Higgs and the new fundamental interactions

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THE ROYAL SOCIETY

Quantum Universe Kickoff Meeting DESY, Hamburg



"big answerable questions" and how we go about answering them

"big unanswered questions" about fundamental particles & their interactions (dark matter, matter-antimatter asymmetry, nature of dark energy, hierarchy of scales...)

V.























Higgs boson existence long known to be consistent with older e⁺e⁻ collider data (cf. LEP, 1989–2000 + SLD).

Tested through the small effect of virtual Higgs bosons on high-precision *(per-mil)* measurements.

Could be interpreted as a weak Higgs mass constraint.





hep-ex/0509008



Higgs boson existence long known to be consistent with older e+e- collider data (cf. LEP, 1989-2000 + SLD).

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> Could be interpreted as a weak Higgs mass constraint.





ATLAS and CMS collaborations at CERN's Large Hadron Collider (LHC):

2012 discovery of a Higgs-like boson

plot shows more recent data





The Higgs boson (2012)

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Success!

"The Standard Model is complete"





The Higgs boson (2012)

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Success!

"The Standard Model is complete"

Crisis!

No supersymmetry, no extra dimensions, there's nothing left for us to do...





The New York Eines

By DENNIS OVERBYE JUNE 19, 2017

|...| a cloud hanging over the physics community. [...]



What if there is nothing new to discover? That prospect is now

https://www.nytimes.com/2017/06/19/science/cern-large-hadron-collider-higgs-physics.html





what is the Standard Model?



particles



what is the Standard Model?



particles



interactions





Z = - FALFALFAL + iFDY + X: Jij X; \$+h.c. $+ D_{M} / - V(D)$

This equation neatly sums up our current understanding of fundamental particles and forces.

STANDARD MODEL — KNOWABLE UNKNOWNS

This is what you get when you buy one of those famous CERN T-shirts







L= - FALFALFAL + iFDY + X: Jij X; \$+h.c. + D g (-V(d))

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"understanding" = knowledge ? "understanding" = assumption ?















NOTATION

- A_{μ} : gauge field
 - ψ : fermion field

photons, gluons, W,Z

quarks & leptons

 ϕ : Higgs field $= \phi_0(\text{VEV}) + H(\text{Higgs})$

 $D_{\mu} = \partial_{\mu} + ieA_{\mu}$ etc. $F_{\mu\nu} \sim [D_{\mu}, D_{\nu}]$



= - + FAL F + X: Jij X; Ø+h.C. + Dg(-V(d))

e.g. $\psi D\psi \rightarrow \psi A_{\mu}\psi \rightarrow$ fermion-fermion-gauge vertex i.e. terms of \mathcal{L} map to particle interactions

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2 = - + Fmu + i FN + 4: 5; 4; \$+h.c. $+ \left| \mathcal{D} \mathcal{P} \right|^{<} - \mathcal{V} \left(\mathcal{O} \right)$

e.g. $F_{\mu\nu}F^{\mu\nu} \to A_{\mu}A_{\nu}\partial_{\mu}A_{\nu} \to \text{triple-gauge vertex}$ i.e. terms of \mathcal{L} map to particle interactions

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ナ Y: Jii Y, Ø +

This equation neatly sums up our current understanding of fundamental particles and forces.

GAUGE PART

e.g. qqγ, qqZ, qqg, evW, ggg, interactions — well established in ep, e⁺e[−], pp collisions, etc. **≡ KNOWLEDGE**

(also being studied at LHC — e.g. jets, DY/Z/W, V+jets, ttbar, etc.)



 $t \chi: \mathcal{Y}_{ij} \chi_{j} \phi$ $+ |D_{\mathcal{P}}(-V(\mathcal{O}))$

This equation neatly sums up our current understanding of fundamental particles and forces.

GAUGE PART

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(also being studied at LHC — e.g. jets, DY/Z/W, V+jets, ttbar, etc.)

Many SM studies probe this part. In some respects dates back to 1860's, i.e. Maxwell's equations.

If you test another corner of this (as one should), don't be surprised if it works









This equation neatly sums up our current understanding of fundamental particles and forces.

Higgs sector

until 6 years ago none of these terms had ever been directly observed.



 $\blacktriangleright \phi$ is a field at every point in space (plot shows potential vs. 1 of 4 components, at 1 point in space)



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► Our universe sits at minimum of $V(\phi)$, at $\phi = \phi_0 = -\frac{\mu}{\sqrt{1-\mu}}$



 $\blacktriangleright \phi$ is a field at every point in space (plot shows potential vs. 1 of 4 components, at 1 point in space)

► Our universe sits at minimum of $V(\phi)$, at $\phi = \phi_0 = \frac{\mu}{\sqrt{2\lambda}}$

 \blacktriangleright Excitation of the φ field around φ_0 is a Higgs boson ($\phi = \phi_0 + H$)





$\varphi = \varphi_0 + H$

established (2012 Higgs boson discovery)



$\varphi = \varphi_0 + H$

(2012 Higgs boson discovery)



 $\bigvee(\phi) = -\mu^2 \phi^2 + \lambda \phi^4$





what terms are there in the Higgs sector? 2. Gauge-Higgs term



 $\rightarrow g^2 \phi_0^2 Z_\mu Z^\mu + 2g^2 \phi_0 H Z_\mu Z^\mu + \dots$

Z-boson mass term

$$\begin{split} (D_{\mu})^2 &\sim (\partial_{\mu} + igZ_{\mu} + \dots)^2 \sim g^2 Z_{\mu} Z^{\mu} + \cdots \\ (\phi)^2 &= (\phi_0 + H)^2 = \phi_0^2 + 2\phi_0 H + H^2 \end{split}$$

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ZZH interaction term

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ZZH interaction term

> Higgs mechanism predicts specific relation between Z-boson mass and HZZ interaction





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Higgs mechanism predicts specific relation between Z-boson mass and HZZ interaction





what terms are there in the Higgs sector? 3. Fermion-Higgs (Yukawa) term



fermion mass term $m_i = y_{ii}\phi_0$

i	Уi	i	Уi
u	$2 \cdot 10^{-5}$	d	$3 \cdot 10^{-5}$
С	$8 \cdot 10^{-3}$	S	$6 \cdot 10^{-4}$
b	$3 \cdot 10^{-2}$	t	1
$ u_e $		е	$3 \cdot 10^{-6}$
$ u_{\mu}$	$\sim 10^{-13}$	μ	$6 \cdot 10^{-4}$
$ u_{ au}$		au	$1 \cdot 10^{-4}$

