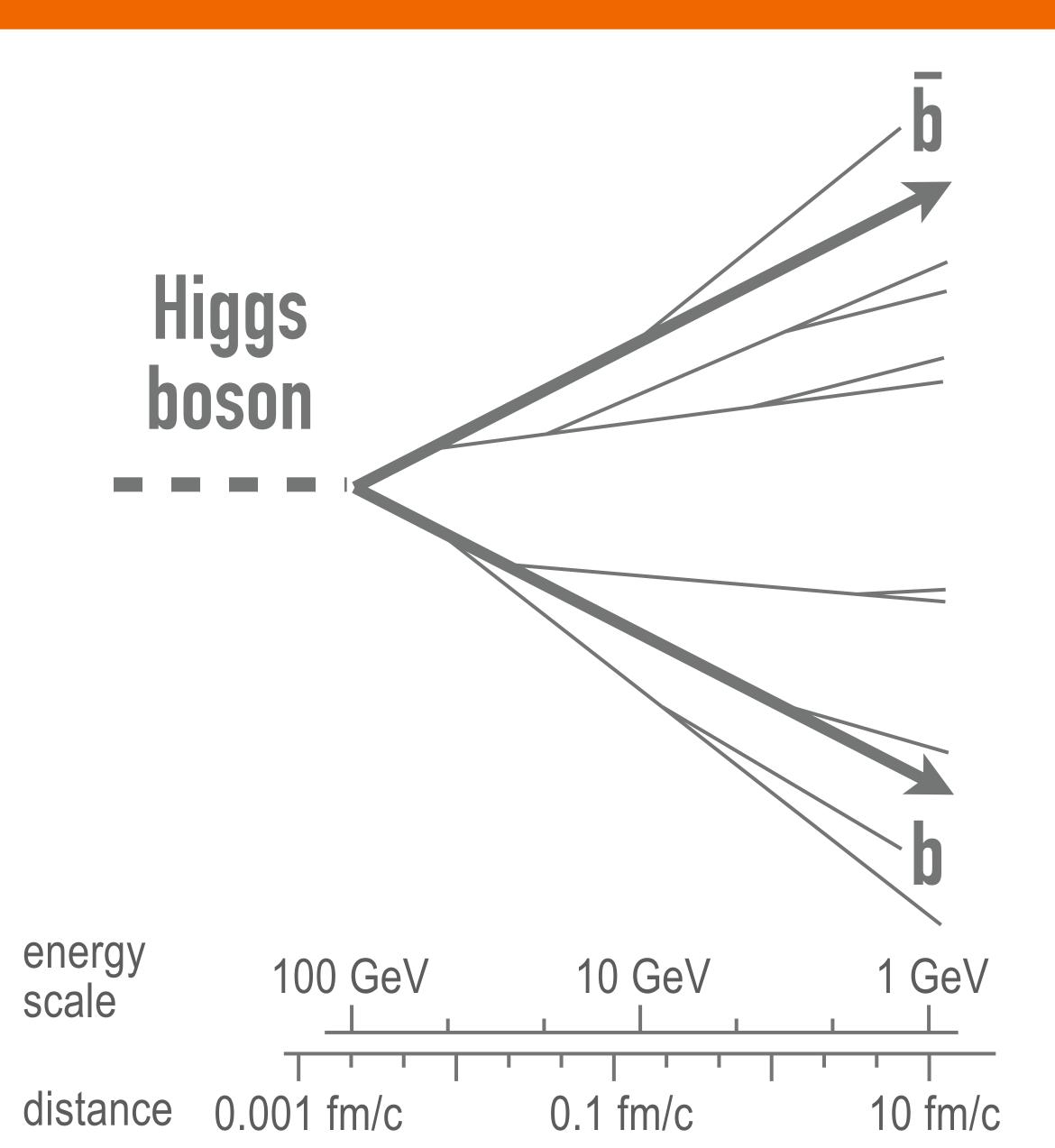








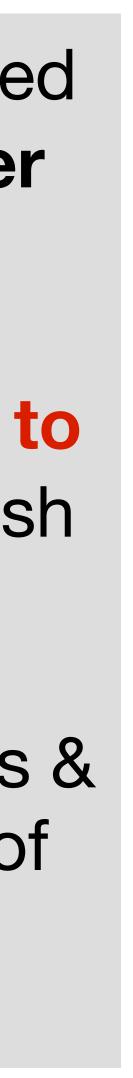




Pattern of branching usually simulated with a Monte Carlo Parton Shower algorithm

