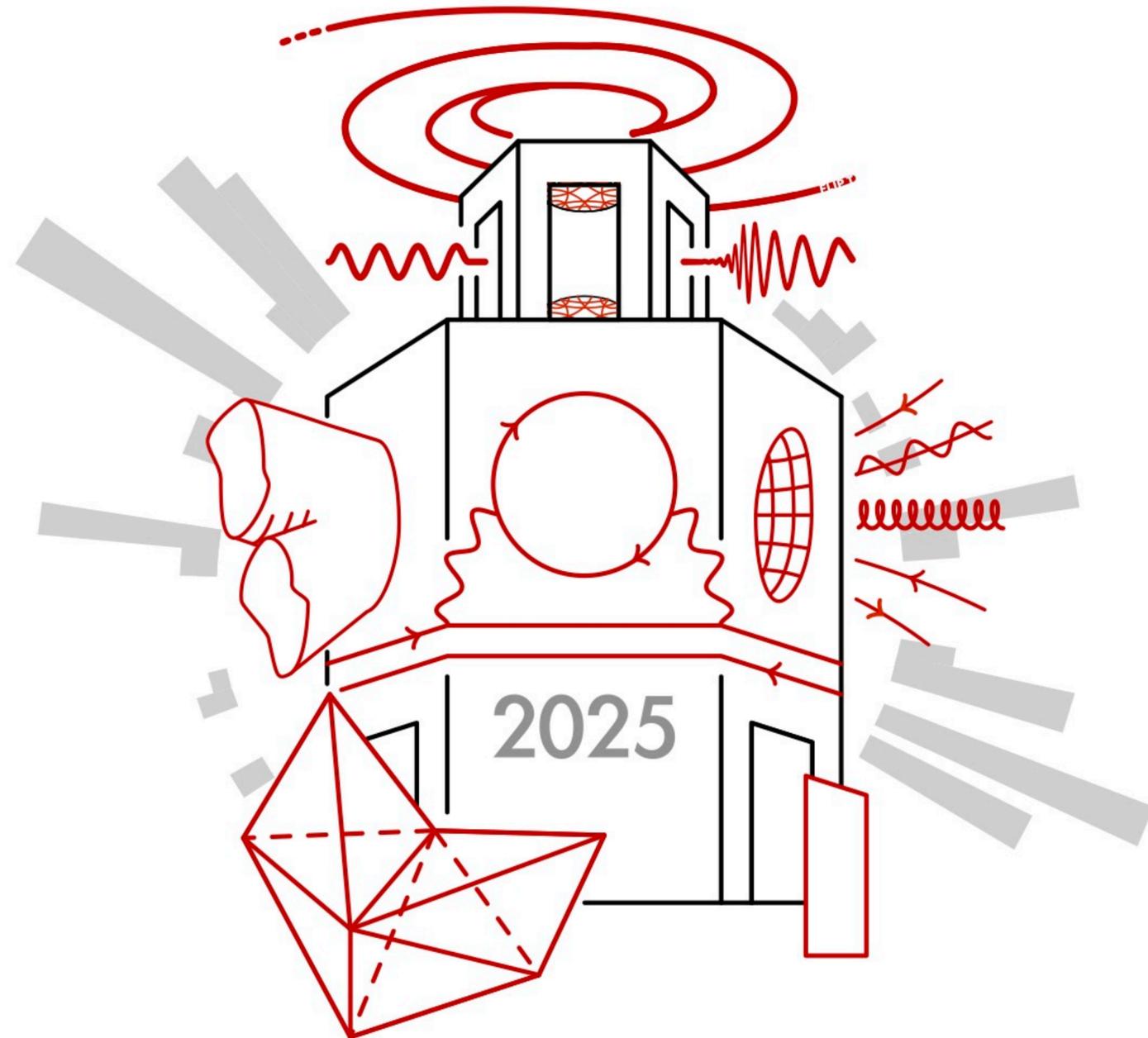


What is particle theory?



What is ~~particle theory~~ physics?

Identifying the fundamental forces and building blocks of the universe

Understanding why they have the properties that we observe



Overview of SUSY results: GMSB / GGM

137 fb⁻¹ (13 TeV)

pp → g \tilde{g}

$\tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0 \rightarrow q\bar{q}\gamma\tilde{G}$ $\gamma + ME_T$: arXiv:1711.08008 (max. exclusion) [36 fb⁻¹]
 $\gamma + H_T$: arXiv:1707.06193 (max. exclusion) [36 fb⁻¹]
 $\gamma\gamma$: arXiv:1903.07070 (max. exclusion) [36 fb⁻¹]

$(q\bar{q}\tilde{\chi}_1^0 \rightarrow q\bar{q}\gamma\tilde{G}/q\bar{q}\tilde{\chi}_1^\pm \rightarrow q\bar{q}W\tilde{G})$

$\gamma + ME_T$: arXiv:1711.08008 (max. exclusion) [36 fb⁻¹]
 $\gamma + H_T$: arXiv:1707.06193 (max. exclusion) [36 fb⁻¹]
 $\gamma + \ell + ME_T$: arXiv:1812.04066 (max. exclusion) [36 fb⁻¹]
combined: arXiv:1907.00857 (max. exclusion) [36 fb⁻¹]

$\tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow (\gamma/H)\tilde{G}$

$\gamma + b + ME_T$: SUS-21-009 (max. exclusion)

$\tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow (\gamma/Z)\tilde{G}$

$\gamma + b + ME_T$: SUS-21-009 (max. exclusion)

$\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow (\gamma/Z)\tilde{G}$

$\gamma + b + ME_T$: SUS-21-009 (max. exclusion)

$\tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}$

2 ℓ opposite-sign: arXiv:2012.08600 (max. exclusion)

pp → q \bar{q}

$\tilde{q} \rightarrow q\tilde{\chi}_1^0 \rightarrow q\gamma\tilde{G}$ $\gamma + ME_T$: arXiv:1711.08008 (max. exclusion) [36 fb⁻¹]
 $\gamma + H_T$: arXiv:1707.06193
 $\gamma\gamma$: arXiv:1903.07070

$\tilde{q} \rightarrow (q\tilde{\chi}_1^0 \rightarrow q\gamma\tilde{G}/q\tilde{\chi}_1^\pm \rightarrow qW\tilde{G})$

$\gamma + ME_T$: arXiv:1711.08008
 $\gamma + H_T$: arXiv:1707.06193
 $\gamma + \ell + ME_T$: arXiv:1812.04066

pp → t \bar{t}

$\tilde{t} \rightarrow t\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow (\gamma/Z)\tilde{G}$ $\gamma + b + ME_T$: SUS-21-009

pp → $\tilde{\chi}_1^0\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \gamma\tilde{G}, \tilde{\chi}_1^\pm \rightarrow W\tilde{G}$

$\gamma + \ell + ME_T$: arXiv:1812.04066
 $\gamma + b + ME_T$: SUS-21-009
 $\gamma + ME_T$: arXiv:1711.08008
 $\gamma + b + ME_T$: SUS-21-009

Higgsino – like NLSPs

pp → $(\tilde{\chi}_1^\pm, \tilde{\chi}_2^0, \tilde{\chi}_1^0)(\tilde{\chi}_1^\pm, \tilde{\chi}_2^0, \tilde{\chi}_1^0)$

$\tilde{h} \rightarrow b\bar{b}$: arXiv:2201.04206
 $h \rightarrow \gamma\gamma$: arXiv:1908.08500 [78 fb⁻¹]
combined: SUS-21-008
2 ℓ opposite-sign: arXiv:2012.08600
 $\geq 3\ell/\tau_h$: arXiv:2106.14246 BF = 50%
 $h \rightarrow \gamma\gamma$: arXiv:1908.08500 BF = 50%
combined: SUS-21-008 BF = 50%
2 ℓ opposite-sign: arXiv:2012.08600
 $\geq 3\ell/\tau_h$: arXiv:2106.14246
combined: SUS-21-008

pp → $\tilde{\chi}_1^0\tilde{\chi}_2^0, \tilde{\chi}_1^0 \rightarrow (h/Z)\tilde{G}, \tilde{\chi}_2^0 \rightarrow \gamma\tilde{G}$

$h \rightarrow b\bar{b}$: arXiv:2201.04206

pp → $\tilde{\chi}_1^0\tilde{\chi}_2^0 \rightarrow (h/Z)\tilde{G}, \tilde{\chi}_2^0 \rightarrow \gamma\tilde{G}$

$h \rightarrow b\bar{b}$: arXiv:2201.04206

pp → $(\tilde{\chi}_2^0, \tilde{\chi}_1^0)$

$\gamma + b + ME_T$: SUS-21-009 BF = 50%

pp → $\tilde{\chi}_1^0\tilde{\chi}_2^0 \rightarrow hh\tilde{G}\tilde{G}$

$h \rightarrow b\bar{b}$: arXiv:2201.04206

pp → $(\tilde{\chi}_2^0, \tilde{\chi}_1^0)$

$\gamma + b + ME_T$: SUS-21-009 BF = 50%

pp → $\tilde{\chi}_1^0\tilde{\chi}_2^0 \rightarrow hh\tilde{G}\tilde{G}$

$h \rightarrow b\bar{b}$: arXiv:2201.04206

mass scale [GeV]

of observed limits at 95% C.L. (theory uncertainties are not included). Probe up to the quoted mass limit for light LSPs unless stated otherwise.

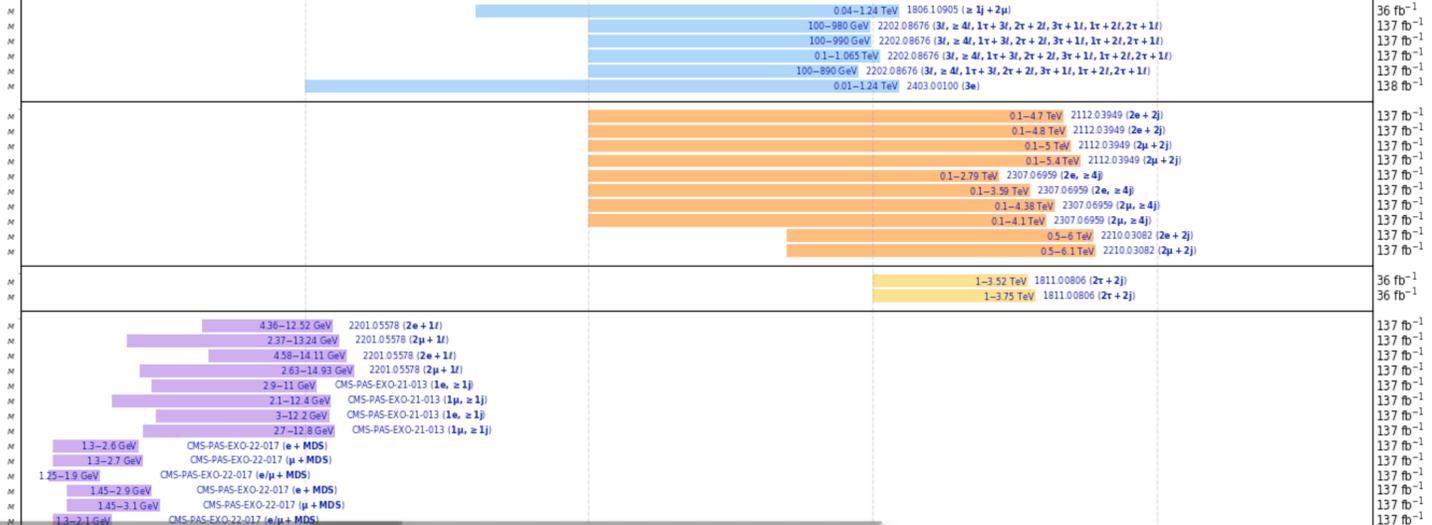
LHC explores in huge number of directions.

SUSY RPV

SUSY RPC



Overview of CMS HNL results



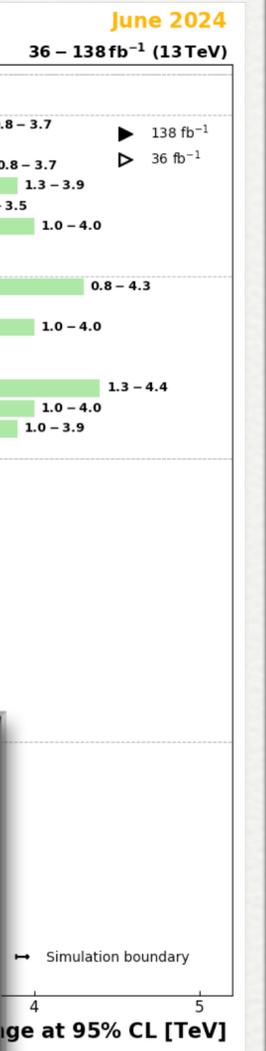
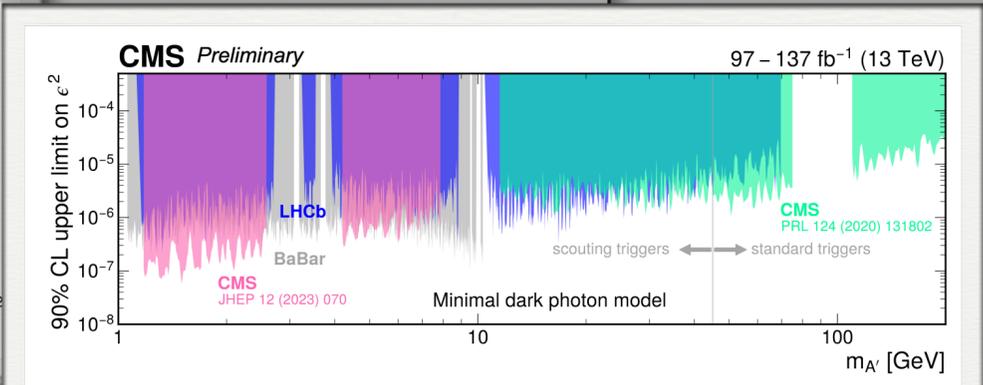
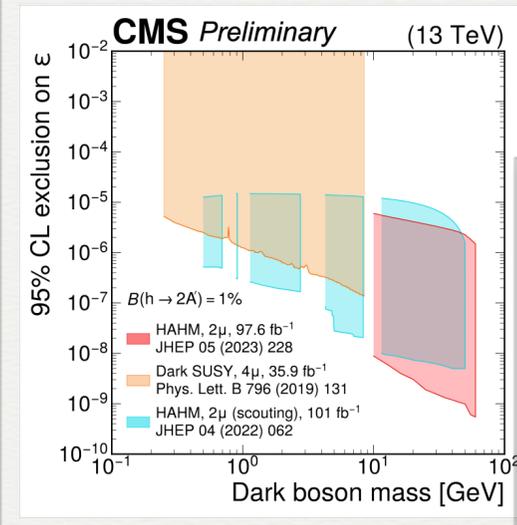
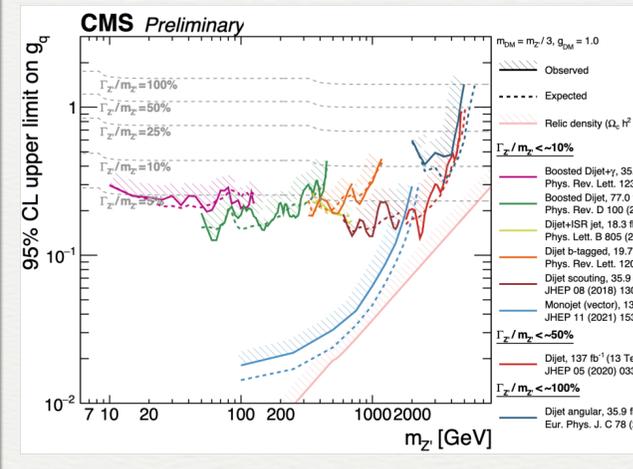
Multiplication
 vMSM, $|V_{cb}|^2 = 1.0$, $|V_{ub}|^2 = 1.0$
 Type-III Seesaw Heavy Fermions, Flavor Democratic
 Type-III Seesaw Heavy Fermions, $B_\mu = 1.0$, $B_\tau = 0.0$
 Type-II Seesaw Heavy Fermions, $B_\mu = 1.0$, $B_\tau = 0.0$
 Type-II Seesaw Heavy Fermions, $B_\mu = 1.0$, $B_\tau = 0.0$
 vMSM, $|V_{cb}|^2 = 1.0$, $|V_{ub}|^2 = 0.0$

Direction jets
 LRSM $W(eN)$, $M_{N_i} < M_{N_j}$ ($= 200\text{GeV}$)
 LRSM $W(eN)$, $M_{N_i} = 0.5M_{N_j}$
 LRSM $W(\mu N)$, $M_{N_i} < M_{N_j}$ ($= 200\text{GeV}$)
 LRSM $W(\mu N)$, $M_{N_i} = 0.5M_{N_j}$
 LRSM $Z(eN)$, $M_{N_i} < 0.5M_{N_j}$ ($= 100\text{GeV}$)
 LRSM $Z(eN)$, $M_{N_i} = 0.25M_{N_j}$
 LRSM $Z(\mu N)$, $M_{N_i} < 0.5M_{N_j}$ ($= 100\text{GeV}$)
 LRSM $Z(\mu N)$, $M_{N_i} = 0.25M_{N_j}$
 Composite Fermions N_i , $M_{N_i} < \Lambda$
 Composite Fermions N_i , $M_{N_i} < \Lambda$

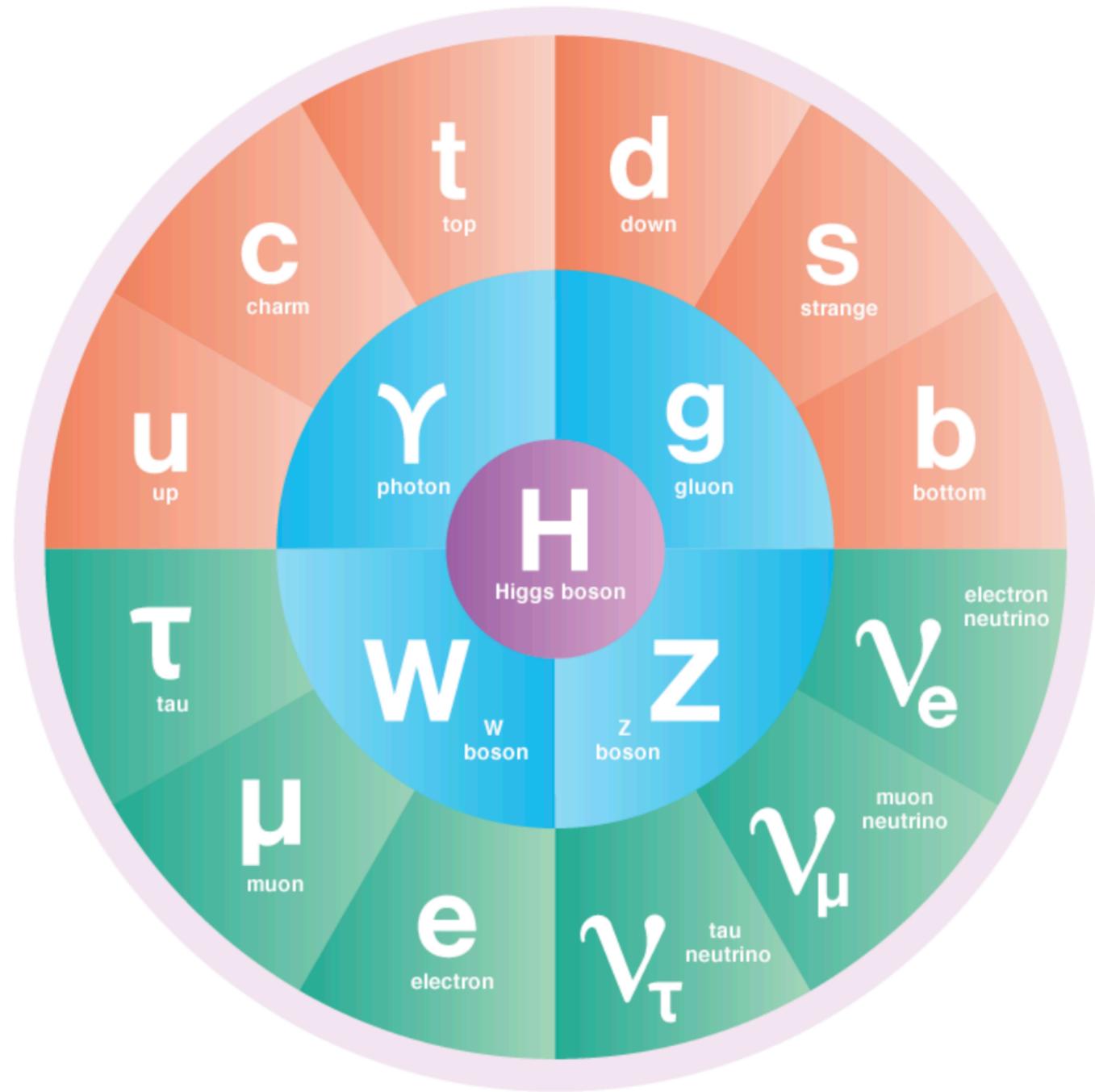
Self Interactions
 LRSM $W(eN)$, $M_{N_i} = 0.8M_{N_j}$
 LRSM $W(eN)$, $M_{N_i} = 0.2M_{N_j}$

Displaced
 Displaced Majorana HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$
 Displaced Majorana HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$
 Displaced Dirac HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$
 Displaced Dirac HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$
 Displaced Majorana HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$
 Displaced Majorana HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$
 Displaced Dirac HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$
 Displaced Dirac HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$
 Displaced Majorana HNL, $|V_{cb}|^2 = 5.0 \times 10^{-3}$
 Displaced Majorana HNL, $|V_{cb}|^2 = 5.0 \times 10^{-3}$
 Displaced Dirac HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$
 Displaced Dirac HNL, $|V_{cb}|^2 = 5.0 \times 10^{-3}$
 Displaced Dirac HNL, $|V_{cb}|^2 = 5.0 \times 10^{-3}$
 Displaced Dirac HNL, $|V_{cb}|^2 = 1.0 \times 10^{-3}$

Selector



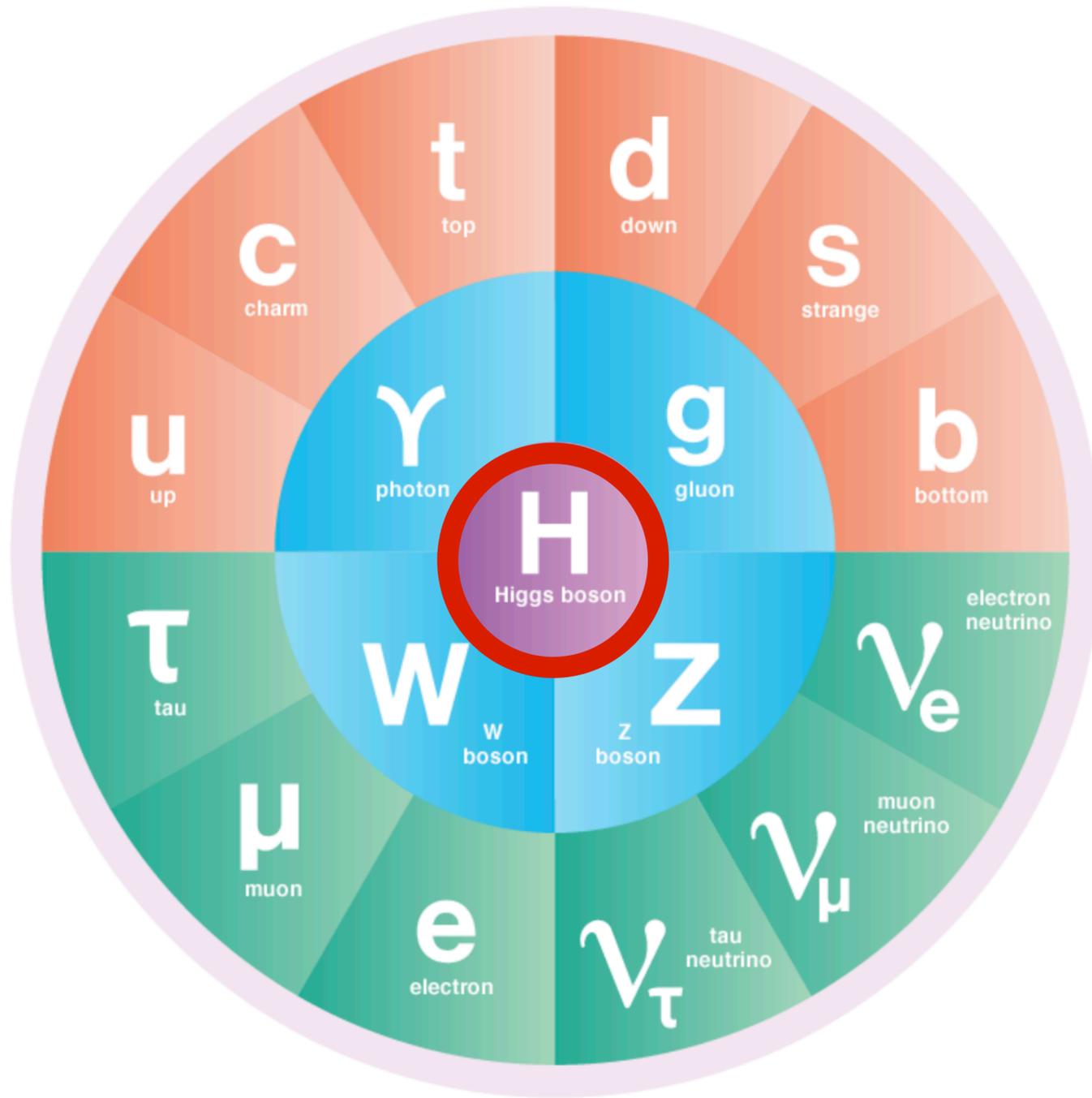
The Standard Model (SM)



particles

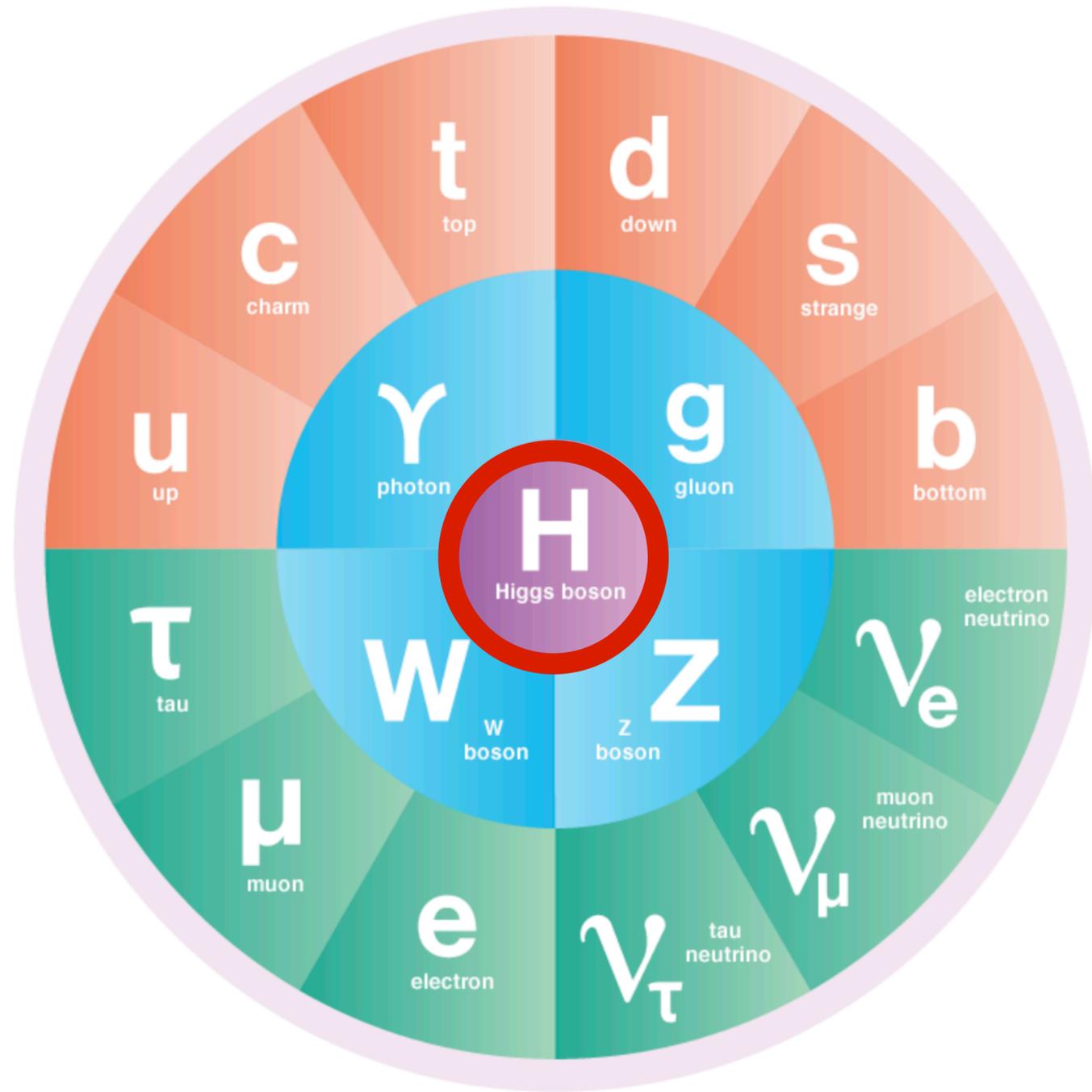
The Standard Model (SM)

**“the standard-model (SM)
is complete”**



particles

The Standard Model (SM)

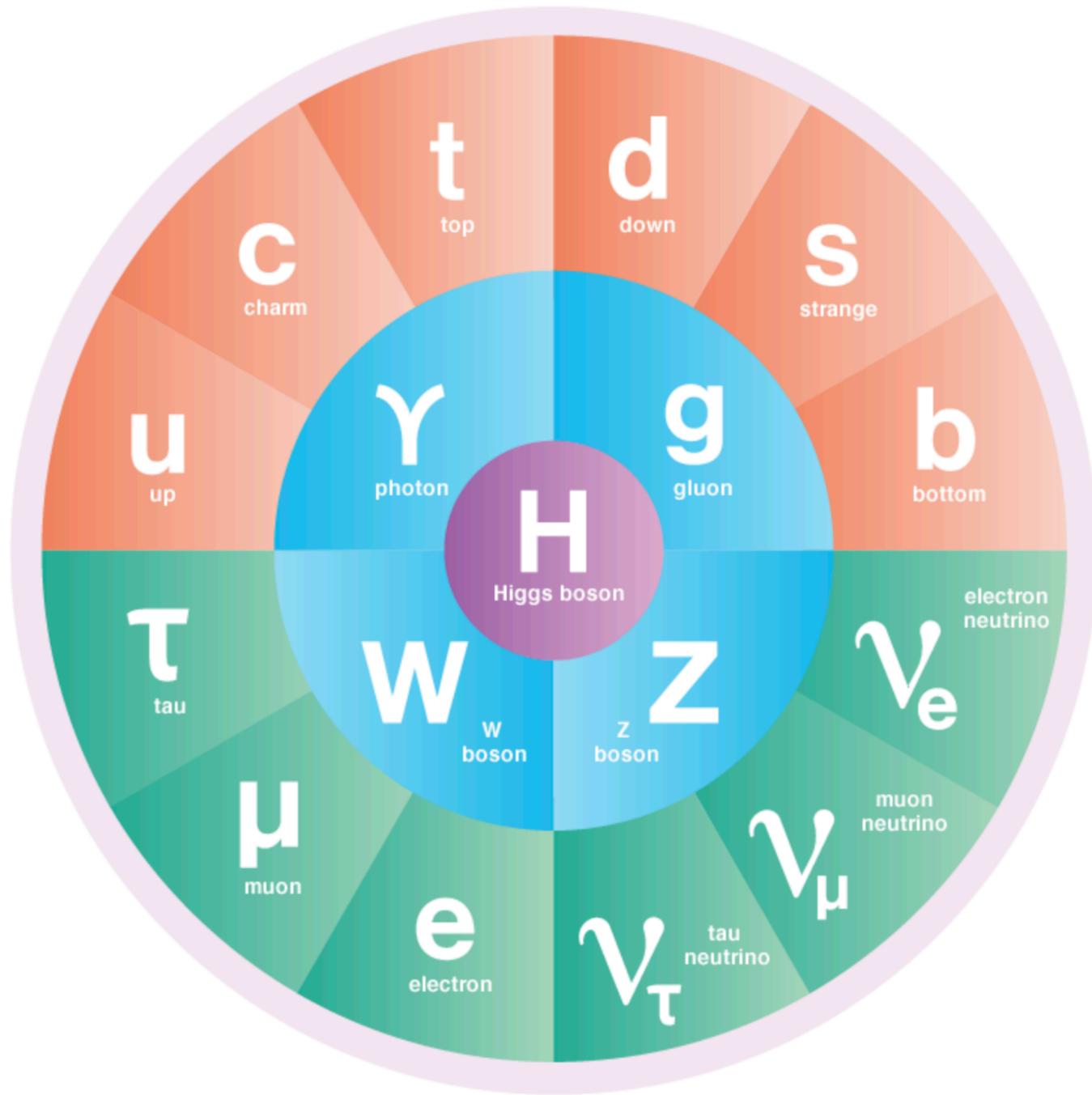


particles

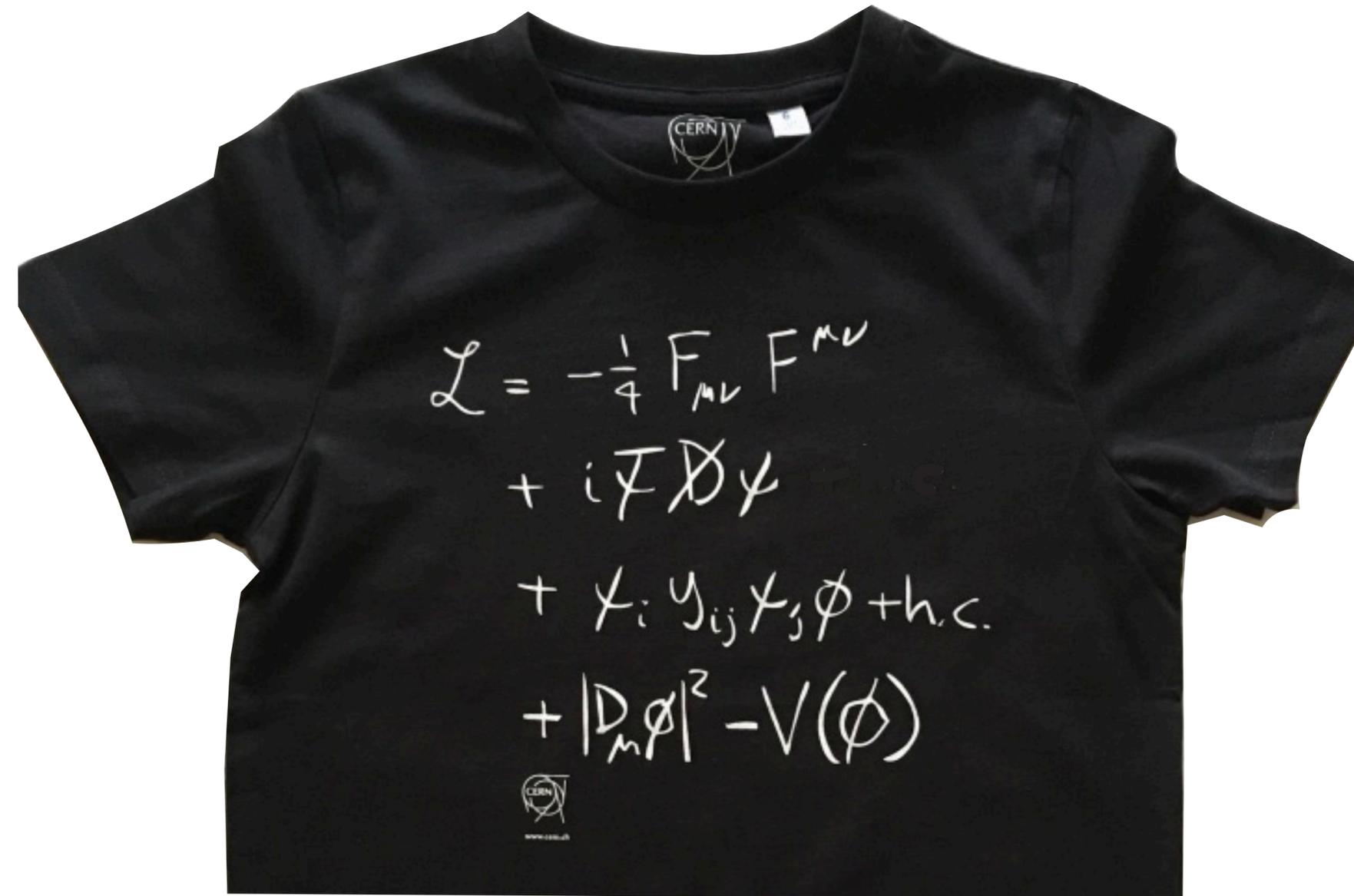
**“the standard-model (SM)
is complete”**



The Standard Model (SM)



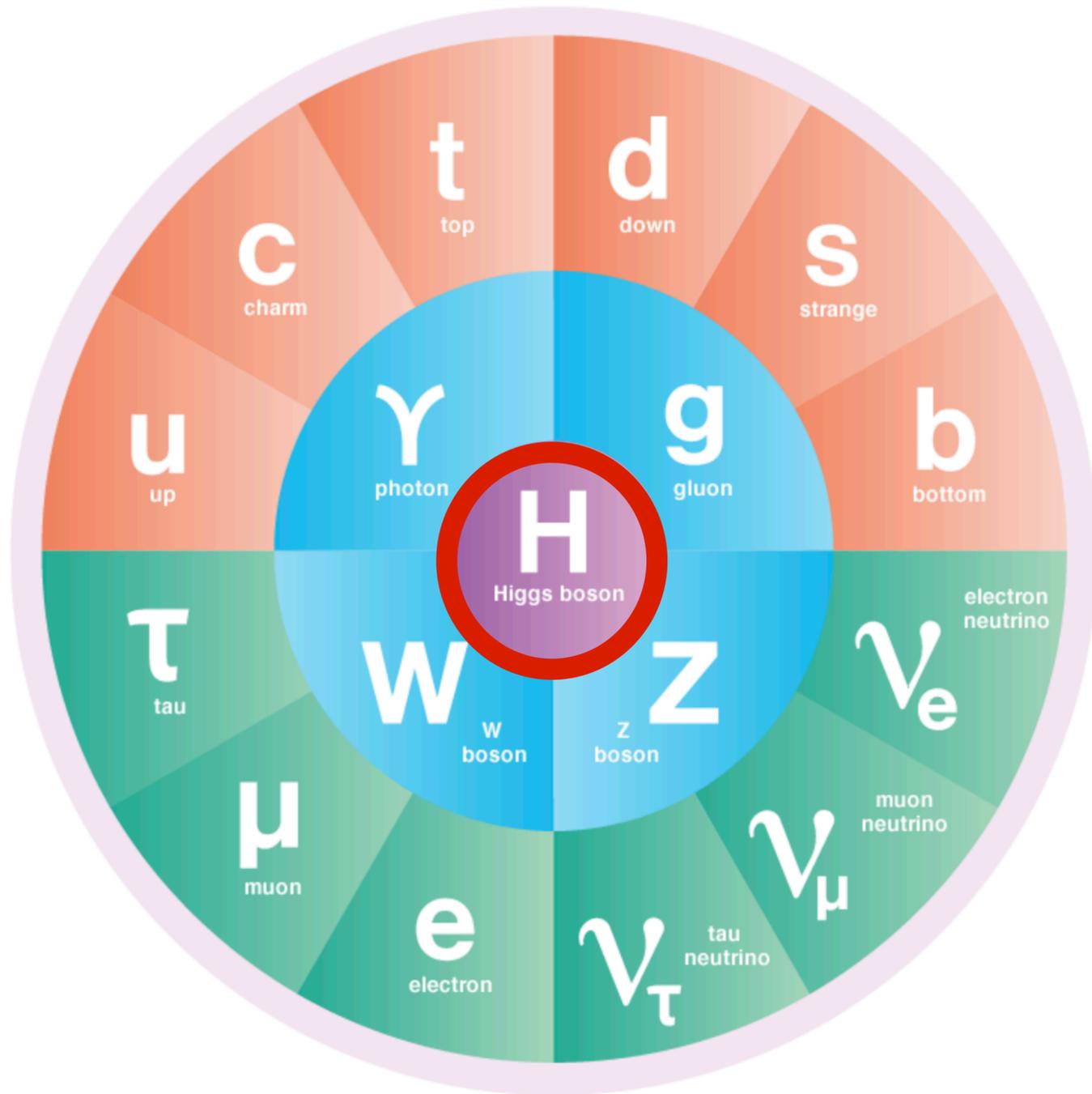
particles



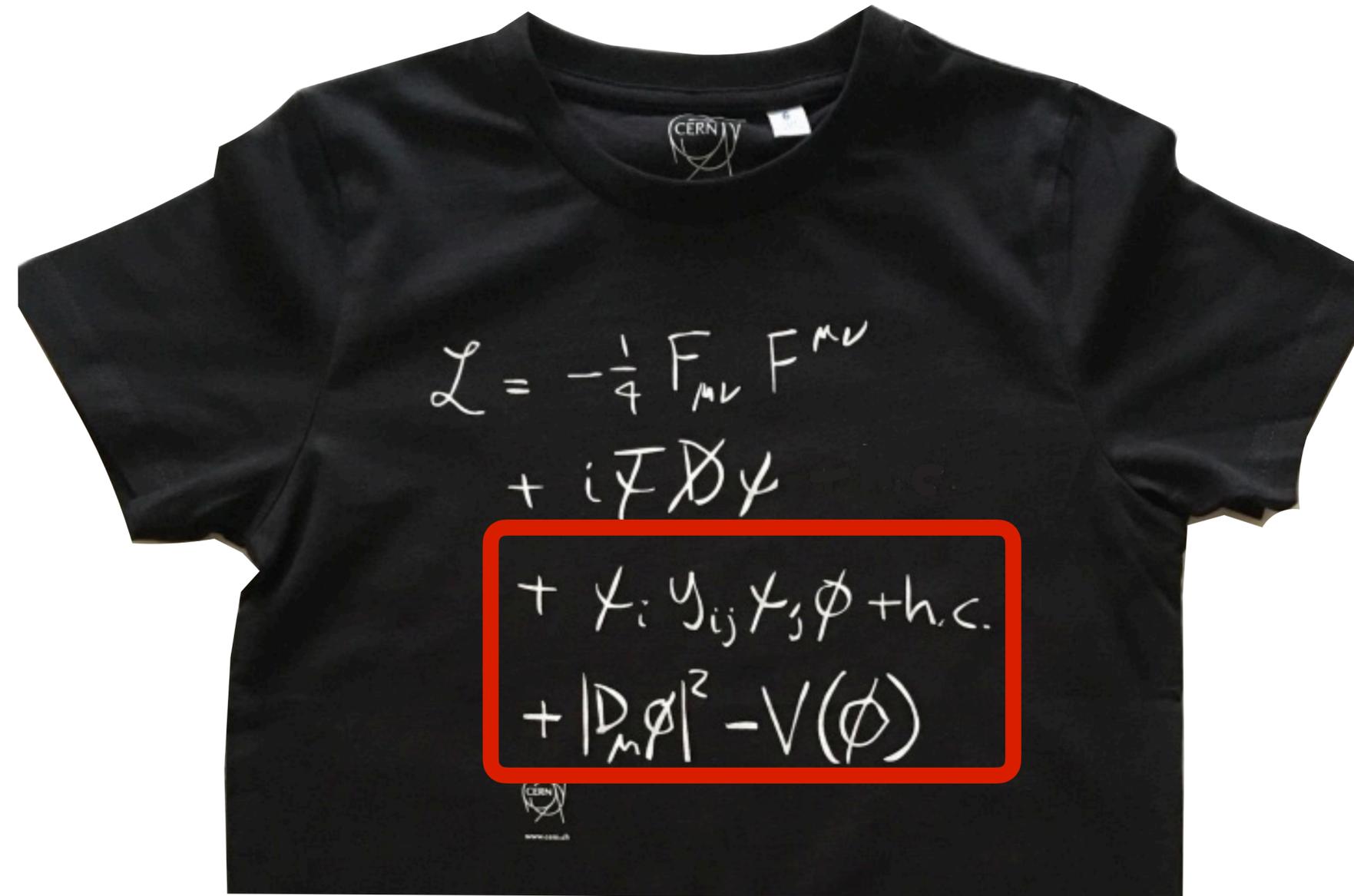
interactions

<https://www.symmetrymagazine.org/standard-model/>

The Standard Model (SM)



particles



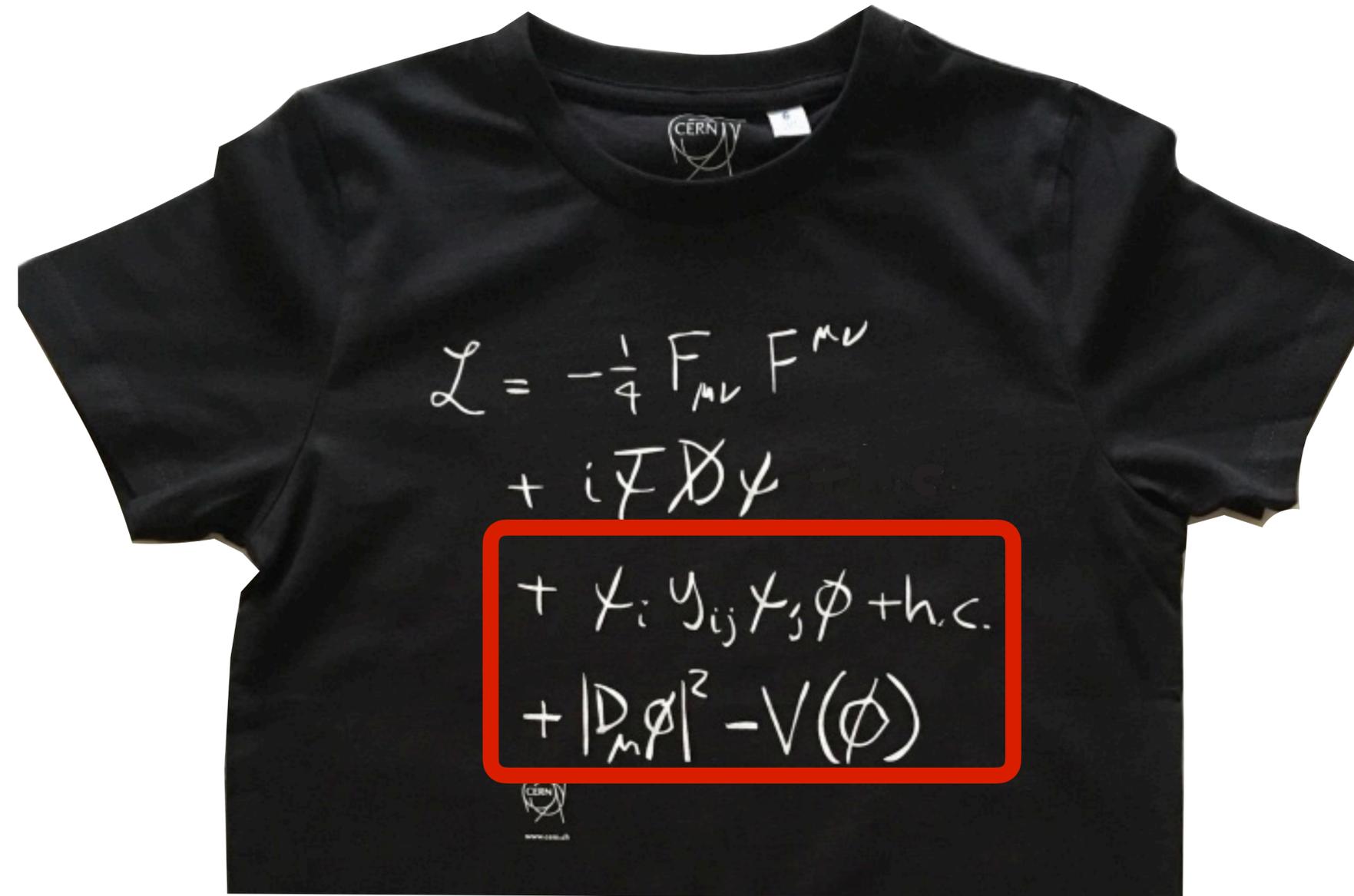
interactions

The Standard Model (SM)

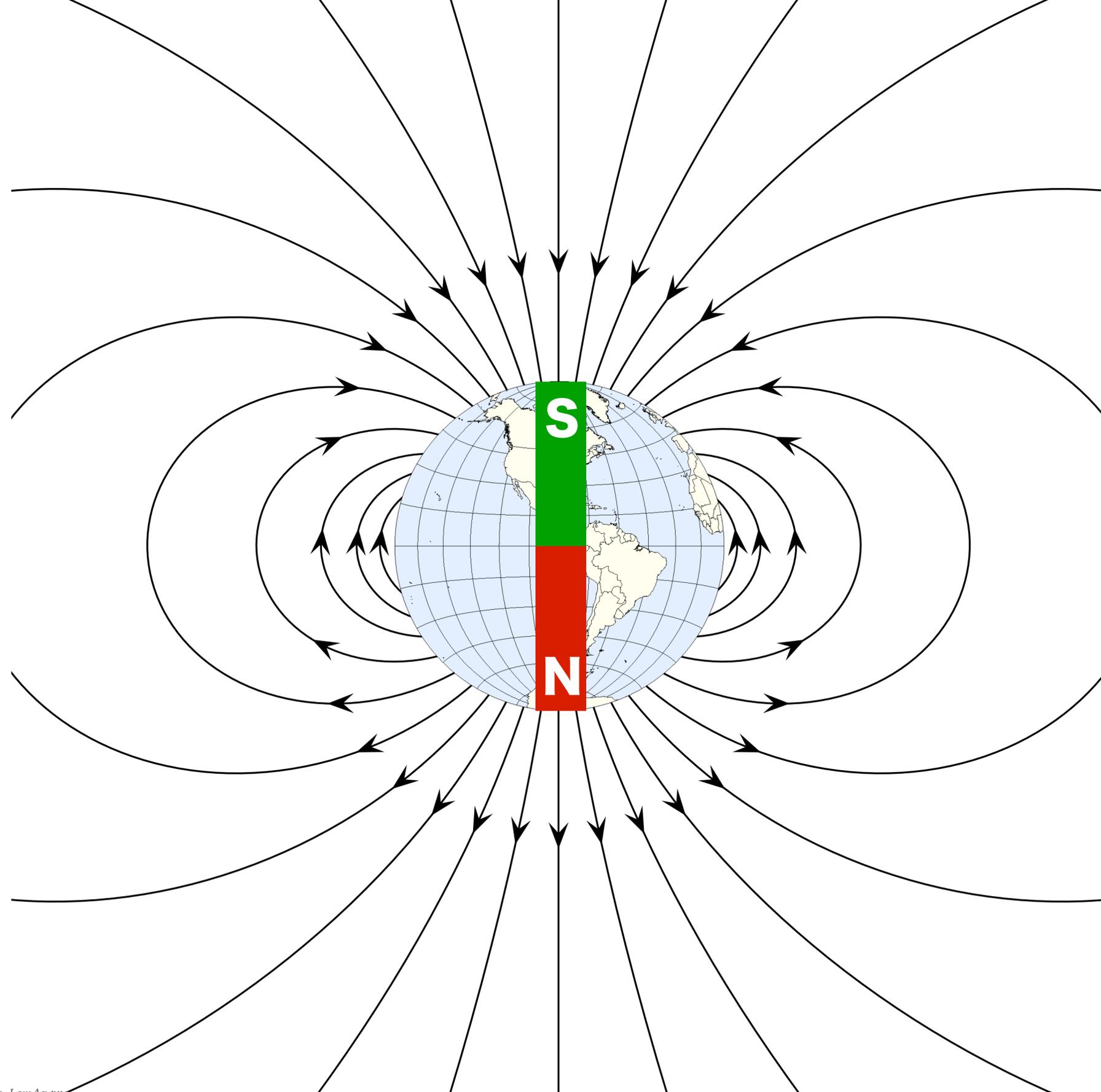
**our experimental exploration of
the Higgs-related SM
interactions is only just starting**