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$\mathcal{Y}_{ij} \mathcal{Y}_{ij} \mathcal{Y}_{ij} \phi \rightarrow y_{ij} \phi_0 \psi_i \psi_j + y_{ij} H \psi_i \psi_i$

fermion-fermion-Higgs interaction term; coupling $\sim \gamma_{ii}$



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fermion-fermion-Higgs interaction term; coupling $\sim y_{ii}$

 $\psi_{ii}H\psi_i\psi_j$

 $\phi = \phi_0 + H$

concentrate on Yukawa interaction hypothesis

Yukawa couplings ~ fermion mass

first fundamental interaction that we probe at the quantum level where interaction strength is not quantised (i.e. no underlying unit of charge across particles)





Why do Yukawa couplings matter? (1) Because, within SM conjecture, they're what give masses to all quarks

Up quarks (mass ~ 2.2 MeV) are lighter than down quarks (mass ~ 4.7 MeV)

proton **neutron** (up+down+down): 2.2 + **4.7** + 4.7 + ... = **939.6** MeV

> So protons are **lighter** than neutrons, \rightarrow protons are stable.

Which gives us the hydrogen atom, & chemistry and biology as we know it



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(up+up+down): 2.2 + 2.2 + 4.7 + ... = 938.3 MeV

proton mass = 938.3 MeV

neutron mass = 939.6 MeVC

Why do Yukawa couplings matter? (2) Because, within SM conjecture, they're what give masses to all leptons

Bohr radius

electron mass determines size of all atoms

it sets energy levels of all chemical reactions

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1st generation (us) has low mass because of weak interactions with Higgs field (and so with Higgs bosons): too weak to test today





1st generation (us) has low mass because of weak interactions with Higgs field (and so with Higgs bosons): too weak to test today 3rd generation (us) has high
mass because of strong
interactions with Higgs field
(and so with Higgs bosons):
can potentially be tested



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ALICE



LHC 7 TeV + 7 TeV 27 km

 \sim





Copyright CERN

ALICE



LHC 7 TeV + 7 TeV 27 km

 \sim







ATLAS & CMS **@LHC**

~up to 2 billion collisions/second

(+ lower rates at LHCb and ALICE)









what underlying processes tell us about Yukawa interactions?







Higgs production: the dominant channel



Expected to happen once for every ~2 billion inelastic proton-proton collisions

LHC data consistent with that already at discovery in 2012









Higgs boson is really being that you're actually seeing a Yukawa coupling?







Higgs production: the ttH channel Higgs out If SM top-Yukawa hypothesis is correct, expect 1 Higgs for every 1600 top-quark pairs.

(rather than 1 Higgs for every 2 billion pp collisions)



















vents with top-quarks & Higgs simultaneously

ATLAS > 5-sigma ttH









vents with top-quarks & Higgs simultaneously

ATLAS > 5-sigma ttH











Yukawa coupling:

1 in every 16 Higgs bosons \sim decays to $\tau^+\tau^-$



observation of $H \rightarrow \tau \tau$

~2 years ago: CMS >5-sigma H $\rightarrow \tau \tau$

35.9 fb⁻¹ (13 TeV)



l year ago: ATLAS >5-sigma H $\rightarrow \tau \tau$

35.9 fb⁻¹ (13 TeV)



 $m_{\tau\tau}^{MMC}$ [GeV]







Yukawa coupling:

~ 58% of Higgs bosons should decay to bb

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six months ago, observation of $H \rightarrow bb$

CMS >5-sigma H \rightarrow **bb**

77.2 fb⁻¹ (13 TeV)



Analysis includes key idea from Butterworth, Davison, Rubin, GPS (PRL 100 (2008) 242001)

ATLAS > 5-sigma $H \rightarrow bb$





what could one be saying about it?

The $>5\sigma$ observations of the ttH process and of H $\rightarrow \tau\tau$ and H \rightarrow bb decays, independently by ATLAS and CMS, firmly establish the existence of a new kind of fundamental interaction, Yukawa interactions.

Yukawa interactions are important because they are:

(1) qualitatively unlike any quantum interaction probed before (effective charge not quantised), (2) hypothesized to be responsible for the stability of hydrogen, and for determining the size of atoms and the energy scales of chemical reactions.

Establishing the pattern of Yukawa couplings across the full remaining set of quarks and charged leptons is one of the major challenges for particle physics today.





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Is this any less important than the discovery of the Higgs boson itself? My opinion: no, because fundamental interactions are as important as fundamental particles

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what could one be saying about it?

This is a **fifth force, the "Higgs force"** (up to you to decide whether you prefer to talk about new interactions or new force)

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today: no evidence yet (1 in 4570 decays)

observable at the LHC within about 10 years.







today: no evidence yet (1 in 35 decays)

needs an e⁺e⁻ or ep collider



Yukawas

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overall normalisation (related to Higgs width): needs an e⁺e⁻ collider

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EFT approach

Well-defined theoretical approach Assumes New Physics states are heavy Write Effective Lagrangian with only light (SM) particles BSM effects can be incorporated as a momentum expansion



BSM effects SM particles







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for much of Higgs sector, we know what to do to get answers. What about other "big" questions

Nature of dark m Fine-tuning (e.g. sup Matter-antimatter as

- Nature of dark matter (& dark energy)
- Fine-tuning (e.g. supersymmetry and similar)
- Matter-antimatter asymmetry of the universe
 - [...]



Finding dark matter and studying it will be the biggest challenge for the Large Hadron Collider's second run

https://www.pbs.org/newshour/science/largehadron-collider-gears-find-dark-matter-newparticles-second-run

-a large LHC experiment's spokesperson [2015]





cark mater

Velocity (km s⁻¹)

50

100

Observations from starlight

Rotation curve of spiral galaxy Messier 33

10,000

Mario De Leo 💿 CC BY-SA 4.0

Observations from 21 cm hydrogen

Expected from the visible disk

20,000 30,000 40,000

Distance (light years)



Looking beyond the SM: searches for dark matter at LHC & elsewhere

Classic dark-matter candidate: a weaklyinteracting massive particle (WIMP, e.g. from supersymmetry).