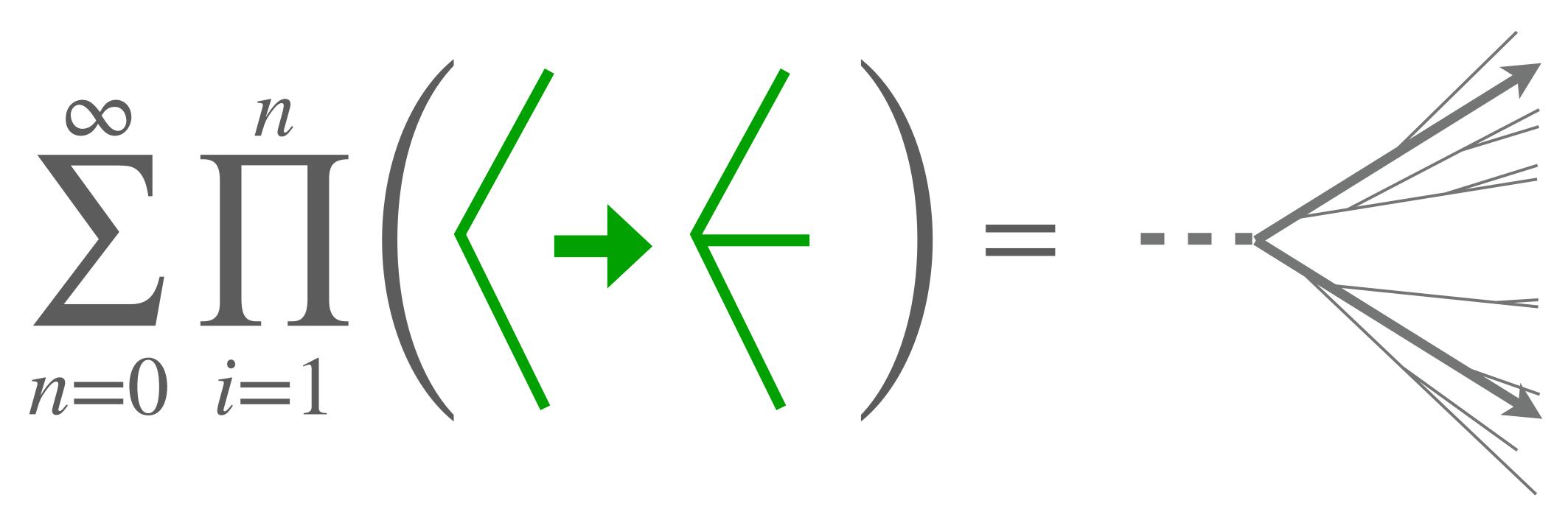
Experiments always compare data to Monte Carlo simulations to establish fundamental hypotheses

Little is known about their robustness & accuracy of multi-scale properties of these simulations: a weak link in the chain





Parton shower = iteration of $2 \rightarrow 3$ (or $1 \rightarrow 2$) splitting kernel



- > in what sense should the distribution of final *n*-particle states be correctly described?
- can it even be correctly described, and with what constraints on the splitting kernels?

Dasgupta, Dreyer, Hamilton, Monni & GPS, JHEP 1809 (2018) 033

and topic of "PanScales" ERC project











I personally expect supersymmetry to be discovered at the LHC

http://cerncourier.com/cws/article/cern/35456

-a Nobel prize-winning theorist [2008]





The New Hork Times

pinion

GRAY MATTER

A Crisis at the Edge of Physics

By Adam Frank and Marcelo Gleiser

June 5, 2015

dead end. It offers no path forward [...]"

"the standard model, despite the glory of its vindication, is also a





The New York Times

pinion

GRAY MATTER

A Crisis at the Edge of Physics

By Adam Frank and Marcelo Gleiser

June 5, 2015

"the standard model, despite the glory of its vindication, is also a dead end. It offers no path forward [...]"

I disagree. **Because the non**gauge part of the standard model is far from being fully explored.

3 Yukawas out of 9 We know nothing about the self coupling





it would be so much more exciting if we'd discovered new physics, right?

Beyond the Standard Model IV



John F Gunion Tao Han James Ohnemus

World Scientific

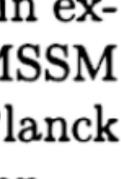
Back in 1995:

PLAN AND ADDRESS OF A ______

1. The Desert. A fun aspect of supersymmetry is that it allows us to obtain exact results about strongly interacting gauge theories. However in the MSSM we have nothing but boring perturbative physics to explore below the Planck scale and the interesting dynamics of supersymmetry breaking is hidden.

not everyone would agree







we must not underestimate our ignorance about the Higgs sector, nor the value of exploring and establishing it

e.g. accessing Yukawa couplings beyond the 3rd generation, the triple-Higgs coupling \rightarrow Higgs-field potential, SM keystone, & the pathway from discovery to precision



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meanwhile, the search for new physics continues

(And finding other things to do with the particles we have)

with much scope for inventing ingenious search techniques, identifying novel models that could be probed



searches, Higgs & other SM physics share in common

the need to think a underlying Lagran with observations of ~10⁷

- the need to think about how we relate the
- underlying Lagrangian of particle physics
- with observations of $\sim 10^{16}$ high-energy proton collisions