*parts of this talk adapted from
“The Higgs boson turns ten”,
GPS, Zanderighi and Wang*

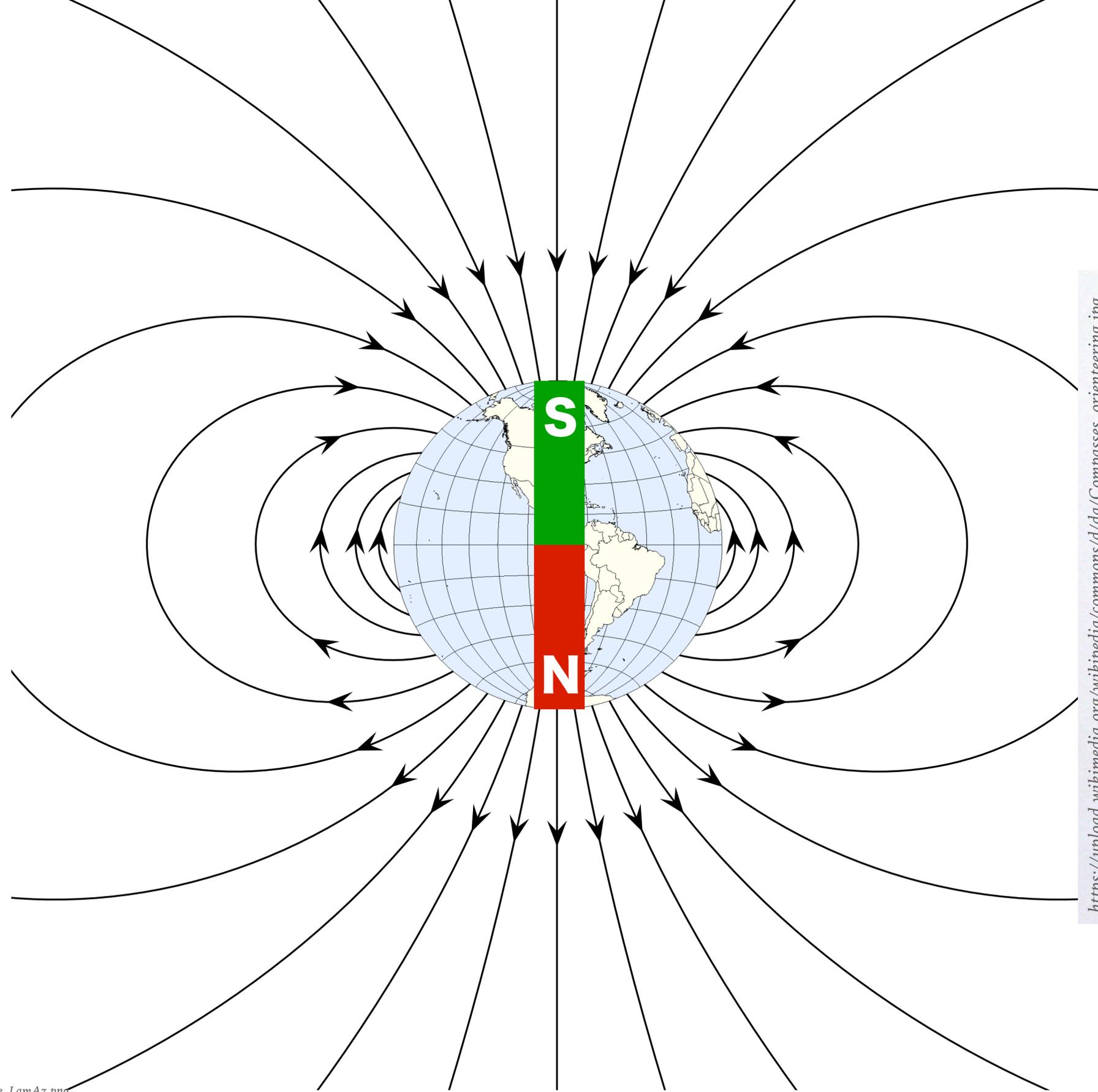
Nature 607 (2022) 7917, 41-47



interactions



https://commons.wikimedia.org/wiki/File:VFpt_Dipole_field.svg
https://en.wikipedia.org/wiki/Western_Hemisphere#/media/File:Western_Hemisphere_LamAz.png



https://commons.wikimedia.org/wiki/File:VFpt_Dipole_field.svg
https://en.wikipedia.org/wiki/Western_Hemisphere#/media/File:Western_Hemisphere_LamAz.png



https://upload.wikimedia.org/wikipedia/commons/d/da/Compasses_orienteering.jpg

HIGGS
FIELD

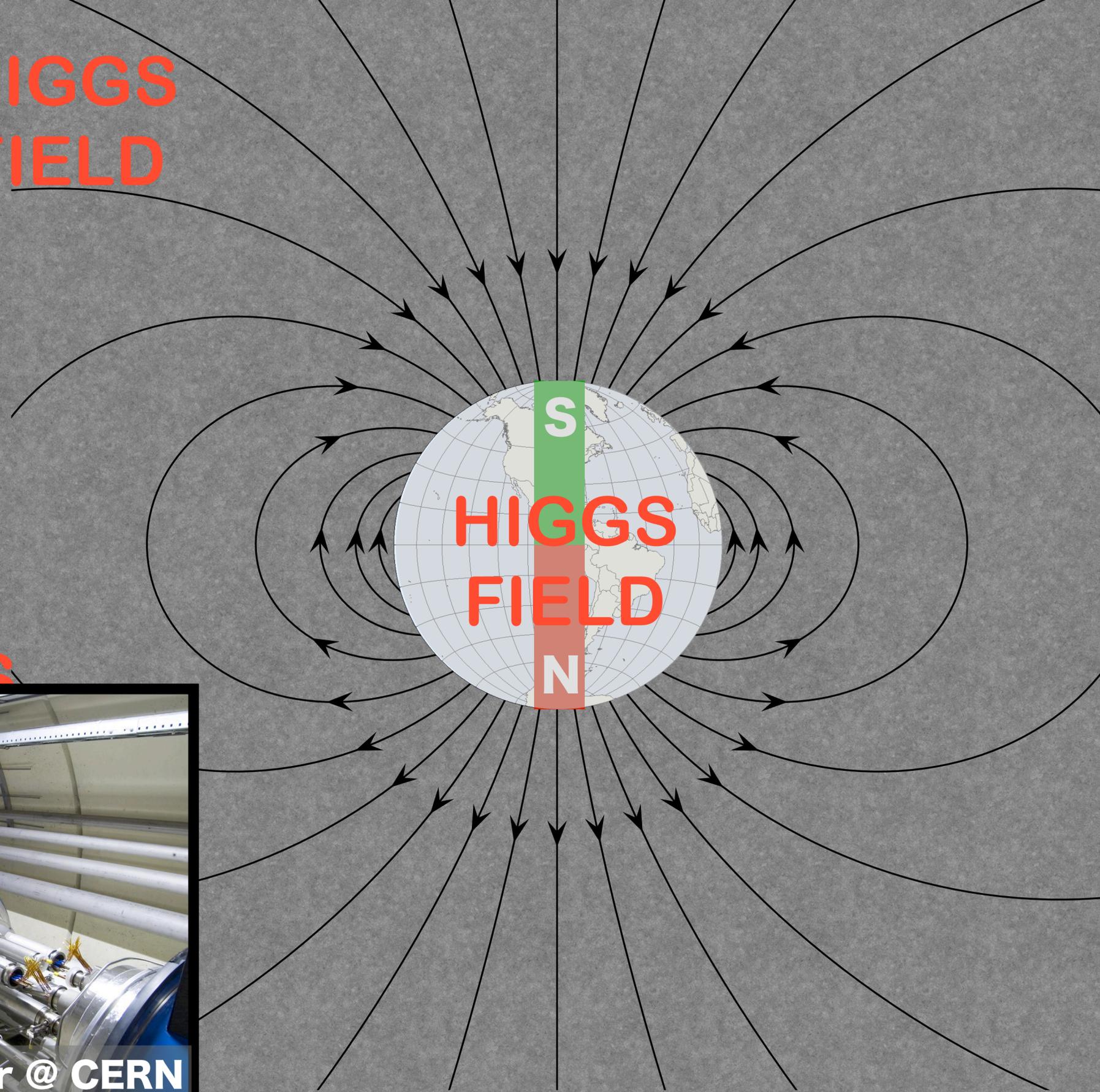
HIGGS
FIELD

HIGGS
FIELD

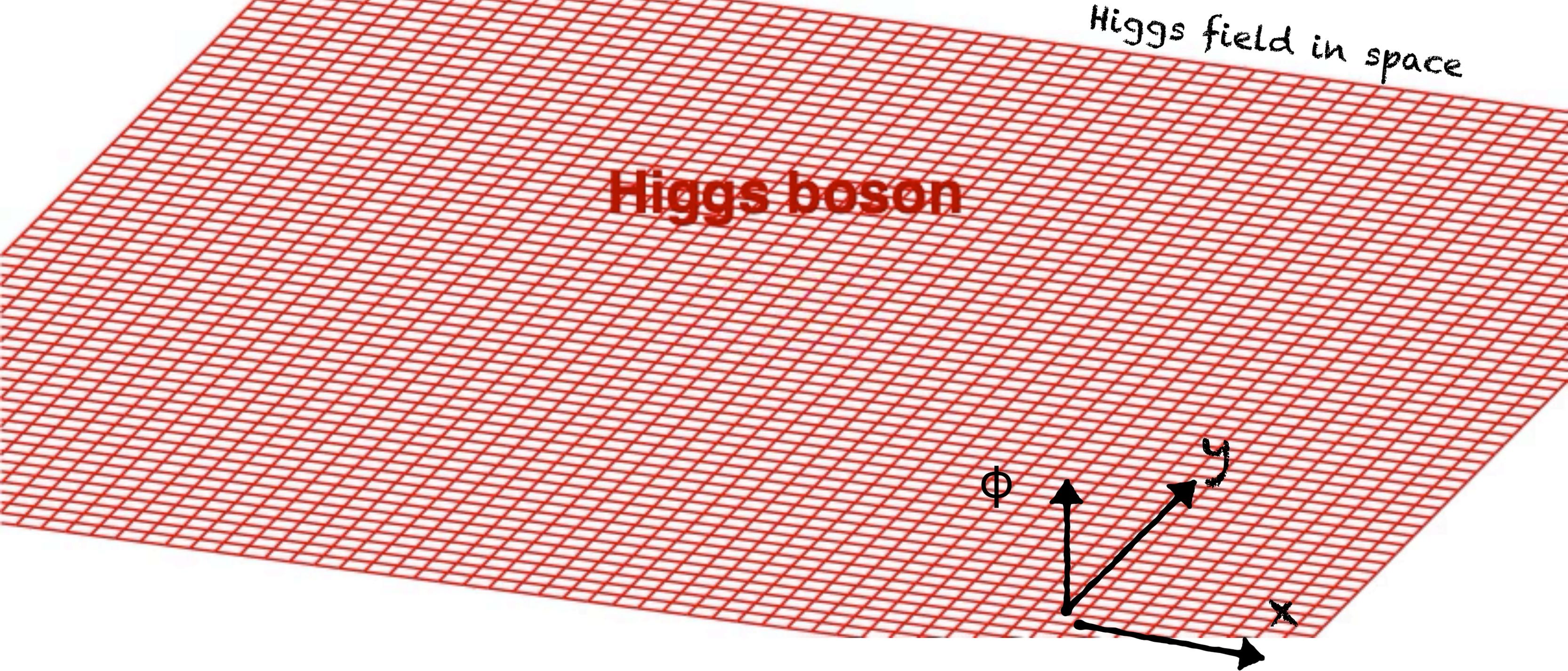
HIGGS

HIGGS
FIELD

HIGGS
FIELD

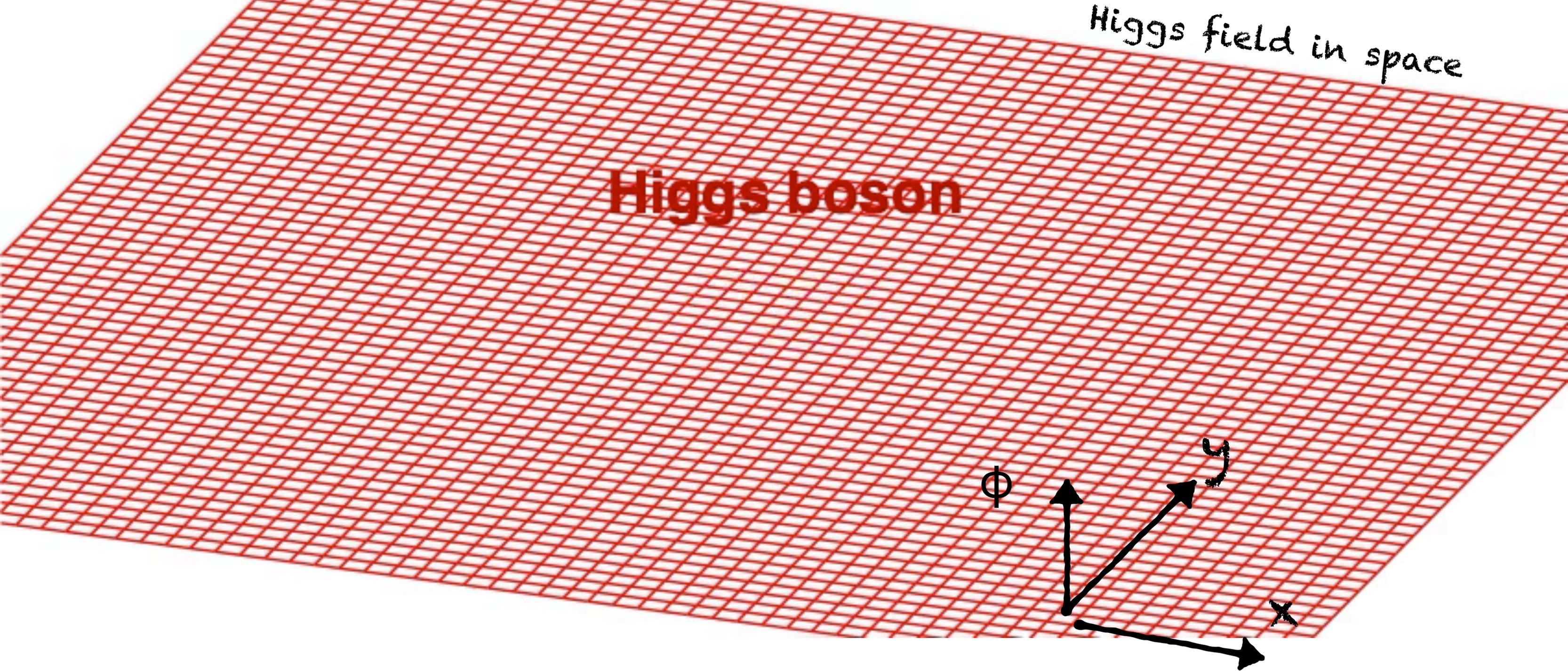


Large Hadron Collider @ CERN



Higgs field (ϕ) can be different at each point in space

A Higgs boson at a given point in space is a fluctuation of the field



Higgs field (ϕ) can be different at each point in space

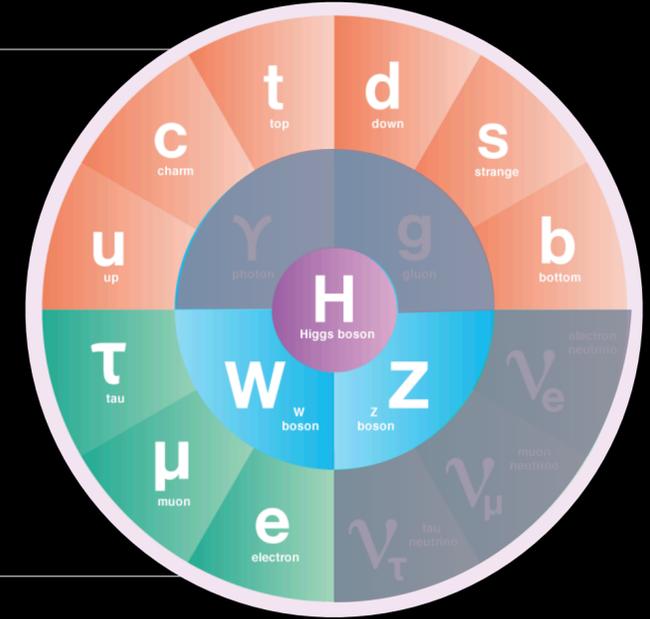
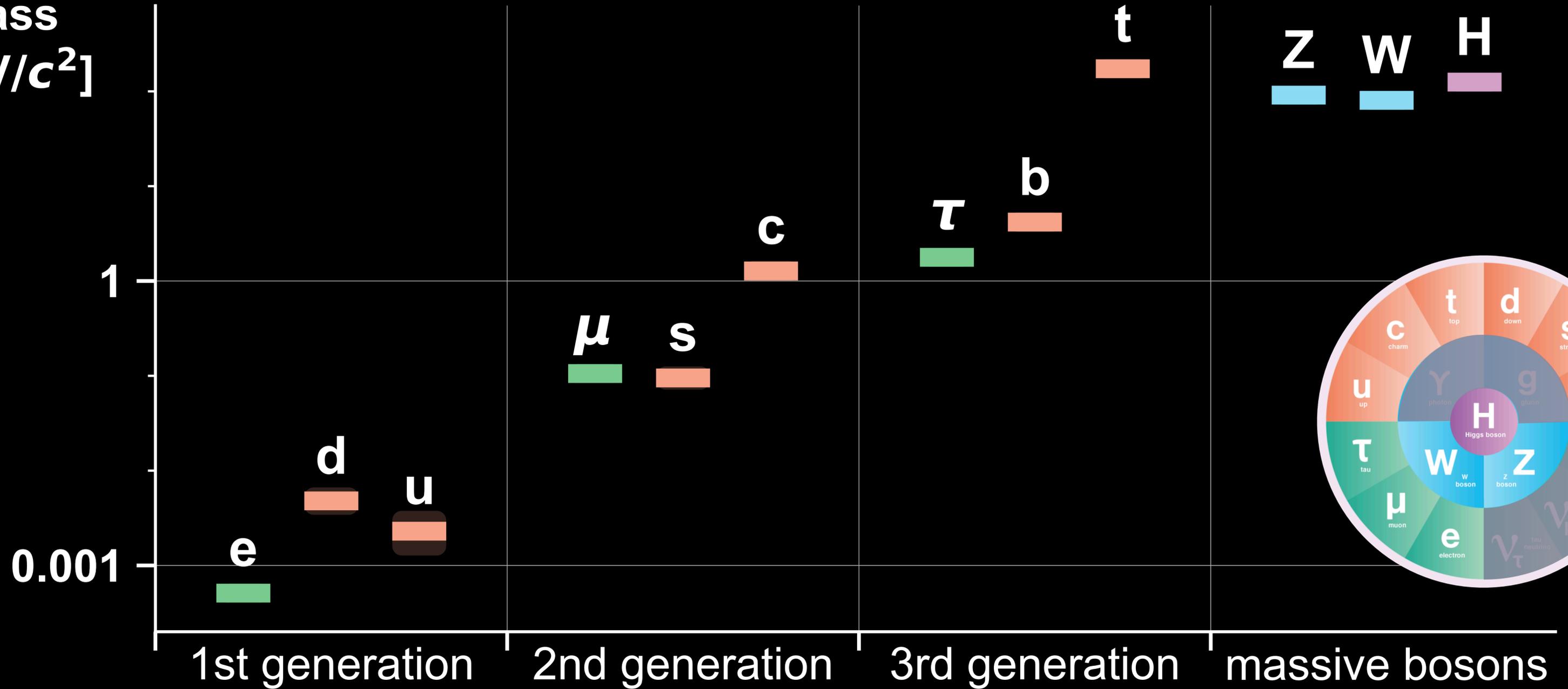
A Higgs boson at a given point in space is a fluctuation of the field

a core hypothesis of Standard Model
fundamental particles get their mass
from interaction with the Higgs field

$$+ \mathcal{L}_i: y_{ij} \bar{\psi}_i \psi_j \phi + h.c.$$

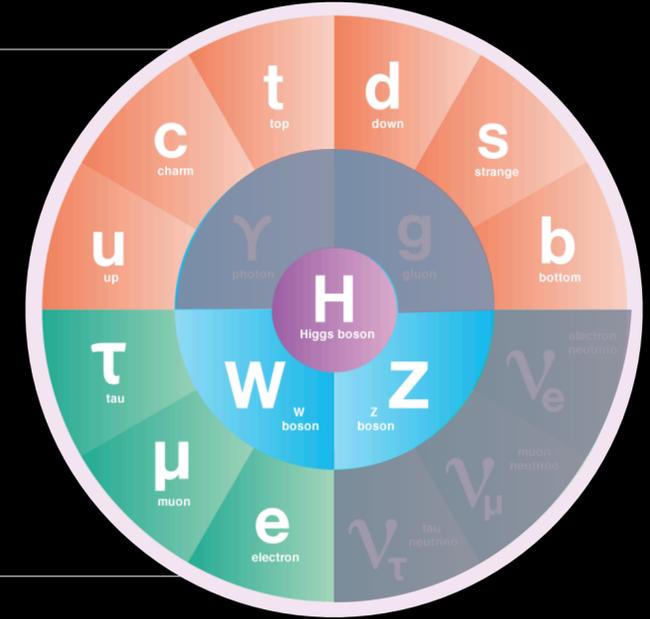
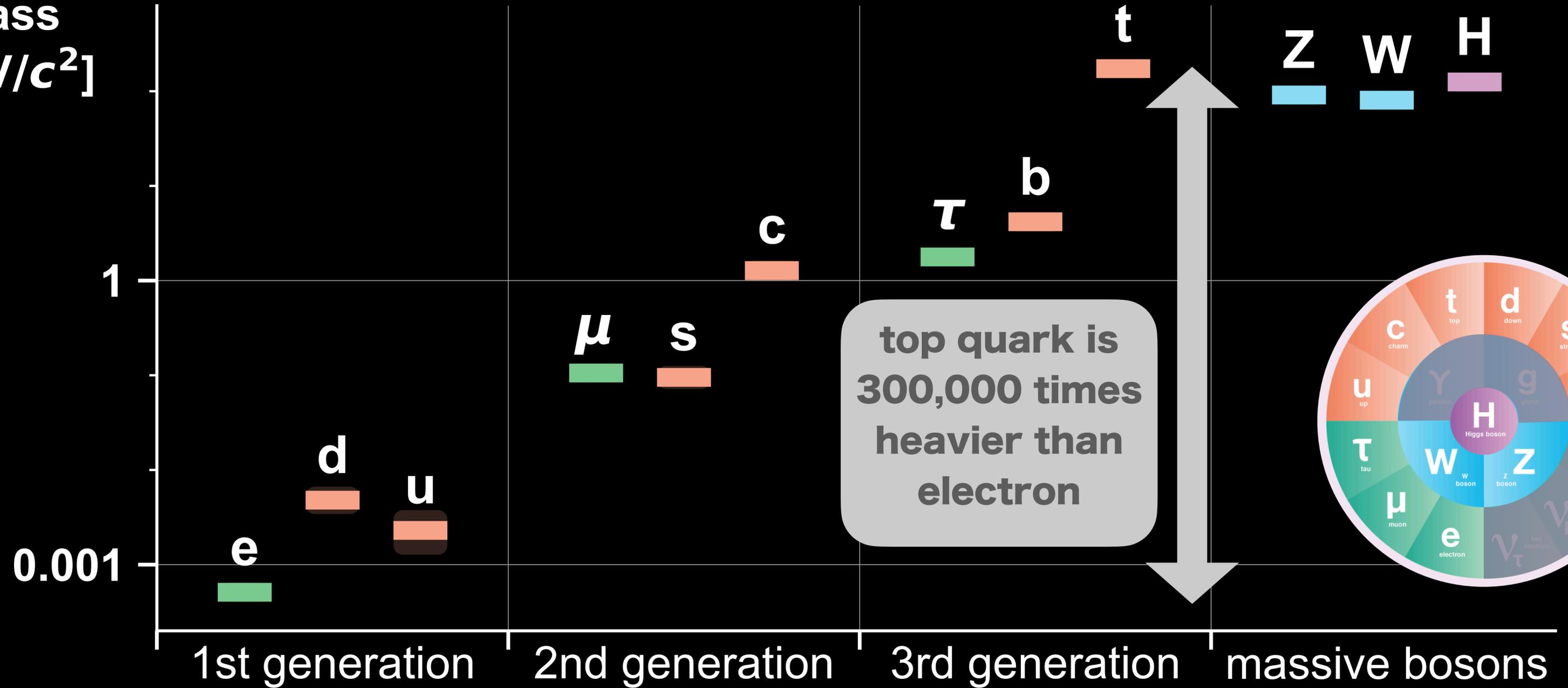
Standard Model massive particles (except ν)

mass
[GeV/c²]



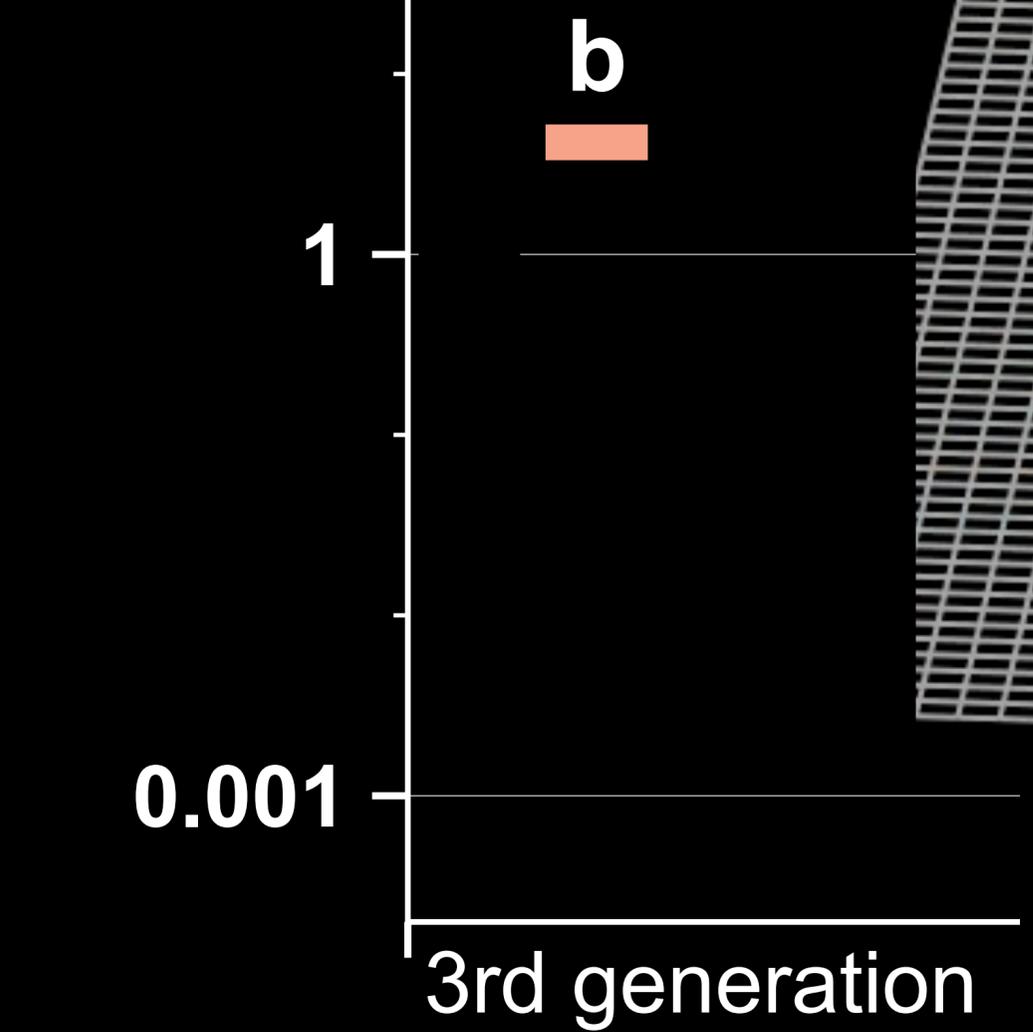
Standard Model massive particles (except ν)

mass
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Higgs field

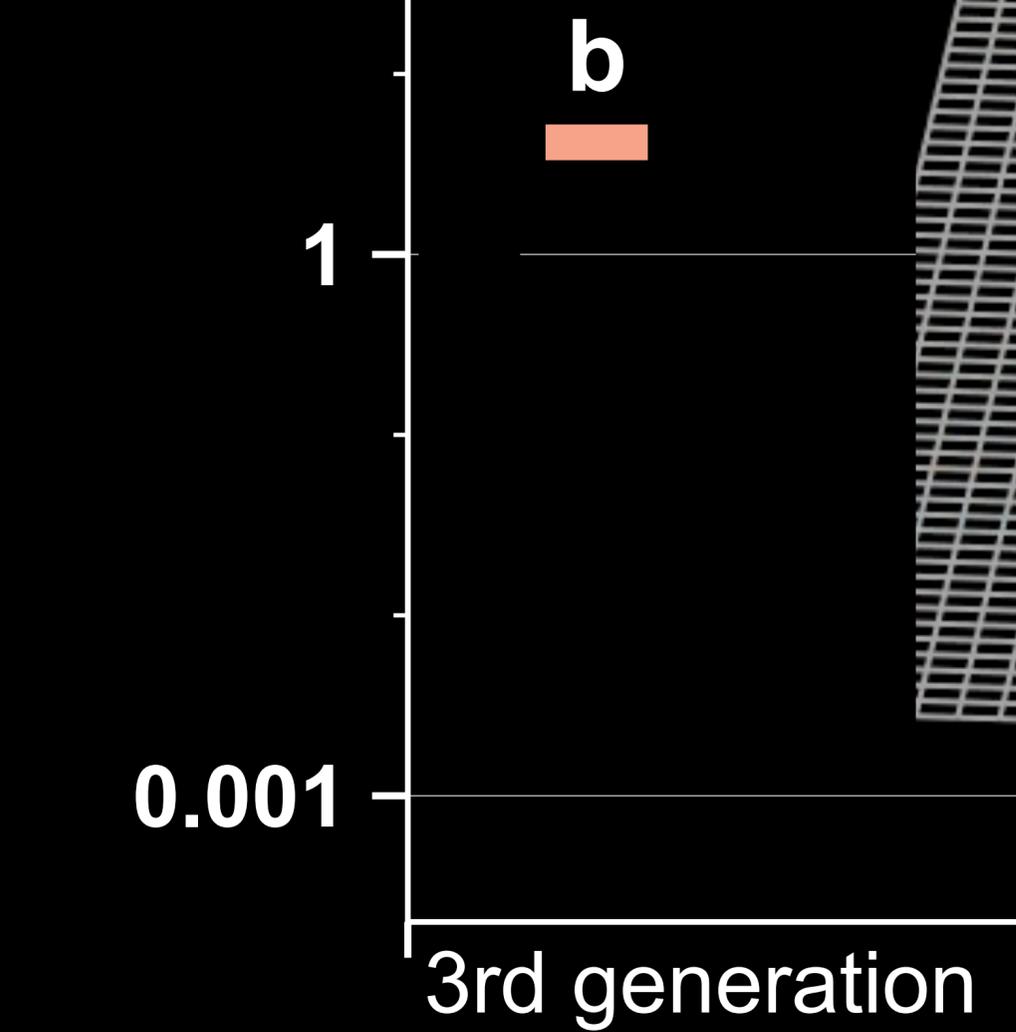
mass
[GeV/c²]



SM: larger mass of top comes from stronger interaction with Higgs field

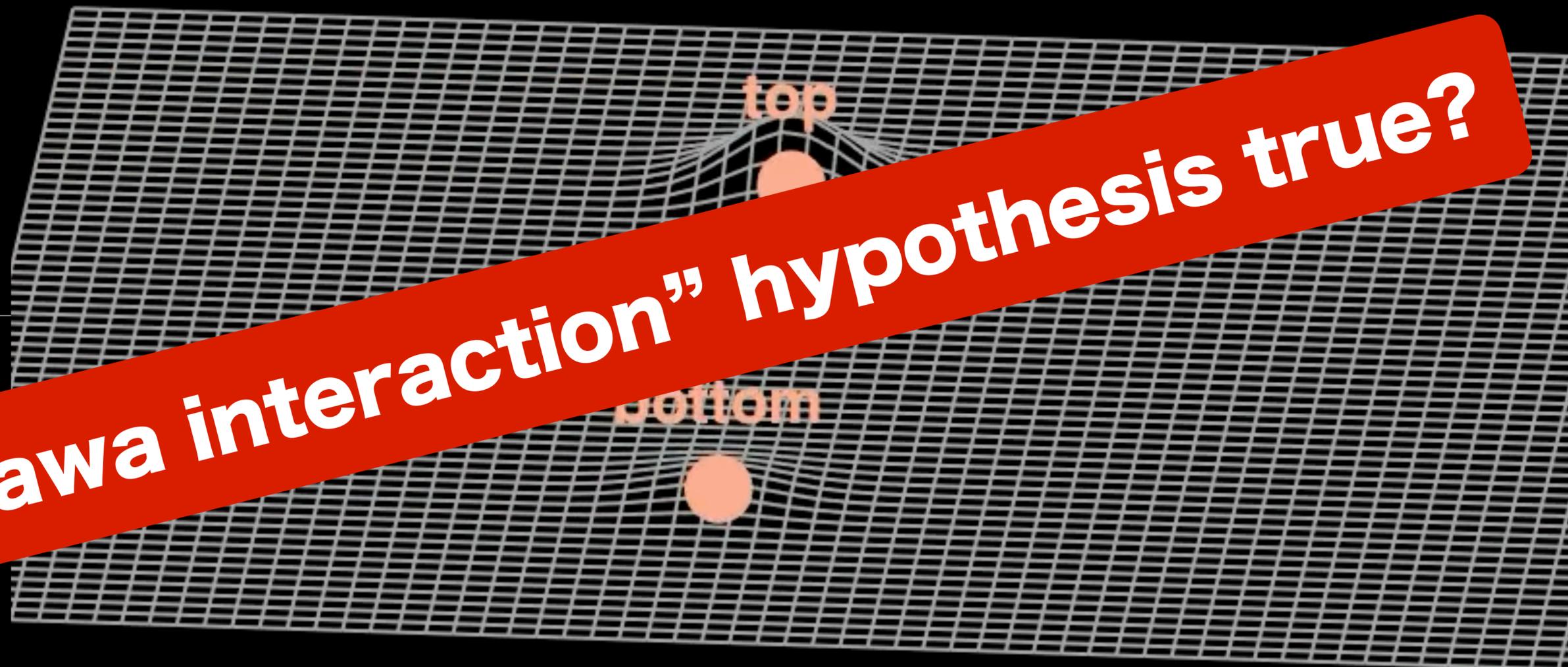
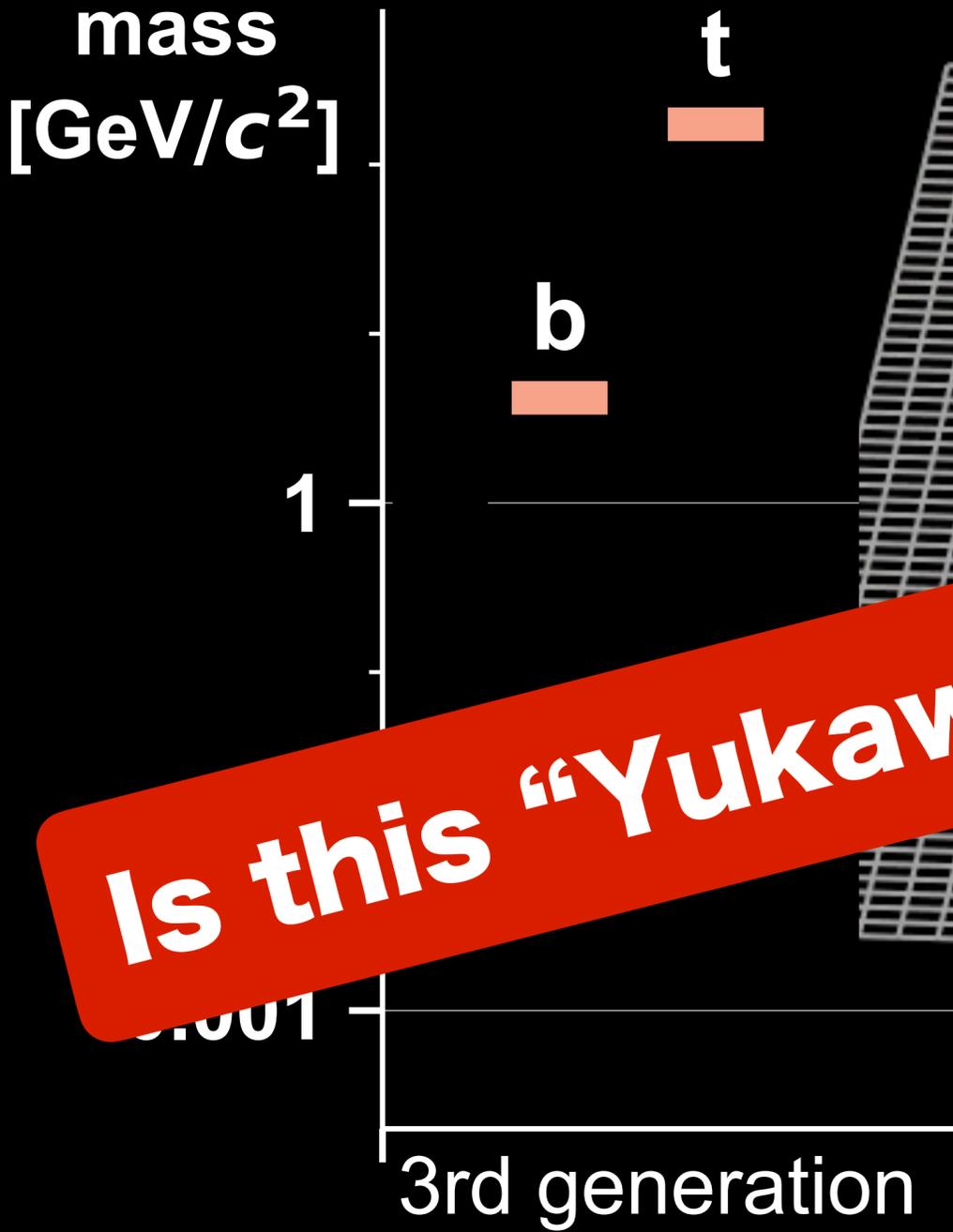
Higgs field

mass
[GeV/c²]



SM: larger mass of top comes from stronger interaction with Higgs field

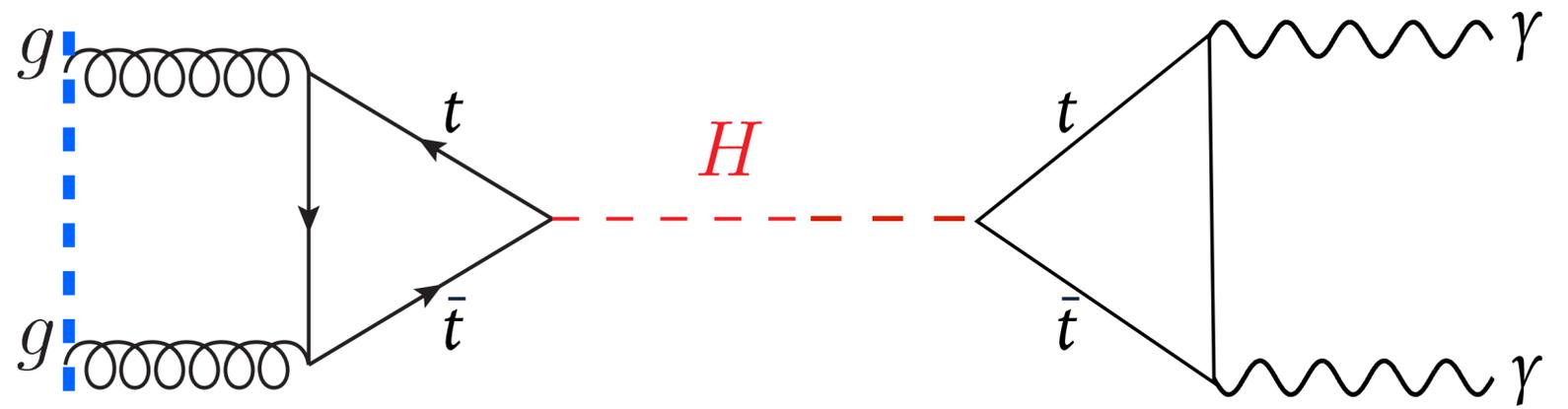
Higgs field



Is this "Yukawa interaction" hypothesis true?

SM: larger mass of top comes from stronger interaction with Higgs field

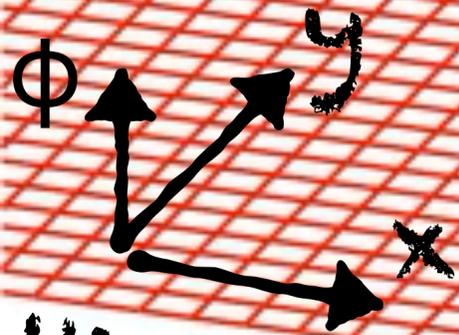
An LHC collision of the kind that led to the Higgs boson discovery



quon

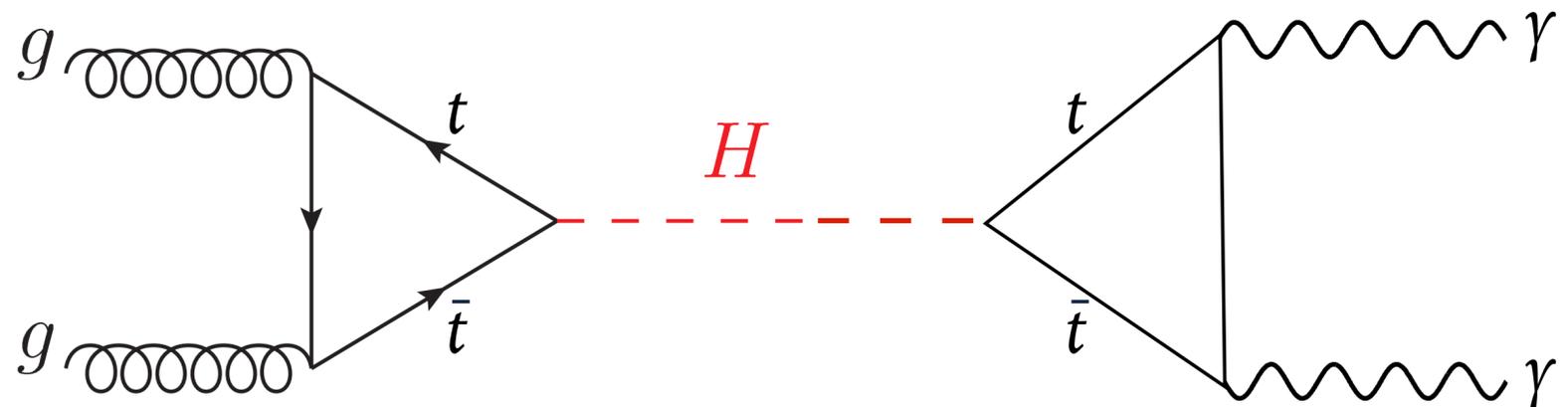


gluo

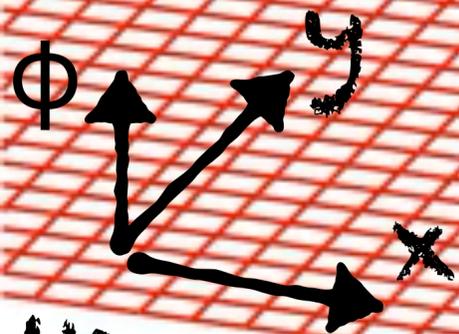


Higgs field in space

An LHC collision of the kind that led to the Higgs boson discovery



quon

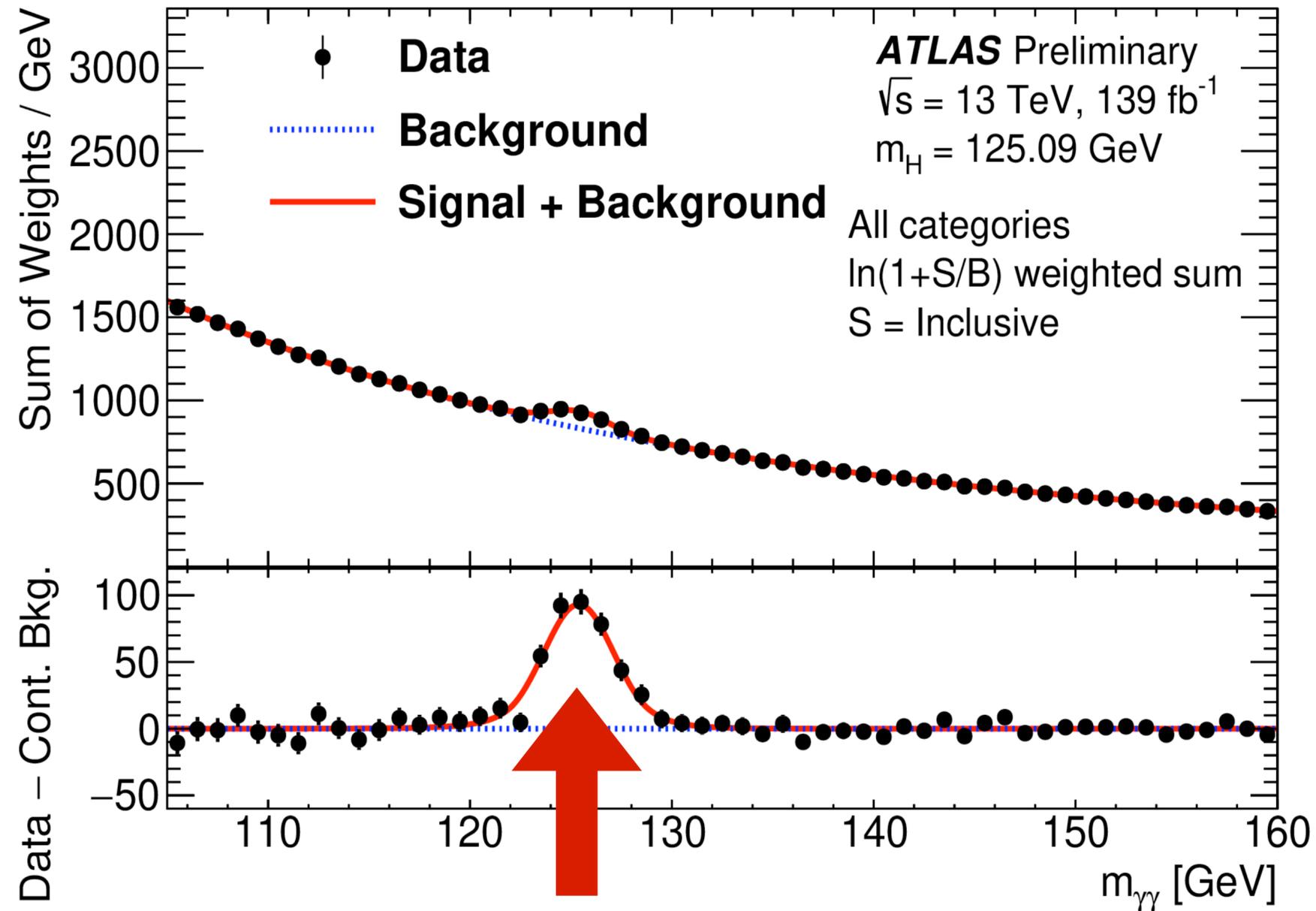


Higgs field in space

gluo

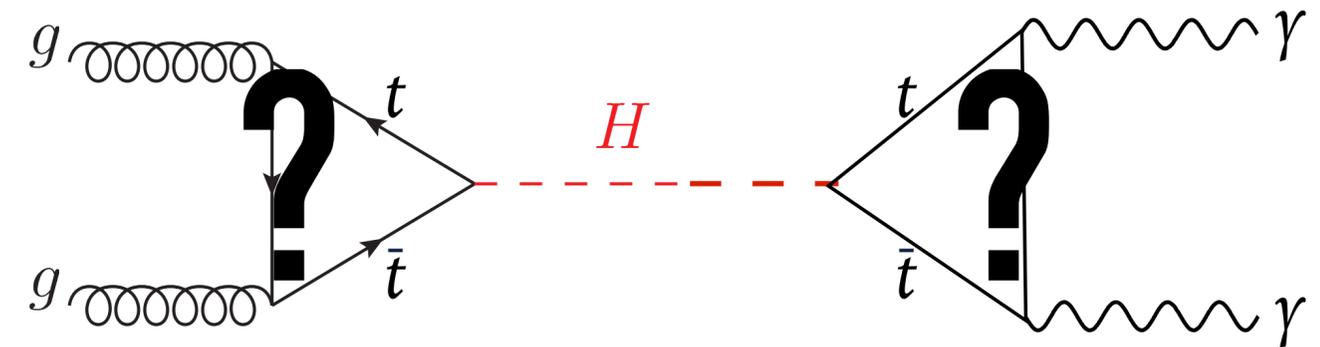
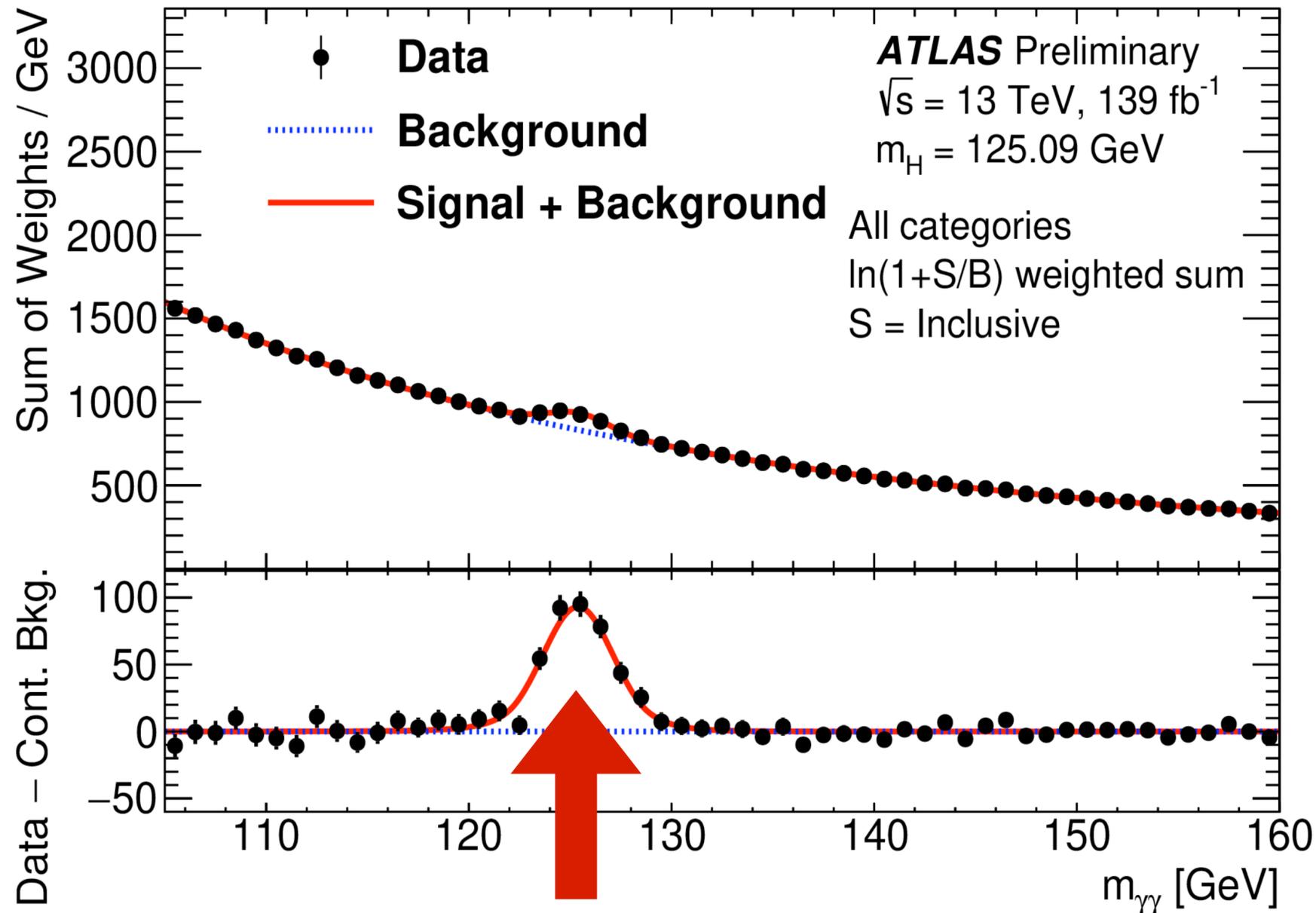


- Record events with two photons;
- classify and count them according to the invariant mass of the two photons (γ)



**more events at specific energy
= Higgs bosons**

- Record events with two photons;
- classify and count them according to the invariant mass of the two photons (γ)



rate of events consistent
with SM to $\sim 10\%$

but how can you be sure
it's a top-quark that's in
the intermediate stages?

more events at specific energy
= Higgs bosons

Press release

Summary

Laureates

Russell A. Hulse

Joseph H. Taylor Jr.