Masses ~ GeV upwards

(search interpretations strongly model dependent)





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musn't be (too) disappointed at lack of dark matter signal at LHC

Evidence for dark matter exists since the 1930s.

Today we know that

- there are many possible models
- > the range of parameters they span is large

We must deploy full ingenuity in searching for dark matter, including at LHC.

But must also recognise that it has remained elusive for 80–90 years, and chances of finding it in any given year are small!

Snowmass non-WIMP dark matter report, 1310.8642



Figure 1. Graphical representation of the (incomplete) landscape of candidates. Above, the landscape of dark matter candidates due to T. Tait. Below, the range of dark matter candidates' masses and interaction cross sections with a nucleus of Xe (for illustrative purposes) compiled by L. Pearce. Dark matter candidates have an enormous range of masses and interaction cross sections.





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ATLAS SUSY Searches* - 95% CL Lower Limits

A D	TLAS SUSY Seal	rches*	- 95%	6 C	L Lov	ver Limits			ATLAS Preliminary $\sqrt{s} = 7, 8, 13$ TeV
	Model	e, μ, τ, γ	Jets	$E_{ m T}^{ m miss}$	∫ <i>L dt</i> [fh	-1] Mass limit	$\sqrt{s} = 7, 8$	TeV $\sqrt{s} = 13$ TeV	Reference
Inclusive Searches	$\begin{array}{l} \tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0} \\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0} \text{ (compressed)} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{1} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{\pm} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qq(\ell\ell/\gamma\nu)\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_{1}^{0} \\ \text{GMSB}(\tilde{\ell} \text{ NLSP}) \\ \text{GGM (bino NLSP)} \\ \text{GGM (higgsino-bino NLSP)} \\ \text{Gravitino LSP} \end{array}$	0 mono-jet 0 ee, μμ 3 e, μ 0 1-2 τ + 0-1 ℓ 2 γ γ 0	2-6 jets 1-3 jets 2-6 jets 2 jets 4 jets 7-11 jets 0-2 jets - 2 jets mono-jet	Yes Yes Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 36.1 36.1 14.7 36.1 36.1 36.1 36.1 36.1 20.3		1.57 TeV 2.02 TeV 2.01 TeV 1.7 TeV 1.87 TeV 1.87 TeV 2.0 TeV 2.15 TeV 2.05 TeV	$\begin{split} &m(\tilde{x}_{1}^{0}) < 200 \mathrm{GeV}, m(1^{\mathrm{st}} \mathrm{gen}, \tilde{\mathfrak{q}}) = m(2^{\mathrm{st}} \mathrm{gen}, \tilde{\mathfrak{q}}) \\ &m(\tilde{q}) - m(\tilde{x}_{1}^{0}) < 5 \mathrm{GeV} \\ &m(\tilde{x}_{1}^{0}) < 200 \mathrm{GeV} \\ &m(\tilde{x}_{1}^{0}) < 200 \mathrm{GeV}, m(\tilde{x}^{\pm}) = 0.5 (m(\tilde{x}_{1}^{0}) + m(\tilde{g})) \\ &m(\tilde{x}_{1}^{0}) < 300 \mathrm{GeV}, \\ &m(\tilde{x}_{1}^{0}) = 0 \mathrm{GeV} \\ &m(\tilde{x}_{1}^{0}) = 0 \mathrm{GeV} \\ &m(\tilde{x}_{1}^{0}) = 0 \mathrm{GeV} \\ &m(\tilde{x}_{1}^{0}) = 10 \mathrm{GeV} \\ &m(\tilde{x}_{1}^{0}) = 1700 \mathrm{GeV}, c\tau(NLSP) < 0.1 mm, \mu > 0 \\ &m(\tilde{\mathcal{G}}) > 1.8 \times 10^{-4} \mathrm{eV}, m(\tilde{g}) = m(\tilde{q}) = 1.5 \mathrm{TeV} \end{split}$	1712.02332 1711.03301 1712.02332 1712.02332 1611.05791 1706.03731 1708.02794 1607.05979 ATLAS-CONF-2017-080 ATLAS-CONF-2017-080 1502.01518
d gen.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\bar{b}\tilde{\chi}_{1}^{0}$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_{1}^{0}$	0 0 -1 <i>c</i> .μ	3b 3h	Yes Yes	36. 1 36.1	ë ë	1 .92 T eV 1.97 TeV	$m(\tilde{\chi}_{1}^{0}) < 600 \text{ GeV}$ $m(\tilde{\chi}_{1}^{0}) < 200 \text{ GeV}$	1711.01901 1711.01901
3 rd gen. squarks 3 rd direct production 3	$\begin{split} \tilde{b}_{1}\tilde{b}_{1}, \tilde{b}_{1} \to b\tilde{\chi}_{1}^{0} \\ \tilde{b}_{1}\tilde{b}_{1}, \tilde{b}_{1} \to t\tilde{\chi}_{1}^{\pm} \\ \tilde{b}_{1}\tilde{b}_{1}, \tilde{b}_{1} \to t\tilde{\chi}_{1}^{\pm} \\ \tilde{i}_{1}\tilde{i}_{1}, \tilde{i}_{1} \to b\tilde{\chi}_{1}^{\pm} \\ \tilde{i}_{1}\tilde{i}_{1}, \tilde{i}_{1} \to Wb\tilde{\chi}_{1}^{0} \text{ or } t\tilde{\chi}_{1}^{0} \\ \tilde{i}_{1}\tilde{i}_{1}, \tilde{i}_{1} \to c\tilde{\chi}_{1}^{0} \\ \tilde{i}_{1}\tilde{i}_{2}, \tilde{i}_{2} \to \tilde{i}_{1} + Z \\ \tilde{i}_{2}\tilde{i}_{2}, \tilde{i}_{2} \to \tilde{i}_{1} + h \end{split}$	0 2 e, μ (SS) 0-2 e, μ 0 2 e, μ (Z) 3 e, μ (Z) 1-2 e, μ	2 b 1 b 1-2 b 0-2 jets/1-2 b mono-jet 1 b 1 b 4 b	Yes Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 4.7/13.3 20.3/36.1 36.1 20.3 36.1 36.1 36.1			$\begin{split} & m(\tilde{x}_{1}^{0}) < 420 \mathrm{GeV} \\ & m(\tilde{x}_{1}^{0}) < 200 \mathrm{GeV}, m(\tilde{x}_{1}^{\pm}) = m(\tilde{x}_{1}^{0}) + 100 \mathrm{GeV} \\ & m(\tilde{x}_{1}^{0}) < 2m(\tilde{x}_{1}^{0}), m(\tilde{x}_{1}^{0}) = 55 \mathrm{GeV} \\ & m(\tilde{x}_{1}^{0}) = 1 \mathrm{GeV} \\ & m(\tilde{x}_{1}^{0}) = 1 \mathrm{GeV} \\ & m(\tilde{x}_{1}^{0}) = 150 \mathrm{GeV} \\ & m(\tilde{x}_{1}^{0}) = 0 \mathrm{GeV} \\ & m(\tilde{x}_{1}^{0}) = 0 \mathrm{GeV} \\ & m(\tilde{x}_{1}^{0}) = 0 \mathrm{GeV} \end{split}$	1708.09266 1706.03731 1209.2102, ATLAS-CONF-2016-077 1506.08616, 1709.04183, 1711.11520 1711.03301 1403.5222 1706.03986 1706.03986
EV direct	$\begin{array}{l} \tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\ell}\nu(\ell \tilde{\nu}) \\ \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{2}^{0} \rightarrow \tilde{\tau}\tau(\nu \tilde{\nu}), \tilde{\chi}_{2}^{0} \rightarrow \tilde{\tau}\tau(\nu \tilde{\nu}) \\ \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{L}\nu \tilde{\ell}_{L}\ell(\tilde{\nu}\nu), \ell \tilde{\nu} \tilde{\ell}_{L}\ell(\tilde{\nu}\nu) \\ \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} Z \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} h \tilde{\chi}_{1}^{0}, h \rightarrow b \bar{b} / W W / \tau \tau / \gamma \gamma \\ \tilde{\chi}_{2}^{0} \tilde{\chi}_{3}^{0}, \tilde{\chi}_{2,3}^{0} \rightarrow \tilde{\ell}_{R} \ell \\ \text{GGM (wino NLSP) weak prod., } \tilde{\chi}_{1}^{0} \rightarrow \end{array}$	2 e,μ 2 e,μ 2 τ 3 e,μ 2-3 e,μ e,μ,γ 4 e,μ γĞ 1 e,μ + γ γĞ 2 γ	0 0 - 0-2 jets 0-2 <i>b</i> 0 -	Yes Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 36.1 36.1 20.3 20.3 20.3 36.1	\tilde{l} 90-500 GeV $\tilde{\chi}_1^{\pm}$ 750 GeV $\tilde{\chi}_1^{\pm}$ 760 GeV $\tilde{\chi}_1^{\pm}, \tilde{\chi}_2^{0}$ 1.13 Te $\tilde{\chi}_1^{\pm}, \tilde{\chi}_2^{0}$ 580 GeV $\tilde{\chi}_{1,x}^{\pm}, \tilde{\chi}_2^{0}$ 635 GeV $\tilde{\chi}_{2,3}^{0}$ 635 GeV \tilde{W} 115-370 GeV \tilde{W} 1.06 TeV	m($ ilde{ extsf{X}}_1^{\pm})=n$ m($ ilde{ extsf{X}}_2^{0})=n$	$\begin{split} & m(\tilde{\chi}_{1}^{0}) \!=\! 0 \\ & m(\tilde{\chi}_{1}^{0}) \!=\! 0, m(\tilde{\ell}, \tilde{\nu}) \!=\! 0.5(m(\tilde{\chi}_{1}^{\pm}) \!+\! m(\tilde{\chi}_{1}^{0})) \\ & m(\tilde{\chi}_{1}^{0}) \!=\! 0, m(\tilde{\tau}, \tilde{\nu}) \!=\! 0.5(m(\tilde{\chi}_{1}^{\pm}) \!+\! m(\tilde{\chi}_{1}^{0})) \\ & m(\tilde{\chi}_{2}^{0}), m(\tilde{\chi}_{1}^{0}) \!=\! 0, m(\tilde{\ell}, \tilde{\nu}) \!=\! 0.5(m(\tilde{\chi}_{1}^{\pm}) \!+\! m(\tilde{\chi}_{1}^{0})) \\ & m(\tilde{\chi}_{1}^{\pm}) \!=\! m(\tilde{\chi}_{2}^{0}), m(\tilde{\chi}_{1}^{0}) \!=\! 0, \tilde{\ell} \text{ decoupled} \\ & m(\tilde{\chi}_{1}^{\pm}) \!=\! m(\tilde{\chi}_{2}^{0}), m(\tilde{\chi}_{1}^{0}) \!=\! 0, \tilde{\ell} \text{ decoupled} \\ & m(\tilde{\chi}_{3}^{\pm}), m(\tilde{\chi}_{1}^{0}) \!=\! 0, m(\tilde{\ell}, \tilde{\nu}) \!=\! 0.5(m(\tilde{\chi}_{2}^{0}) \!+\! m(\tilde{\chi}_{1}^{0})) \\ & c\tau \!<\! 1 mm \\ c\tau \!<\! 1 mm \end{split}$	ATLAS-CONF-2017-039 ATLAS-CONF-2017-039 1708.07875 ATLAS-CONF-2017-039 ATLAS-CONF-2017-039 1501.07110 1405.5086 1507.05493 ATLAS-CONF-2017-080
Long-lived particles	Direct $\tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{-}$ prod., long-lived $\tilde{\chi}_{1}^{\pm}$ Direct $\tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{-}$ prod., long-lived $\tilde{\chi}_{1}^{\pm}$ Stable, stopped \tilde{g} R-hadron Stable \tilde{g} R-hadron Metastable \tilde{g} R-hadron Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow qq \tilde{\chi}_{1}^{0}$ GMSB, stable $\tilde{\tau}, \tilde{\chi}_{1}^{0} \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e, \mu)$ GMSB, $\tilde{\chi}_{1}^{0} \rightarrow \gamma \tilde{G}$, long-lived $\tilde{\chi}_{1}^{0}$ $\tilde{g}\tilde{g}, \tilde{\chi}_{1}^{0} \rightarrow eev/e\mu v/\mu\mu v$	Disapp. trk dE/dx trk 0 trk dE/dx trk displ. vtx $1-2 \mu$ 2γ displ. $ee/e\mu/\mu$	1 jet - 1-5 jets - - - - - -	Yes Yes - Yes - Yes -	36.1 18.4 27.9 3.2 32.8 19.1 20.3 20.3		1.58 TeV 1.57 TeV 2.37	$\begin{split} & m(\tilde{\chi}_{1}^{\pm}) - m(\tilde{\chi}_{1}^{0}) \sim 160 \; MeV, \; \tau(\tilde{\chi}_{1}^{\pm}) = 0.2 \; ns \\ & m(\tilde{\chi}_{1}^{\pm}) - m(\tilde{\chi}_{1}^{0}) \sim 160 \; MeV, \; \tau(\tilde{\chi}_{1}^{\pm}) < 15 \; ns \\ & m(\tilde{\chi}_{1}^{0}) = 100 \; GeV, \; 10 \; \mu s < \tau(\tilde{g}) < 1000 \; s \\ & m(\tilde{\chi}_{1}^{0}) = 100 \; GeV, \; 10 \; \mu s < \tau(\tilde{g}) < 1000 \; s \\ & m(\tilde{\chi}_{1}^{0}) = 100 \; GeV, \; \tau > 10 \; ns \\ & TeV \; \; \tau(\tilde{g}) = 0.17 \; ns, \; m(\tilde{\chi}_{1}^{0}) = 100 \; GeV \\ & 10 < lan\beta < 50 \\ & 1 < \tau(\tilde{\chi}_{1}^{0}) < 3 \; ns, \; SPS8 \; \mathsf{model \\ & 7 < cr(\tilde{\chi}_{1}^{0}) < 740 \; mm, \; m(\tilde{g}) = 1.3 \; TeV \end{split}$	1712.02118 1506.05332 1310.6584 1606.05129 1604.04520 1710.04901 1411.6795 1409.5542 1504.05162
RPV	LFV $pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e\mu/e\tau/\mu\tau$ Bilinear RPV CMSSM $\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow W\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow eev, e\mu\nu, \mu\mu\nu$ $\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow W\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow \tau\tau\nu_{e}, e\tau\nu_{\tau}$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow qqq$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow qqq$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_{1}t, \tilde{t}_{1} \rightarrow bs$ $\tilde{t}_{1}\tilde{t}_{1}, \tilde{t}_{1} \rightarrow b\ell$	$e\mu,e au,\mu au$ 2 e,μ (SS) 4 e,μ 3 $e,\mu + au$ 0 4 1 e,μ 8 0 2 e,μ	- 0-3 b - -5 large- <i>R</i> je -10 jets/0-4 8-10 jets/0-4 2 jets + 2 b 2 b	- Yes Yes ts- b- b- -	3.2 20.3 13.3 20.3 36.1 36.1 36.1 36.7 36.1	\tilde{v}_{τ} \tilde{v}_{τ} \tilde{q}, \tilde{g} 1.14 Te $\tilde{\chi}_{1}^{\pm}$ 450 GeV \tilde{g} \tilde{g} \tilde{g} 100-470 GeV \tilde{i}_{1} 100-470 GeV \tilde{i}_{1} 0.4-	1.9 TeV 1.45 TeV 20 1.875 TeV 2.1 TeV 1.65 TeV -1.45 TeV	$\begin{split} \lambda'_{311} = &0.11, \ \lambda_{132/133/233} = &0.07 \\ m(\tilde{q}) = &m(\tilde{q}), \ c\tau_{LSP} < 1 \ mm \\ m(\tilde{\chi}_1^0) > &400 \text{GeV}, \ \lambda_{12k} \neq 0 \ (k = 1, 2) \\ m(\tilde{\chi}_1^0) > &0.2 \times m(\tilde{\chi}_1^{\pm}), \ \lambda_{133} \neq 0 \\ m(\tilde{\chi}_1^0) = &1075 \ \text{GeV} \\ m(\tilde{\chi}_1^0) = &1 \ \text{TeV}, \ \lambda_{112} \neq 0 \\ m(\tilde{t}_1) = &1 \ \text{TeV}, \ \lambda_{323} \neq 0 \\ \end{split}$	1607.08079 1404.2500 ATLAS-CONF-2016-075 1405.5086 SUSY-2016-22 1704.08493 1704.08493 1710.07171 1710.05544
Other	Scalar charm, $\tilde{c} \rightarrow c \tilde{\chi}_1^0$	0	2 <i>c</i>	Yes	20.3	č 510 GeV		$m(\tilde{\chi}_1^0)$ <200 GeV	1501.01325
Only	a selection of the available may	ss limits on i limits are ba	new states	s or	1	0 ⁻¹ 1	ı	Mass scale [TeV]	