Press release

[Speed read](#)

Award ceremony speech

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13 October 1993

[The Royal Swedish Academy of Sciences](#) has decided to award the Nobel Prize Physics for 1993 jointly to **Russell A. Hulse** and **Joseph H. Taylor, Jr**, both of Princeton University, New Jersey, USA **for the discovery of a new type of pulsar, a discovery that has opened up new possibilities for the study of gravitation**

Gravity investigated with a binary pulsar

A very important observation was made when the system had been followed for some years [...] **reduction of the orbit period by about 75 millionths of a second per year** [...] **because the system is emitting energy in the form of gravitational waves** in accordance with what Einstein in 1916 predicted should happen to masses moving relatively to each other. [...] the theoretically calculated value from the relativity theory agrees to within about one half of a percent with the observed value. The first report of this effect was made by Taylor and co-workers at the end of 1978, four years after the discovery of the binary pulsar was reported.

Press release

**indirect
observation**

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Prize announcement

Press release

Popular information

Advanced information

Award ceremony video

Award ceremony speech

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3 October 2017

[The Royal Swedish Academy of Sciences](#) has decided to award the Nobel Prize in Physics 2017 with one half to

Rainer Weiss
LIGO/VIRGO Collaboration

and the other half jointly to

Barry C. Barish
LIGO/VIRGO Collaboration

and

Kip S. Thorne
LIGO/VIRGO Collaboration

“for decisive contributions to the LIGO detector and the observation of gravitational waves”

Gravitational waves finally captured

On 14 September 2015, the universe’s gravitational waves were observed for the very first time. The waves, which were predicted by Albert Einstein a hundred years ago, came from a collision between two black holes. It took 1.3 billion years for the waves to arrive at the LIGO detector in the USA.

Nobel Prize in Physics 2017

Press release

direct
observation

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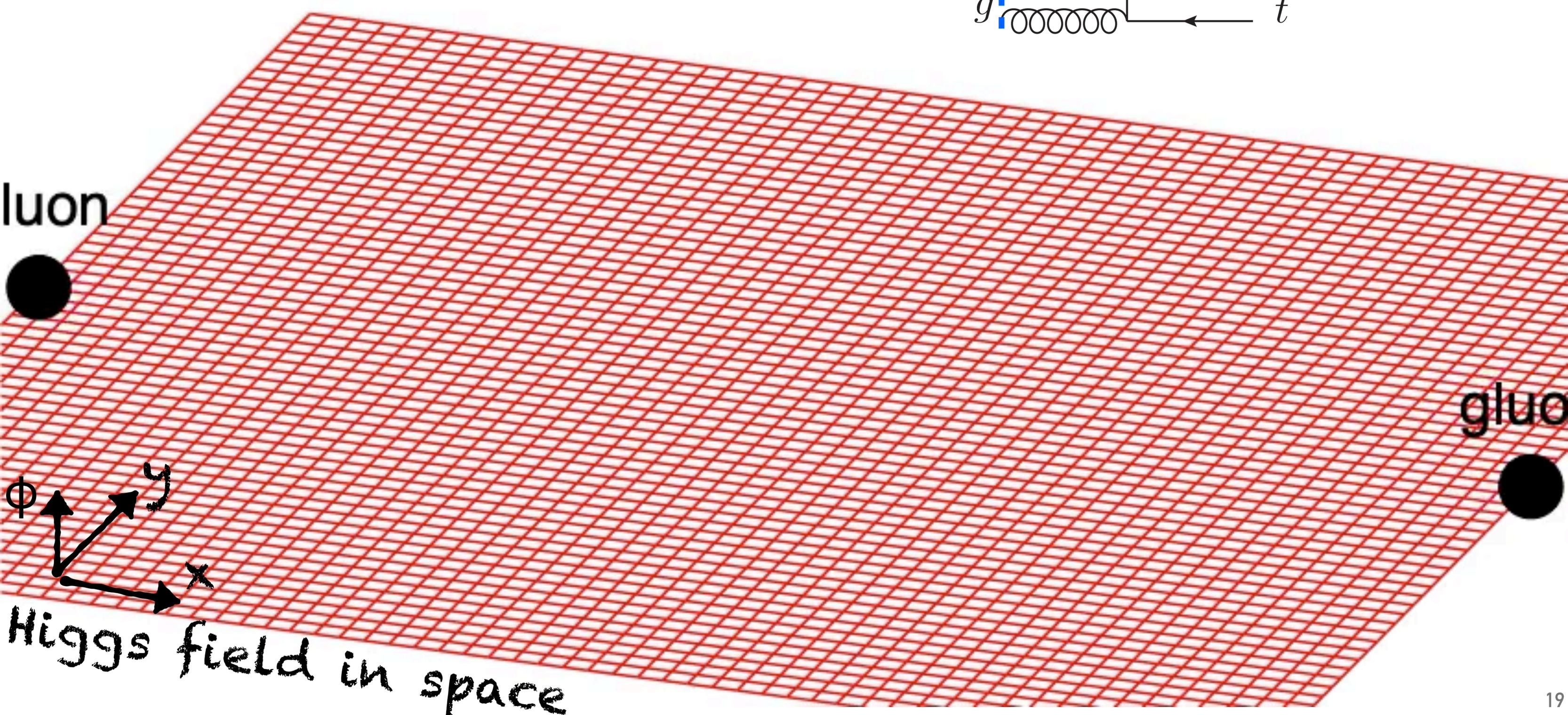
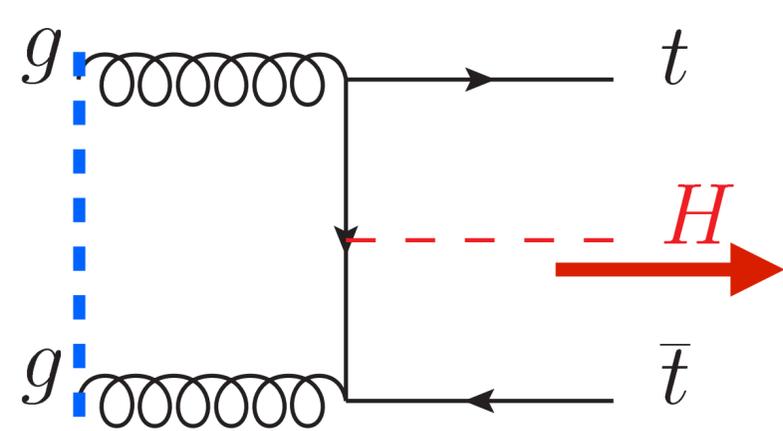
and

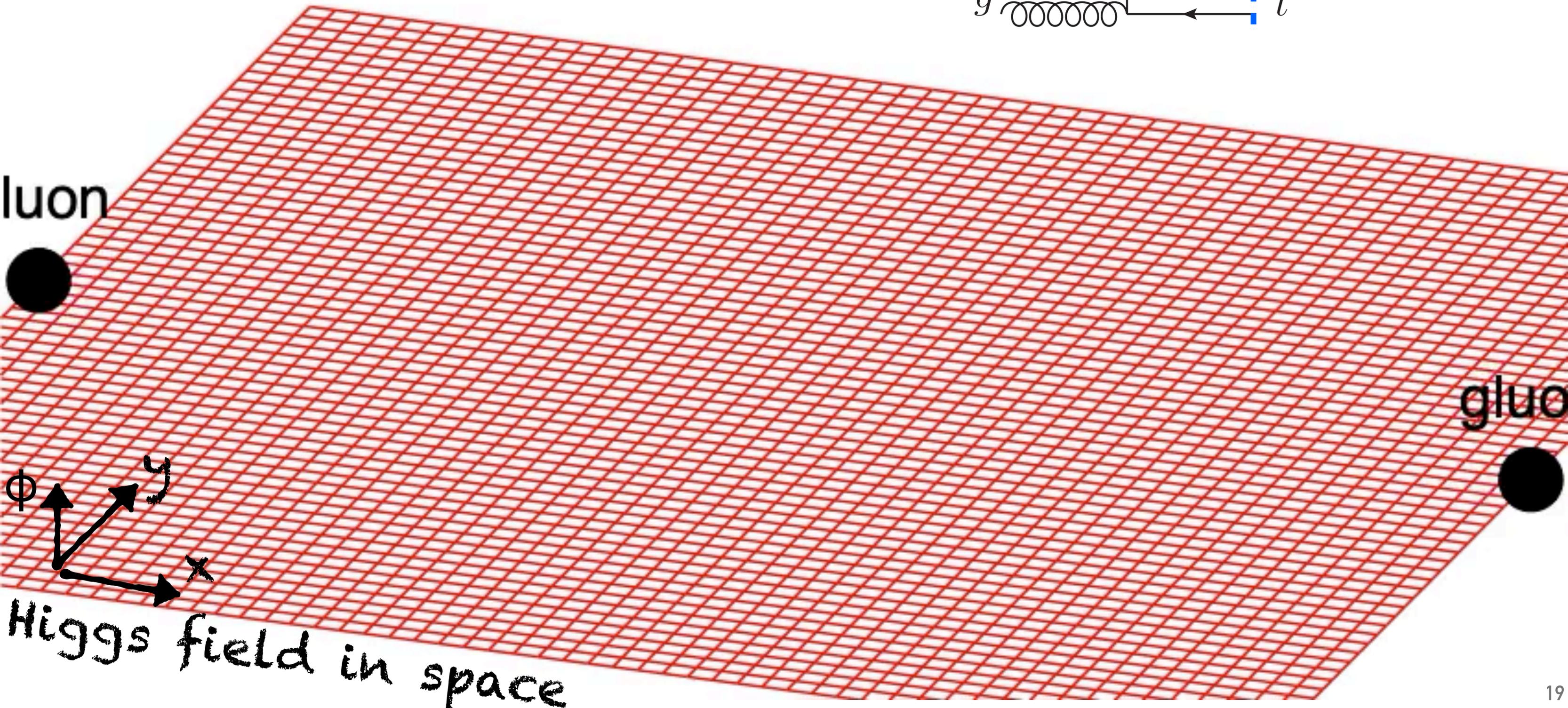
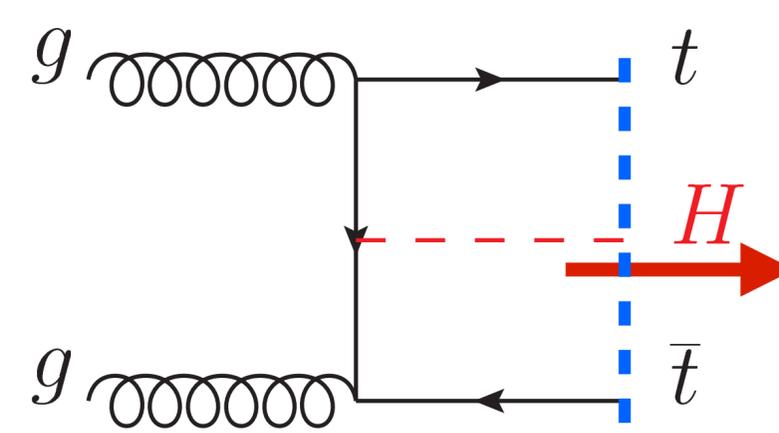
Kip S. Thorne
LIGO/VIRGO Collaboration

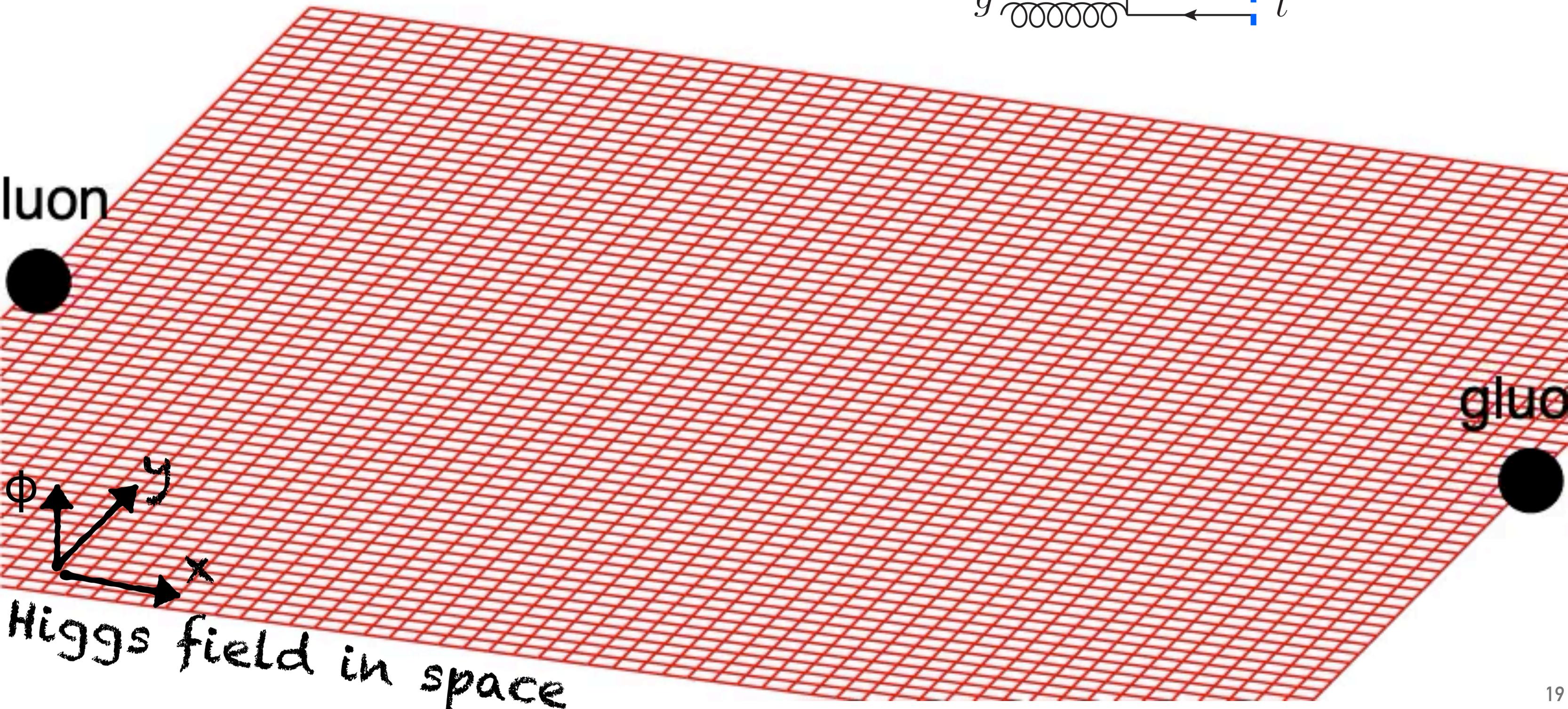
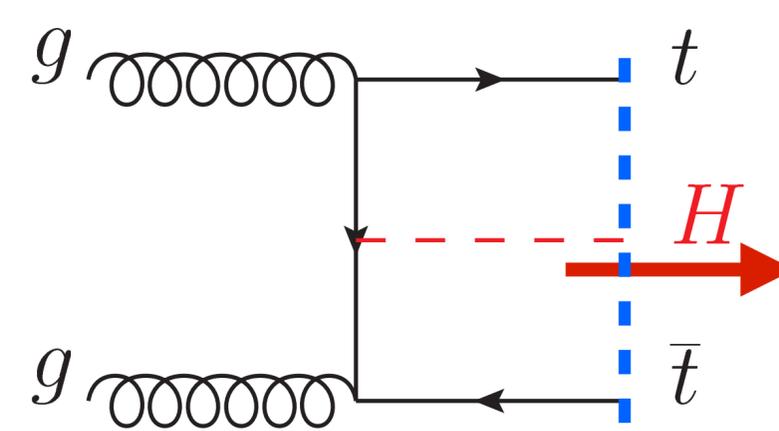
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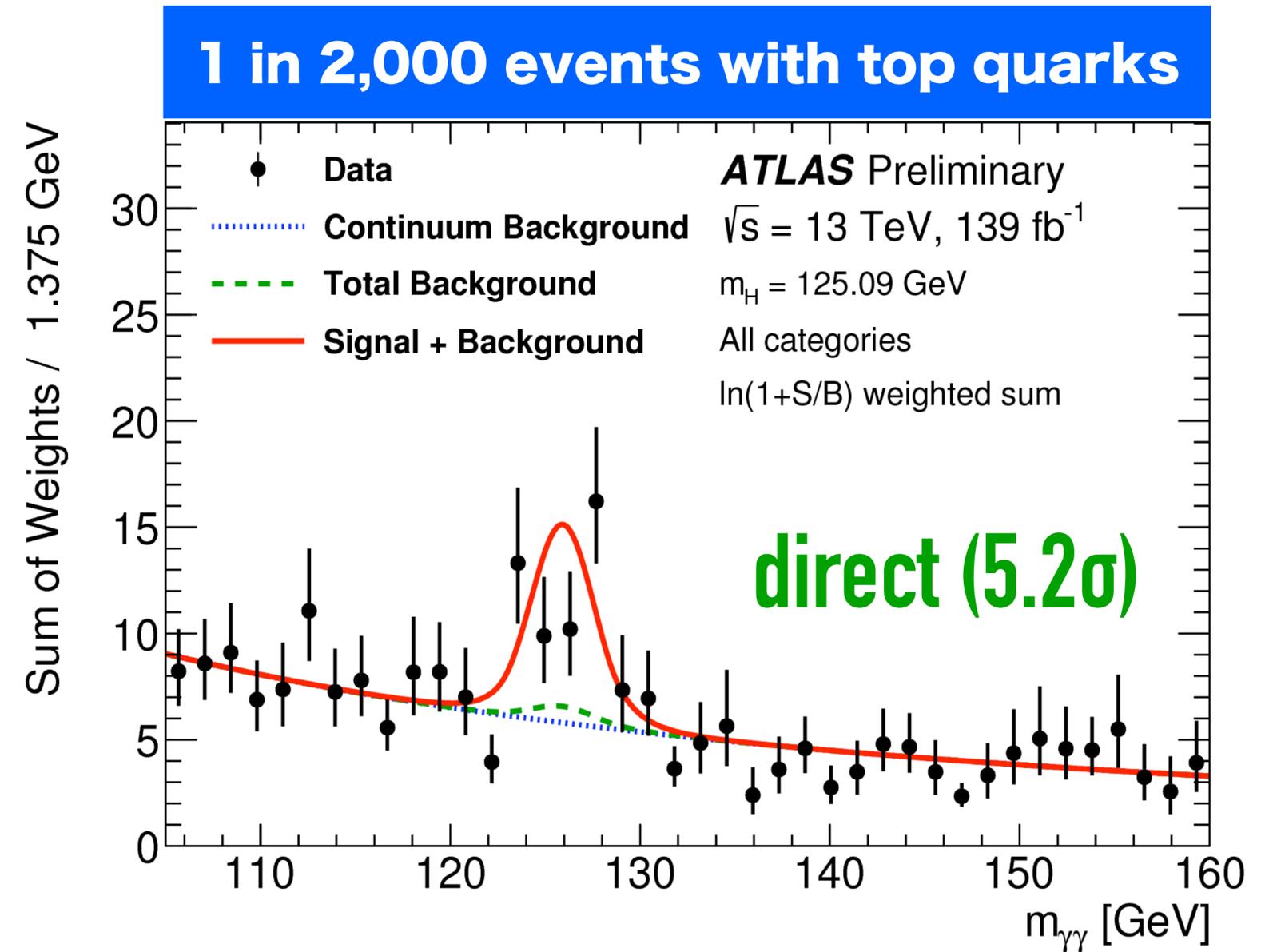
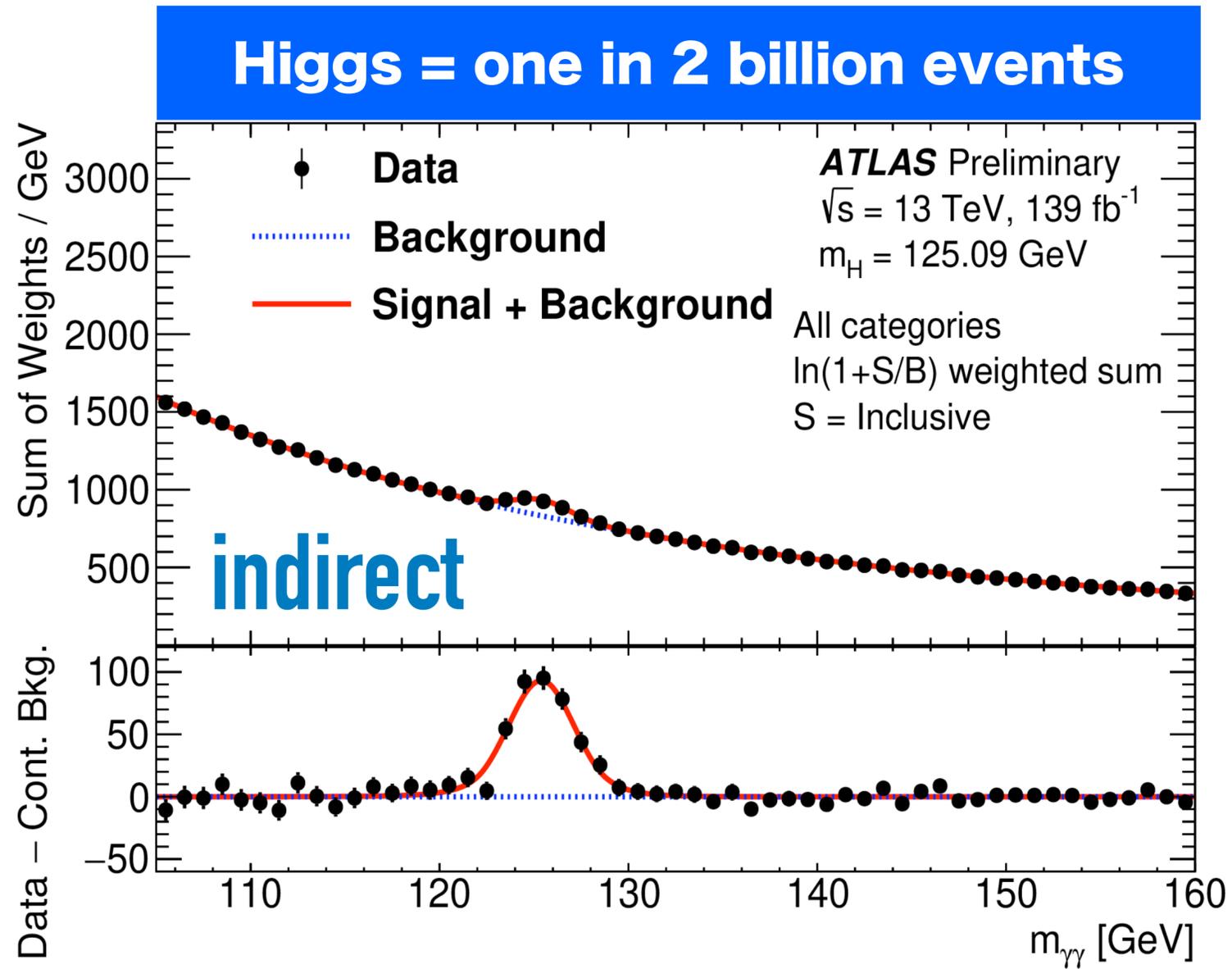


Situation at start of LHC (2009)

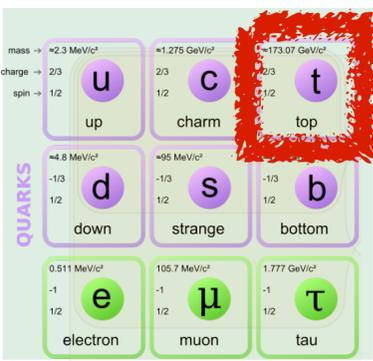
“Due to a (too) low signal-to-background ratio $S/B \sim 1/9$ [ttH] channel might not reach a 5σ significance for any luminosity.”

[from introduction to arXiv:0910.5472,
summarising ATLAS and CMS ttH(\rightarrow bb) studies at that point]

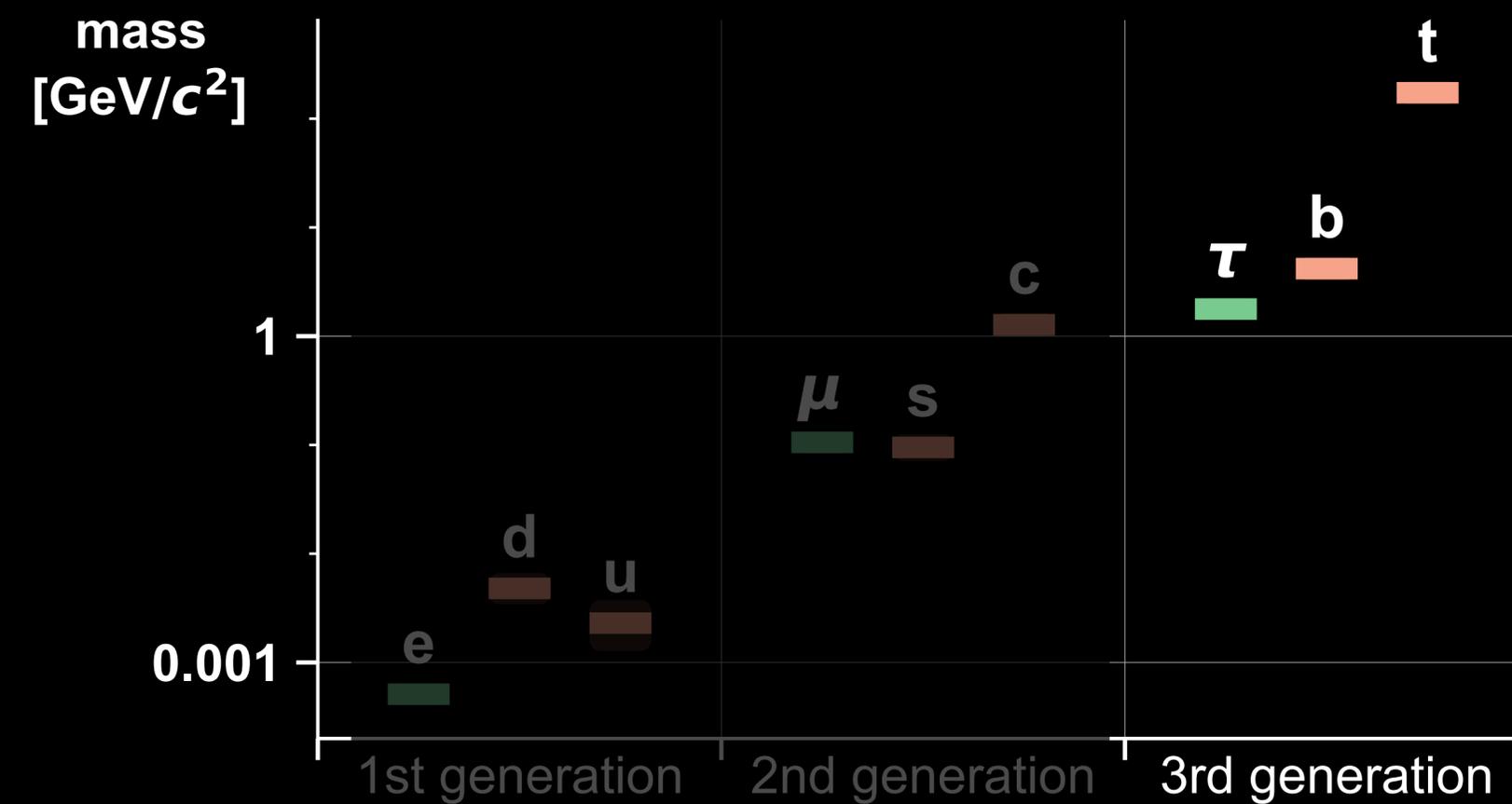
since 2018: ATLAS & CMS see (at $>5\sigma$) events with top-quarks & Higgs simultaneously



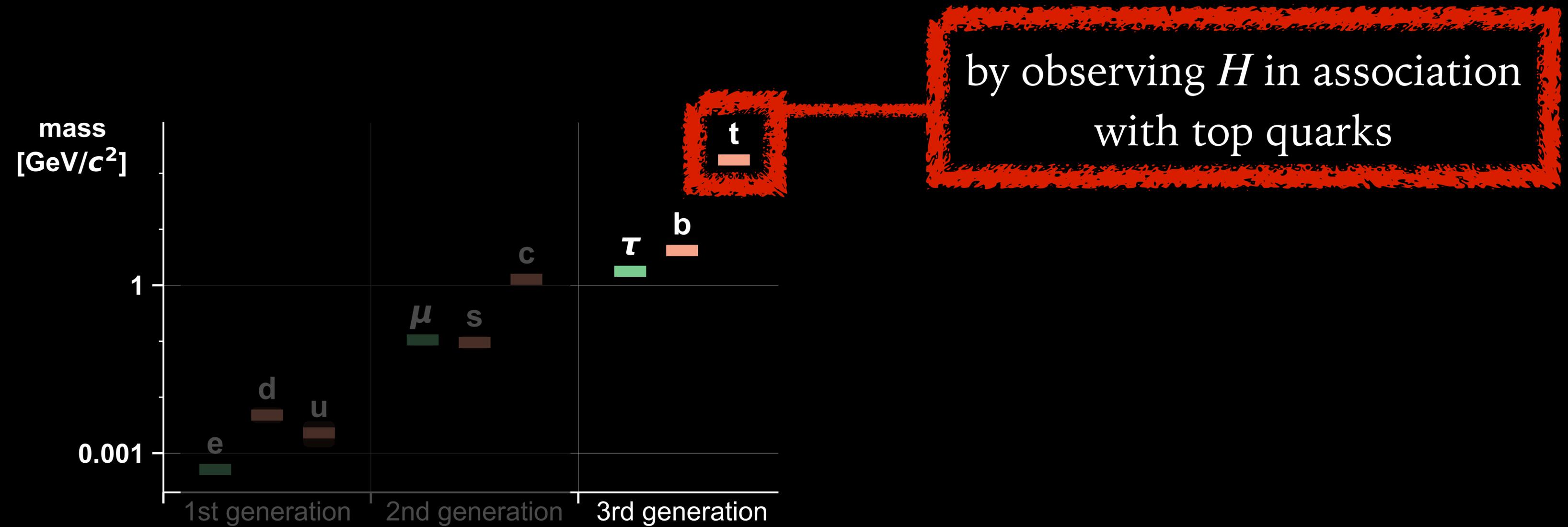
enhanced fraction of Higgs bosons in events with top quarks
 → direct observation of Higgs interaction with tops
 (consistent with SM to c. $\pm 25\%$)



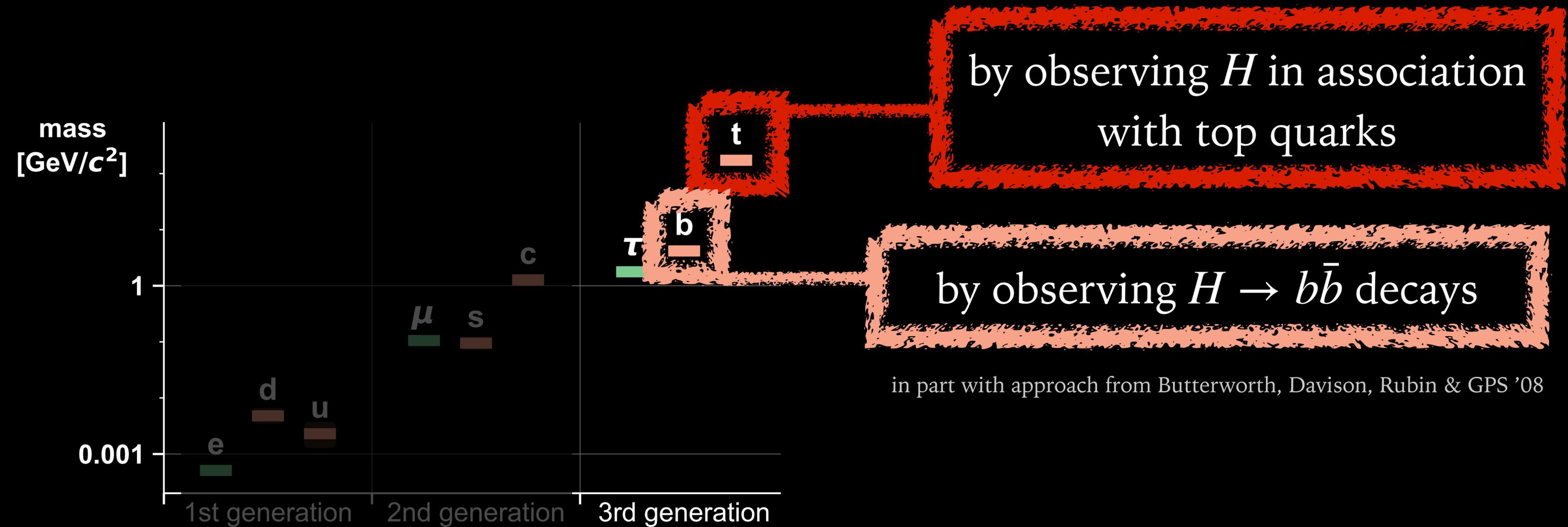
Discovery of 3rd generation—Higgs field interactions by ATLAS & CMS ~ 2018



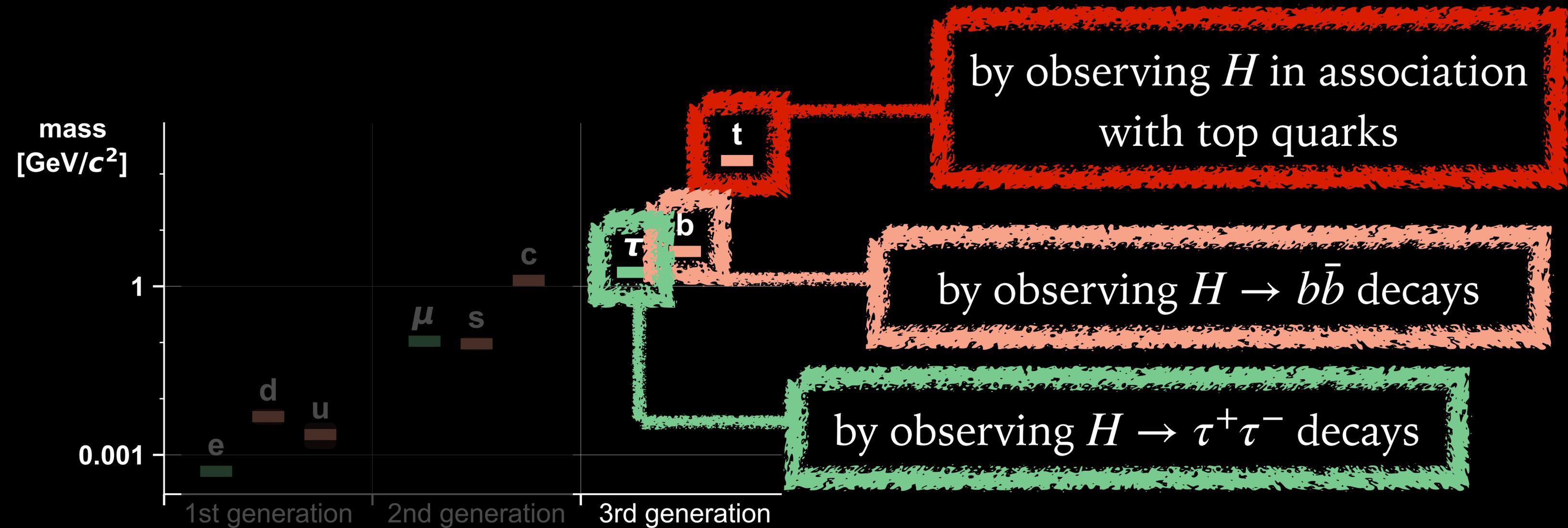
Discovery of 3rd generation—Higgs field interactions by ATLAS & CMS ~ 2018



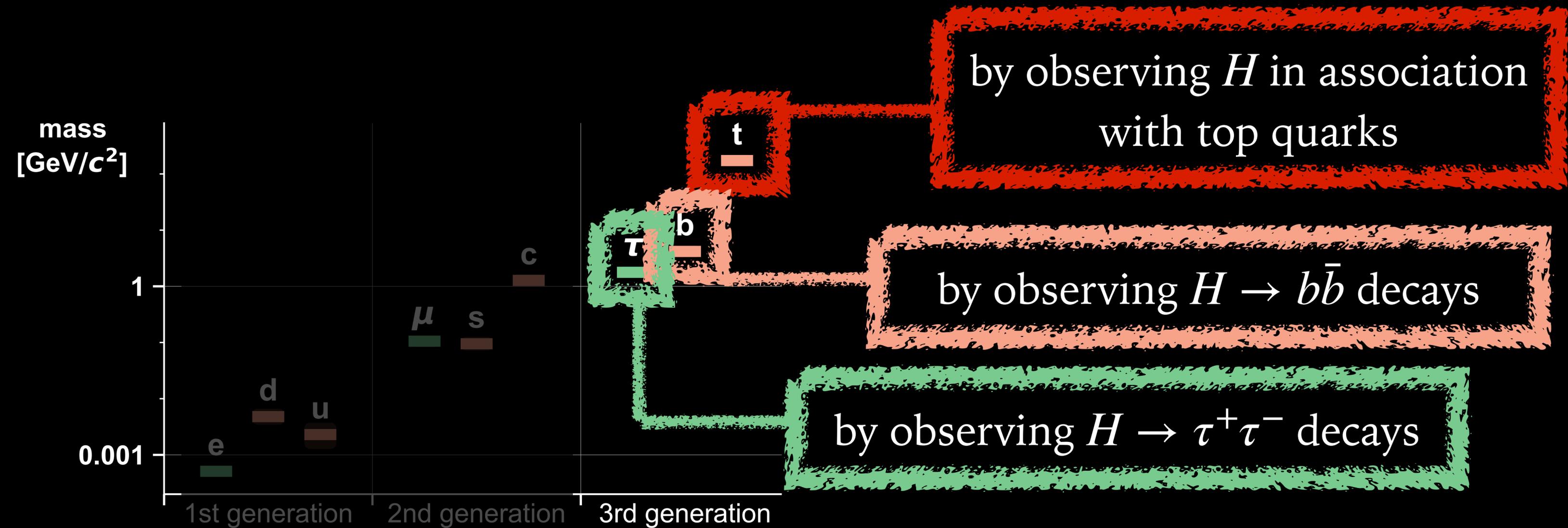
Discovery of 3rd generation—Higgs field interactions by ATLAS & CMS ~ 2018



Discovery of 3rd generation—Higgs field interactions by ATLAS & CMS ~ 2018

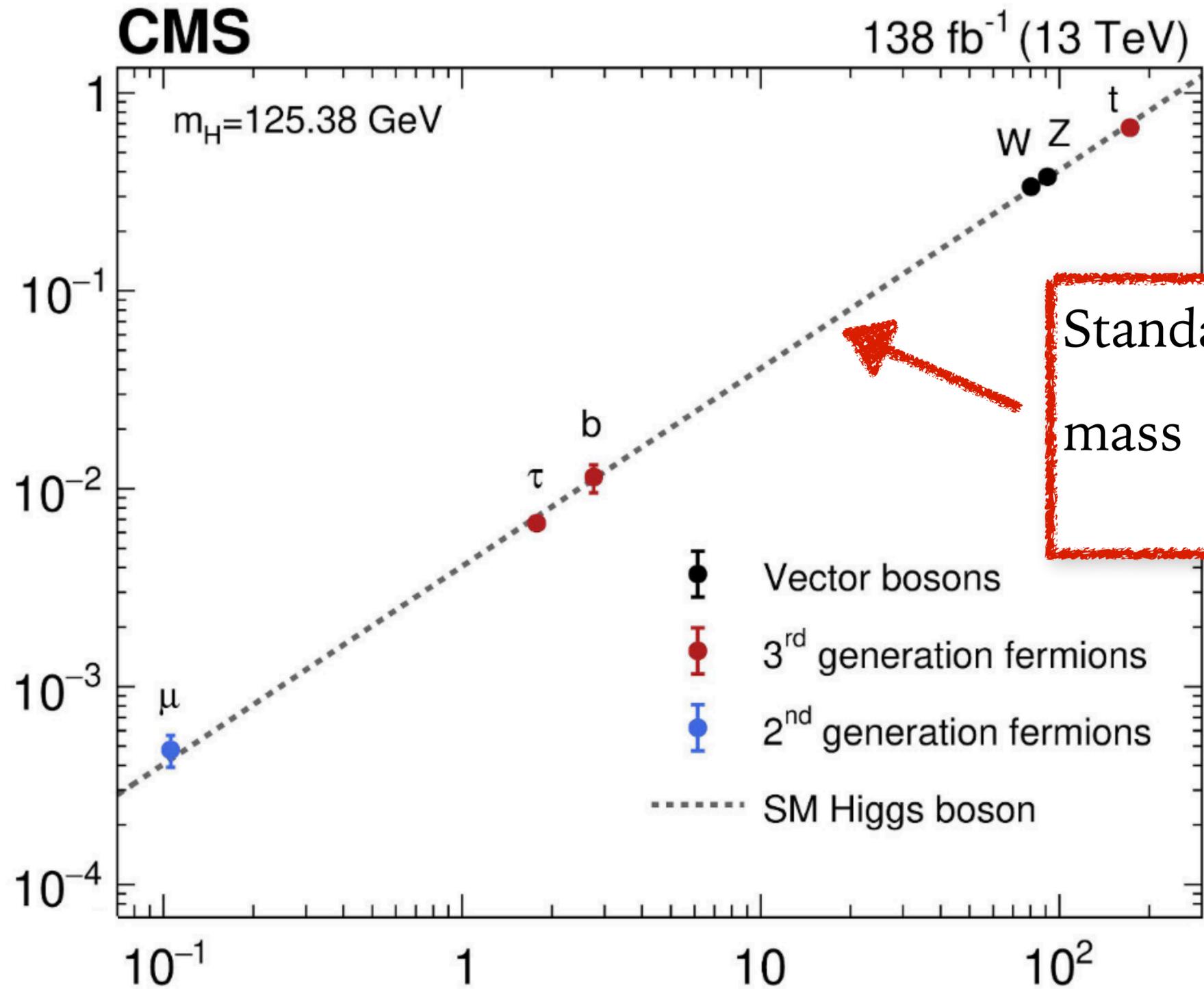


Discovery of 3rd generation—Higgs field interactions by ATLAS & CMS ~ 2018



**Full 3rd generation Yukawas were not part of the LHC design case.
Amazing achievement of LHC experiments to have directly observed them**

↑
Particle's strength of interaction with Higgs field



Standard Model prediction:
mass = higgs-field-value
× interaction-strength

Particle's mass [GeV] →

what could one be saying about it?

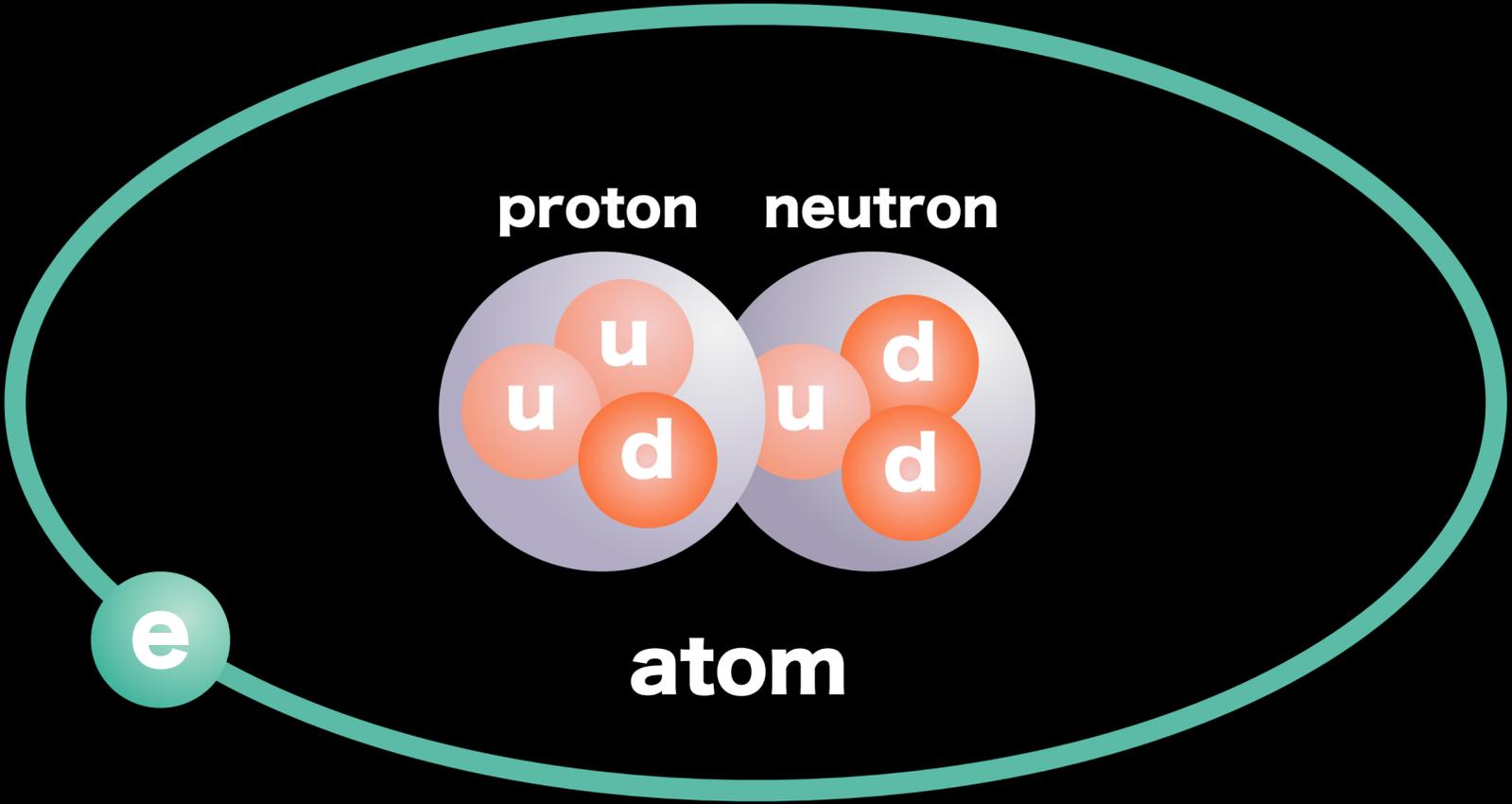
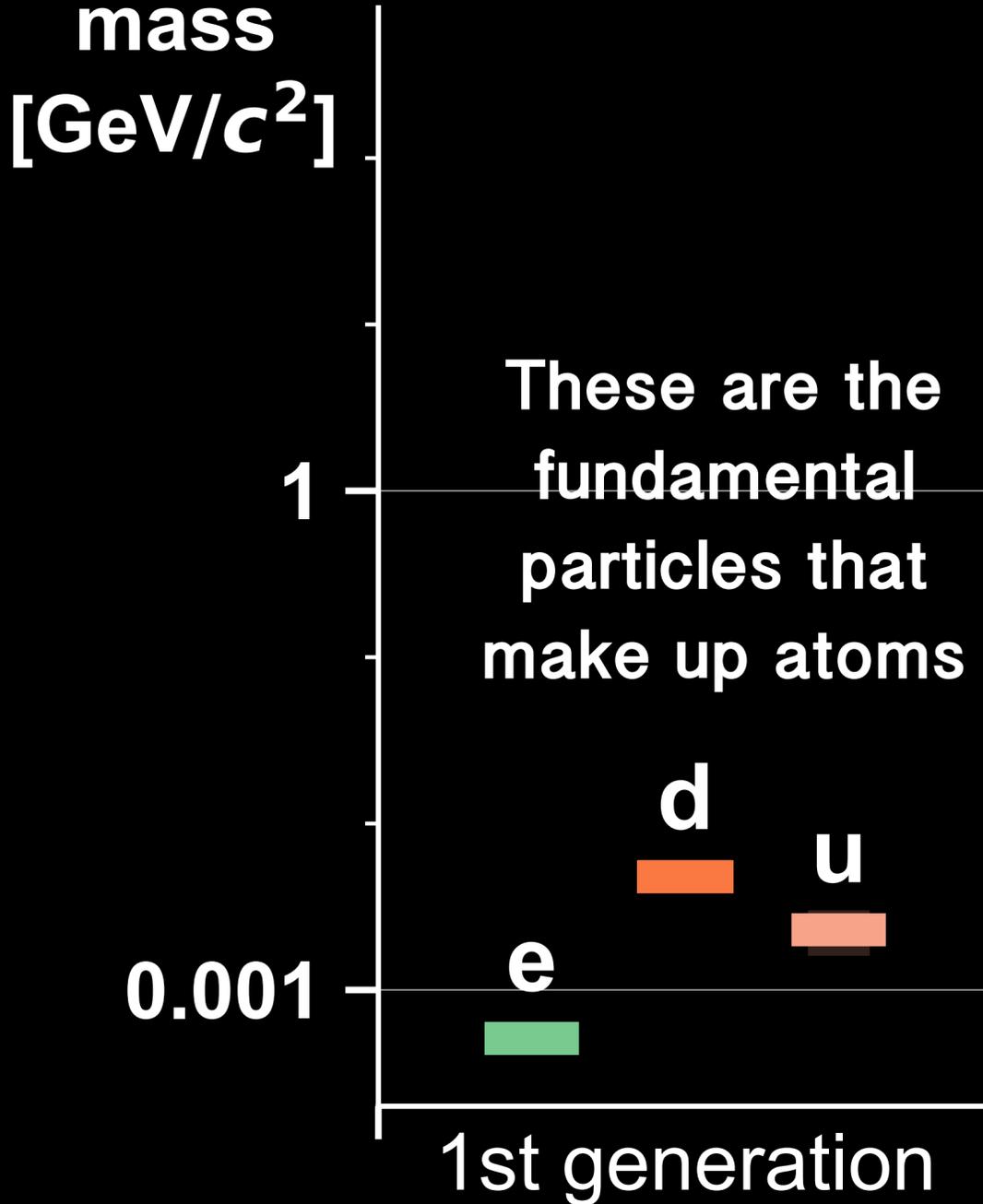
For a full set of particles (3rd generation) that are like the ones we're made of, the LHC has demonstrated that their mass is not an intrinsic property, but is generated by an interaction with a non-zero Higgs field.