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10⁻¹

Mass scale [TeV]



Excited

Contact

ä

SSM Z'({{})
SSM Z'(gā)
LFV Z', BR($e\mu$) = 10%
SSM W'(<i>t</i> v)
SSM W'(qq)
SSM W'(τν)
LRSM $W_R(lN_R)$, $M_{N_R} = 0.5M_{W_R}$
LRSM W _R (τN_R), $M_{N_R} = 0.5 M_{W_R}$
Axigluon, Coloron, $cot \theta = 1$

scalar LQ (pair prod.), coupling to 1^{st} gen. fermions, $\beta = 1$ scalar LQ (pair prod.), coupling to 1^{st} gen. fermions, $\beta = 0.5$ scalar LQ (pair prod.), coupling to 2^{nd} gen. fermions, $\beta = 1$ scalar LQ (pair prod.), coupling to 2^{nd} gen. fermions, $\beta = 0.5$ scalar LQ (pair prod.), coupling to 3^{rd} gen. fermions, $\beta = 1$ scalar LQ (single prod.), coup. to 3^{rd} gen. ferm., $\beta = 1, \lambda = 1$

excited light quark ($q\bar{q}$), $\Lambda = m_q^*$ excited light quark (qy), $f_5 = f = f' = 1$, $\Lambda = m_a^*$ excited b quark, $f_{s} = f = f' = 1$, $\Lambda = m_{n}^{*}$ excited electron, $f_{\rm S} = f = f' = 1$, $\Lambda = m_{\rm e}^*$ excited muon, $f_{S} = f = f' = 1$, $\Lambda = m_{\mu}^{*}$

quark compositeness ($q \bar{q}$), $\eta_{
m LL/BR} = 1$ quark compositeness (ℓl), $\eta_{LURR} = 1$ quark compositeness ($q \tilde{q}$), $\eta_{
m LL/RR} = -1$ quark compositeness (ℓl), $\eta_{\rm LL/RR} = -1$

ADD (jj) HLZ, $n_{ED} = 3$ ADD $(\gamma\gamma, \ell\ell)$ HLZ, $n_{ED} = 3$ ADD G_{KK} emission, n = 2ADD QBH (jj), $n_{ED} = 6$ ADD QBH ($e\mu$), $n_{ED} = 6$ RS G_{KK}($q\bar{q}, gg$), $k/\overline{M}_{\rm Pl} = 0.1$ RS $G_{KK}(\ell \ell)$, $k/\overline{M}_{Pl} = 0.1$ RS $G_{KK}(\gamma\gamma)$, $k/\overline{M}_{Pl} = 0.1$ RS QBH (jj), $n_{ED} = 1$ RS QBH ($e\mu$), $n_{ED} = 1$ non-rotating BH, $M_D = 4 \text{ TeV}$, $n_{ED} = 6$ split-UED, $\mu \ge 4$ TeV

Matter Dark

Other

(axial-)vector mediator ($\chi\chi$), $g_q = 0.25$, $g_{\rm DM} = 1$, $m_\chi = 1~{
m GeV}$ (axial-)vector mediator ($q\ddot{q}$), $g_{\rm q}$ = 0.25, $g_{\rm DM}$ = 1, m_{χ} = 1 GeV scalar mediator (+ $t/t\bar{t}$), $g_{\rm g} = 1$, $g_{\rm DM} = 1$, $m_{\chi} = 1$ GeV pseudoscalar mediator (+ $t/t\bar{t}$), $g_{\rm q}$ = 1, $g_{\rm DM}$ = 1, m_{χ} = 1 GeV scalar mediator (fermion portal), $\lambda_u = 1$, $m_\chi = 1$ GeV complex sc. med. (dark QCD), $m_{\text{KDE}} = 5 \text{ GeV}$, $c\tau_{\text{XDE}} = 25 \text{ mm}$

Type III Seesaw, $B_e = B_\mu = B_\tau$ string resonance

CMS

M _{Z'}	1803.06292 (2 ℓ)	
$M_{Z'}$	1806.00843 (2 j)	
$M_{Z'}$	1802.01122 (е µ)	
$M_{W'}$	1803.11133 (<i>l</i> + E	niss)
$M_{W'}$	1806.00843 (2 j)	
$M_{W'}$	1807.11421 (τ + Ε	niss)
$M_{W_{\rm P}}$	1803.11116 (2ℓ + 2	2j)
$M_{\rm Wp}$	1811.00806 (2 τ +	2j)
$M_{\rm C}$	1806.00843 (2 j)	
M_{LQ}	1811.01197 (2e +	2j)
M_{LO}	1811.01197 (2e +	2j; e + 2j + E ^{miss})
M_{LQ}	1808.05082 (2µ +	2j)
i M _{LQ}	1808.05082 (2µ +	2j; μ + 2j + Ε τ ^{miss})
M_{LQ}	1811.00806 (2τ +	2j)
M_{LQ}	1806.03472 (2 τ +	b) 0.74
-		
Ma-	1806.00843 (2 j)	
Ma	1711.04652 (y + j)	
Mb-	1711.04652 (y + j)	
Me-	1811.03052 (y + 2	e)
<i>M</i> u-	1811.03052 (y + 2	μ)
-		
Δ^+_{UUDD}	1803.08030 (2j)	
ALLIER	1812.10443 (2l)	
A _{LUBB}	1803.08030 (2 j)	
ALLER	1812.10443 (2l)	
LLIN		
$M_{\rm S}$	1803.08030 (2 j)	
$M_{\rm S}$	1812.10443 (2y, 2	<i>t</i>)
$M_{\rm D}$	1712.02345 (≥1j	+ E ^{miss})
M _{QEH}	1803.08030 (2j)	
MOBIL	1802.01122 (eµ)	
MGer	1806.00843 (2 j)	
MGr	1803.06292 (2 ℓ)	
MGer	1809.00327 (2y)	
MOBH	1803.08030 (2j)	
M _{QBH}	1802.01122 (eµ)	
M _{BH}	1805.06013 (≥ 7j(<i>ℓ</i> , γ))
1/R	1803.11133 (<i>l</i> + E	niss)
$M_{ m med}$	1712.02345 (≥1j	+ E ^{miss})
M _{med}	1806.00843 (2 j)	
M _{med}	1901.01553 (0, 1 ℓ	$+ \geq 3j + E_T^{miss}$) 0.29
M _{med}	1901.01553 (0, 1 ℓ	$+ \geq 3j + E_T^{\text{miss}}$ 0.3
$M_{\Phi_{\alpha}}$	1712.02345 (≥1j	+ E ^{miss})
Mx	1810.10069 (4j)	
M _{Sigma}	1708.07962 (≥3ℓ)	0.84
Ms	1806.00843 (2j)	
	0.	1

Selection of observed exclusion limits at 95% C.L. (theory uncertainties are not included).



CMS exotics searches

January 2019





anomalies

the current place where there are hints of something happening



charged current

$$R(D^*) \equiv rac{\mathcal{B}(B^0
ightarrow D^{*-} au^+
u_ au)}{\mathcal{B}(B^0
ightarrow D^{*-} \mu^+
u_\mu)}$$

$R(D^*)$ and R(D) combination Combine LHCb's $R(D^*)$ results with results from B factories:



(latest SM computation: JHEP 11 (2017) 061)

Humair @ LHCP'18

neutral current

$$R(K^{(*)}) = \frac{\mathcal{B}(B \to K^{(*)}\mu^+\mu^-)}{\mathcal{B}(B \to K^{(*)}e^+e^-)}$$

R(K) and $R(K^*)$ results



All LHCb results below SM expectations:

► *B* factories have less precise but compatible results.





CMS pp \rightarrow b j $\mu^+\mu^-$ + X



https://arxiv.org/abs/1808.01890

ALEPH e⁺e⁻ \rightarrow b b $\mu^+\mu^-$ + X





future progress?