A field is something that can in principle be controlled and modified. Could the masses of elementary particles conceivably also be controlled and modified? Science fiction...

Is this any less important than the discovery of the Higgs boson itself?

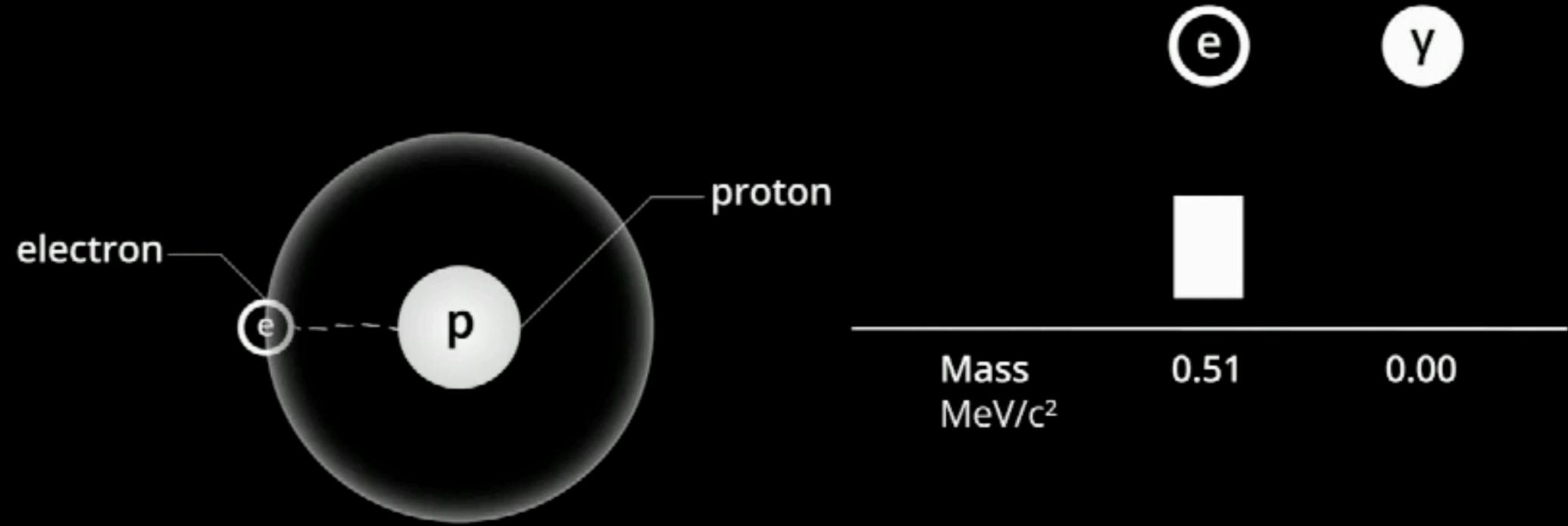
My opinion: no

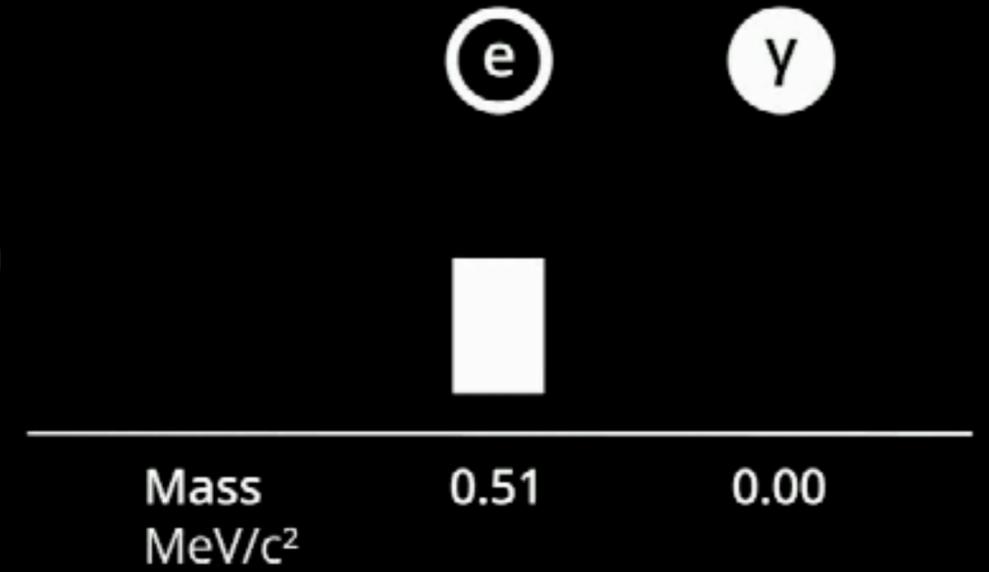
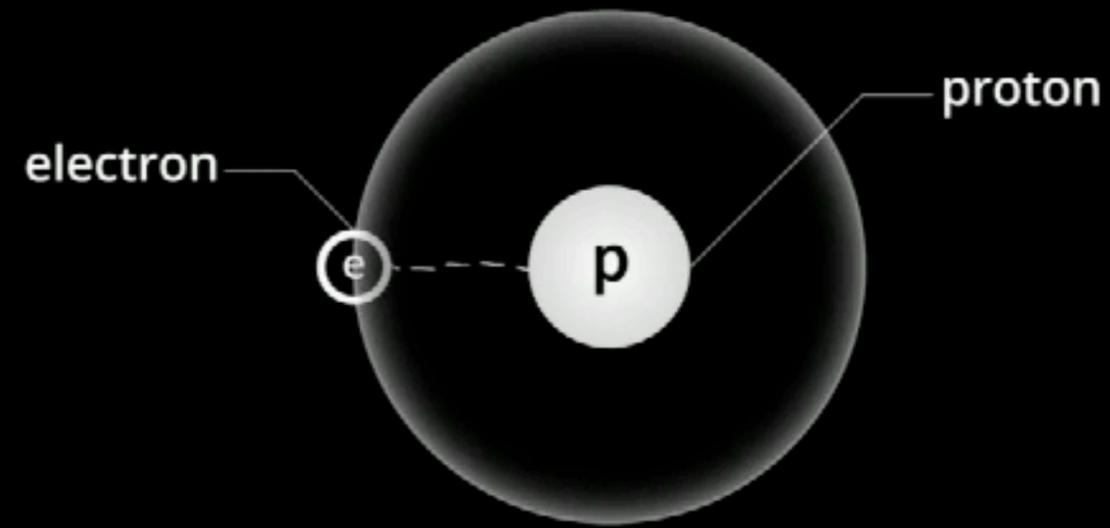
NB: most of mass of proton and neutron comes from other sources



Bohr radius of atom

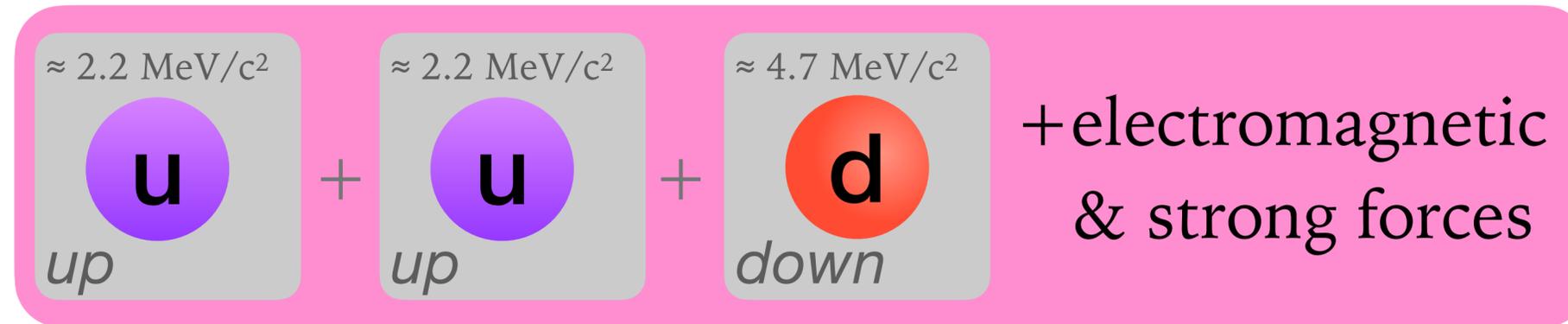
$$a_0 = \frac{4\pi\epsilon_0\hbar^2}{m_e e^2} = \frac{\hbar}{m_e c \alpha} \propto \frac{1}{y_e}$$





2.2 MeV 2.2 MeV 4.7 MeV

proton:



$\approx 938.3 \text{ MeV}$

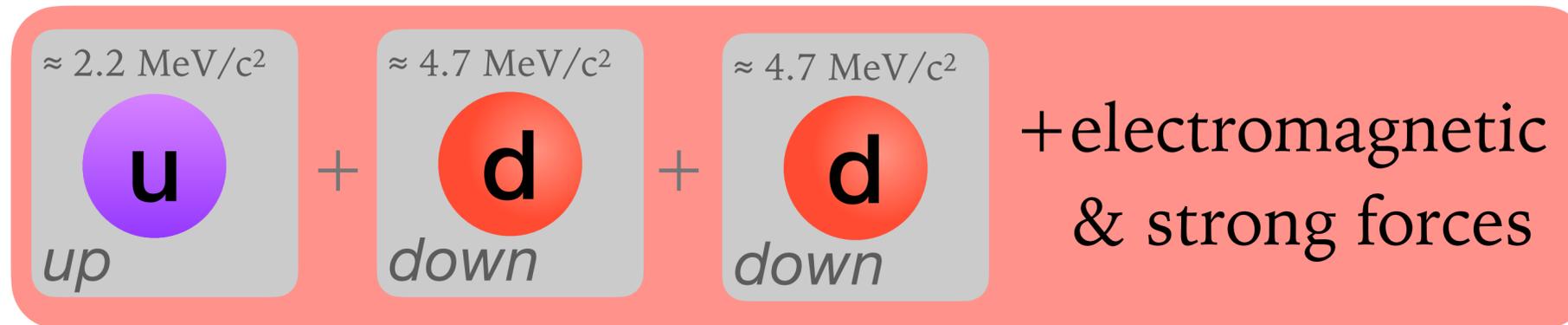
2.2 MeV **2.2 MeV** 4.7 MeV

proton:



$\approx 938.3 \text{ MeV}$

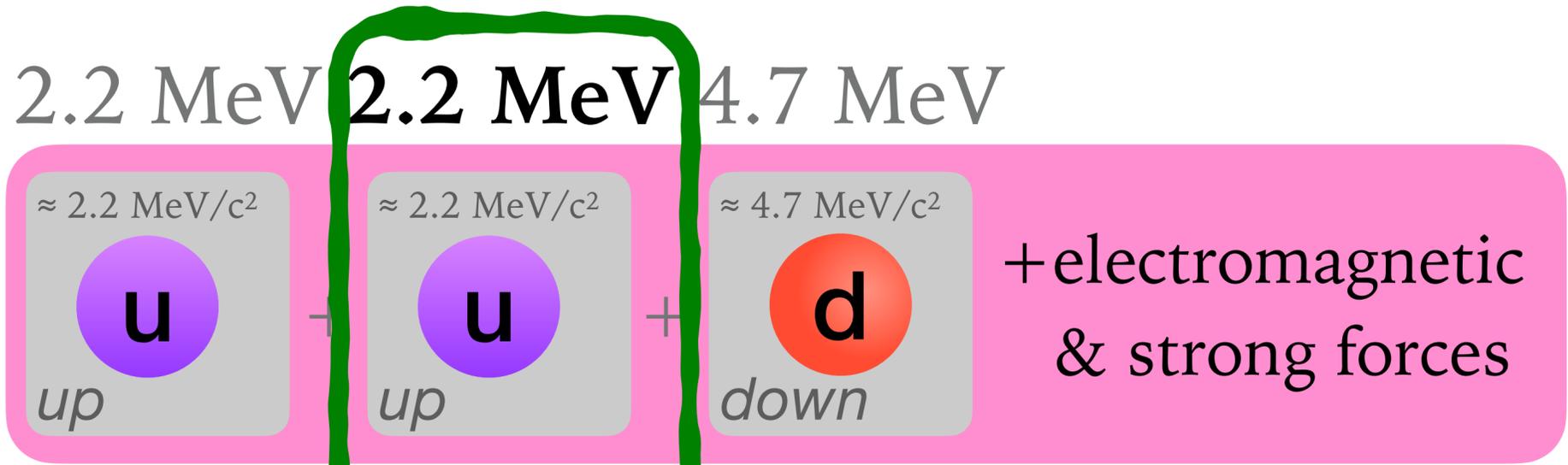
neutron:



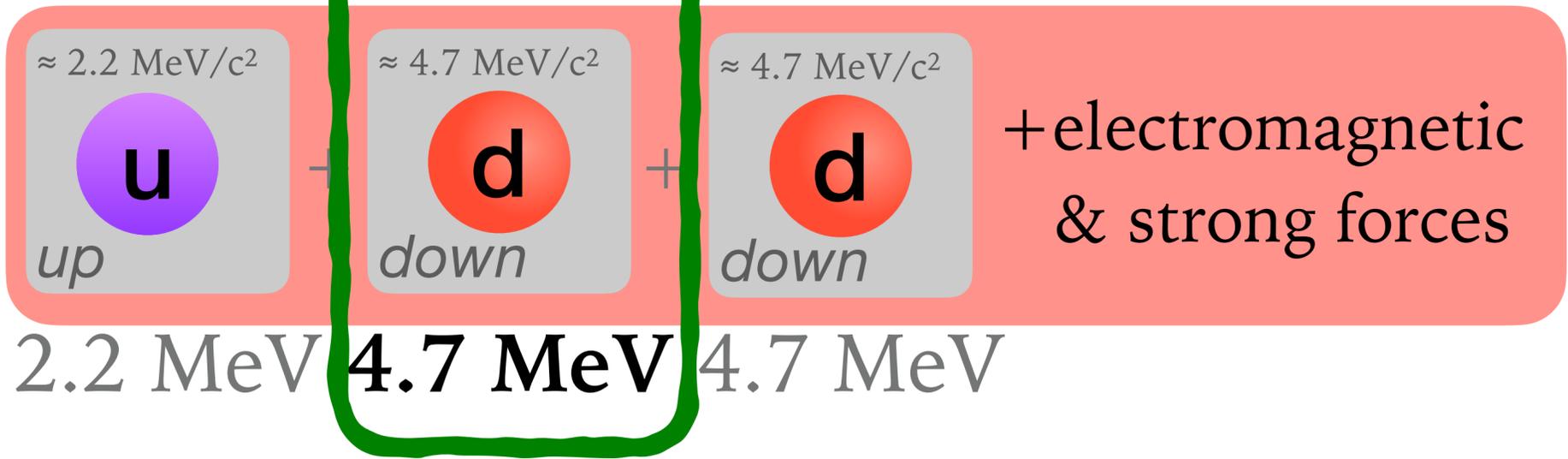
$\approx 939.6 \text{ MeV}$

2.2 MeV **4.7 MeV** 4.7 MeV

proton:

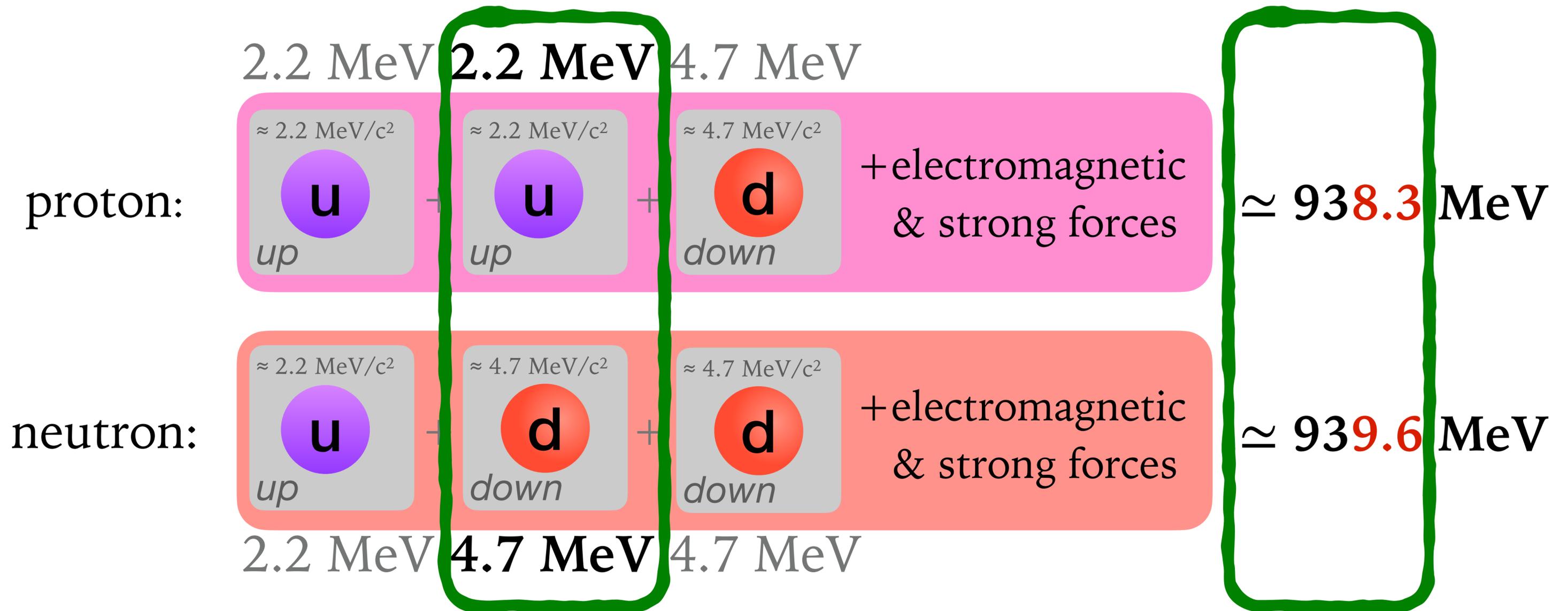


neutron:



$\approx 938.3 \text{ MeV}$

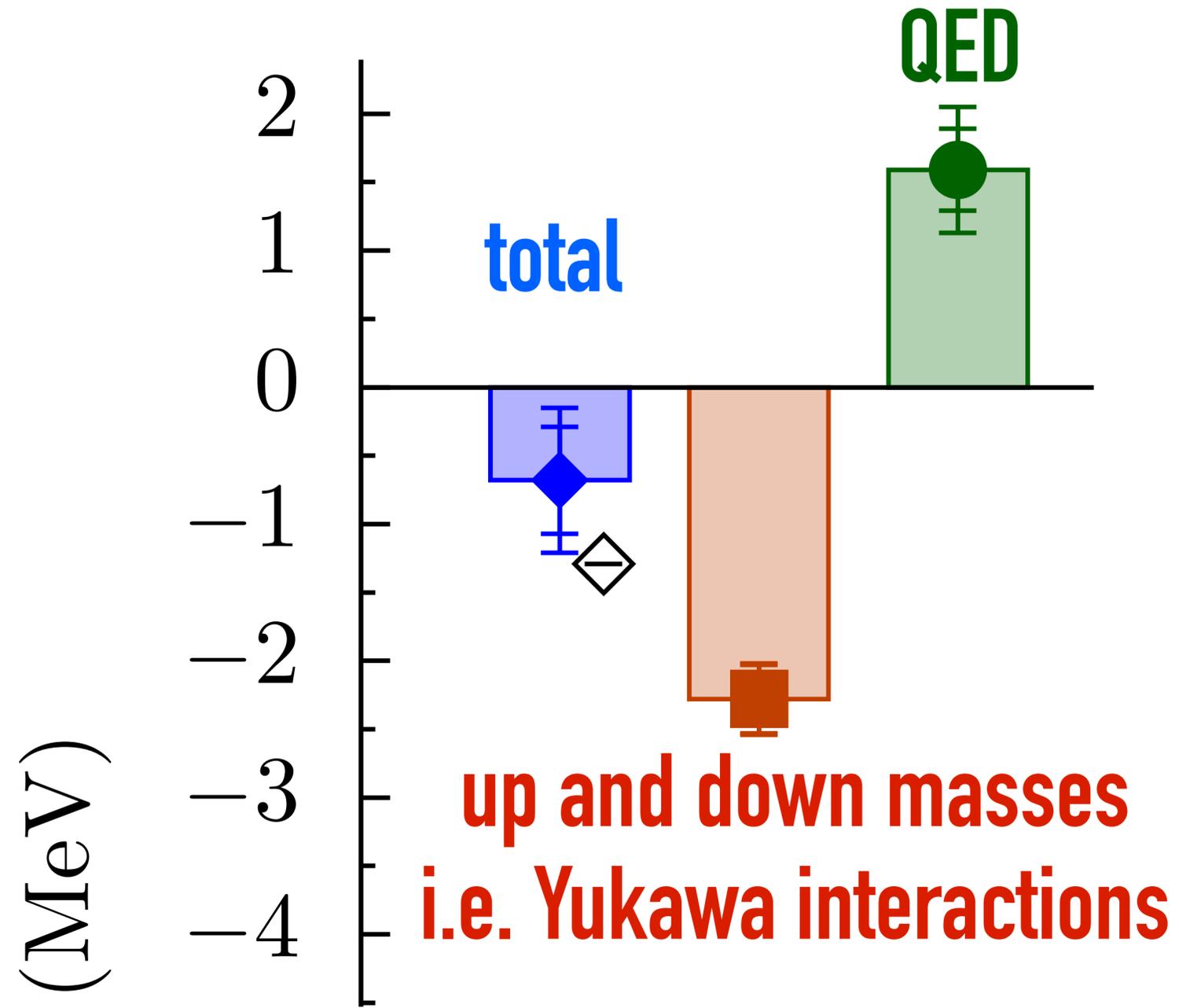
$\approx 939.6 \text{ MeV}$



Protons are **lighter** than neutrons \rightarrow protons are stable.
 Giving us the hydrogen atom, & chemistry and biology as we know it

**Supposedly because up quarks interact more weakly
 with the Higgs field than down quarks**

proton - neutron mass difference

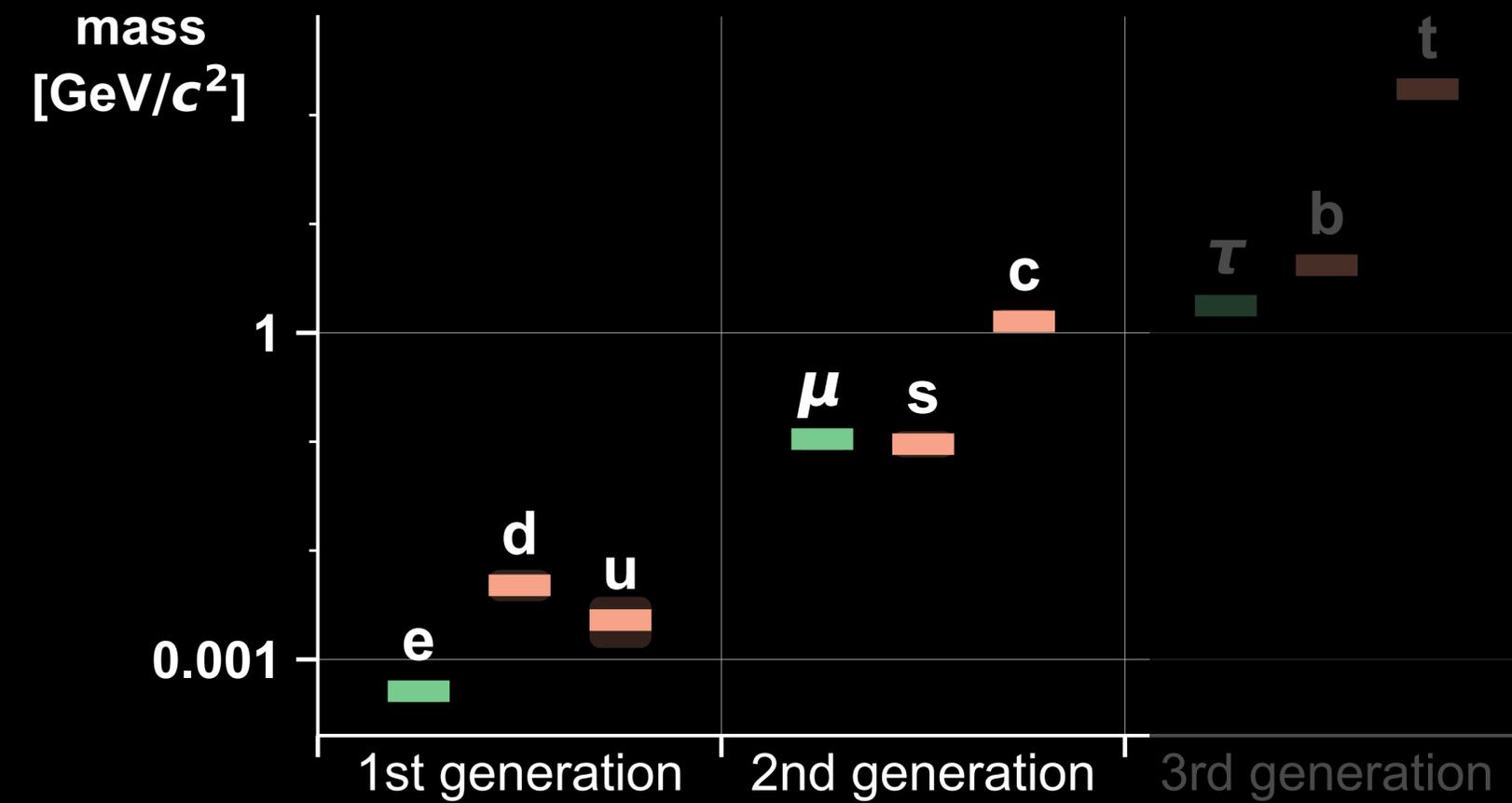


Lattice calculation
(BMW collab.)

1306.2287

1406.4088

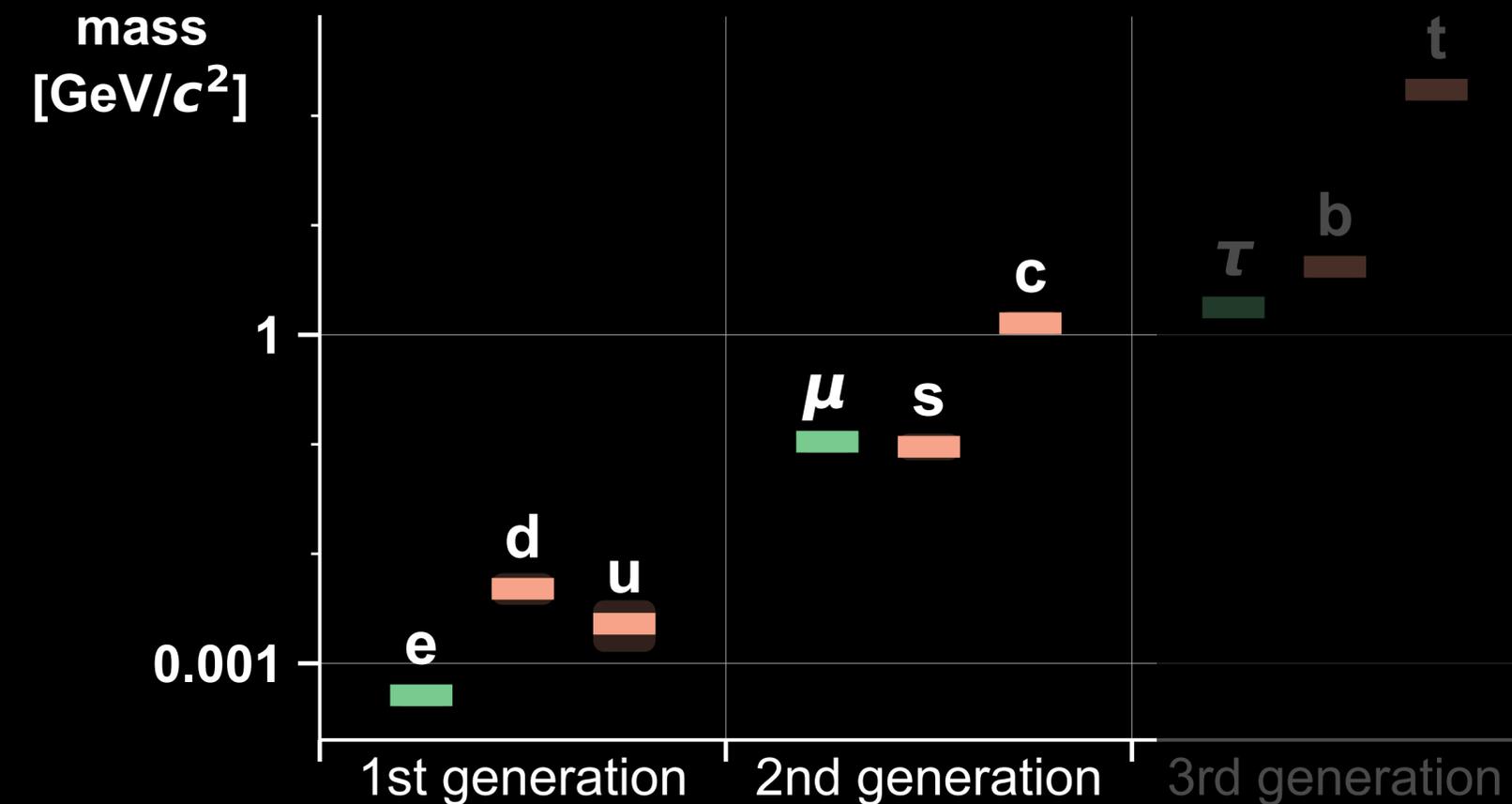
**currently we have no evidence that up and down quarks
and electron get their masses from Yukawa interactions
— it's in textbooks, but is it nature?**



a BIG question of particle physics is whether all of these particles acquire their mass in the same way

In SM hypothesis: the lighter the particle, the less it interacts with the Higgs field

→ the more difficult it is establish if it actually gets mass from interactions with the Higgs field



a BIG question of particle physics is whether all of these particles acquire their mass in the same way



European Strategy Update

2024 to 2026

EUROPEAN STRATEGY FOR PARTICLE PHYSICS

[...] cornerstone of Europe's decision-making process for the long-term future of the field

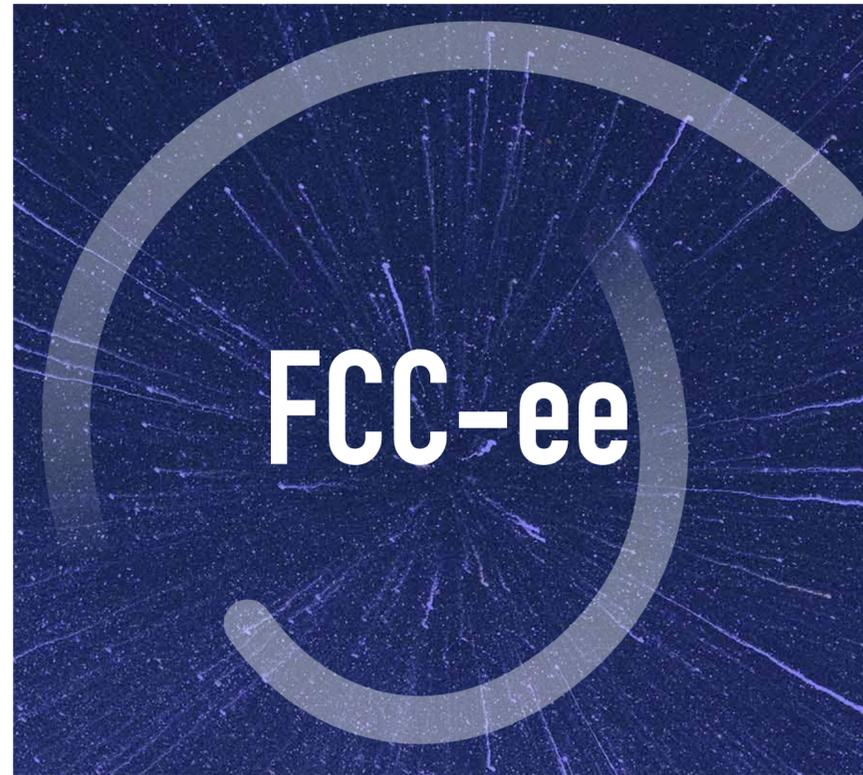
[...] develop a visionary and concrete plan that greatly advances knowledge in fundamental physics through the realisation of the next flagship collider at CERN, and to prioritize alternative options to be pursued if the preferred plan turns out not to be feasible or competitive.



2029–2041

proton–proton
14,000 GeV energy
10× more collisions
than LHC

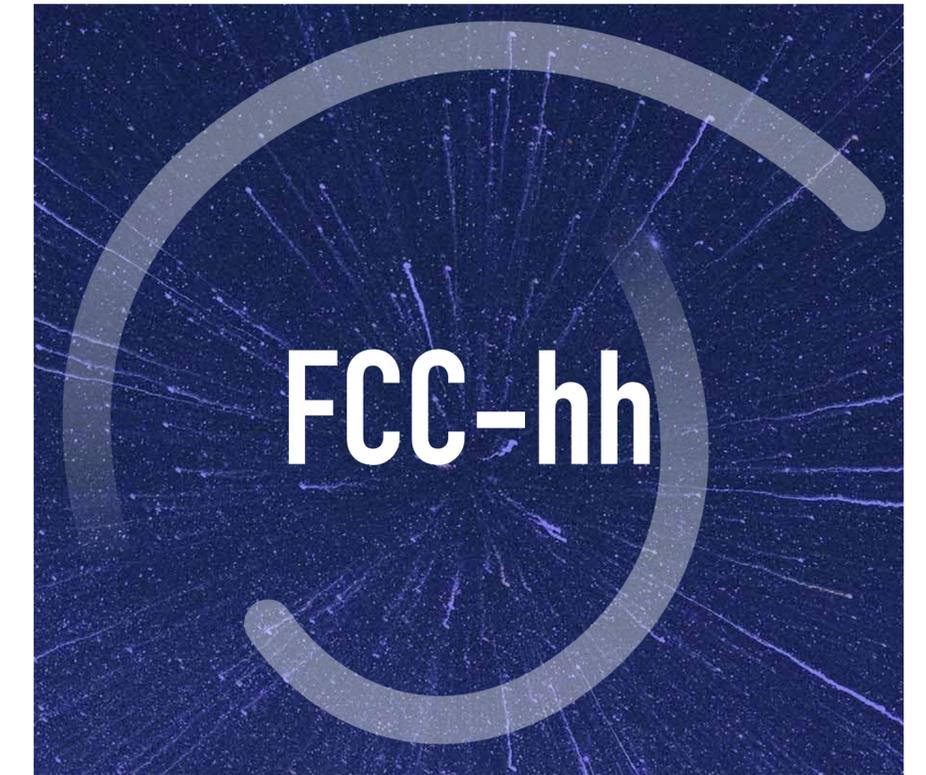
**approved & upgrade
under construction**



2045–2060(c.)

electron–positron
91–365 GeV energy
300,000× more
collisions than LEP

[or CEPC@China,
ILC, CLIC, C³]



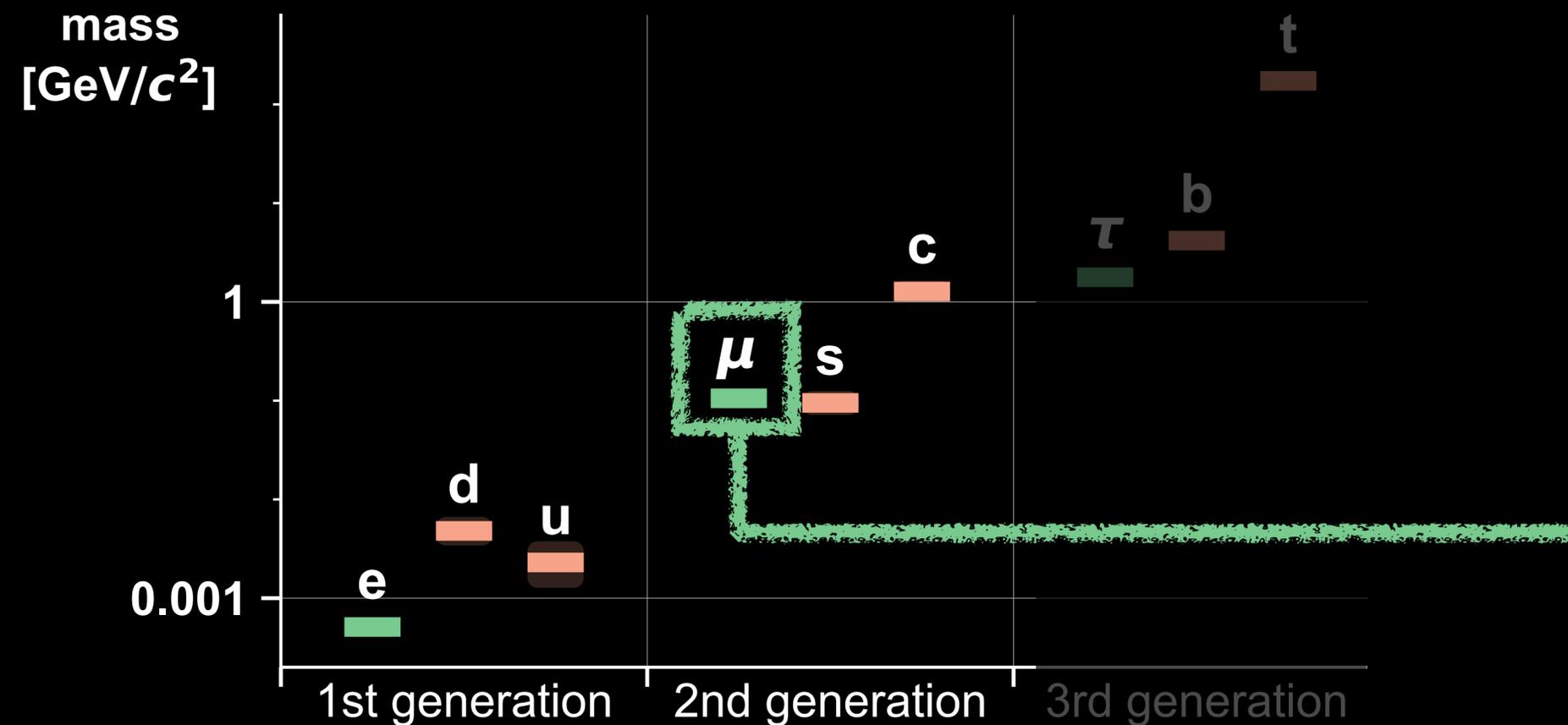
2070–2090(c.)

proton–proton
~100,000 TeV energy
10× more collisions
than HL-LHC

or SppS@China
or muon collider

In SM hypothesis: the lighter the particle, the less it interacts with the Higgs field

→ the more difficult it is establish if it actually gets mass from interactions with the Higgs field



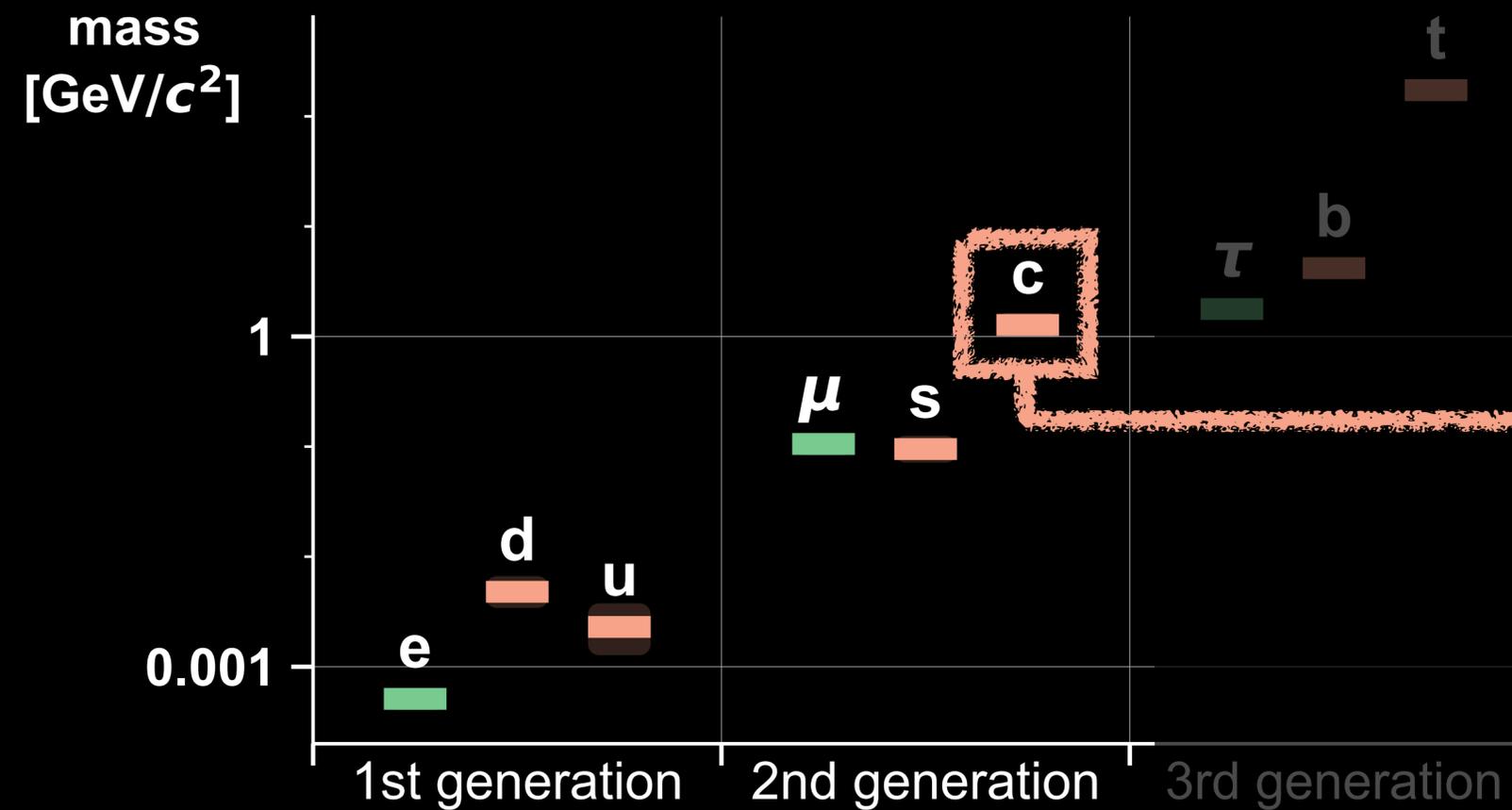
a major LHC goal of the next years (Run-3 or HL-LHC) will be to establish, for the first time, whether a 2nd generation particle also acquires its mass in the same way

[ATLAS/CMS have first indications, but not yet 5σ]

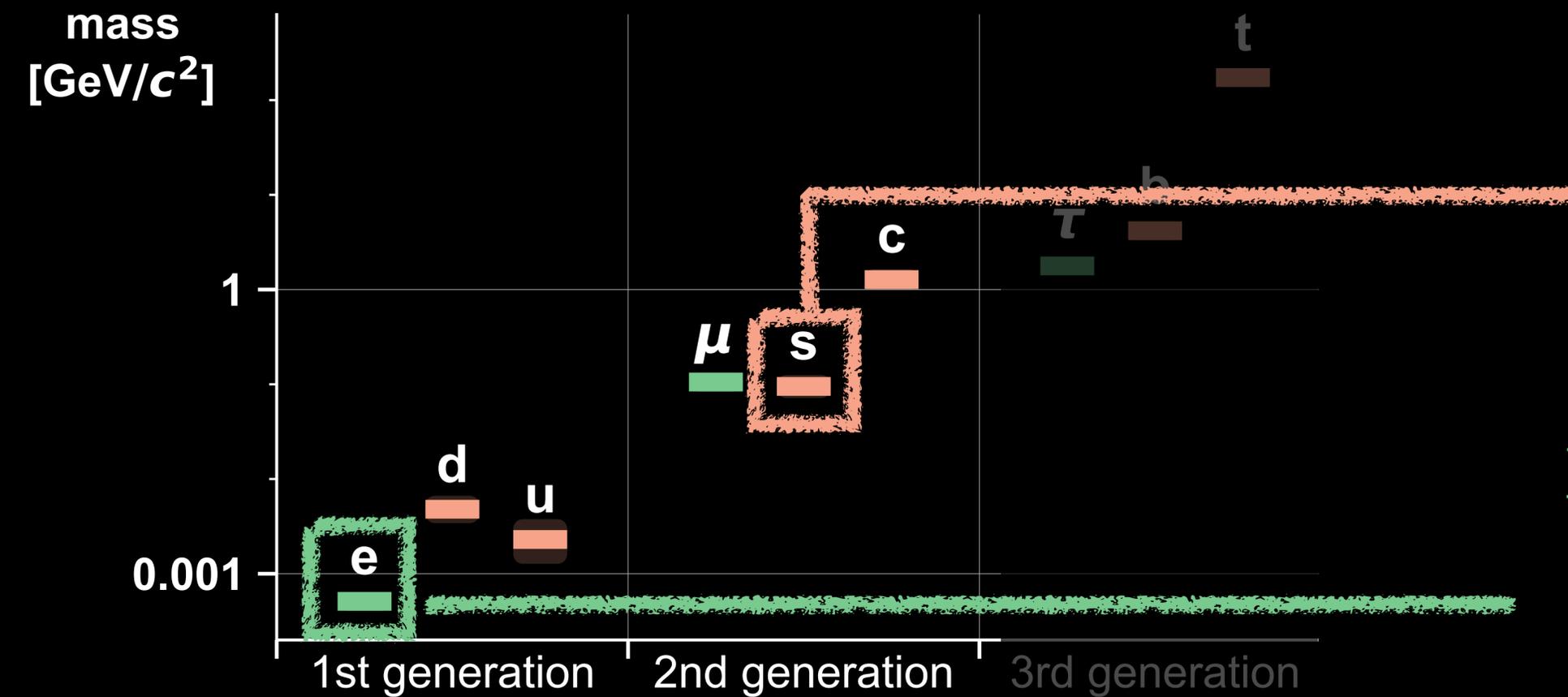
What of future colliders

quarks and yet-lighter particles
are much harder

future e^+e^- collider, if built,
will clearly establish if charm-
quarks get their mass from
Higgs-field interactions



What of future colliders



It's becoming clear that strange quark and electron "Yukawas" are just barely at the edge of reach of FCC-ee

Discovering origin of electron mass would be a huge accomplishment

electron Yukawa: see d'Enterria, Poldaru, Wojcik, [2107.02686](#)

desirable features of the next **major** HEP project(s)?

an important target to be reached ~ guaranteed discovery

?