(1) approved plans

LHC will collect ~ 40 times more data than used for the plots shown so far, though at mostly similar energy (13–14 TeV)



Higgs precision ($H \rightarrow \gamma \gamma$) : optimistic estimate v. luminosity & time

extrapolation of μ_{vv} precision from 7+8 TeV results



 $1 \text{ fb}^{-1} = 10^{14} \text{ collisions}$



Higgs precision (H $\rightarrow \gamma\gamma$) : optimistic estimate v. luminosity & time

extrapolation of μ_{vv} precision from 7+8 TeV results



The LHC has the statistical potential to take Higgs physics from "observation" to 1–2% precision

But only if we learn how to connect experimental observations with theory at that precision

 $1 \text{ fb}^{-1} = 10^{14} \text{ collisions}$







HL-LHC official Higgs coupling projections (by 2036)



Right now, Higgs coupling precisions are in the 10-20% range.

We wouldn't consider electromagnetism established (textbook level) if we only knew it to 10%

HL-LHC can deliver 1-2% for a range of couplings



2nd-generation Yukawas at HL-LHC ($H \rightarrow \mu\mu$)





i	Уi	i	Уi
u	$2 \cdot 10^{-5}$	d	$3 \cdot 10^{-5}$
С	$8 \cdot 10^{-3}$	S	$6 \cdot 10^{-4}$
b	$3 \cdot 10^{-2}$	t	1
ν_e		е	$3 \cdot 10^{-6}$
$ u_{\mu}$	$\sim 10^{-13}$	μ	$6 \cdot 10^{-4}$
$ u_{ au}$		au	$1 \cdot 10^{-4}$



2nd-generation Yukawas at HL-LHC ($H \rightarrow \mu\mu$)







today: no evidence yet (1 in 4570 decays) observable at HL-LHC (within about 10 years)



2nd & 1st generation Yukawas

- the hierarchy of masses between generations remains a mystery (even if it's one that some people consign to the "hopeless" category)
- > Does not necessarily come from hierarchy of dimensionless Yukawa coefficients
- ► E.g. the Giudice-Lebedev mechanism (and follow-up work)

$$-\mathcal{L}_Y = Y_{ij}(\phi)\bar{\psi}_i\psi_j\phi + \text{h.c.}$$

- \blacktriangleright smallness of certain masses is consequence of vev²/M² suppression, not small c_{ij} \blacktriangleright measured Hqq interaction larger by factor (2n_{ii} + 1)
- ► cf. also various more recent discussions, e.g. by Bauer, Carena, Carmona

$$Y_{ij}(\phi) = c_{ij} \left(\frac{\phi^{\dagger}\phi}{M^2}\right)^{n_{ij}}$$

1801.00363

HL-LHC discovery potential



1 fb⁻¹ = 10¹⁴ collisions

Future

> 2018 (recorded):
> 140 fb⁻¹ @ 13 TeV
> 2023: ~400 fb⁻¹ @ 1? TeV
> 2036: 3000 fb⁻¹ @ 14 TeV

Today (analysed) > 20 fb⁻¹ @ 8 TeV > 35-80 fb⁻¹ @ 13 TeV



future progress?

(2) proposed future colliders e+e-: ILC, CLIC, CepC, FCC-ee, LEP3 pp: CppC, HE-LHC, FCC-hh ep: LHeC, FCC-eh



e+e- colliders: luminosity v. CoM energy













e+e-& eh colliders: coupling measurements (precision)

Collider	HL-LHC	ILC ₂₅₀	CLIC ₃₈₀		FCC-ee		FCC-eh
Luminosity (ab^{-1})	3	2	0.5	5 @	+1.5 @	+	2
				240 GeV	365 GeV	HL-LHC	
Years	25	15	7	3	+4		20
$\delta \Gamma_{\rm H} / \Gamma_{\rm H}$ (%)	SM	3.8	6.3	2.7	1.3	1.1	SM
$\delta g_{\mathrm{HZZ}}/g_{\mathrm{HZZ}}$ (%)	1.3	0.35	0.80	0.2	0.17	0.16	0.43
$\delta g_{\mathrm{HWW}}/g_{\mathrm{HWW}}$ (%)	1.4	1.7	1.3	1.3	0.43	0.40	0.26
$\delta g_{ m Hbb}/g_{ m Hbb}$ (%)	2.9	1.8	2.8	1.3	0.61	0.55	0.74
$\delta g_{ m Hcc}/g_{ m Hcc}$ (%)	SM	2.3	6.8	1.7	1.21	1.18	1.35
$\delta g_{\mathrm{Hgg}}/g_{\mathrm{Hgg}}$ (%)	1.8	2.2	3.8	1.6	1.01	0.83	1.17
$\delta g_{\mathrm{H}\tau\tau}/g_{\mathrm{H}\tau\tau}$ (%)	1.7	1.9	4.2	1.4	0.74	0.64	1.10
$\delta g_{\mathrm{H}\mu\mu}/g_{\mathrm{H}\mu\mu}$ (%)	4.4	13	n.a.	10.1	9.0	3.9	n.a.
$\delta g_{\rm H\gamma\gamma}/g_{\rm H\gamma\gamma}$ (%)	1.6	6.4	n.a.	4.8	3.9	1.1	2.3
$\delta g_{\mathrm{Htt}}/g_{\mathrm{Htt}}$ (%)	2.5	_		_		2.4	1.7
BR_{EXO} (%)	SM	< 1.8	< 3.0	< 1.2	< 1.0	< 1.0	n.a.