exploration into the unknown by a significant factor in energy

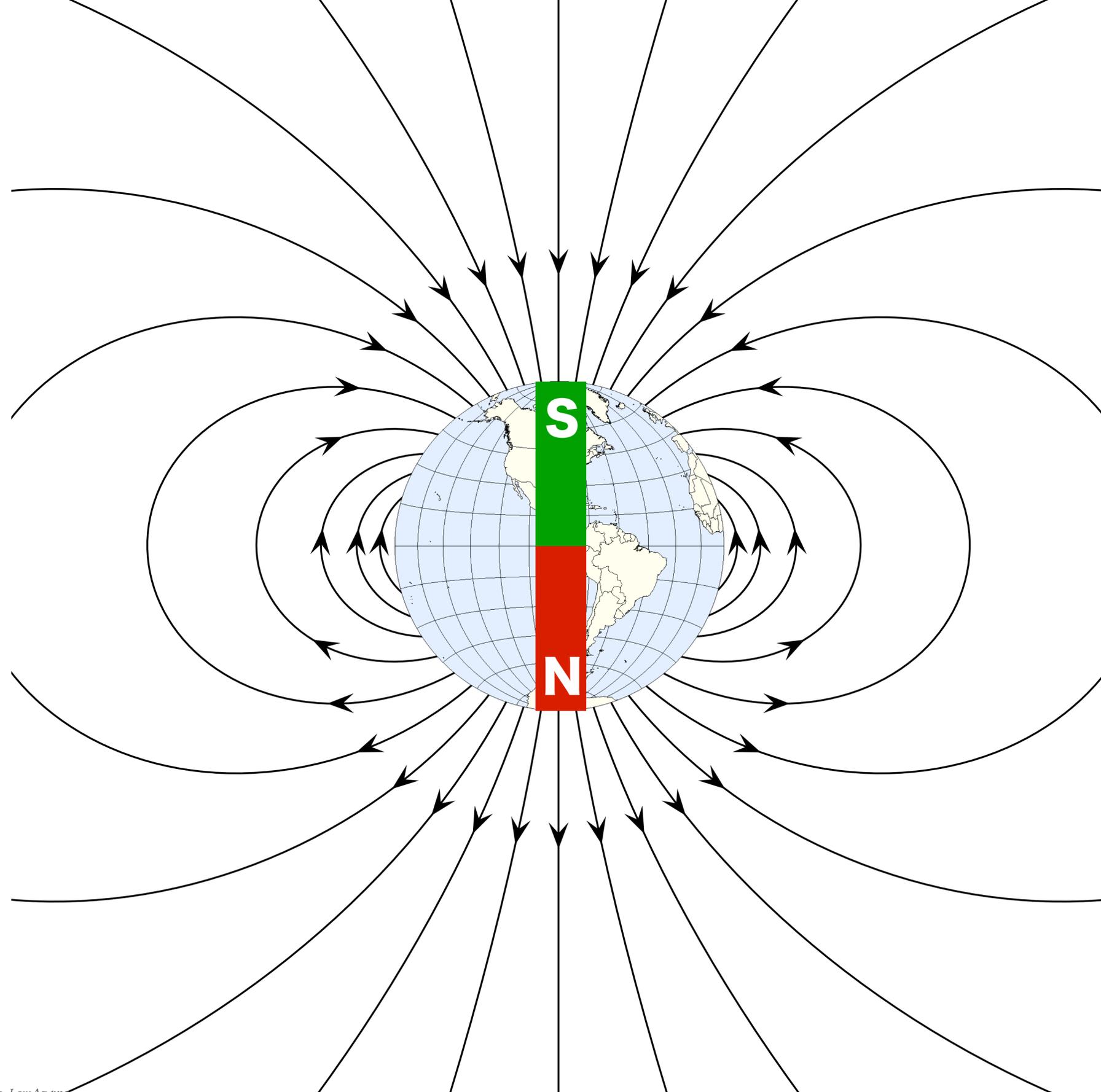
major progress on a broad array of particle physics topics

likelihood of success, robustness (e.g. multiple experiments)

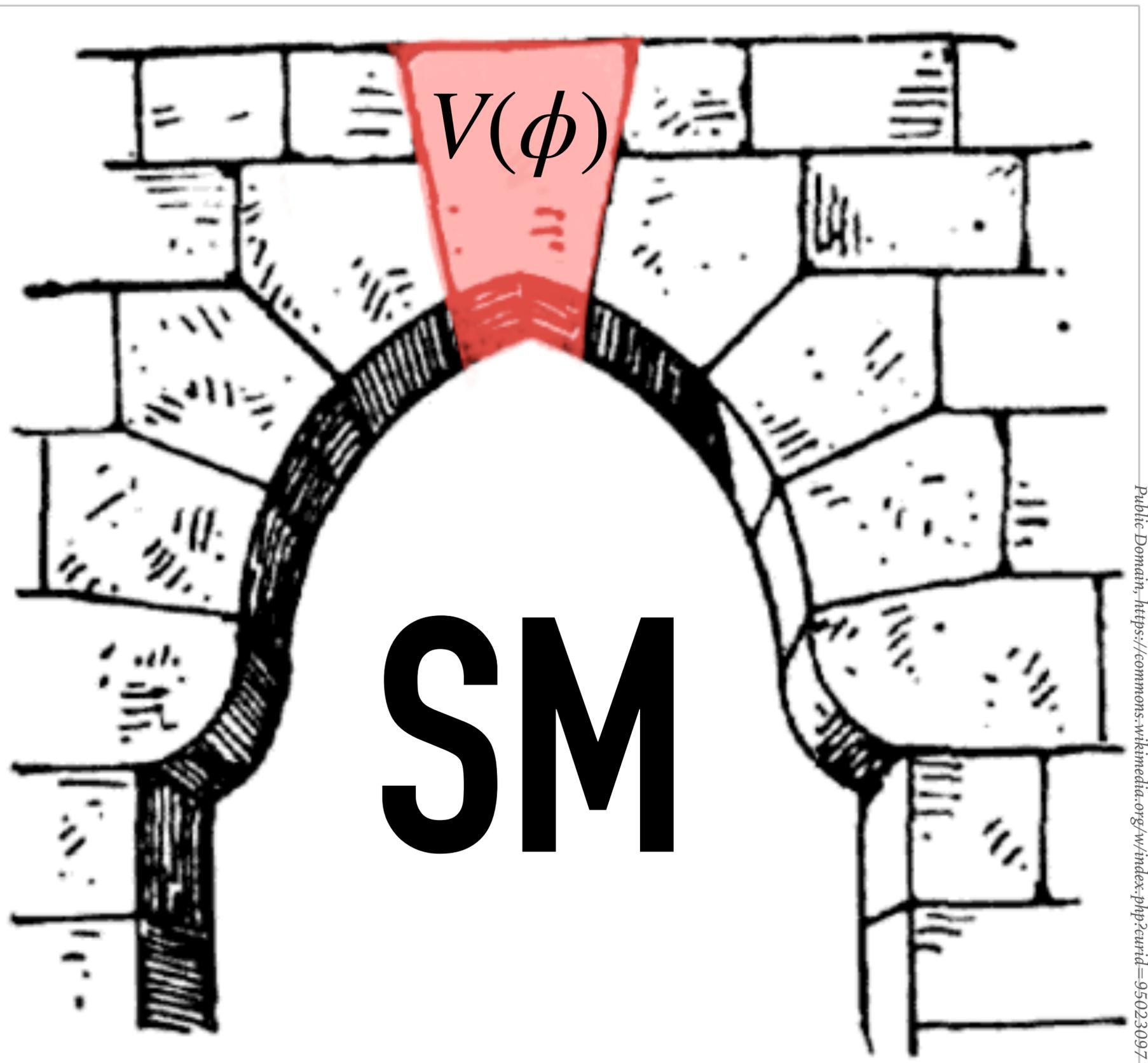
cost-effective construction & operation,
low carbon footprint, novel technologies

**fundamental particles only get
mass if the Higgs field is
non-zero**

Why is the Higgs field non-zero?



https://commons.wikimedia.org/wiki/File:VFpt_Dipole_field.svg
https://en.wikipedia.org/wiki/Western_Hemisphere#/media/File:Western_Hemisphere_LamAz.png



Public Domain, <https://commons.wikimedia.org/w/index.php?curid=95023097>

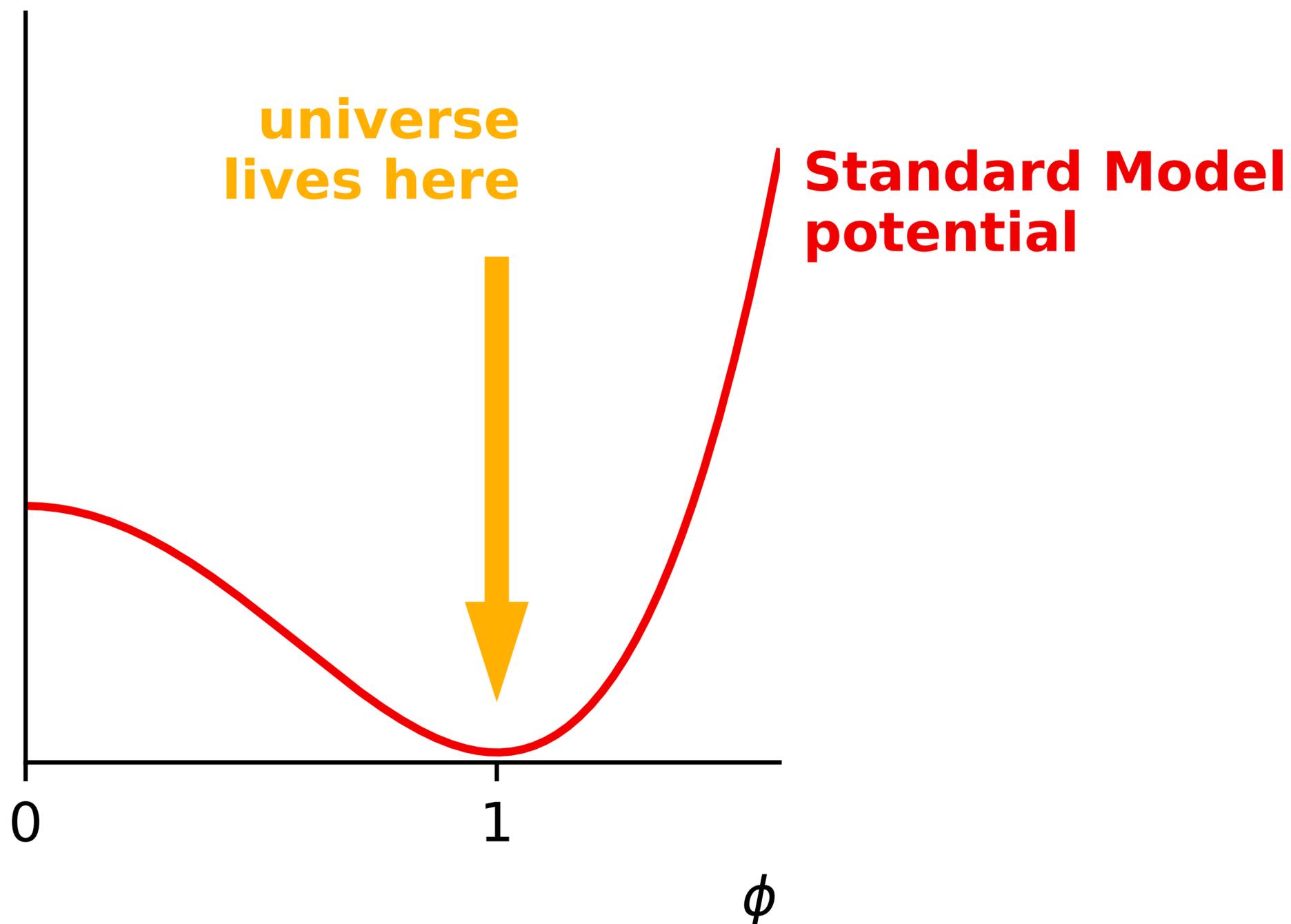
unique among all the fields we know, the Higgs field is the only one that is non-zero “classically”

**Why?
Higgs potential?**

Keystone of SM

Higgs potential

$V(\phi)$, SM



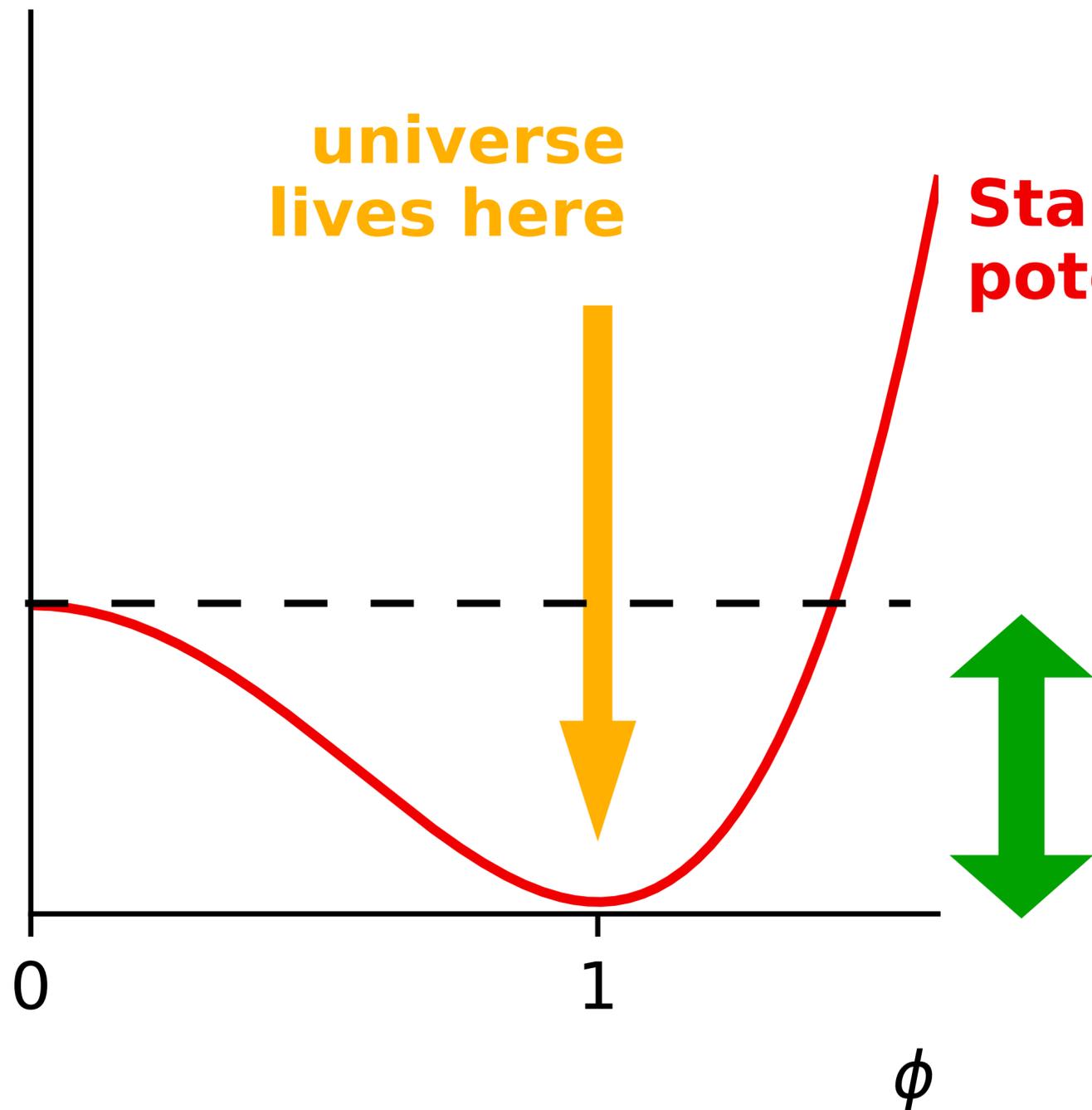
The Higgs field is non-zero because that ensures the lowest potential energy

The SM proposes a very specific form for the potential as a function of the Higgs field

$$V(\phi) \sim -\mu^2\phi^2 + \lambda\phi^4$$

Higgs potential – remember: it's an energy density

$V(\phi)$, SM



Corresponds to an energy density of $1.5 \times 10^{10} \text{ GeV/fm}^3$

$E = mc^2 \rightarrow$ Mass density of $2.6 \times 10^{28} \text{ kg/m}^3$
i.e. >40 billion times nuclear density

https://en.wikipedia.org/wiki/Globe#/media/File:World_Globe_Map.jpg

https://en.wikipedia.org/wiki/Old_fashioned_glass#/media/File:Old_Fashioned_Glass.jpg

Harder Stadium, photo from google street view, terms say OK to

“publicly display content with proper attribution online, in video, and in print.”





https://en.wikipedia.org/wiki/Globe#/media/File:World_Globe_Map.jpg

https://en.wikipedia.org/wiki/Old_fashioned_glass#/media/File:Old_Fashioned_Glass.jpg

Harder Stadium, photo from google street view, terms say OK to

“publicly display content with proper attribution online, in video, and in print.”

Earth at neutron star density



Earth at neutron star density

https://en.wikipedia.org/wiki/Globe#/media/File:World_Globe_Map.jpg
https://en.wikipedia.org/wiki/Old_fashioned_glass#/media/File:Old_Fashioned_Glass.jpg
Harder Stadium, photo from google street view, terms say OK to
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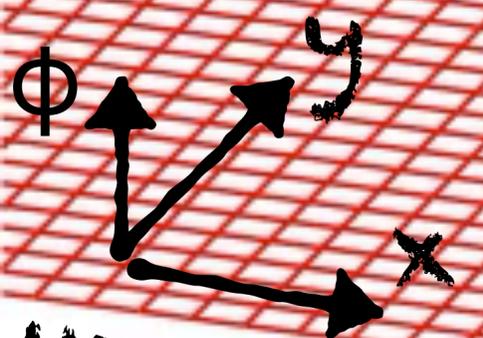


Earth at Higgs potential density

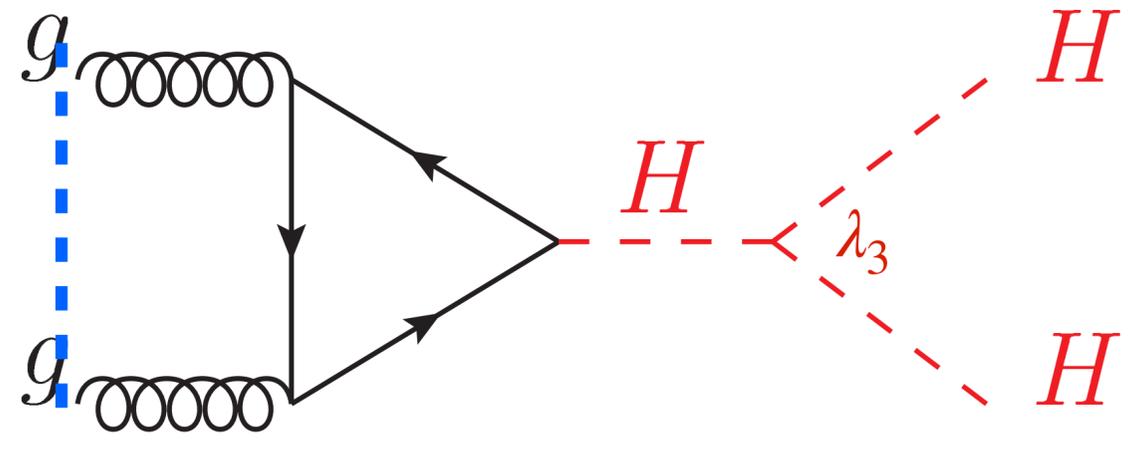
quon



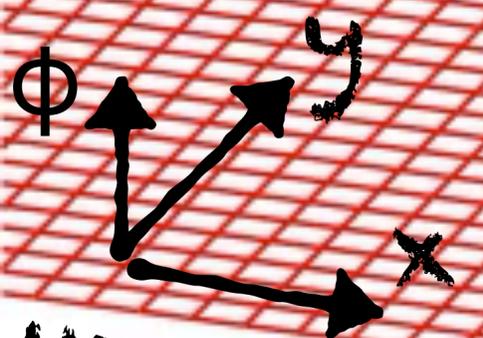
gluon



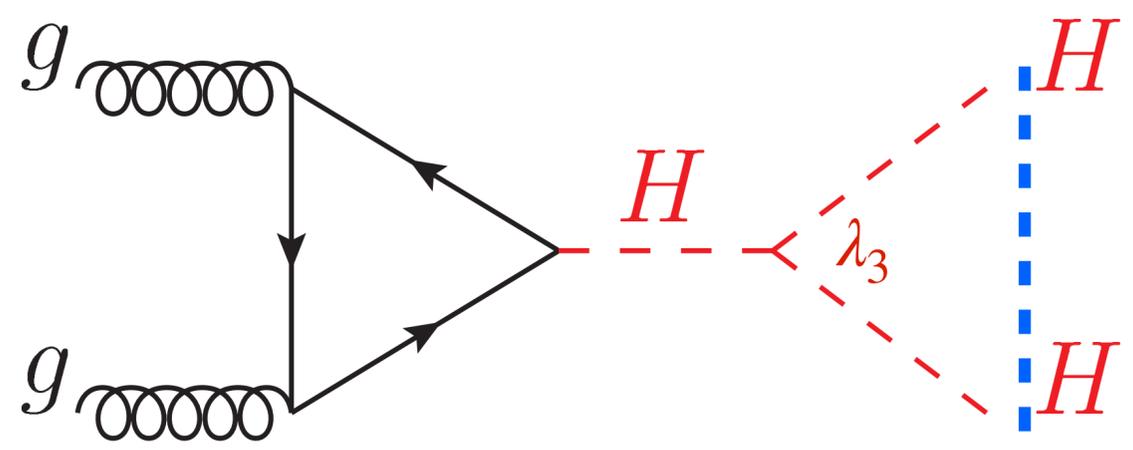
Higgs field in space



quon



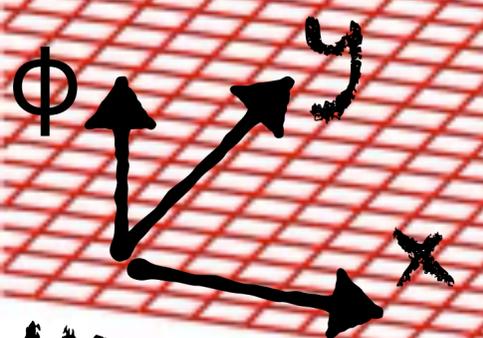
Higgs field in space



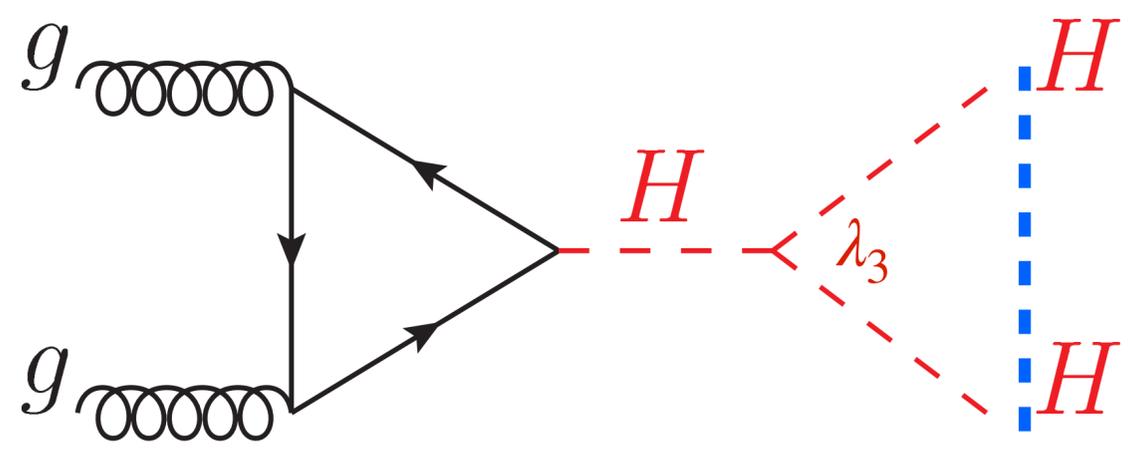
gluon



quon



Higgs field in space

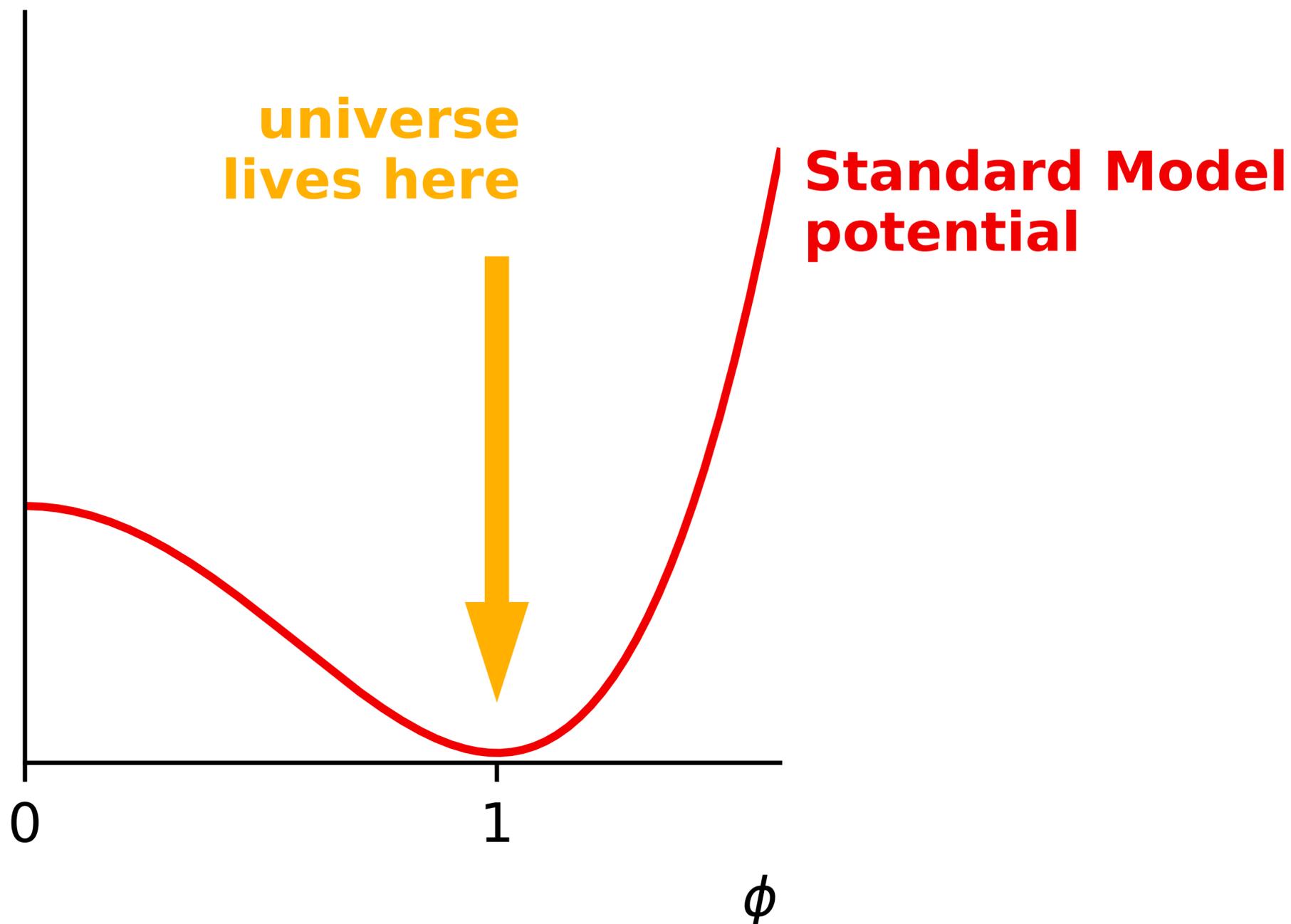


gluon



Higgs potential

$V(\phi)$, SM

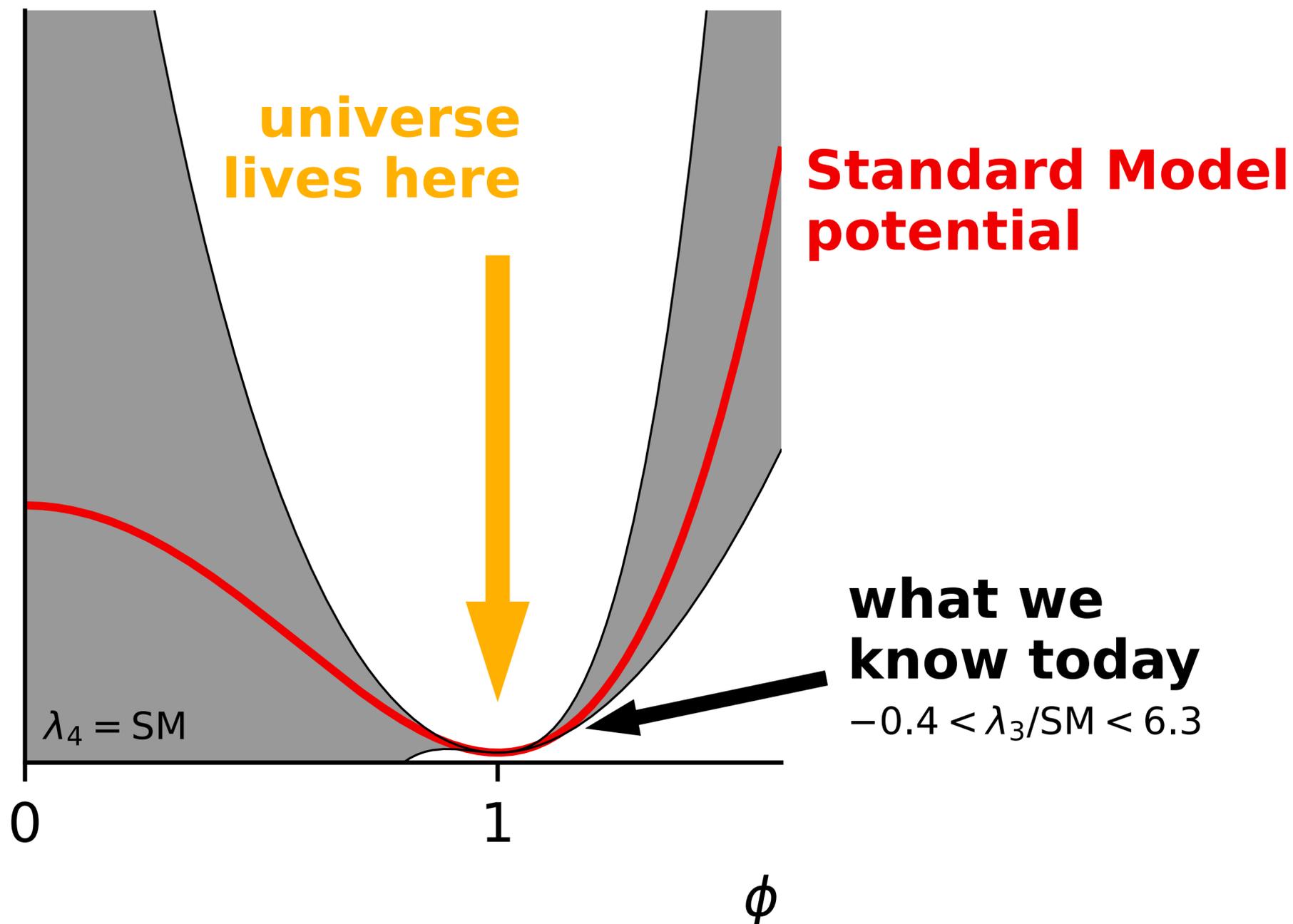


Studying $H \rightarrow HH$ probes specific mathematical property of the potential's shape: its third derivative (λ_3), i.e. how asymmetric it is at the minimum

[reconstruction in plot assumes higher derivatives as in SM]

Higgs potential

$V(\phi)$, today

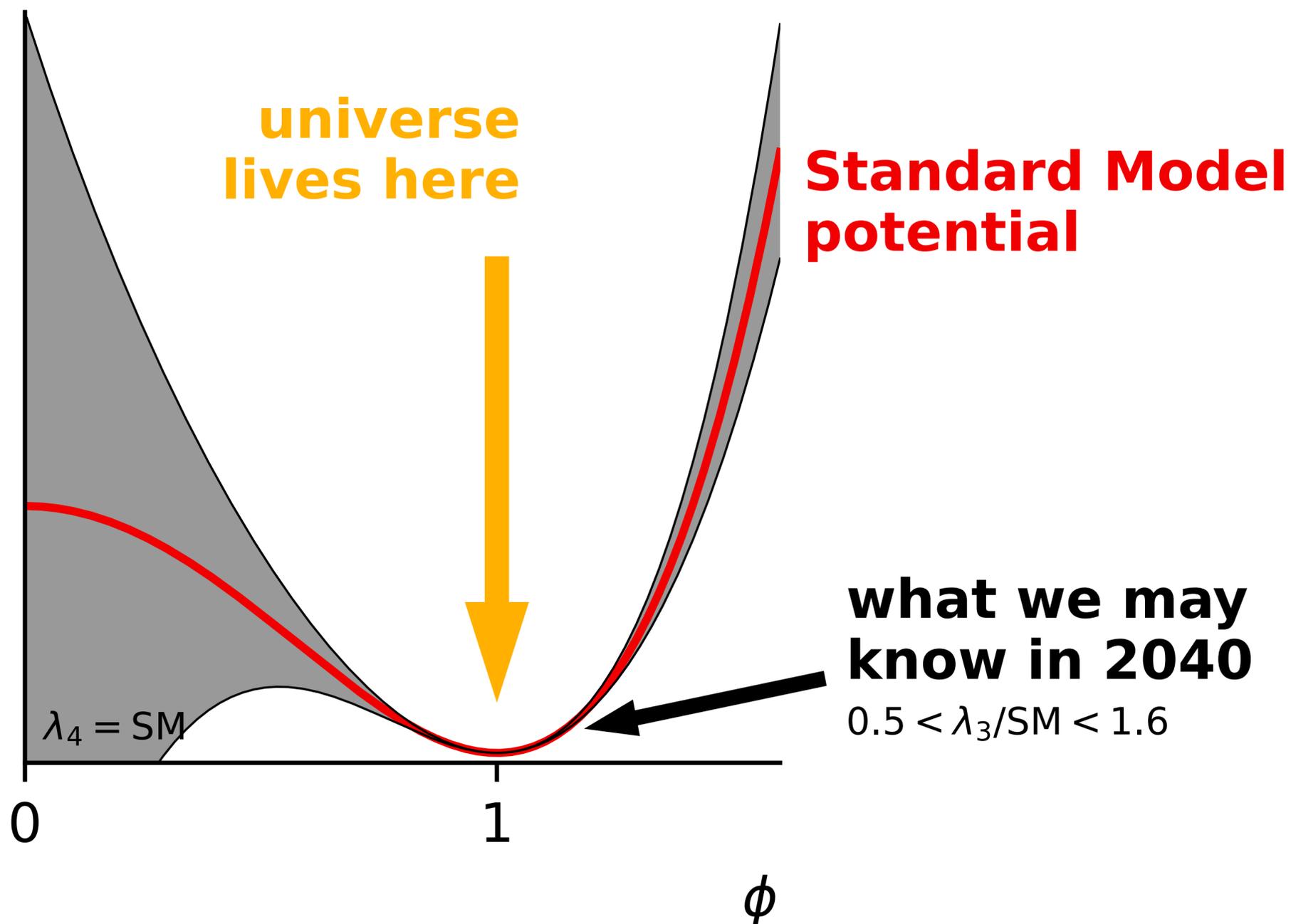


Studying $H \rightarrow HH$ probes specific mathematical property of the potential's shape: its third derivative (λ_3), i.e. how asymmetric it is at the minimum

[reconstruction in plot assumes higher derivatives as in SM]

Higgs potential

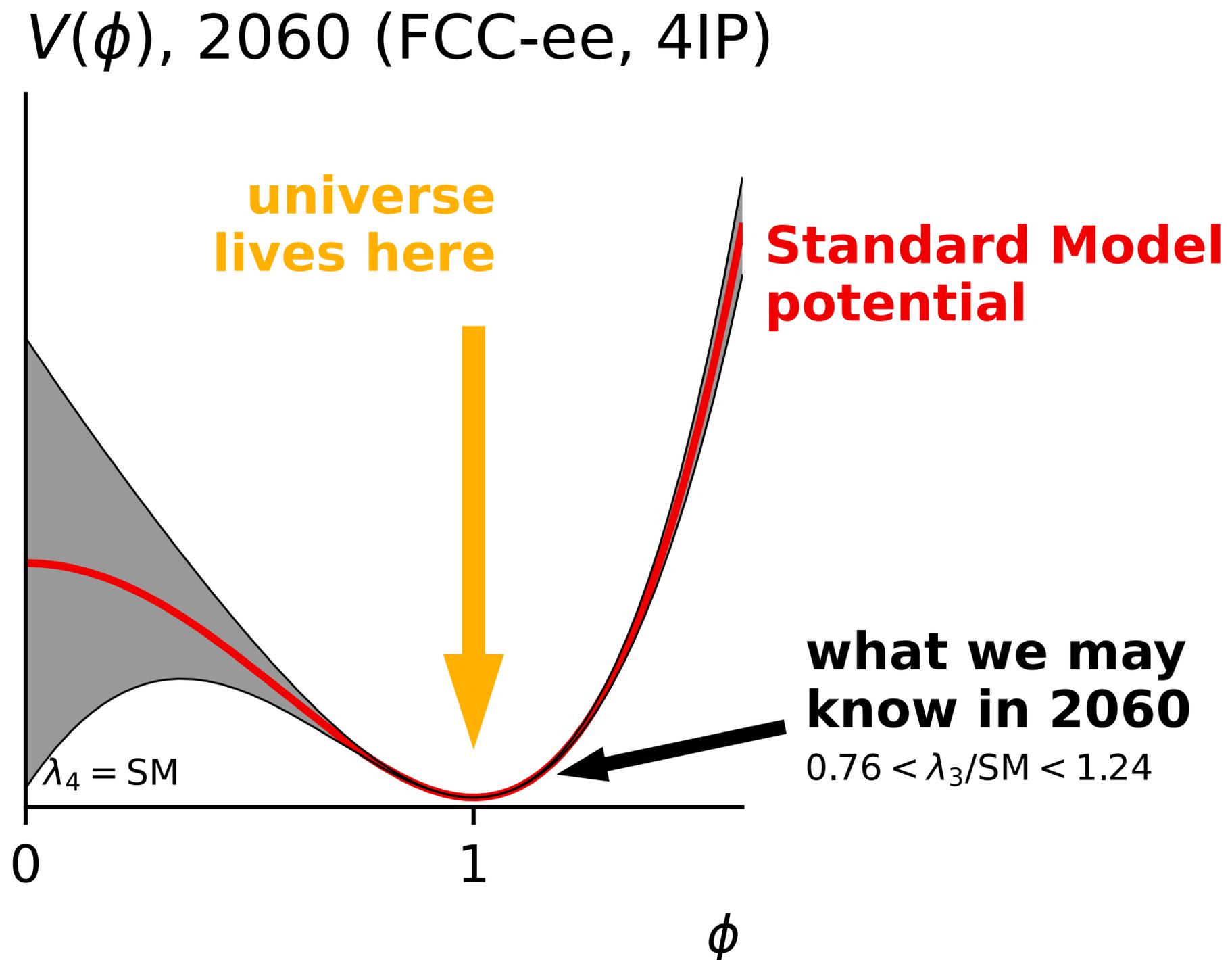
$V(\phi)$, 2040 (HL-LHC)



Studying $H \rightarrow HH$ probes specific mathematical property of the potential's shape: its third derivative (λ_3), i.e. how asymmetric it is at the minimum

[reconstruction in plot assumes higher derivatives as in SM]

Higgs potential

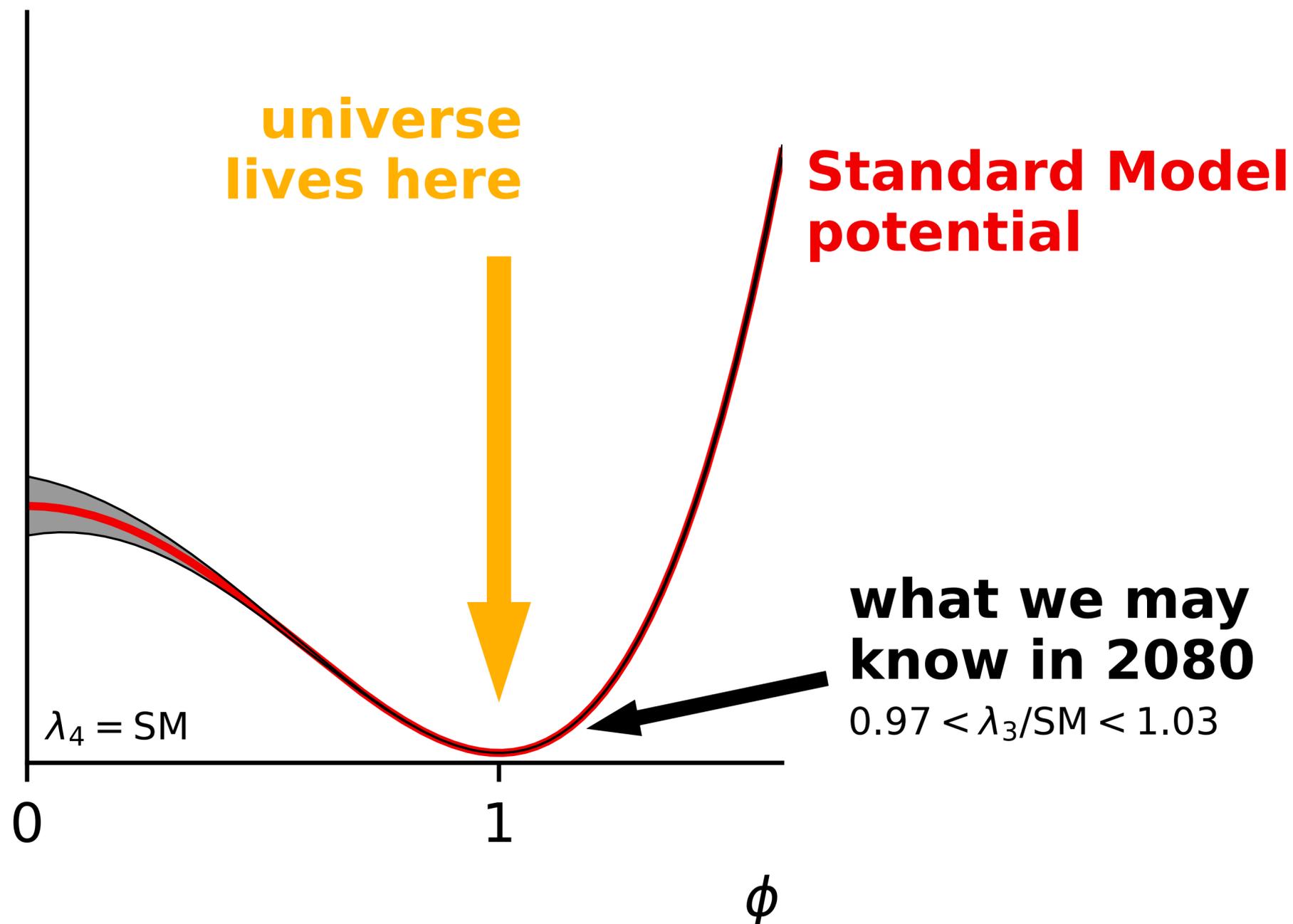


Studying $H \rightarrow HH$ probes specific mathematical property of the potential's shape: its third derivative (λ_3), i.e. how asymmetric it is at the minimum

[reconstruction in plot assumes higher derivatives as in SM]

Higgs potential

$V(\phi)$, 2080 (FCC-hh)

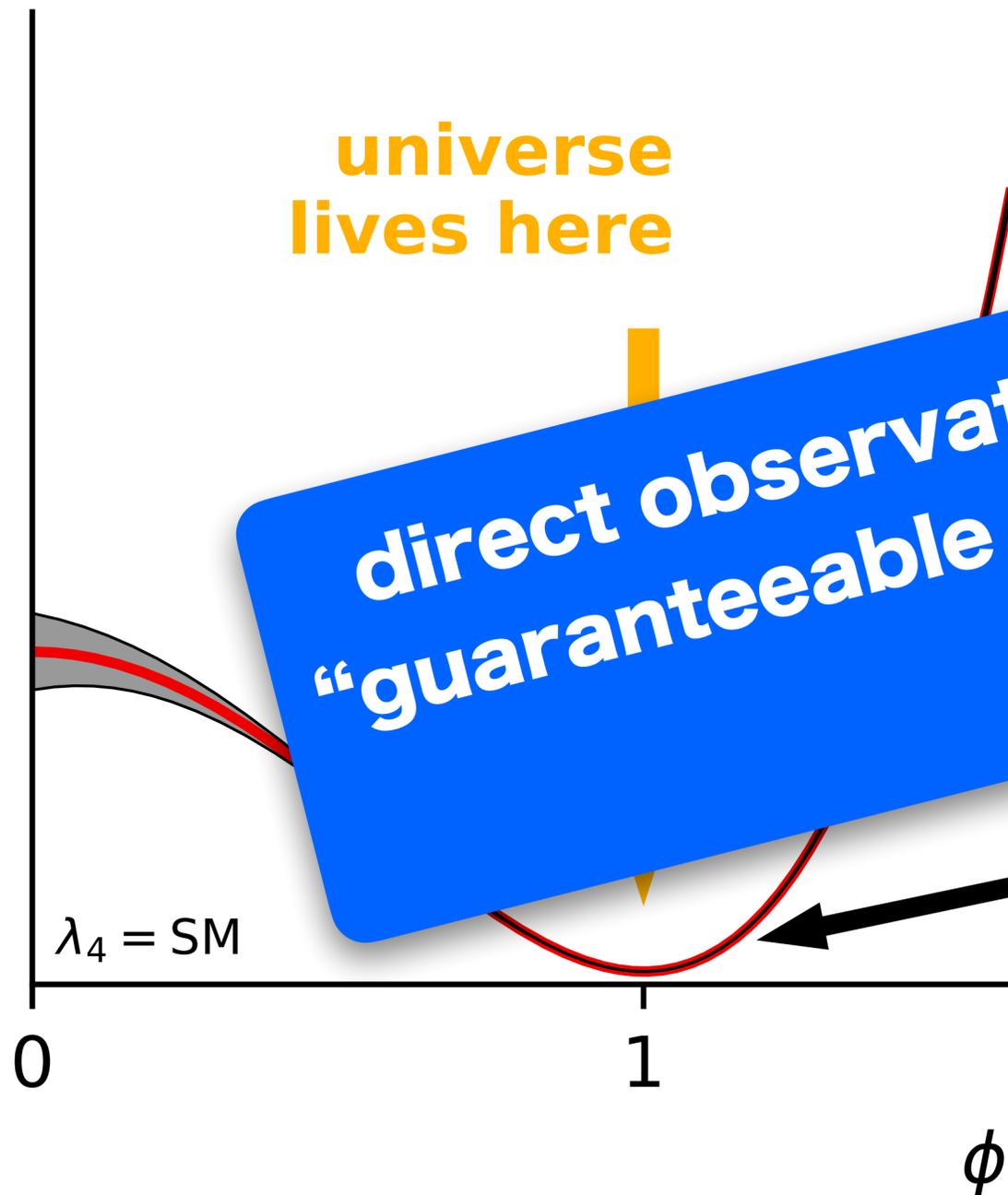


Studying $H \rightarrow HH$ probes specific mathematical property of the potential's shape: its third derivative (λ_3), i.e. how asymmetric it is at the minimum

[reconstruction in plot assumes higher derivatives as in SM]

Higgs potential

$V(\phi)$, 2080 (FCC-hh)



universe lives here

Standard Model

direct observation of $H \rightarrow HH$ interaction is a "guaranteeable discovery" that HEP should be aiming for

what we may know in 2080

$$0.97 < \lambda_3 / \text{SM} < 1.03$$

Studying $H \rightarrow HH$ probes

specific physical property of the potential's shape:

the cubic term (λ_3),

if symmetric it is zero at the minimum

[reconstruction in plot assumes higher derivatives as in SM]

Almost every problem of the Standard Model originates from Higgs interactions

$$\mathcal{L} = y H \psi \bar{\psi} + \mu^2 |H|^2 - \lambda |H|^4 - V_0$$



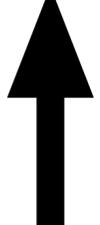
flavour



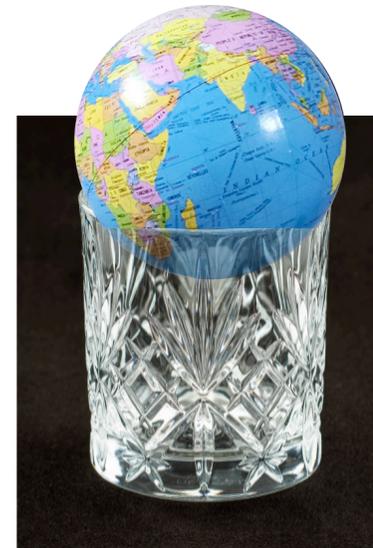
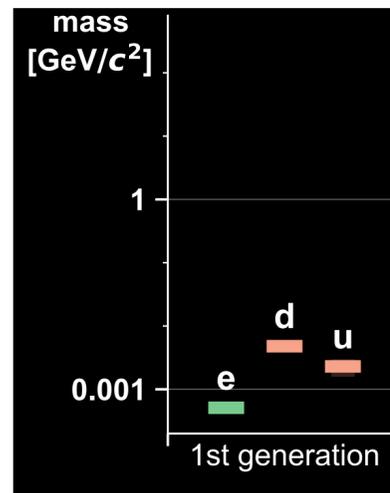
naturalness