e+e- & eh colliders: Higgs-charm (2nd generation) coupling

today: no evidence yet (1 in 35 decays) needs an e⁺e⁻ or ep collider

Collider	HL-LHC	ILC ₂₅₀	CLIC ₃₈₀		FCC-ee		FCC-eh
Luminosity (ab^{-1})	3	2	0.5	5@	+1.5 @	+	2
				240 GeV	365 GeV	HL-LHC	
Years	25	15	7	3	+4		20
$\delta\Gamma_{ m H}/\Gamma_{ m H}$ (%)	SM	3.8	6.3	2.7	1.3	1.1	SM
$\delta g_{\mathrm{HZZ}}/g_{\mathrm{HZZ}}$ (%)	1.3	0.35	0.80	0.2	0.17	0.16	0.43
$\delta g_{ m HWW}/g_{ m HWW}$ (%)	1.4	1.7	1.3	1.3	0.43	0.40	0.26
	2.0			1.3	0.61	<u>055</u>	0.74
$\delta g_{ m Hcc}/g_{ m Hcc}$ (%)	SM	2.3	6.8	1.7	1.21	1.18	1.35
⁹ 9Hgg/9Hgg (70)		L.Z	2.0	1.0		0.00	1.17
$\delta g_{ m H au au}/g_{ m H au au}$ (%)	1.7	1.9	4.2	1.4	0.74	0.64	1.10
$\delta g_{\mathrm{H}\mu\mu}/g_{\mathrm{H}\mu\mu}$ (%)	4.4	13	n.a.	10.1	9.0	3.9	n.a.
$\delta g_{ m H\gamma\gamma}/g_{ m H\gamma\gamma}$ (%)	1.6	6.4	n.a.	4.8	3.9	1.1	2.3
$\delta g_{ m Htt}/g_{ m Htt}$ (%)	2.5					2.4	1.7
BR_{EXO} (%)	SM	< 1.8	< 3.0	< 1.2	< 1.0	< 1.0	n.a.





e^+e^- colliders: total Higgs width (= lifetime)

decay channels, whether observed or not).

Only e⁺e⁻ colliders can measure this directly.



Collider	HL-LHC	ILC ₂₅₀	CLIC ₃₈₀		FCC-ee		FCC-
Luminosity (ab ⁻¹)	3	2	0.5	5 @	+1.5 @	+	
				240 GeV	365 GeV	HL-LHC	
Years	25	15	7	3	+4		
$\delta \Gamma_{\rm H} / \Gamma_{\rm H}$ (%)	SM	3.8	6.3	2.7	1.3	1.1	S
$\delta g_{\mathrm{HZZ}}/g_{\mathrm{HZZ}}$ (%)	1.3	0.35	0.80	0.2	0.17	0.16	0.
$S \sim (01)$	1 /	1 7	1 2	1 2	0.42	0.40	0

- All current fits need to make assumptions about the total Higgs width (sum over all





pp colliders (concentrate on FCC-hh)



Figure 2: Higgs production cross sections versus collision energies normalized to the 14 TeV rates.

ttH HH

VBF ggH ZH WH

TeV

Higgs production rate increases substantially with collider centre-of-mass energy







is Higgs interaction pointlike?



study in events with large momentum transfers high-p_T or offshell Higgs



$V(\Phi) = m^2 \Phi^{\dagger} \Phi + \lambda (\Phi^{\dagger} \Phi)^2$



The Higgs potential holds together the rest of the standard model (keystone)

► so far (as a fundamental potential) only ever seen in

 $\sim -\phi^2 + \phi^4$ implies specific Taylor expansion around $\phi = \phi_0$:

$$(H_0 + H) = V_0 + \frac{1}{2}m_H^2 H^2 + c_3 H^3 + \cdots$$



















$V(\Phi) = m^2 \Phi^{\dagger} \Phi + \lambda (\Phi^{\dagger} \Phi)^2$?



hh channel	bbγγ	$b\bar{b}ZZ^*[\rightarrow 4\ell]$
cision	6.5%	14%

FCC triple Higgs v. LHC and HE-LHC







European Strategy Update

EUROPEAN STRATEGY FOR PARTICLE PHYSICS

The European Strategy for Particle Physics is the cornerstone of Europe's decision-making process for the long-term future of the field. Mandated by the CERN Council, it is formed through a broad consultation of the grass-roots particle physics community, it actively solicits the opinions of physicists from around the world, and it is developed in close coordination with similar processes in the US and Japan in order to ensure coordination between regions and optimal use of resources globally.



ongoing (2018 - 2020)





Figure 9: Overview of implementation timeline for the integral FCC program, starting in 2020. Numbers in the top row indicate the year. Physics operation for FCC-ee would start towards the end-2030s; physics operation for FCC-hh would start in the mid-2060s.





FCC physics CDR, table of contents (one of several volumes)

FCC Physics Opportunities

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I personally expect supersymmetry to be discovered at the LHC

http://cerncourier.com/cws/article/cern/35456

-a Nobel prize-winning theorist [2008]



The New York Times

pinion

GRAY MATTER

A Crisis at the Edge of Physics

By Adam Frank and Marcelo Gleiser

June 5, 2015

dead end. It offers no path forward [...]"

"the standard model, despite the glory of its vindication, is also a





The New York Times

pinion

GRAY MATTER

A Crisis at the Edge of Physics

By Adam Frank and Marcelo Gleiser

June 5, 2015

"the standard model, despite the glory of its vindication, is also a dead end. It offers no path forward [...]"

I disagree. **Because the non**gauge part of the standard model is far from being fully explored.

3 Yukawas out of 9 We know nothing about the self coupling





it would be so much more exciting if we'd discovered new physics, right?

Beyond the Standard Model IV



John F Gunion Tao Han James Ohnemus

World Scientific

Back in 1995:

Photo Charles Share

1. The Desert. A fun aspect of supersymmetry is that it allows us to obtain exact results about strongly interacting gauge theories. However in the MSSM we have nothing but boring perturbative physics to explore below the Planck scale and the interesting dynamics of supersymmetry breaking is hidden.

not everyone would agree







some theorists

it's interesting if it's what everyone is thinking about right

now

experimenter

it's interesting if it's never been observed before



some theorists

it's interesting if it's what everyone is thinking about right

now

both have a point (don't let one side dampen the other side's interest)

experimenter

it's interesting if it's never been observed before




we must not underestimate our ignorance about the Higgs sector, nor the value of exploring and establishing it

e.g. accessing Yukawa couplings beyond the 3rd generation, the triple-Higgs coupling \rightarrow Higgs-field potential, SM keystone, & the pathway from discovery to precision



75

meanwhile, the search for new physics continues

with much scope for inventing ingenious search techniques, and identifying novel models that could be probed

(And finding other things to do with the particles we have)



searches, Higgs & other SM physics share in common

the need to think a underlying Lagran with observations of ~10⁷

- the need to think about how we relate the
- underlying Lagrangian of particle physics
- with observations of $\sim 10^{16}$ high-energy proton collisions