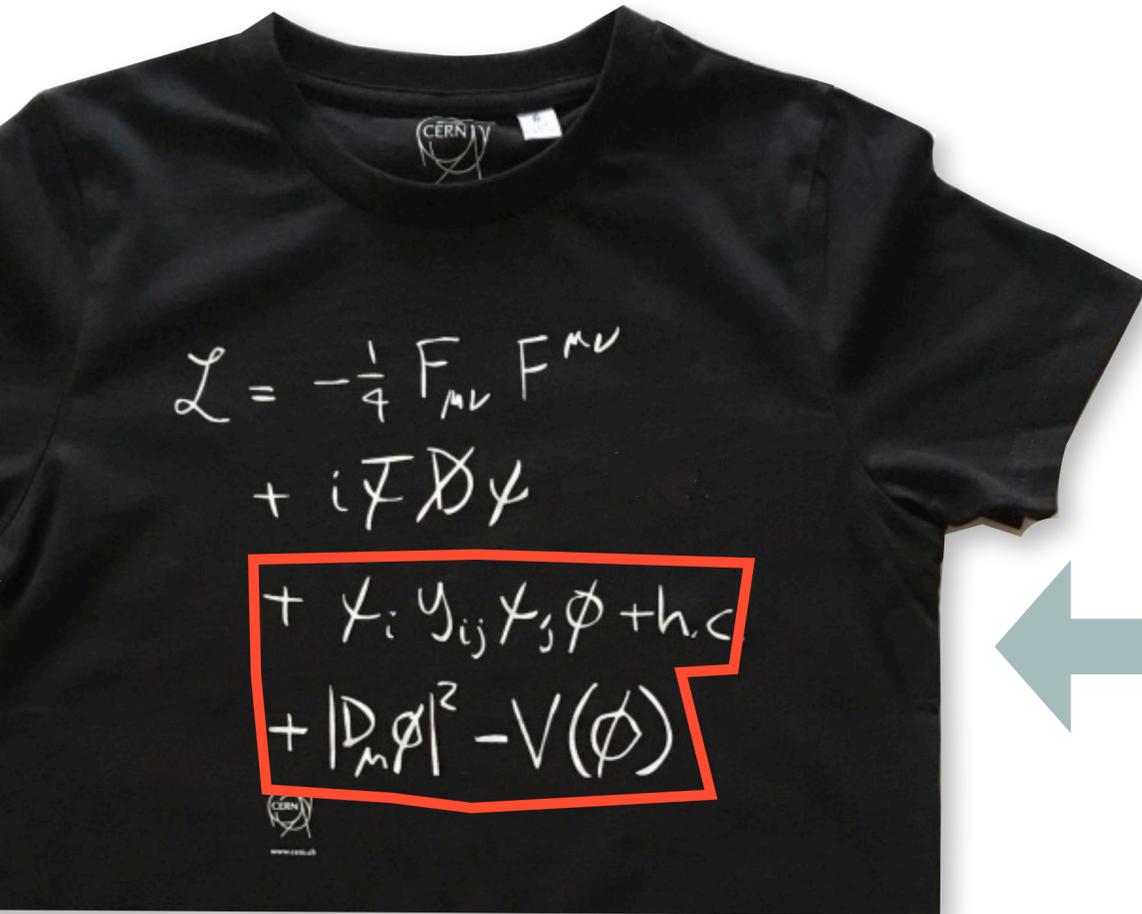
stability



cosmological constant



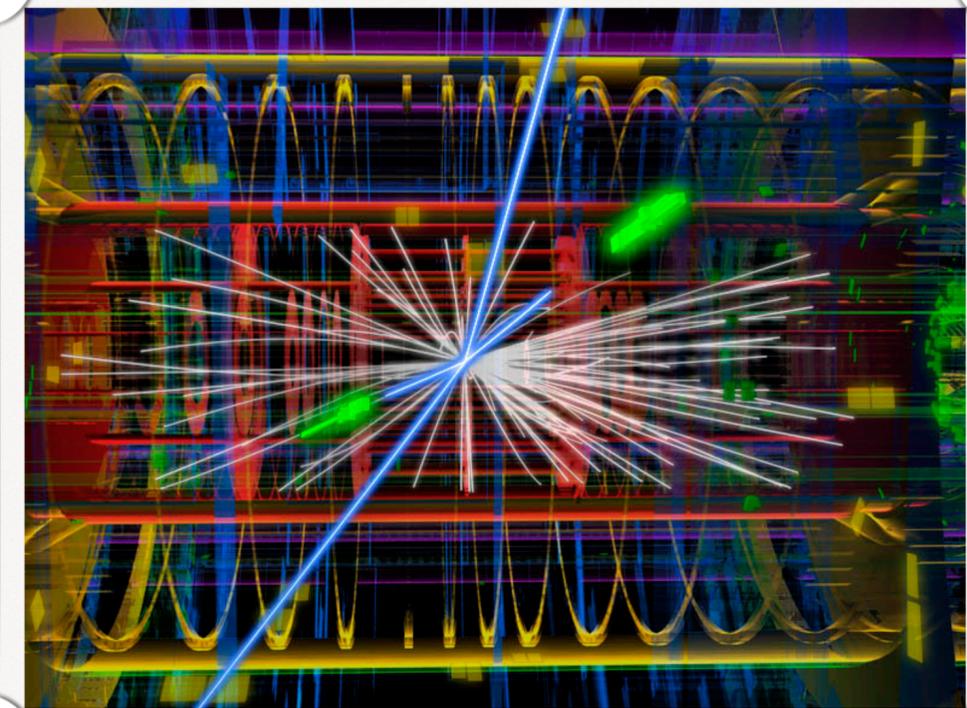
UNDERLYING THEORY



*how do you make
quantitative
connection?*



EXPERIMENTAL DATA



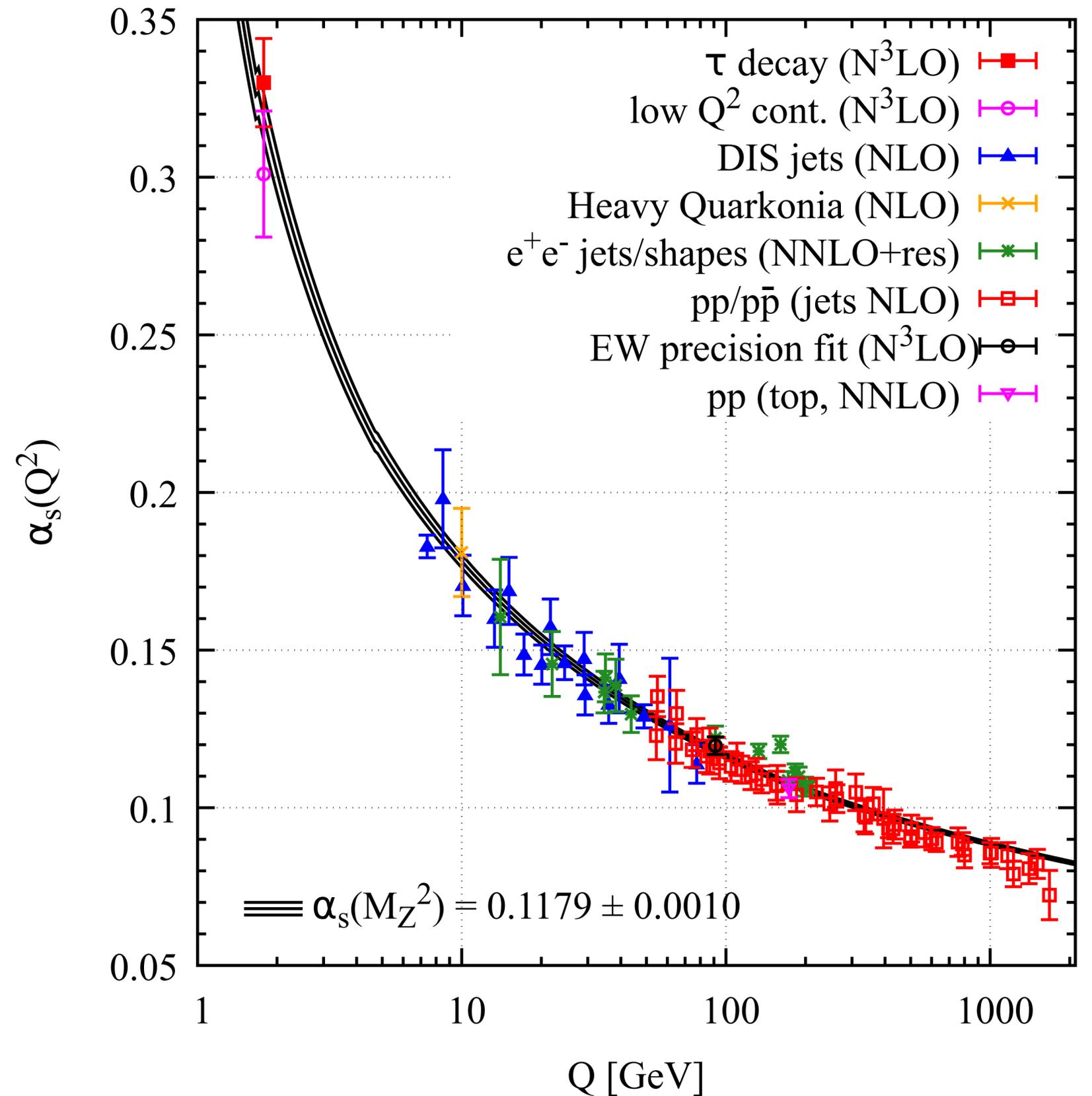
QCD

quantum chromodynamics
the theory of the strong
interaction

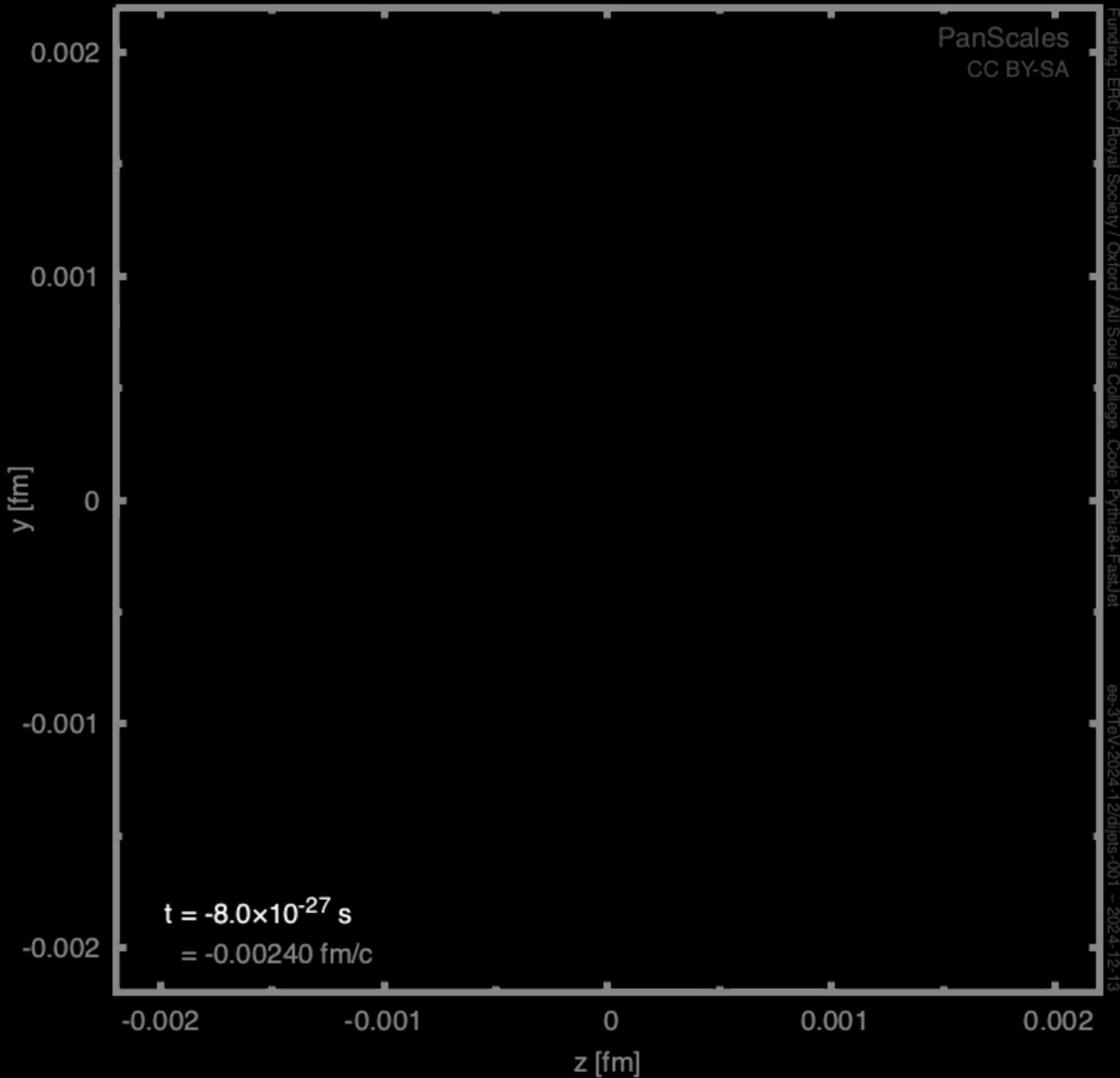
Like QED, with key differences

- Charge comes in three variants (red, green, blue)
- Force carrier (gluon), is charged
- Coupling is larger (and non-perturbative at small momenta)

strong coupling (α_s) v. momentum scale

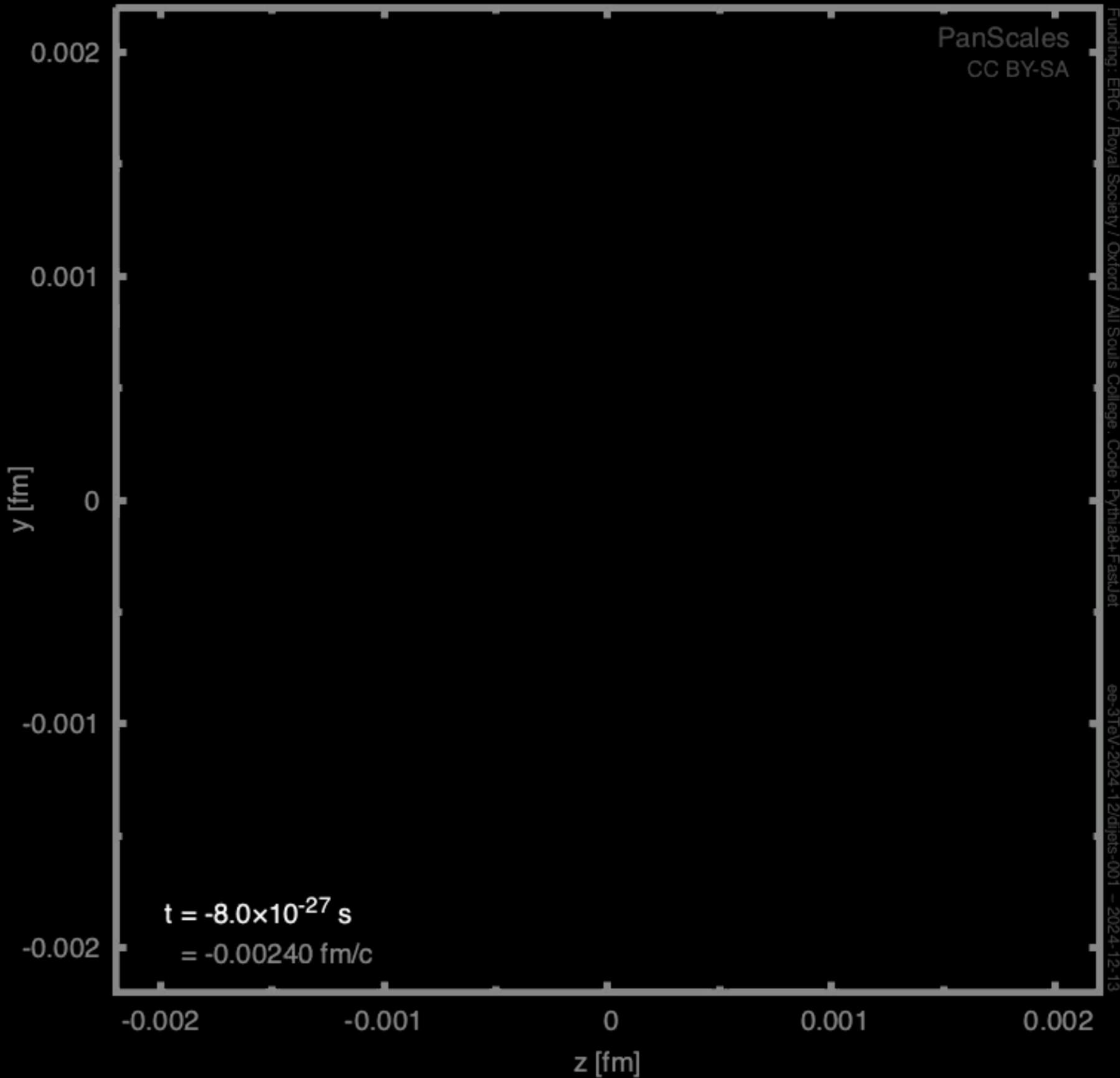


What actually happens in a collision?



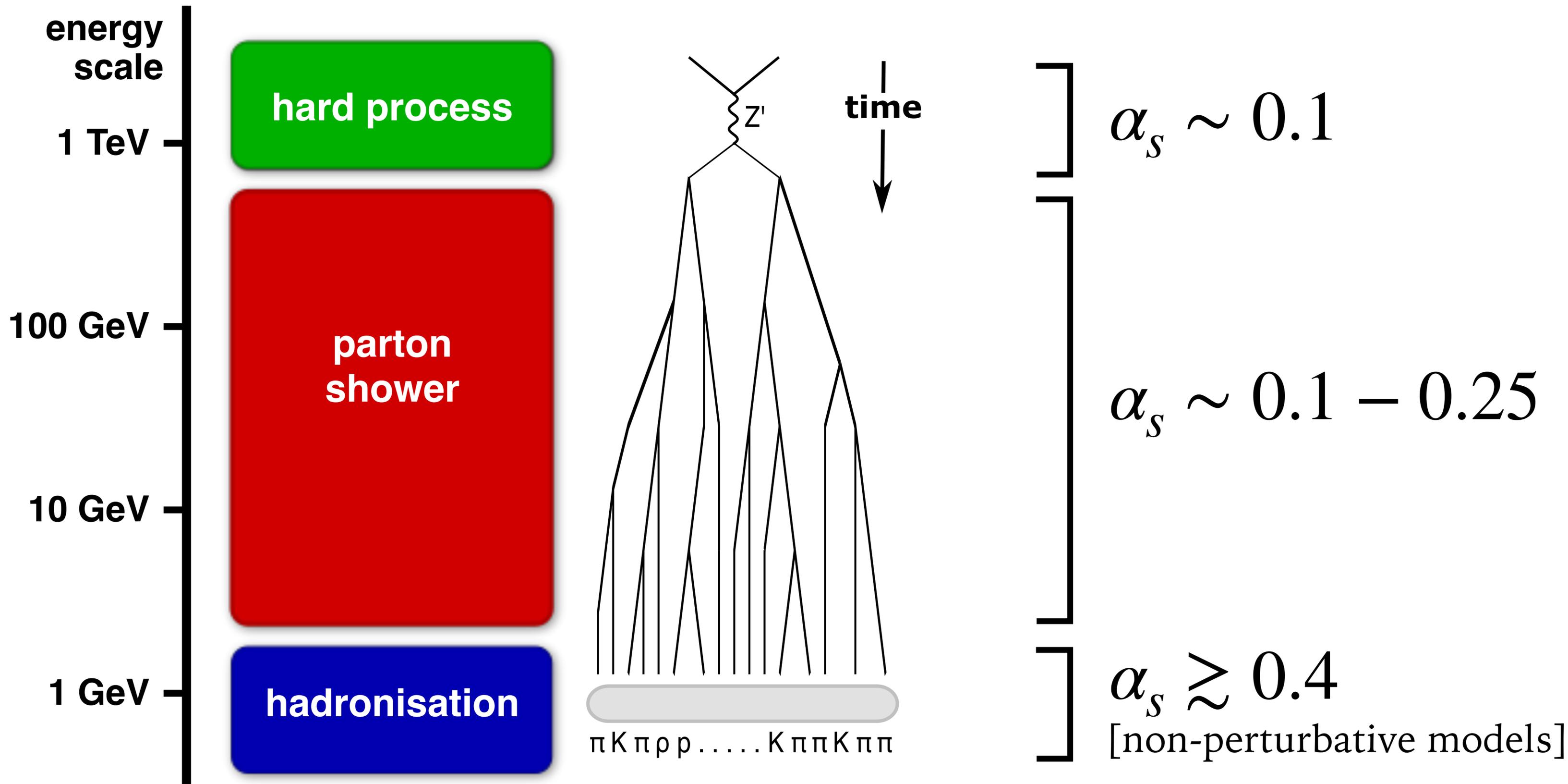
-  incoming beam particle
-  intermediate particle
(quark or gluon)
-  final particle (hadron)

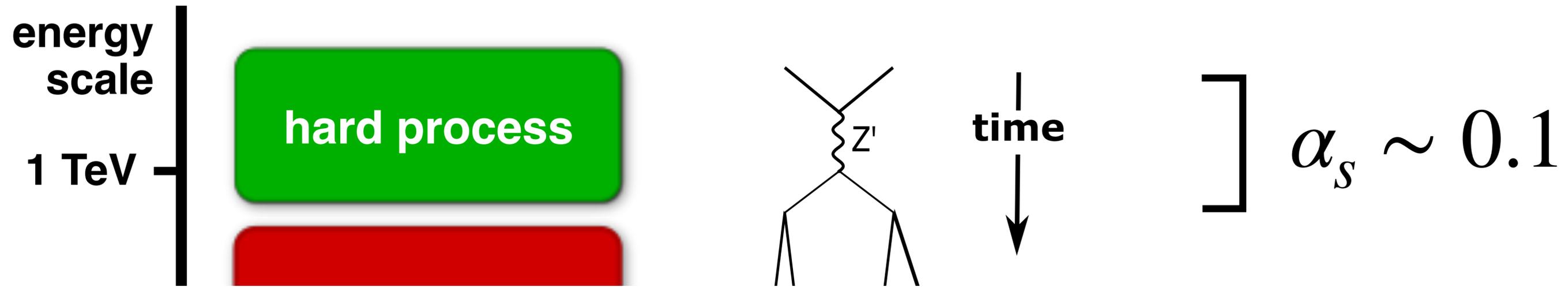
Event evolution spans 7 orders of magnitude in space-time



-  incoming beam particle
-  intermediate particle
(quark or gluon)
-  final particle (hadron)

Event evolution spans 7 orders of magnitude in space-time

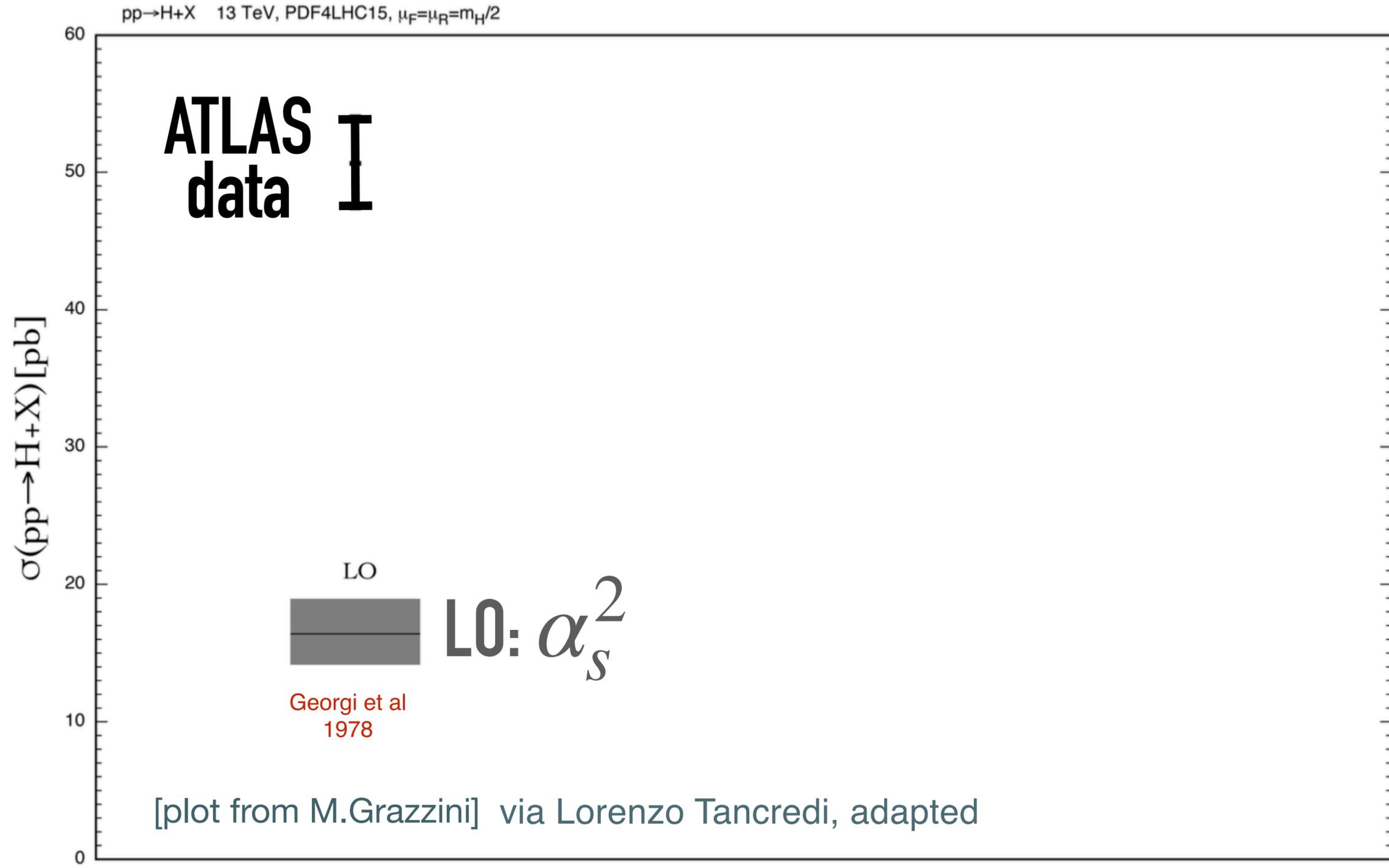




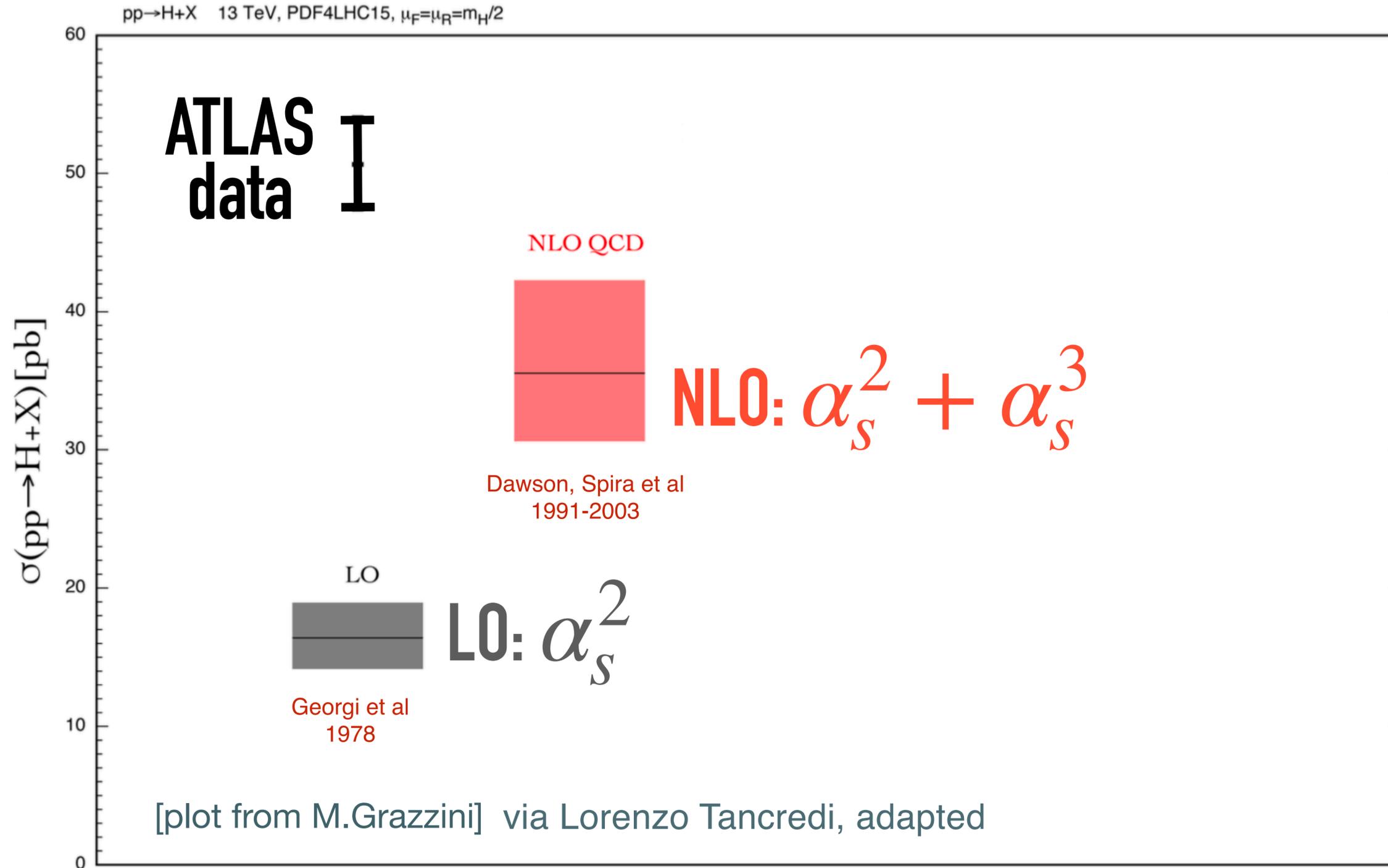
Calculate scattering cross sections as a “perturbative” series expansion in power of the strong coupling α_s

$$\sigma = c_0 + c_1 \alpha_s + c_2 \alpha_s^2 + \dots$$

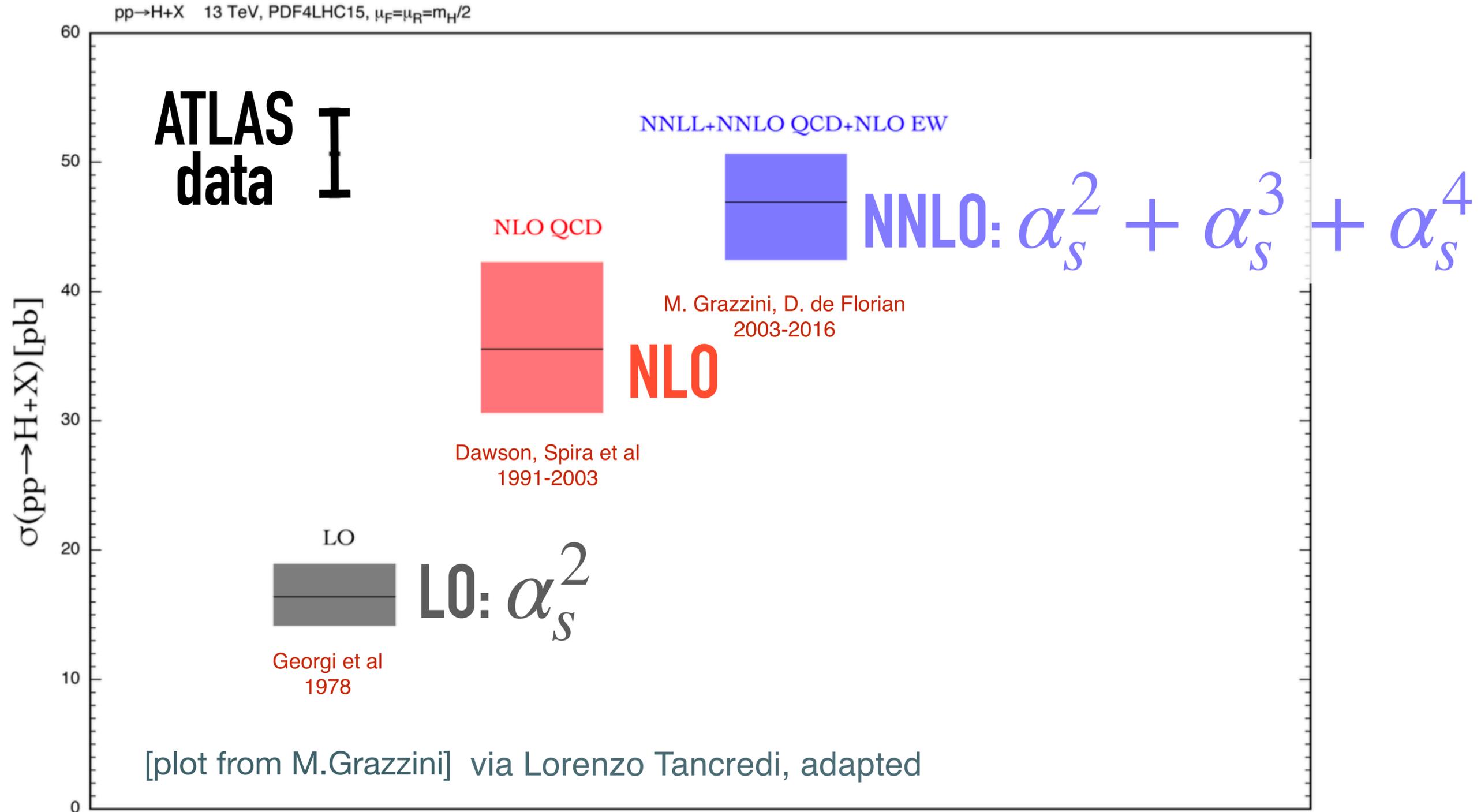
Higgs Boson Cross Section



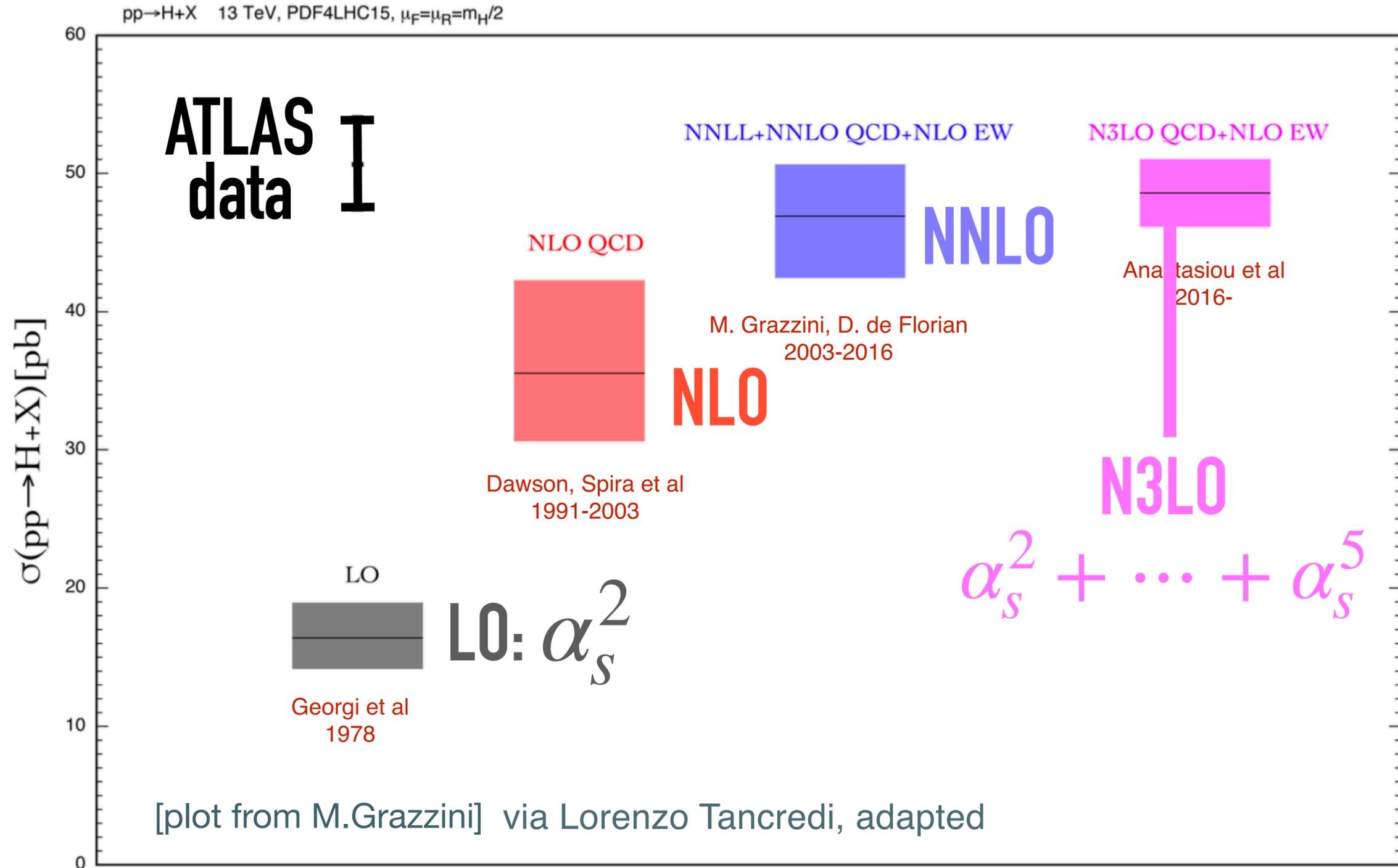
Higgs Boson Cross Section



Higgs Boson Cross Section

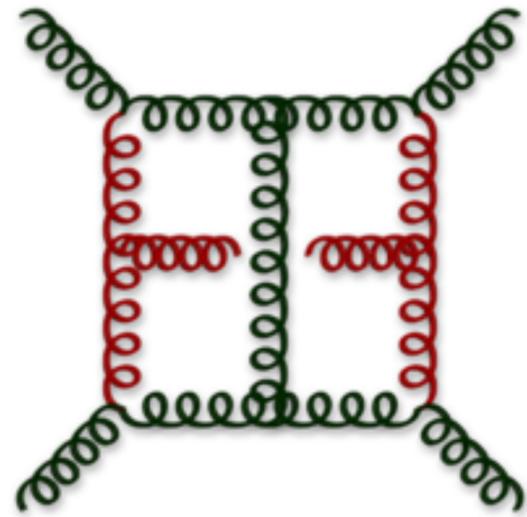


Higgs Boson Cross Section



FOUR-PARTON SCATTERING TO THREE LOOPS

[Caola, Chakraborty, Gambuti, von Manteuffel, JT '21]
Phys.Rev.Lett. 128 (2022) 21, 212001

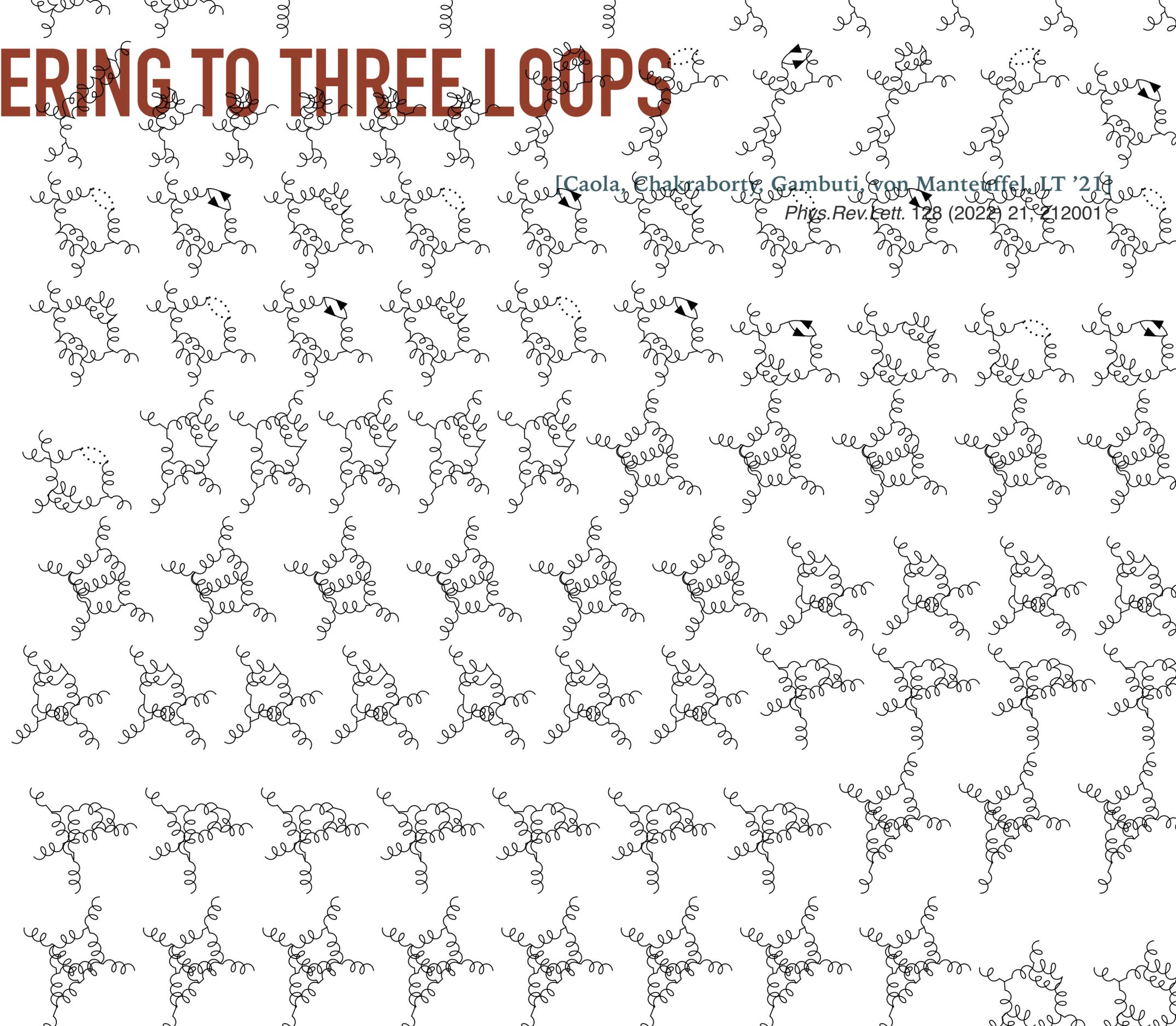


$gg \rightarrow gg$ @ 3 loops in QCD

+ 500 more pages

= 50000 Feynman diagrams

= 10^7 Feynman integrals!



selected collider-QCD accuracy milestones

Drell-Yan (γ/Z) & Higgs production at hadron colliders

LO

NLO

NNLO[.....]

N3LO

1970

1980

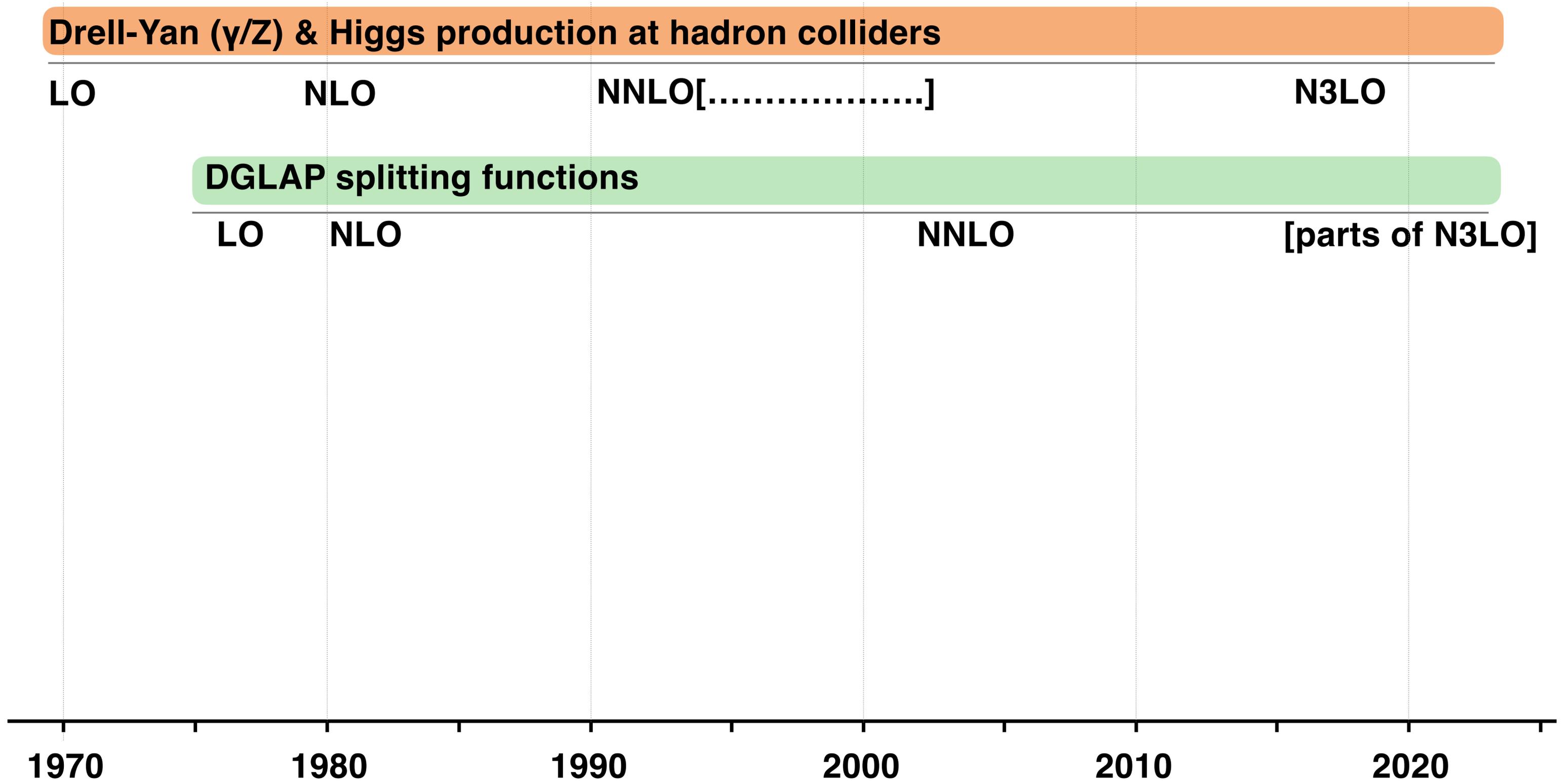
1990

2000

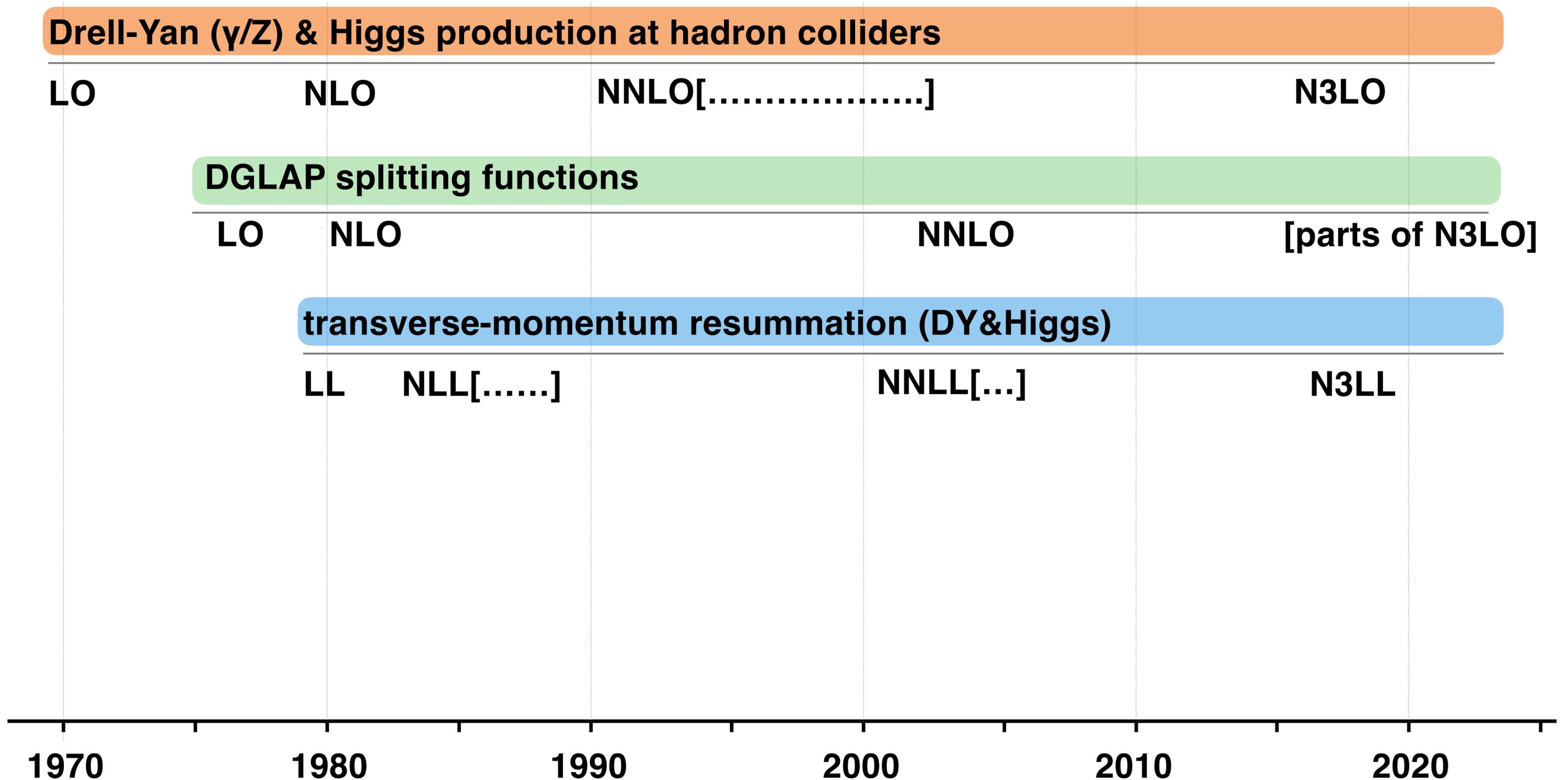
2010

2020

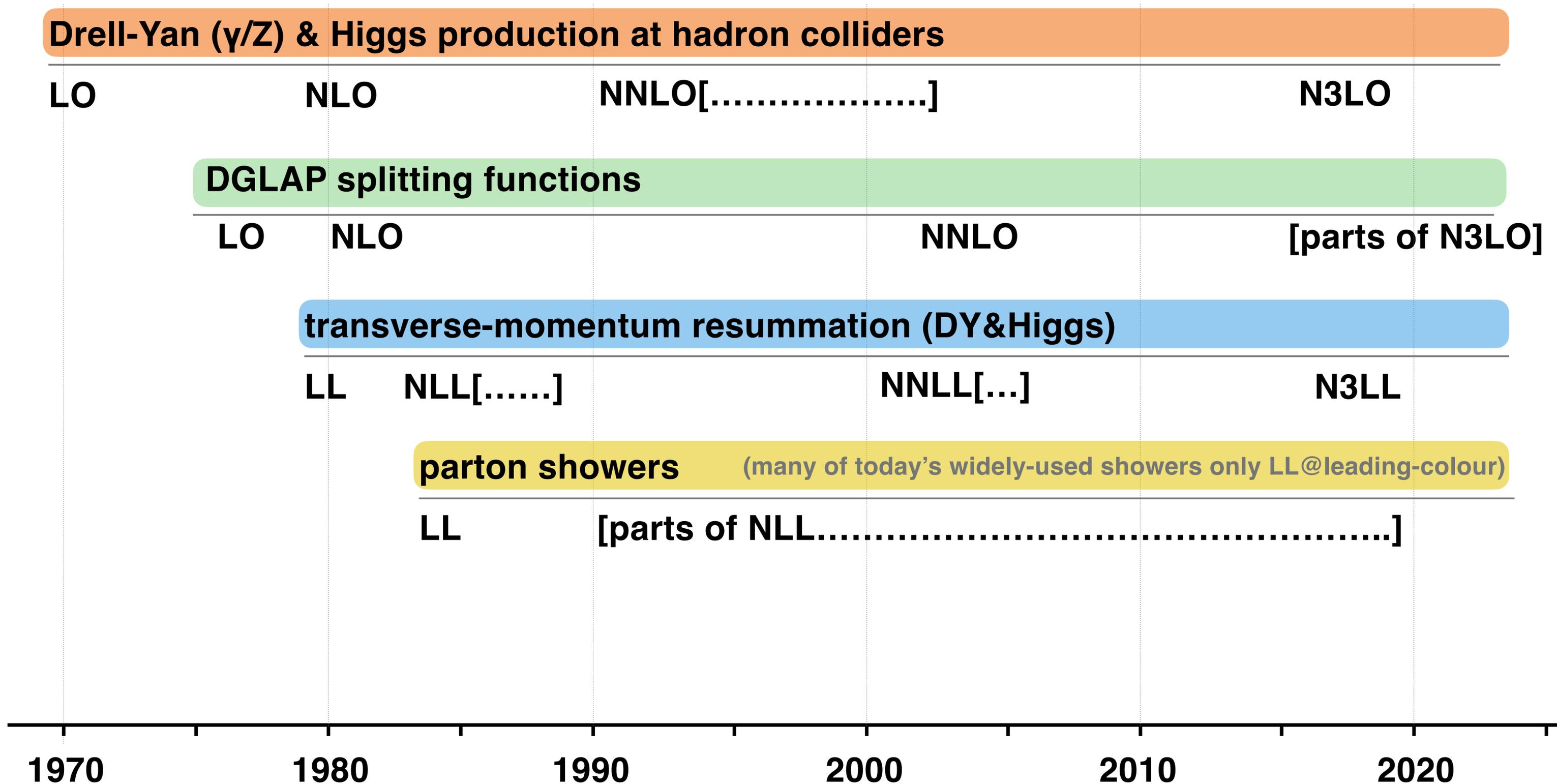
selected collider-QCD accuracy milestones



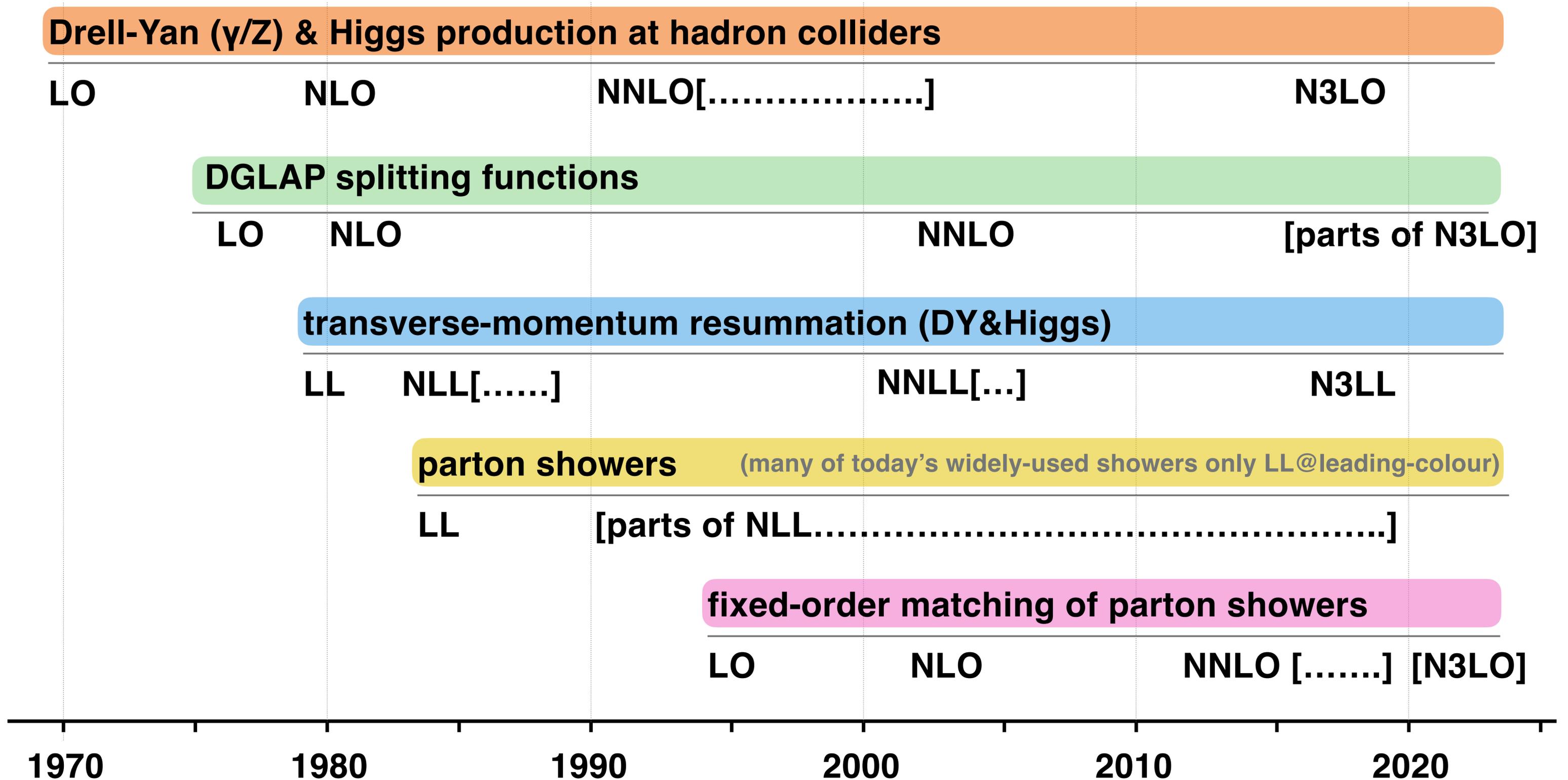
selected collider-QCD accuracy milestones



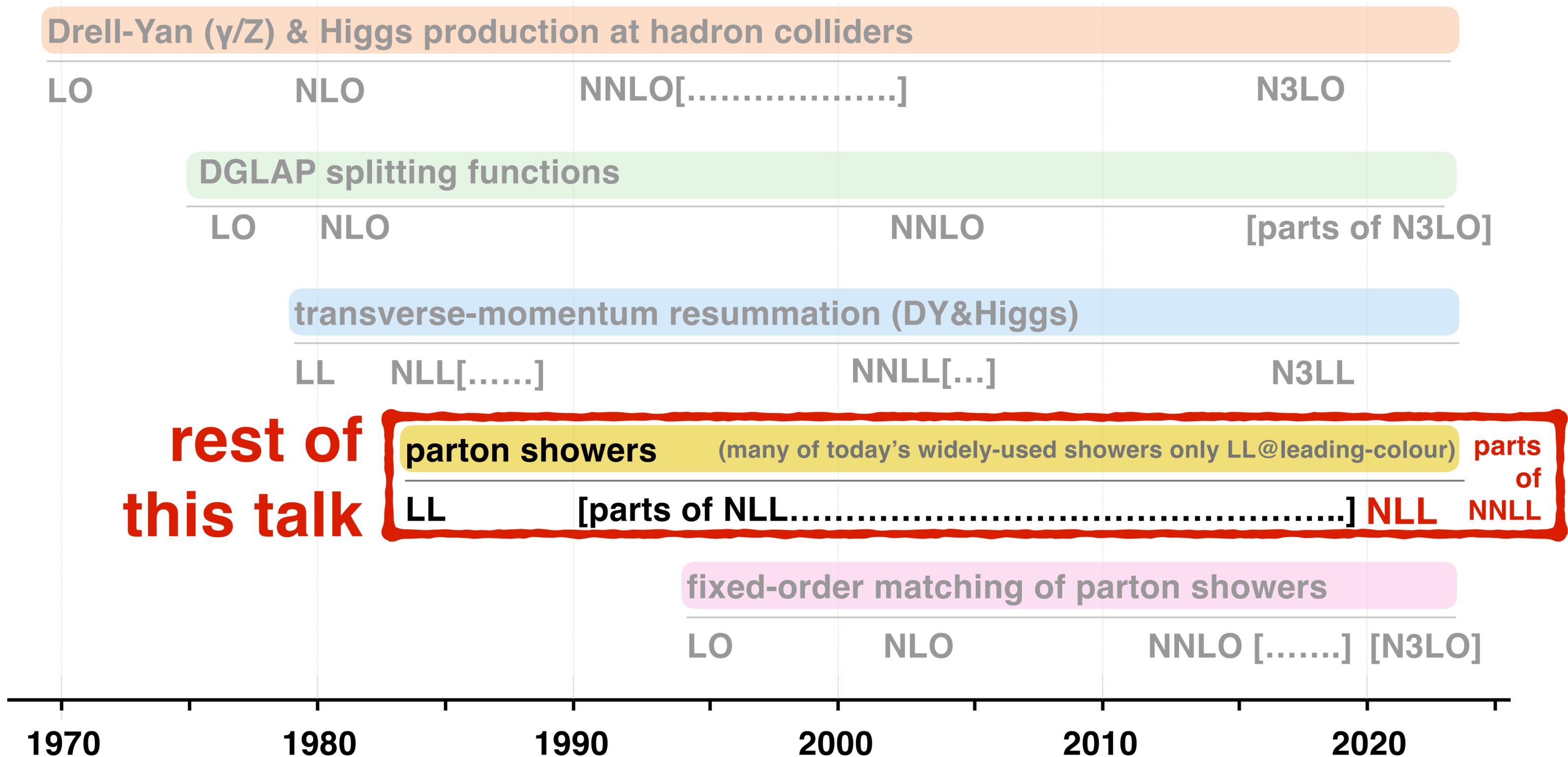
selected collider-QCD accuracy milestones



selected collider-QCD accuracy milestones



selected collider-QCD accuracy milestones



rest of
this talk

parton showers (many of today's widely-used showers only LL@leading-colour) parts of NNLL

All of this is impossible without simulations



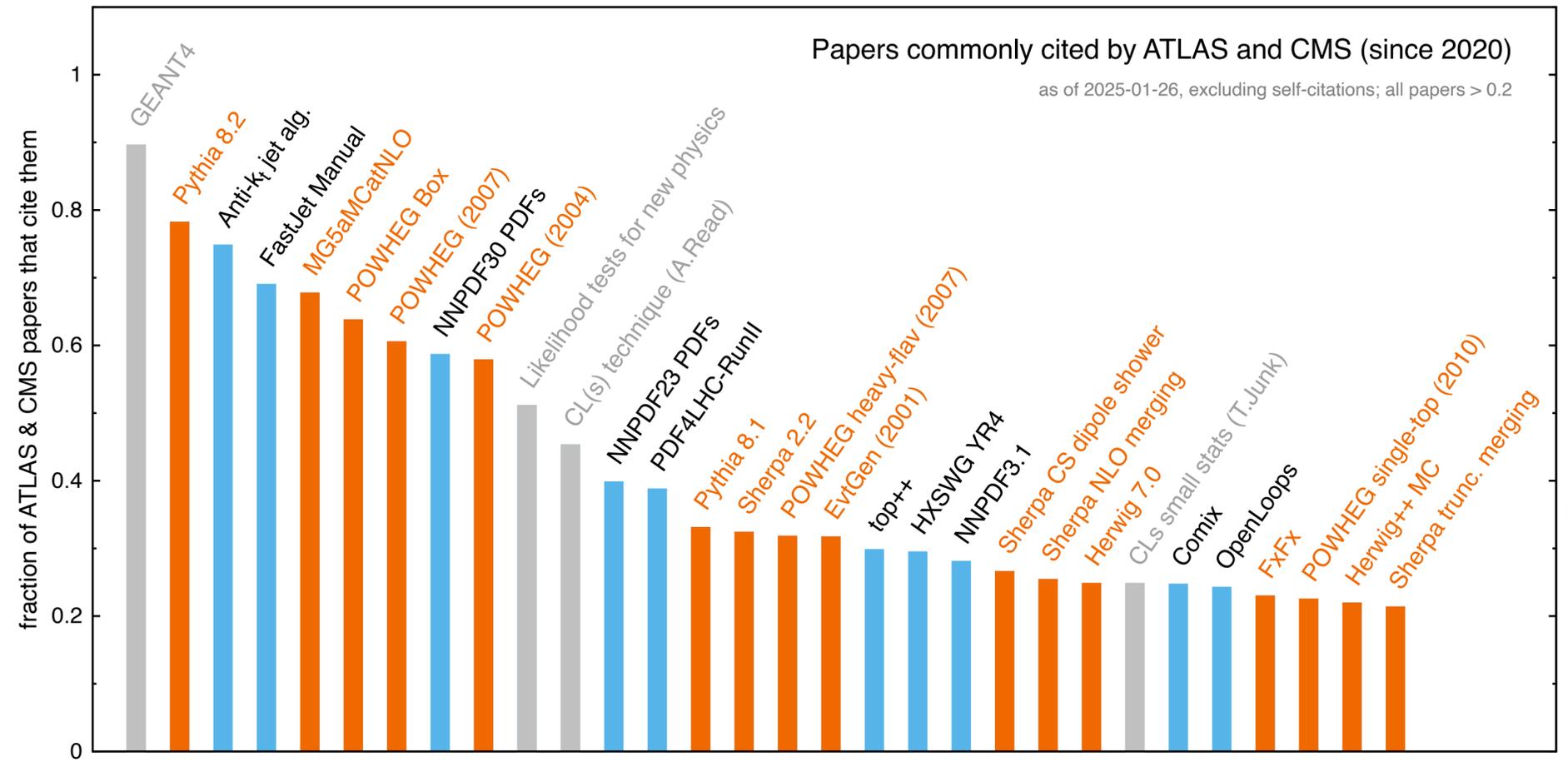
Herwig 7



Pythia 8



Sherpa 3



used in ~95% of ATLAS/CMS publications
they do an amazing job of simulating vast swathes of data;
collider physics would be unrecognisable without them

```

*-----*
|      PPP  Y  Y  TTTTT  H  H  III  A      Welcome to the Lund Monte Carlo!
|      P  P  Y Y  T      H  H  I  A A      This is PYTHIA version 8.303
|      PPP  Y  T      HHHHH  I  AAAAA      Last date of change: 1 Sep 2020
|      P    Y  T      H  H  I  A  A
|      P    Y  T      H  H  III A  A      Now is 19 Jul 2021 at 11:41:40
|
*-----*

```

----- PYTHIA Process Initialization -----

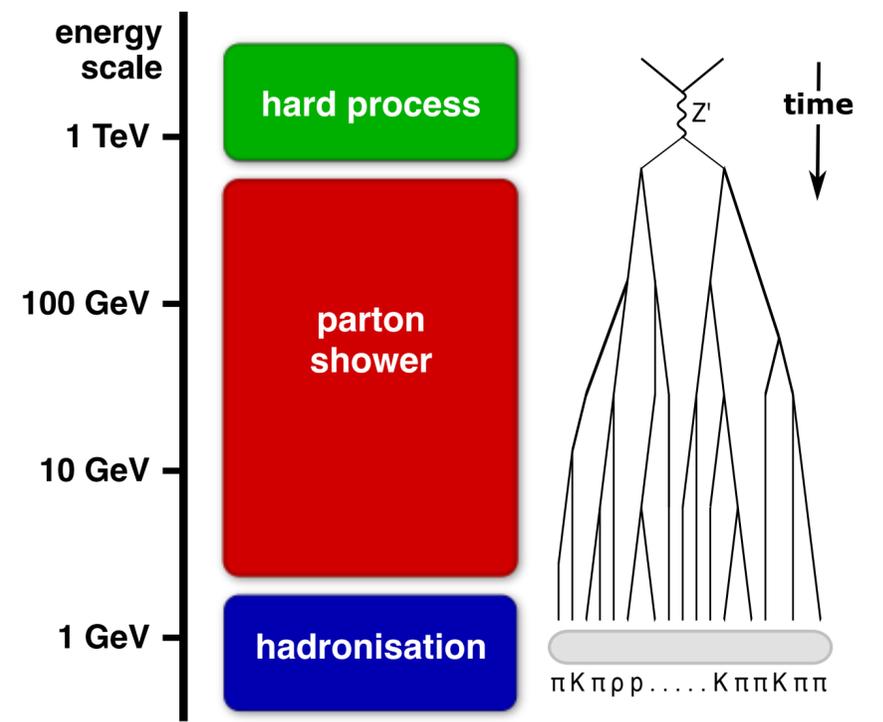
We collide p+ with p+ at a CM energy of 1.400e+04 GeV

Subprocess	Code	Estimated max (mb)
g g -> H (SM)	902	3.065e-07

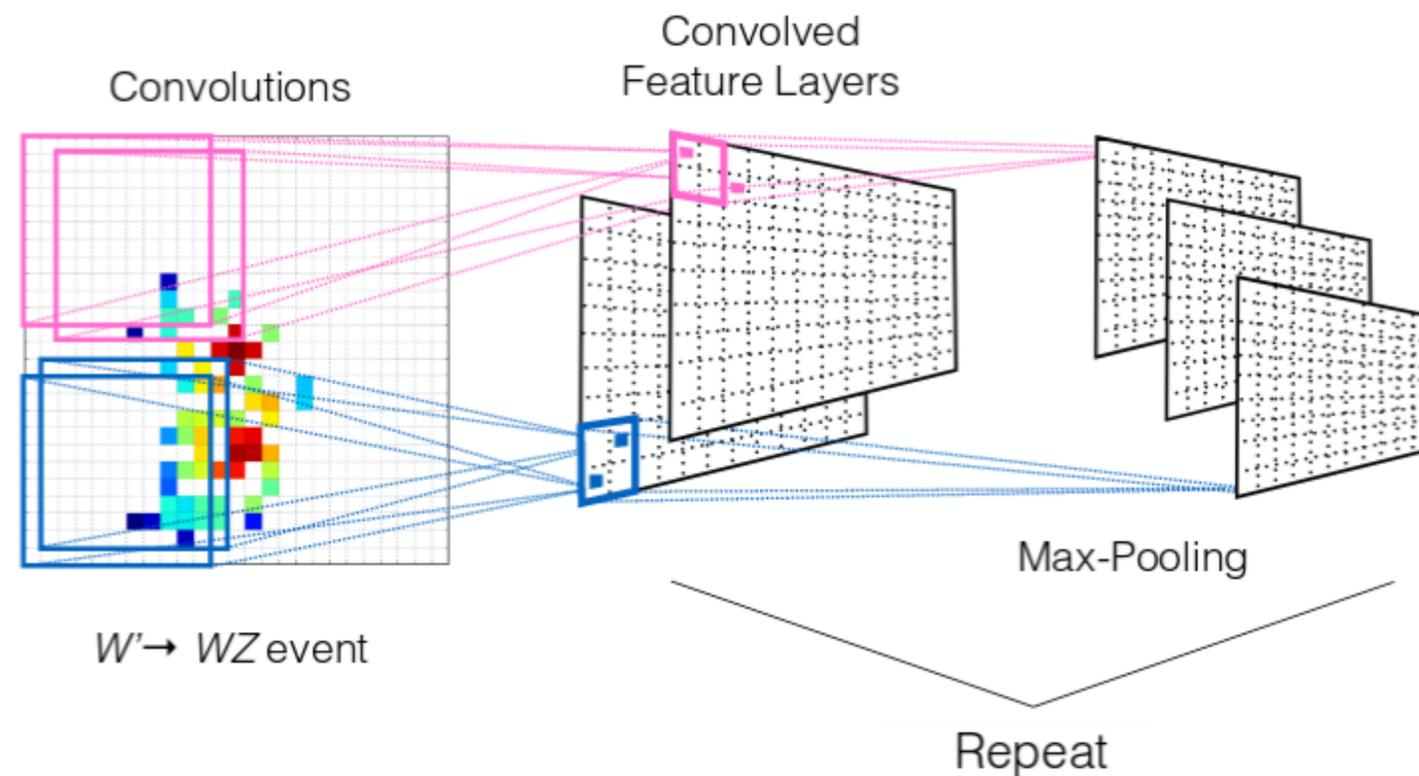
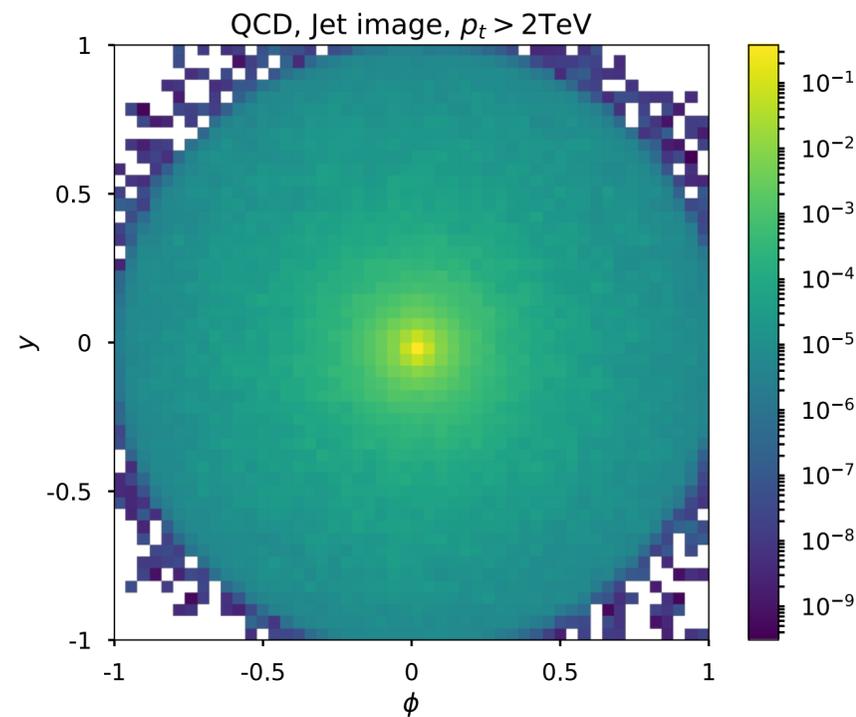
----- PYTHIA Event Listing (complete event) -----

no	id	name	status	mothers	daughters	colours	p_x	p_y	p_z	e	m		
0	90	(system)	-11	0	0	0	0.000	0.000	0.000	14000.000	14000.000		
1	2212	(p+)	-12	0	649	0	0.000	0.000	7000.000	7000.000	0.938		
2	2212	(p+)	-12	0	650	0	0.000	0.000	-7000.000	7000.000	0.938		
3	21	(g)	-21	19	5	0	101	102	10.638	10.638	0.000		
4	21	(g)	-21	20	5	0	102	101	-373.110	373.110	0.000		
5	25	(h0)	-22	3	4	21	21	0	-362.472	383.747	126.000		
6	21	(g)	-31	75	8	9	104	105	162.462	162.462	0.000		
7	21	(g)	-31	76	8	9	106	104	-8.450	8.450	0.000		
8	21	(g)	-33	6	7	42	43	106	107	11.466	0.000		
9	21	(g)	-33	6	7	44	44	107	105	159.447	0.000		
10	21	(g)	-31	14	12	13	108	109	14.037	14.037	0.000		
.....													
1624	111	pi0	91	1516	0	0	0	0	0.081	0.097	-0.757	0.779	0.135
1625	111	pi0	91	1516	0	0	0	0	-0.082	-0.156	-0.614	0.653	0.135
1626	130	K_L0	91	1522	1522	0	0	0	-2.188	0.152	13.925	14.106	0.498
			Charge sum:	2.000	Momentum sum:			-0.000	0.000	-0.000	14000.000	14000.000	

----- End PYTHIA Event Listing -----



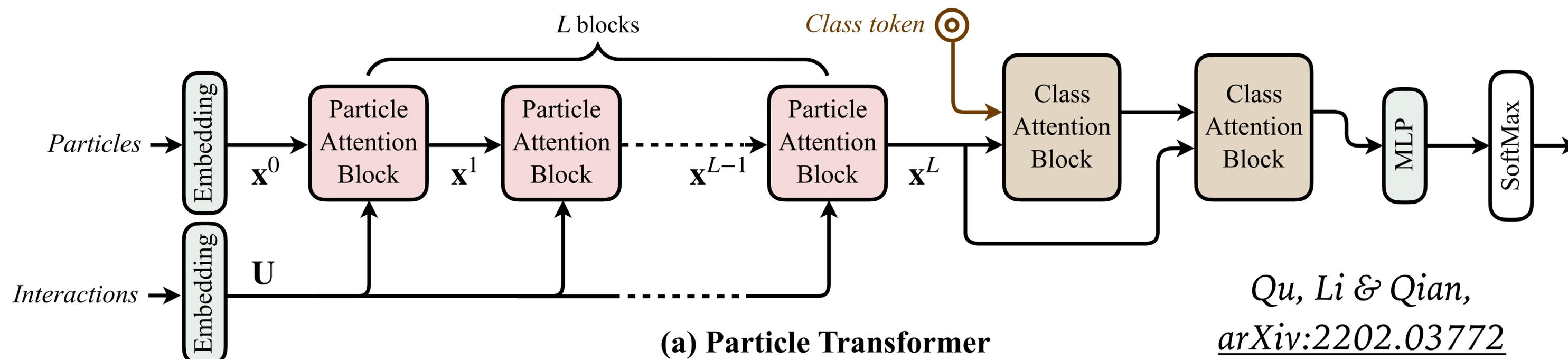
Machine learning and jet/event structure



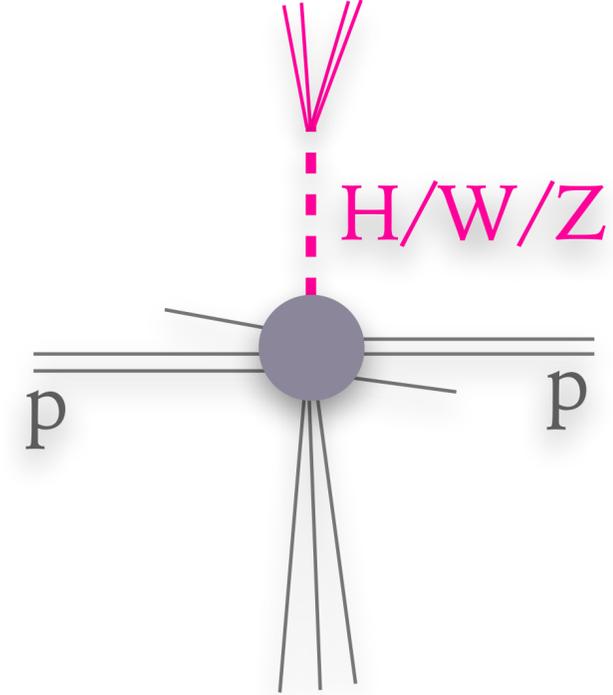
[Cogan, Kagan, Strauss, Schwartzman [JHEP 1502 \(2015\) 118](#)]

[de Oliveira, Kagan, Mackey, Nachman, Schwartzman [JHEP 1607 \(2016\) 069](#)]

2021 Young Experimental Physicist
EPS HEPP prize

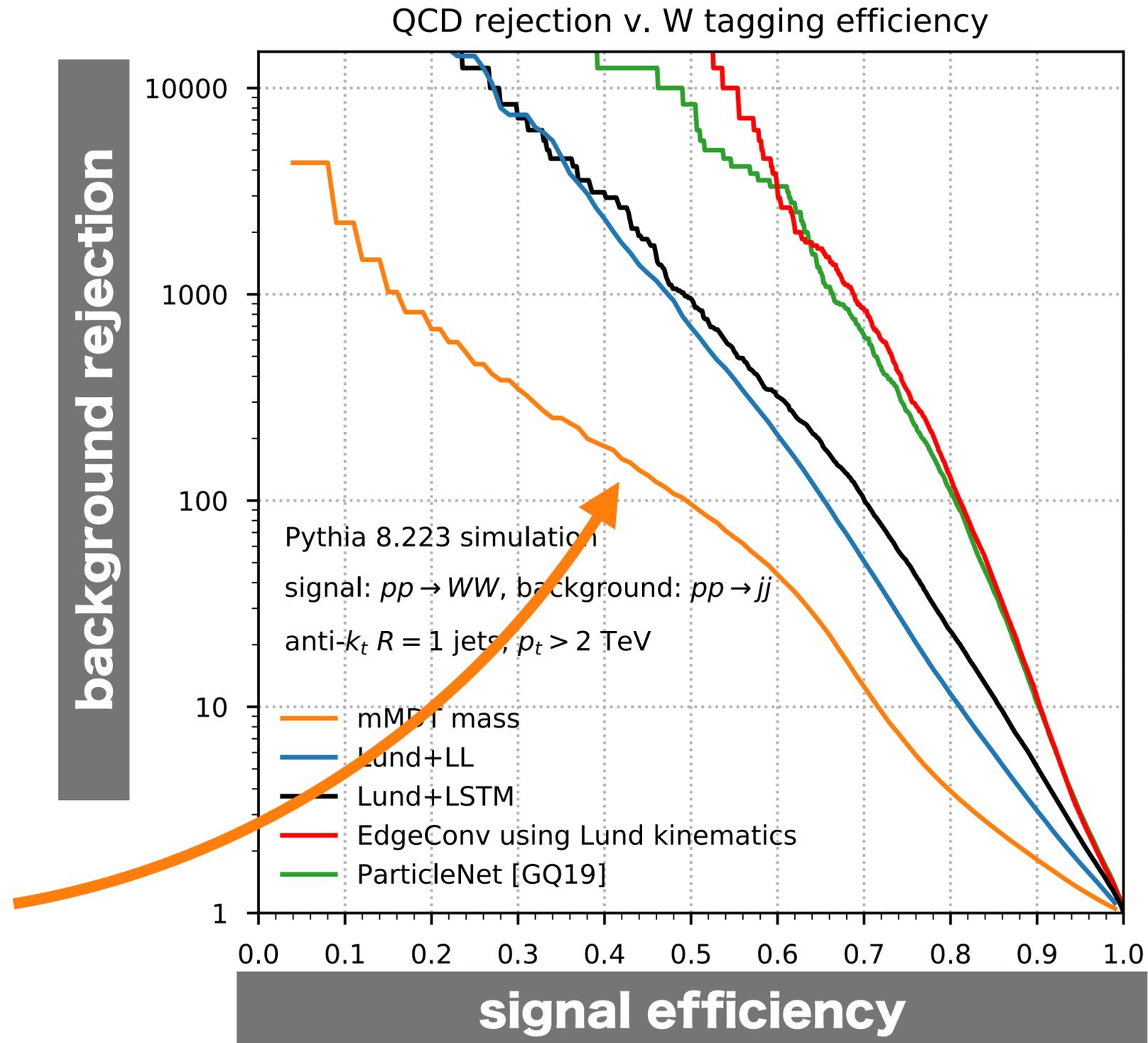


using full jet/event information for H/W/Z-boson tagging

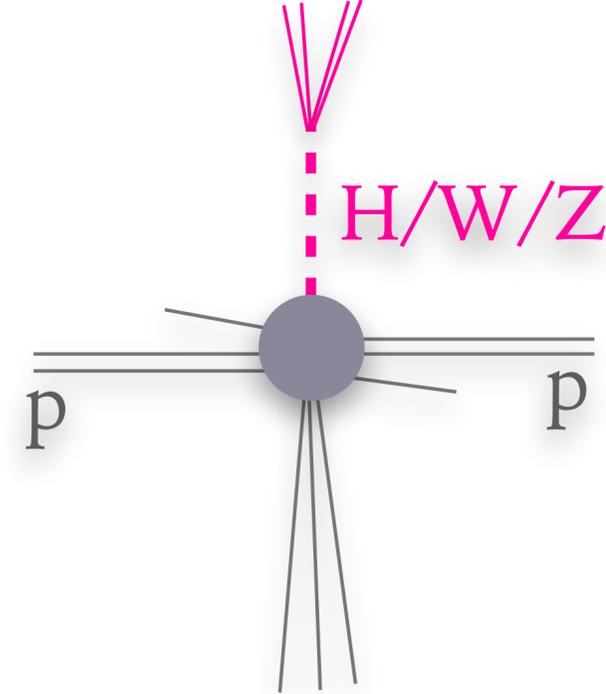


adapted from
Dreyer & Qu
2012.08526

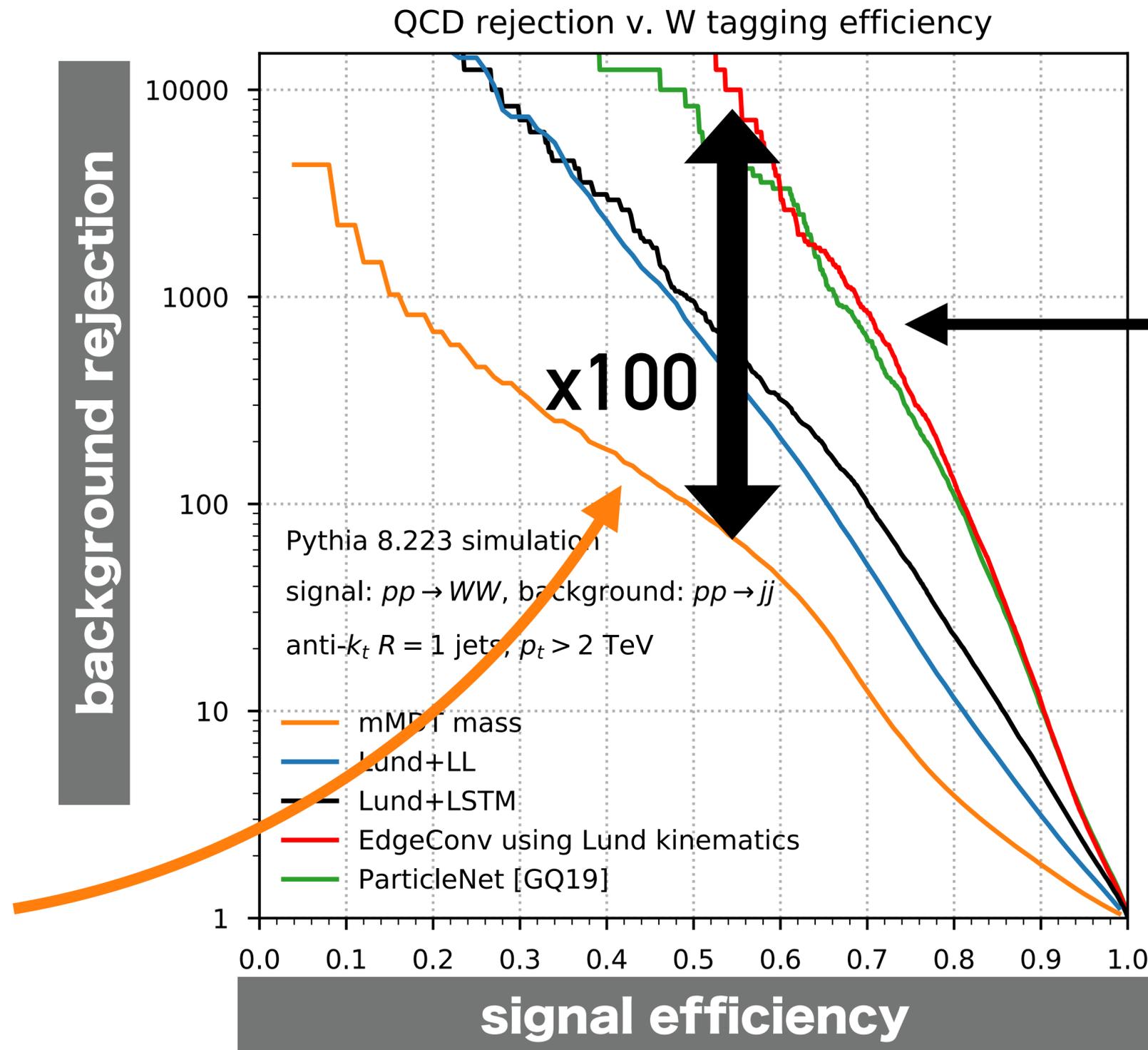
QCD rejection with
just jet mass
(SD/mMDT)
i.e. 2008 tools &
their 2013/14
descendants



using full jet/event information for H/W/Z-boson tagging



adapted from
Dreyer & Qu
2012.08526



QCD rejection
with use of full jet
substructure
(2021 tools)
100x better

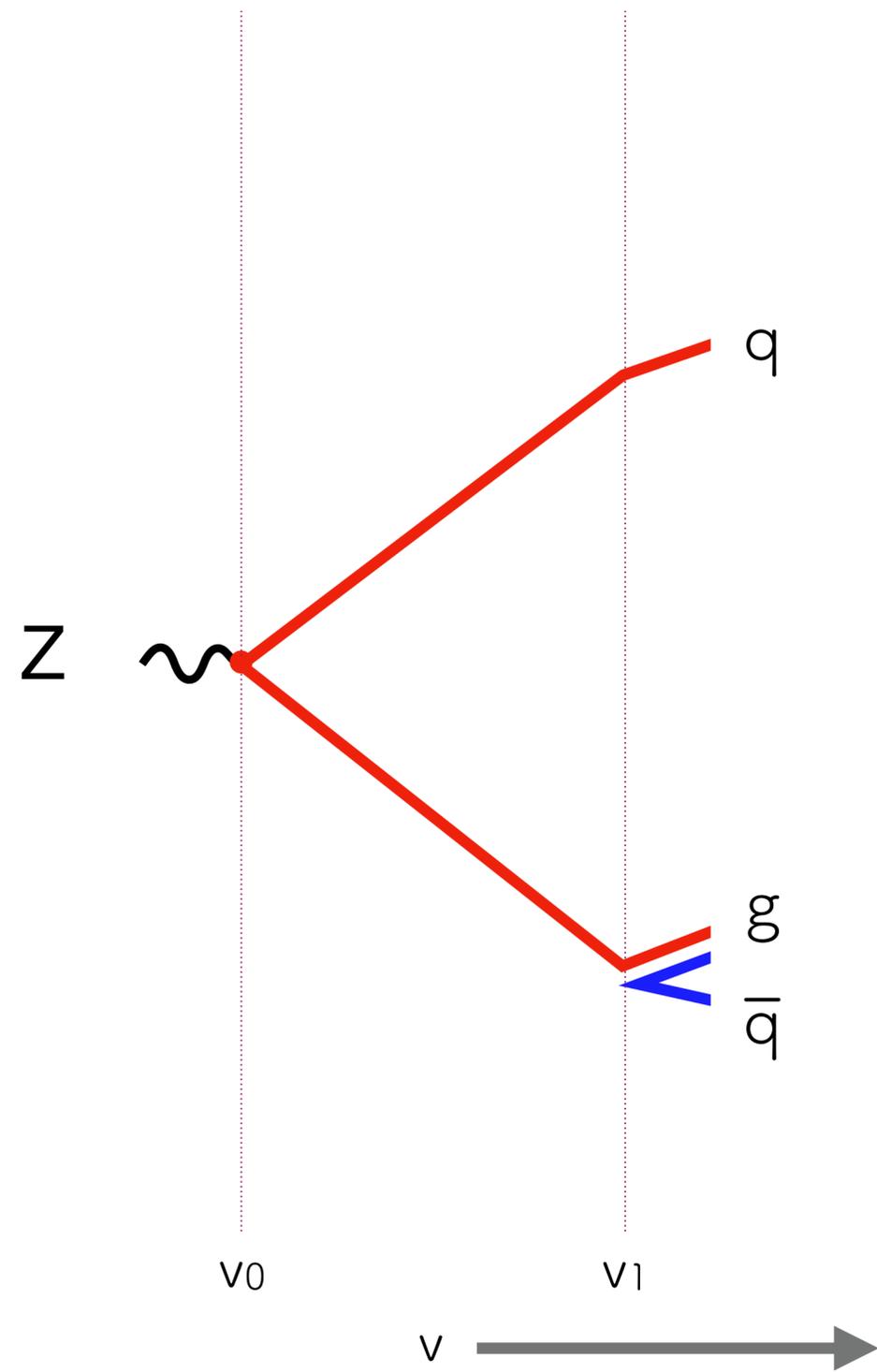
QCD rejection with
just jet mass
(SD/mMDT)
i.e. 2008 tools &
their 2013/14
descendants

First started to be exploited
by Thaler & Van Tilburg with
“N-subjettiness” (2010/11)

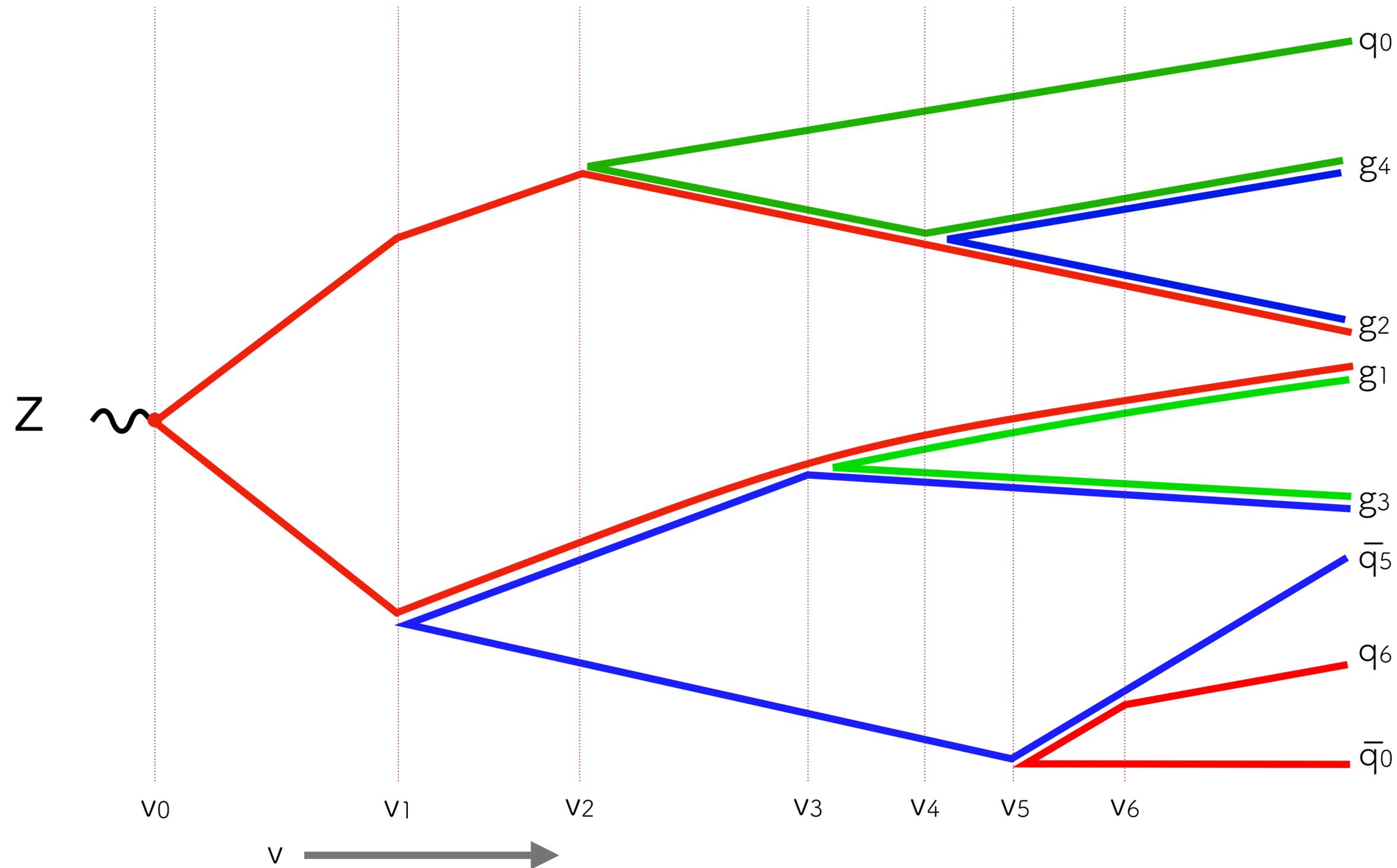
Element #1: what are parton showers trying to achieve?

parton showers span disparate scales
natural language is “logarithmic” accuracy

QCD **parton shower**: an evolution equation (in **evolution scale v** , e.g. transverse momentum)



QCD **parton shower**: an evolution equation (in **evolution scale v** , e.g. transverse momentum)



self-similar
evolution
continues until it
reaches a non-
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scale

**branchings
widely
separated in
space-time
treated as
~independent**

QCD **parton shower**: an evolution equation (in **evolution scale v** , e.g. transverse momentum)

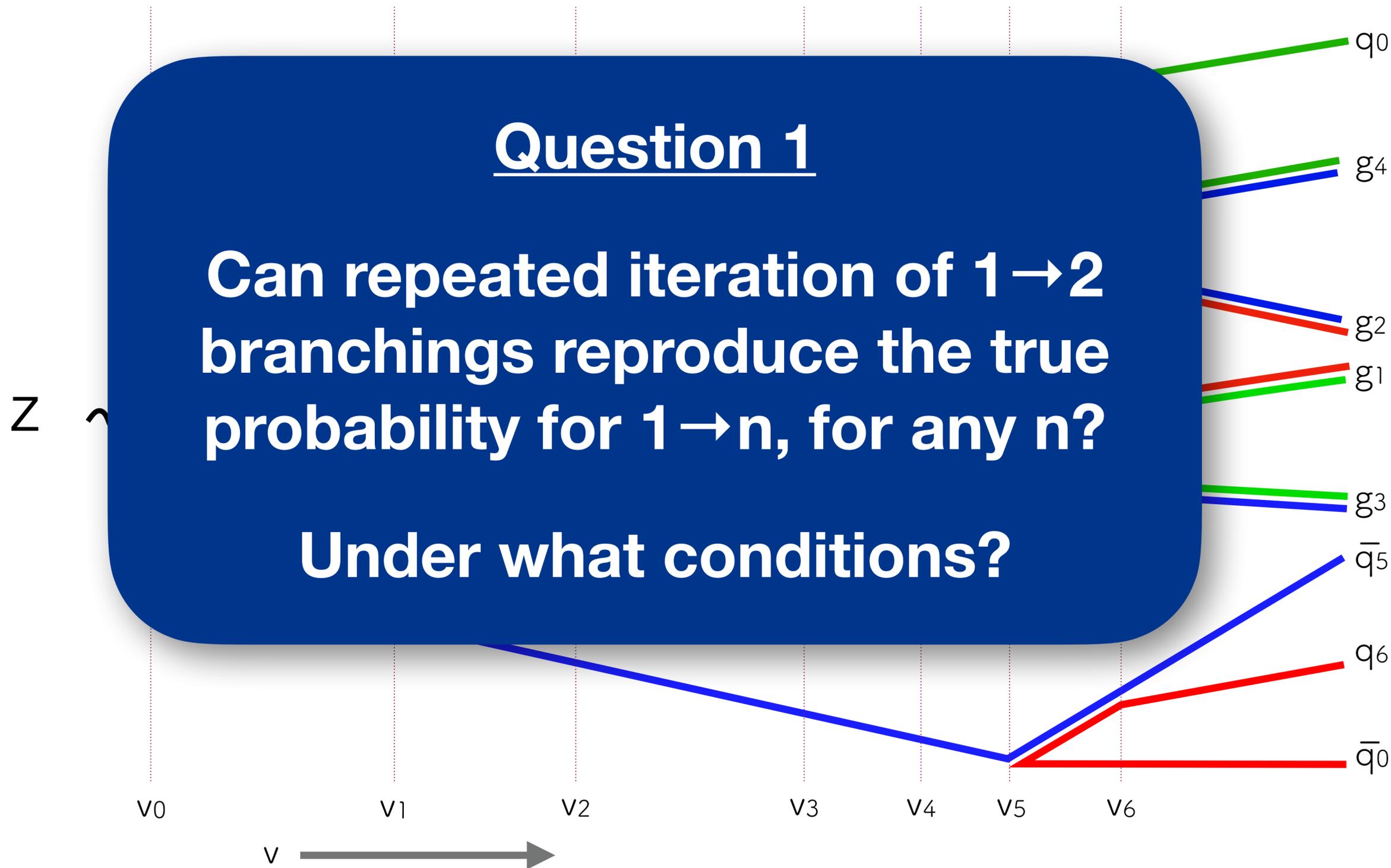
Question 1

Can repeated iteration of $1 \rightarrow 2$ branchings reproduce the true probability for $1 \rightarrow n$, for any n ?

Under what conditions?

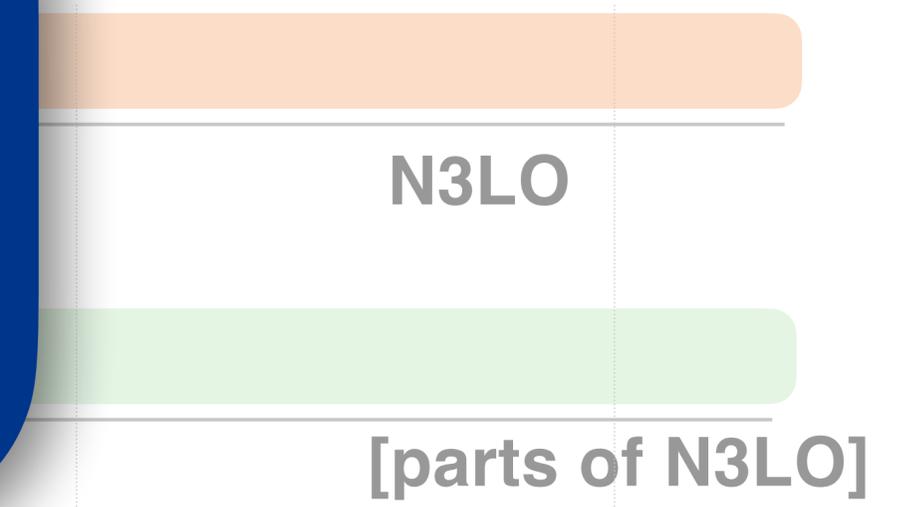
self-similar evolution continues until it reaches a non-perturbative scale

branchings widely separated in space-time treated as \sim independent



Question 2

Can a single parton shower reproduce all known resummations? [perturbation theory across disparate scales]



transverse-momentum resummation (DY&Higgs)

LL NLL[.....] NNLL[...] N3LL

parton showers (many of today's widely-used showers only LL@leading-colour) parts of NNLL

LL [parts of NLL.....] NLL

fixed-order matching of parton showers

LO NLO NNLO [.....] [N3LO]

1970

1980

1990

2000

2010

2020



Melissa van Beekveld
NIKEHF



Mrinal Dasgupta
Manchester



Basem El-Menoufi
Monash



Silvia Ferrario Ravasio
CERN



Keith Hamilton
Univ. Coll. London



Jack Helliwell
Monash



Alexander Karlberg
CERN



Pier Monni
CERN



GPS
Oxford



Nicolas Schalch
Oxford



Ludovic Scyboz
Monash



Alba Soto-Ontoso
Granada



Grégory Soyez
IPhT, Saclay



Silvia Zanoli
Oxford

PanScales

A project to bring logarithmic understanding and accuracy to parton showers



ERC funded
2018-2024



Frédéric Drever



Rob Verheyen



Rok Medves

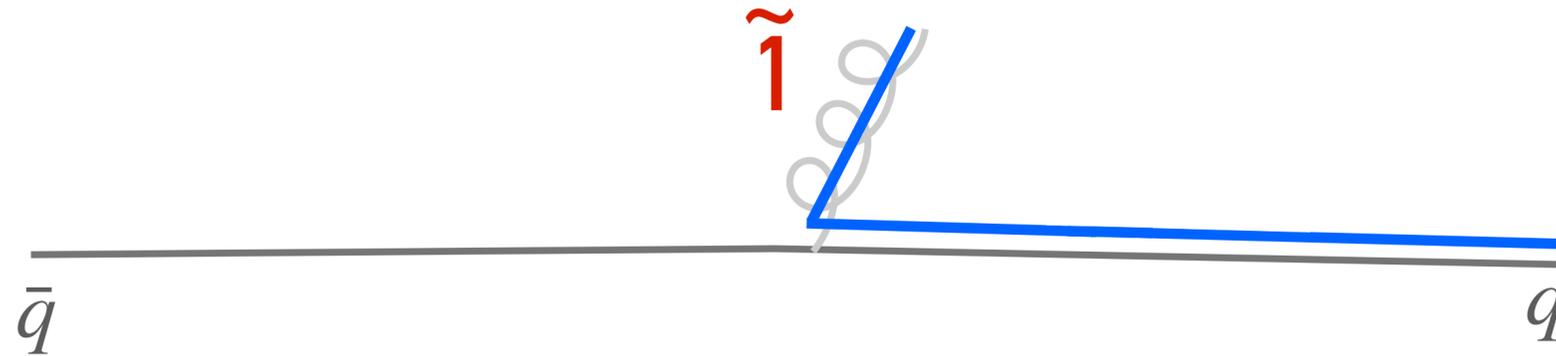


Emma Slade

former members

1. Momentum conservation: the core of any shower

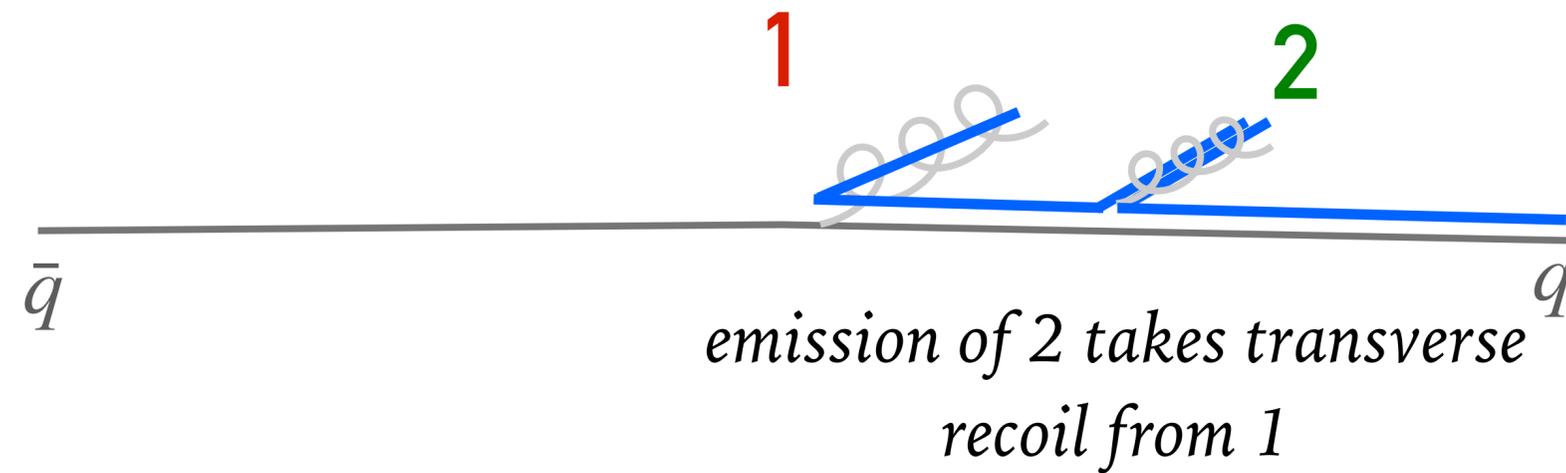
Dipole showers conserve momentum at each step. Traditional dipole-local recoil:



$$d\mathcal{P}_{\tilde{i} \rightarrow ik}^{\text{FS}} = \frac{\alpha_s(k_{\perp}^2)}{2\pi} \frac{dk_{\perp}^2}{k_{\perp}^2} \frac{dz}{z} \frac{d\varphi}{2\pi} N_{ik}^{\text{sym}} [z P_{\tilde{i} \rightarrow ik}(z)]$$

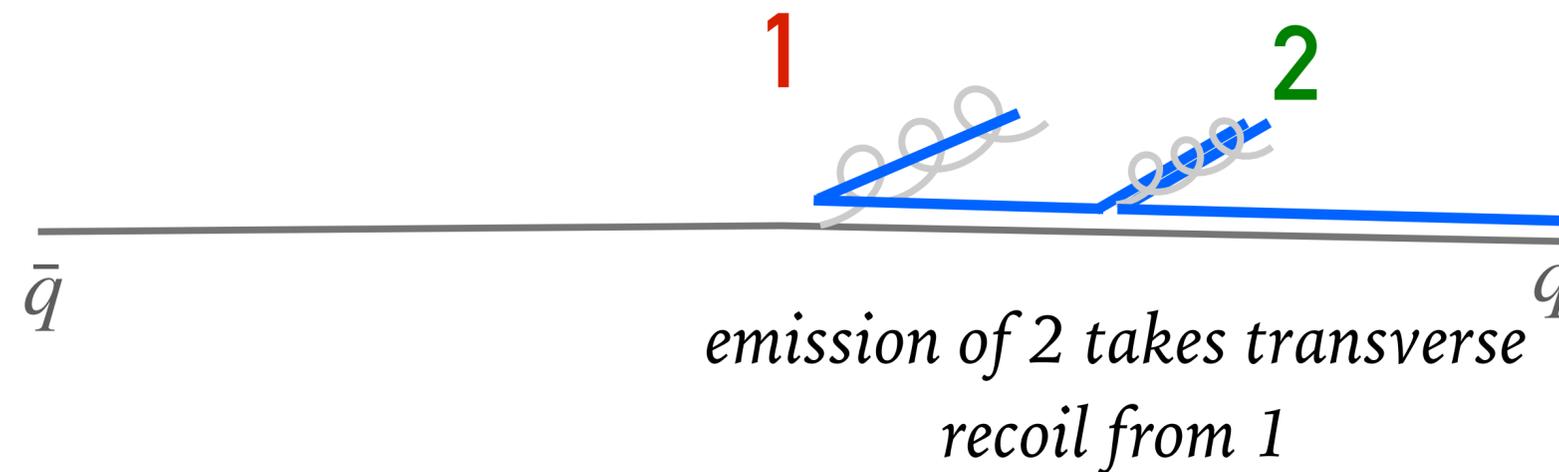
1. Momentum conservation: the core of any shower

Dipole showers conserve momentum at each step. Traditional dipole-local recoil:



1. Momentum conservation: the core of any shower

Dipole showers conserve momentum at each step. Traditional dipole-local recoil:



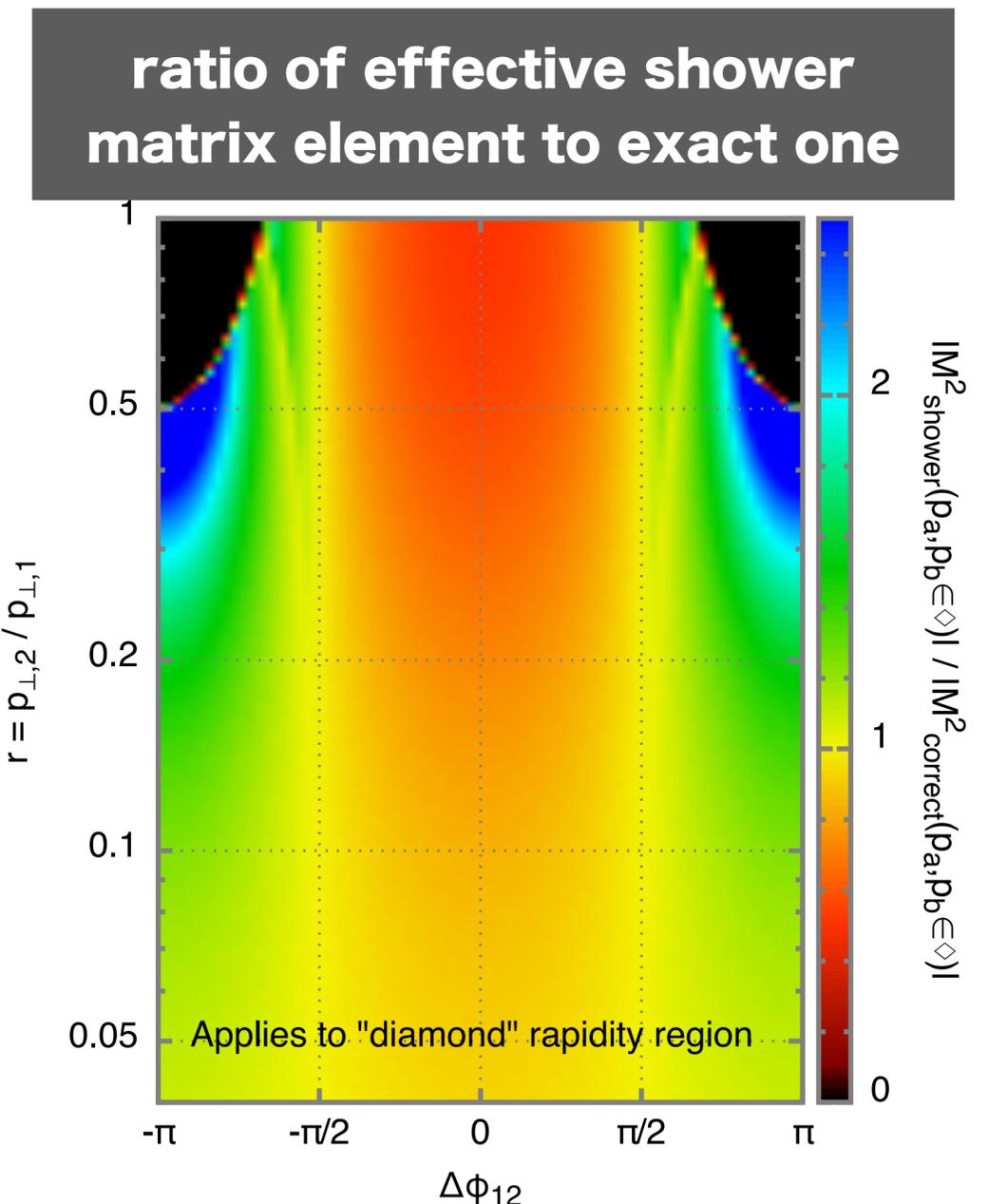
Shower initially generated matrix element for particle $\tilde{1}$, whose momentum differs (by $\sim 50\%$) from final particle 1.

Matrix element is incorrect wrt final momentum 1.

First observed: Andersson, Gustafson, Sjogren '92

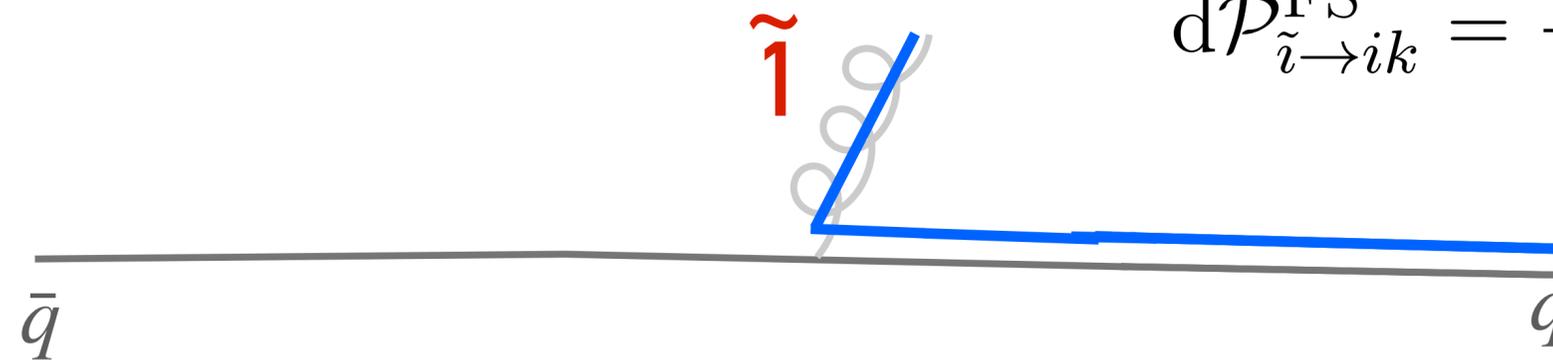
Closely related effect present for Z p_t : Nagy & Soper [0912.4534](#)

Impact on log accuracy across many observables: Dasgupta, Dreyer, Hamilton, Monni, GPS, [1805.09327](#)



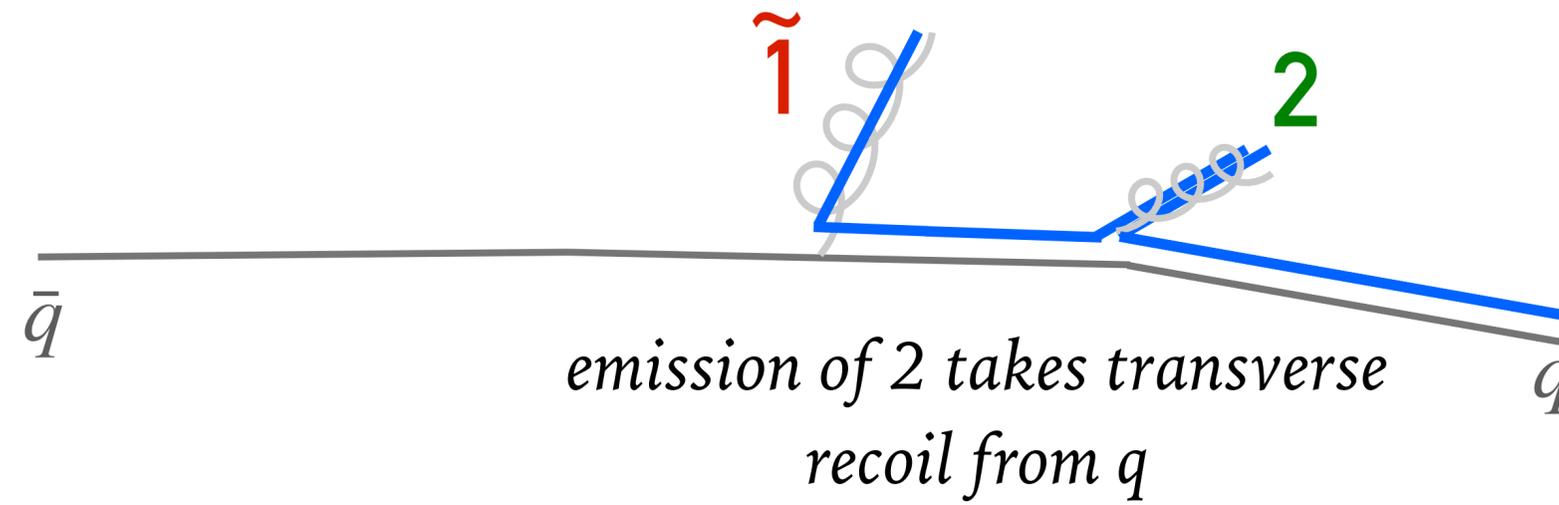
1. Correct recoil rule: **no side effects on other distant emissions**

One approach


$$d\mathcal{P}_{\tilde{i} \rightarrow ik}^{\text{FS}} = \frac{\alpha_s(k_{\perp}^2)}{2\pi} \frac{dk_{\perp}^2}{k_{\perp}^2} \frac{dz}{z} \frac{d\varphi}{2\pi} N_{ik}^{\text{sym}} [z P_{\tilde{i} \rightarrow ik}(z)]$$

1. Correct recoil rule: **no side effects on other distant emissions**

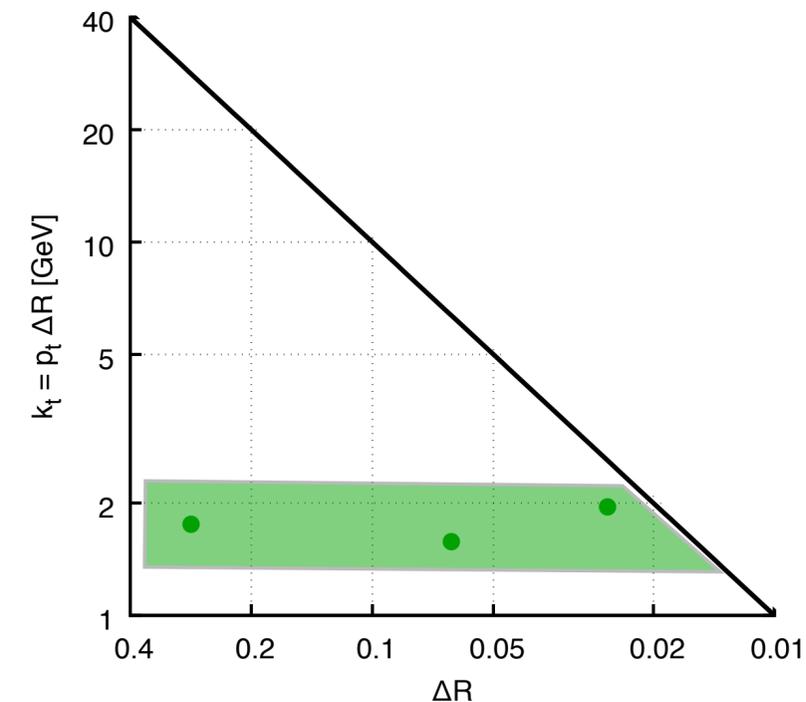
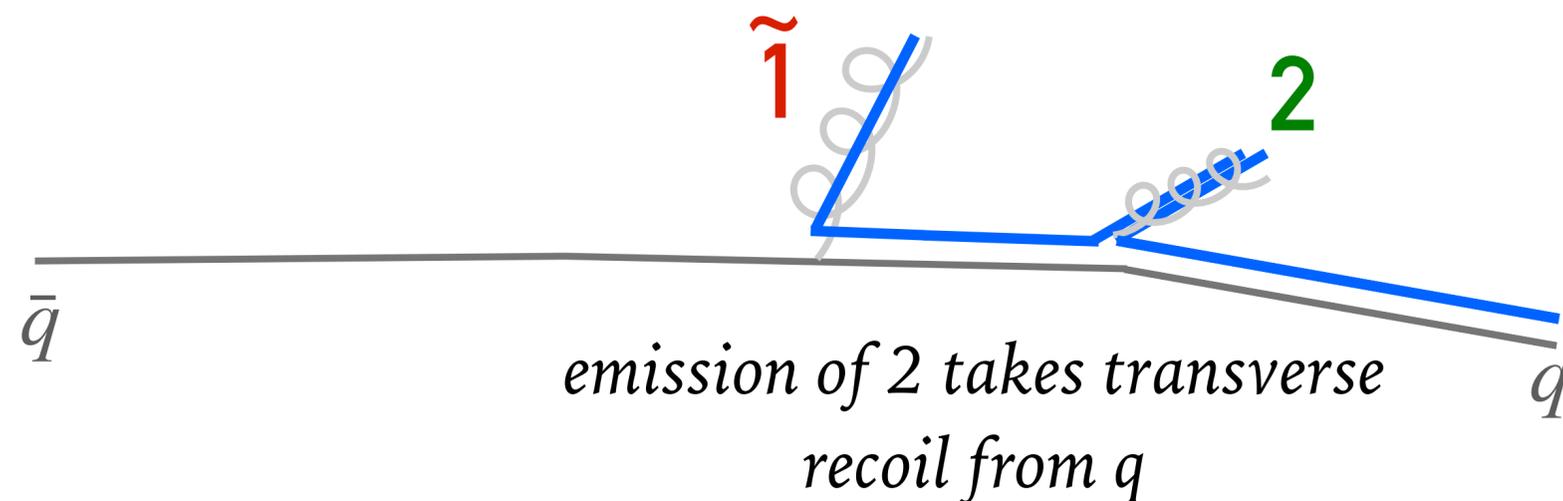
One approach



θ_{1q} left almost unchanged if \perp recoil from emission of 2 taken by (much harder) q

1. Correct recoil rule: **no side effects on other distant emissions**

One approach



θ_{1q} left almost unchanged if \perp recoil from emission of 2 taken by (much harder) q

Can be achieved in multiple ways:

► global transverse recoil

(Dasgupta et al [2002.11114](#), “PanGlobal”; Holguin Seymour & Forshaw [2003.06400](#); Alaric [2208.06057](#) + ..., Apollo, [2403.19452](#))

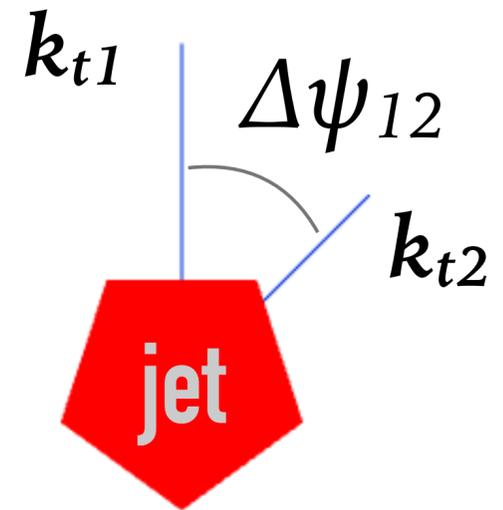
► local transverse recoil, with non-standard shower ordering & dipole partition

([2002.11114](#) “PanLocal”; Nagy & Soper [0912.4534](#) + ..., “Deductor”)

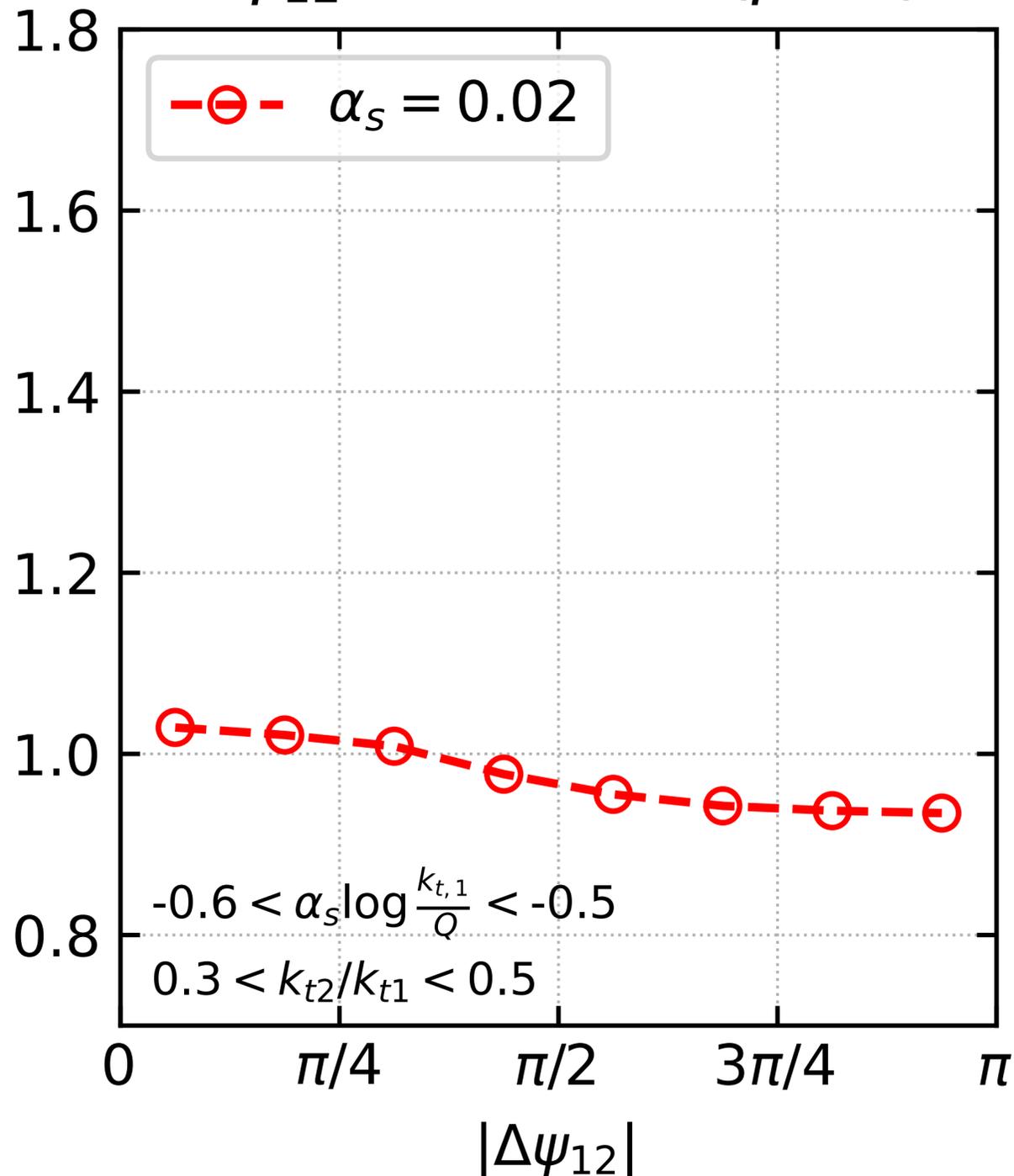
Element #2: testing correctness

Parton showers operate at all orders and mix many effects. How can you separate out just the orders you aim to control to test they're correct?

Test class 2: full shower v. all-order NLL



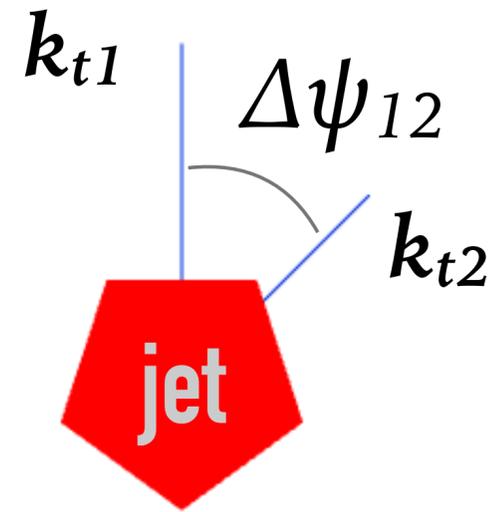
$\Delta\psi_{12}$, PanGlobal($\beta=0$)



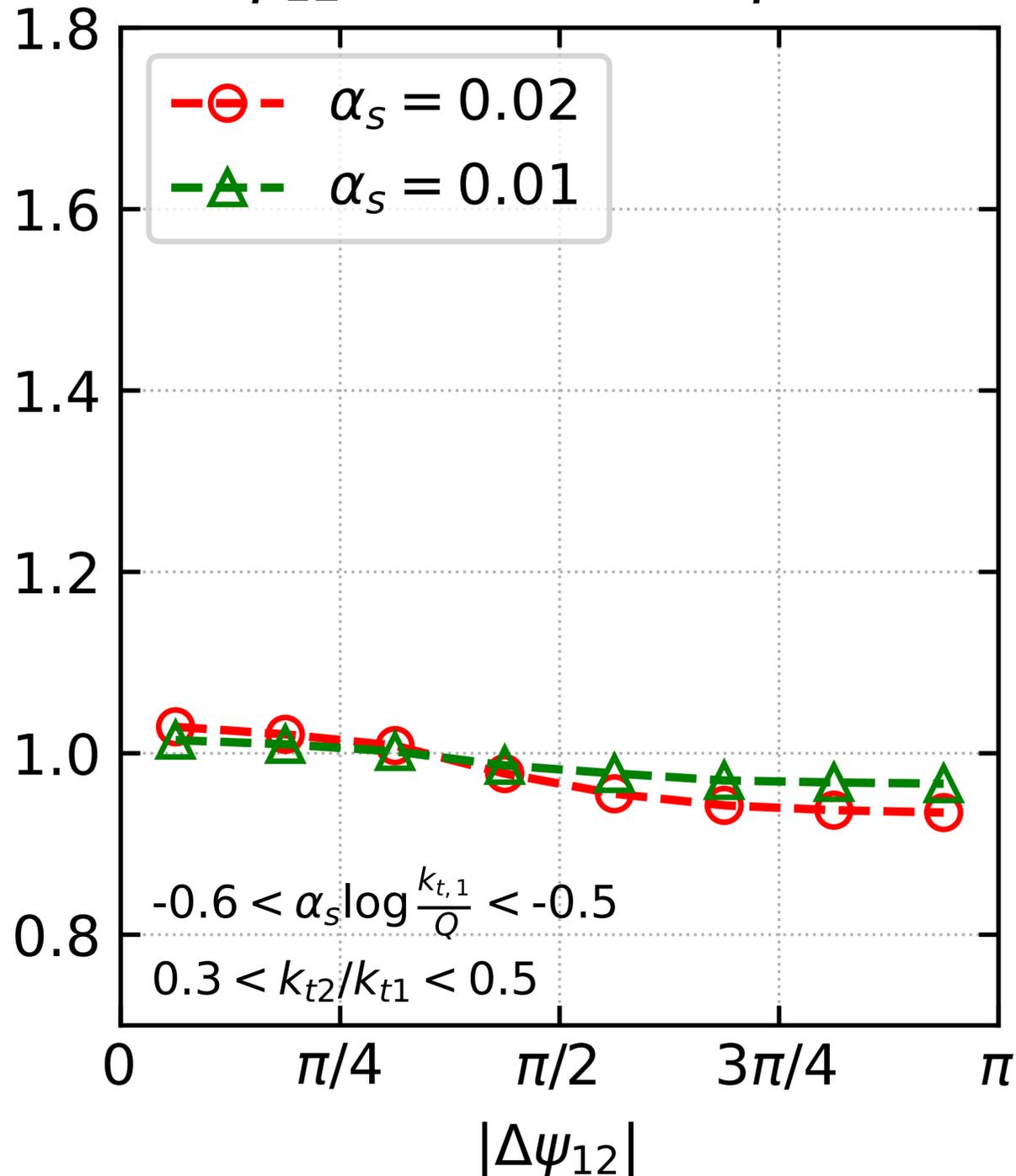
- ▶ run full shower & measure specific observable: azimuth between two highest- k_t emissions (soft-collinear)
- ▶ Normal QCD: $\alpha_s \simeq 0.1$ and two orders of magnitude in momentum
- ▶ Focus on “logarithmic” part by taking smaller $\alpha_s = 0.02$ and 10 orders of magnitude
- ▶ ratio to NLL should be flat $\equiv 1$
- ▶ it isn't: have we got an NLL mistake? Or a residual subleading (NNLL) term?

ratio
to
NLL

Test class 2: full shower v. all-order NLL



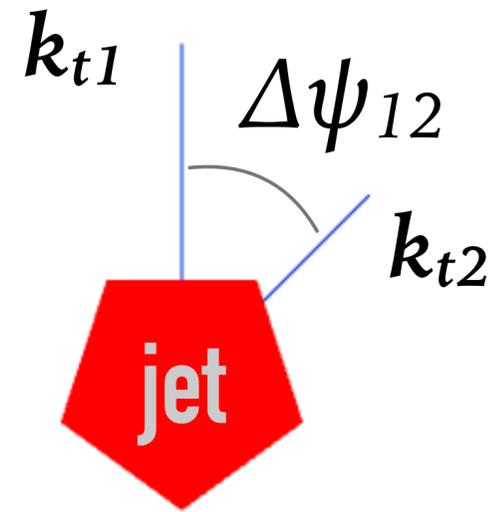
$\Delta\psi_{12}$, PanGlobal($\beta=0$)



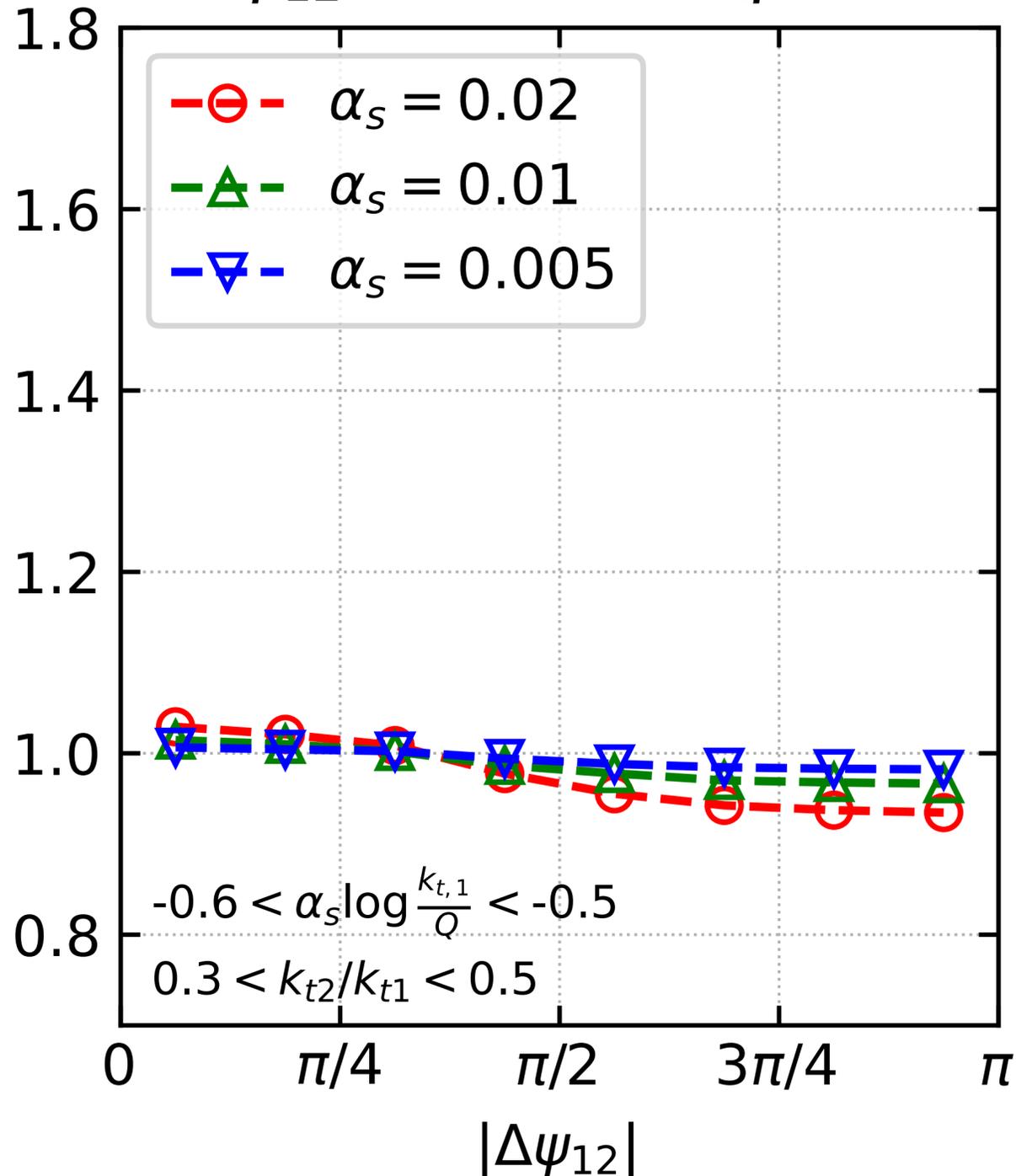
ratio
to
NLL

- ▶ run full shower & measure specific observable: azimuth between two highest- k_t emissions (soft-collinear)
- ▶ Normal QCD: $\alpha_s \simeq 0.1$ and two orders of magnitude in momentum
- ▶ Focus on “logarithmic” part by taking smaller $\alpha_s = 0.01$ and 20 orders of magnitude
- ▶ ratio to NLL should be flat $\equiv 1$
- ▶ it isn't: have we got an NLL mistake? Or a residual subleading (NNLL) term?

Test class 2: full shower v. all-order NLL



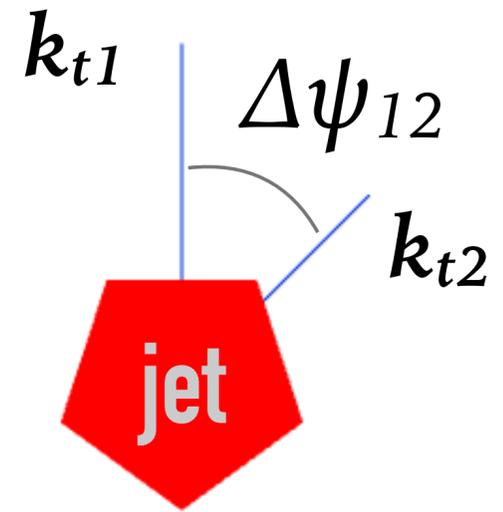
$\Delta\psi_{12}$, PanGlobal($\beta=0$)



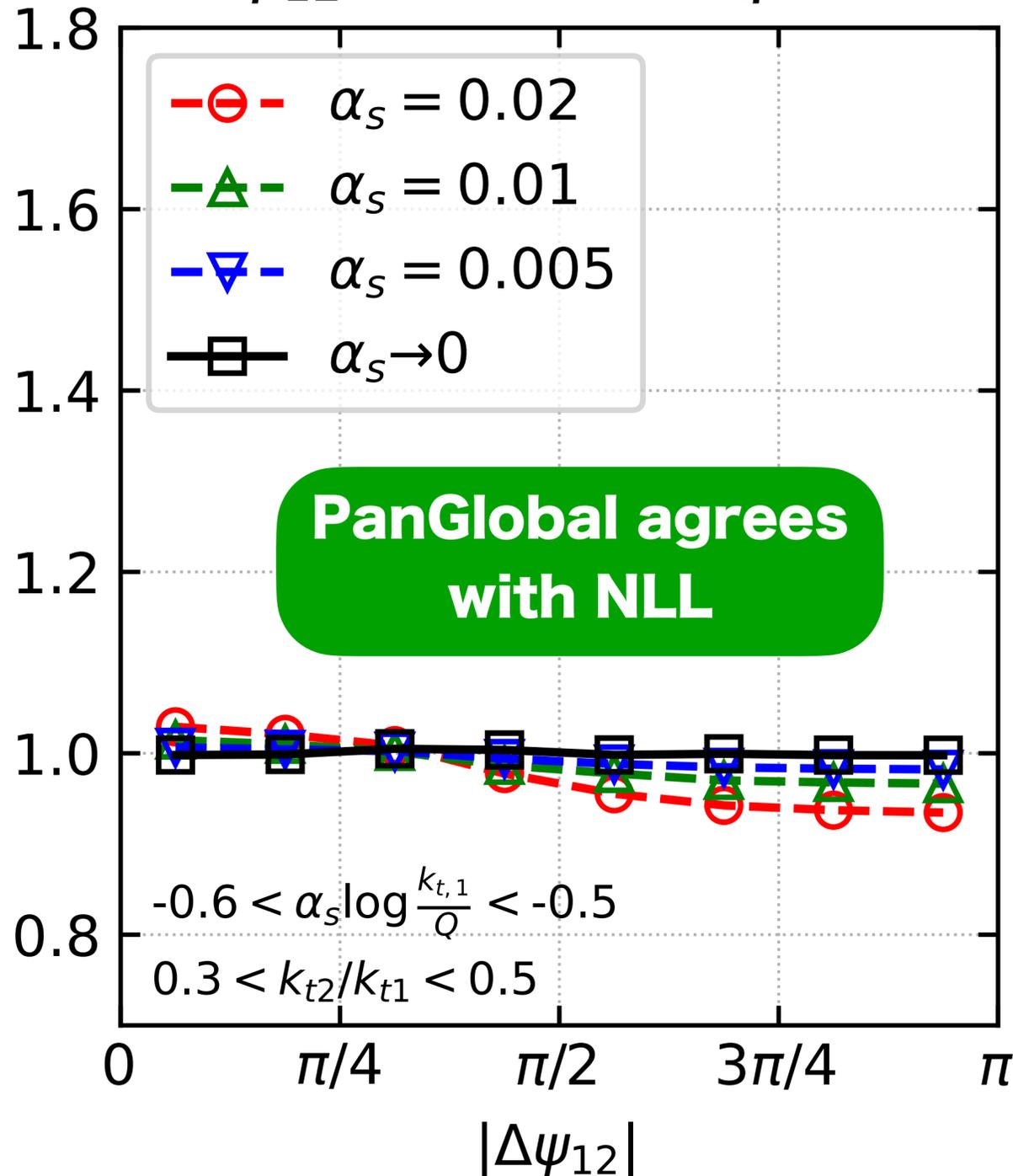
ratio
to
NLL

- ▶ run full shower & measure specific observable: azimuth between two highest- k_t emissions (soft-collinear)
- ▶ Normal QCD: $\alpha_s \simeq 0.1$ and two orders of magnitude in momentum
- ▶ Focus on “logarithmic” part by taking smaller $\alpha_s = 0.005$ and 40 orders of magnitude
- ▶ ratio to NLL should be flat $\equiv 1$
- ▶ it isn't: have we got an NLL mistake? Or a residual subleading (NNLL) term?

Test class 2: full shower v. all-order NLL



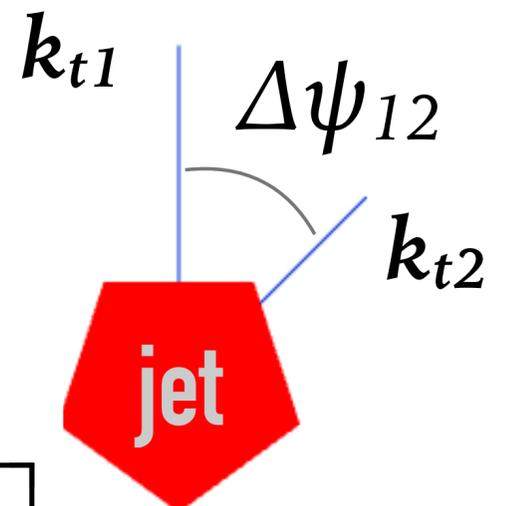
$\Delta\psi_{12}$, PanGlobal($\beta=0$)



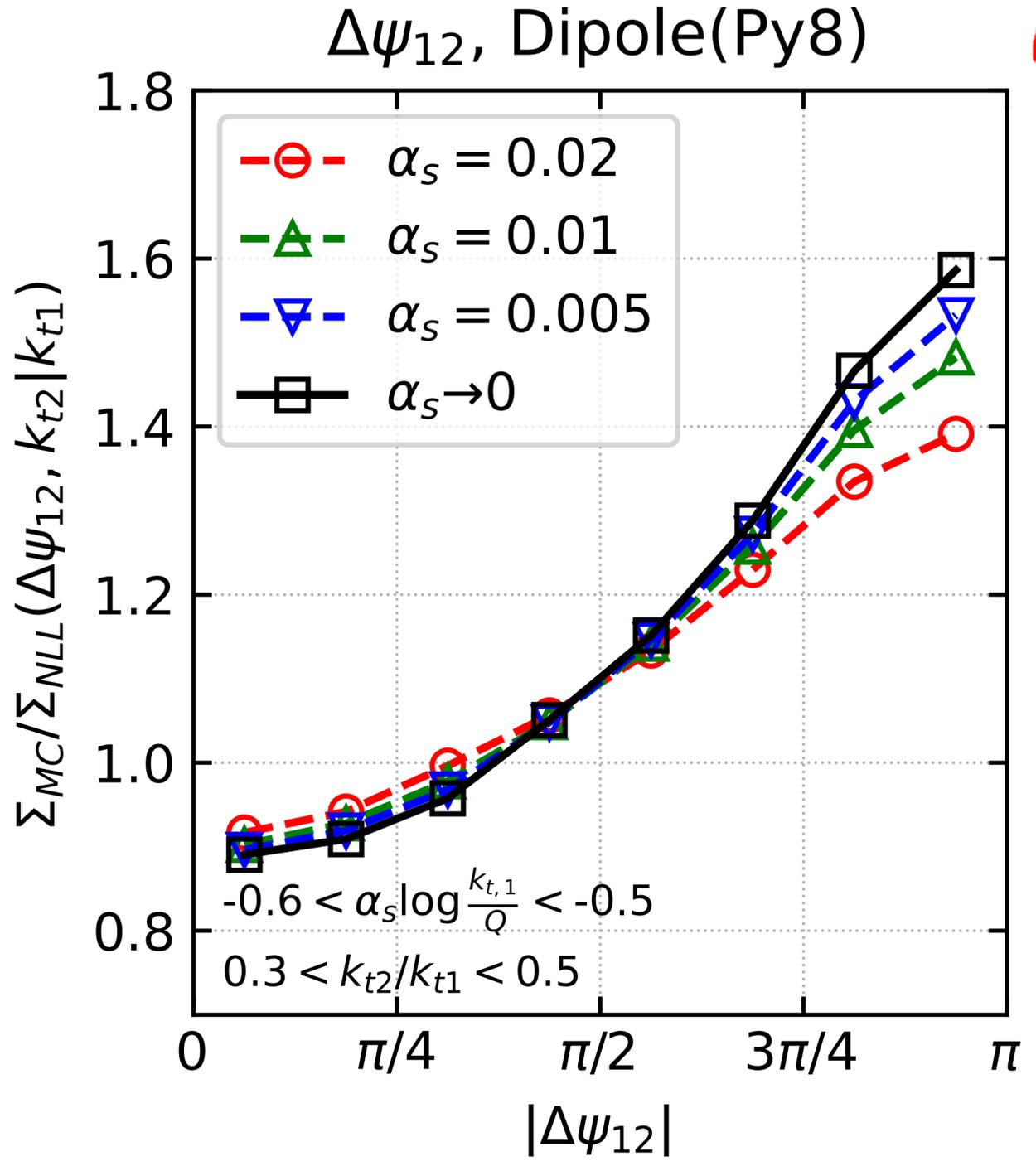
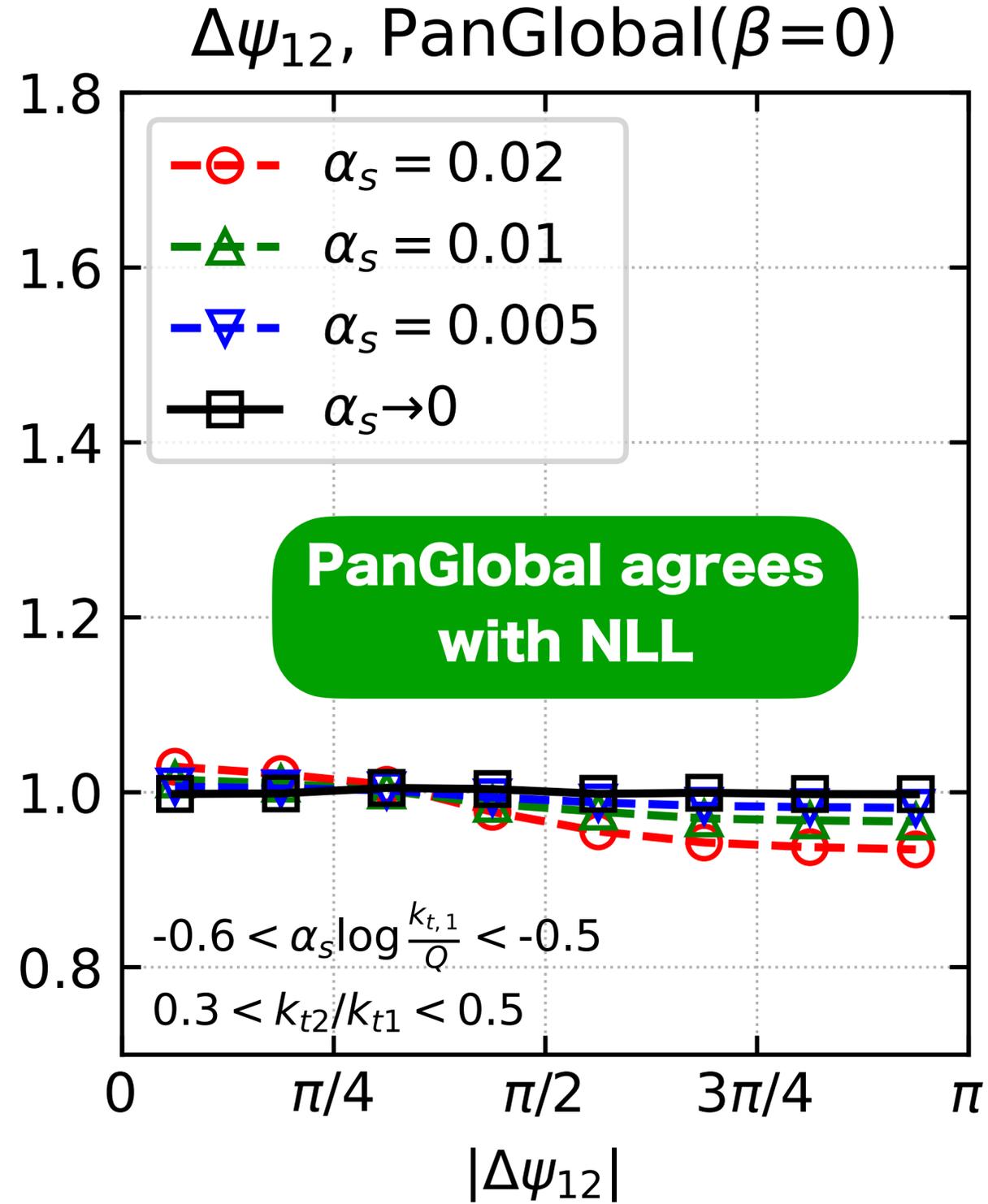
- run full shower & measure specific observable: azimuth between two highest- k_t emissions (soft-collinear)
- Normal QCD: $\alpha_s \simeq 0.1$ and two orders of magnitude in momentum
- Focus on “logarithmic” part by taking smaller $\alpha_s = 0.005$ and 40 orders of magnitude
- ratio to NLL should be flat $\equiv 1$
- extrapolation $\alpha_s \rightarrow 0$ agrees with NLL

ratio
to
NLL

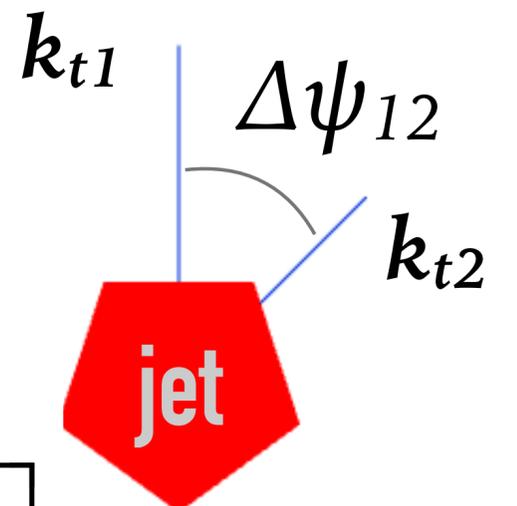
Test class 2: full shower v. all-order NLL



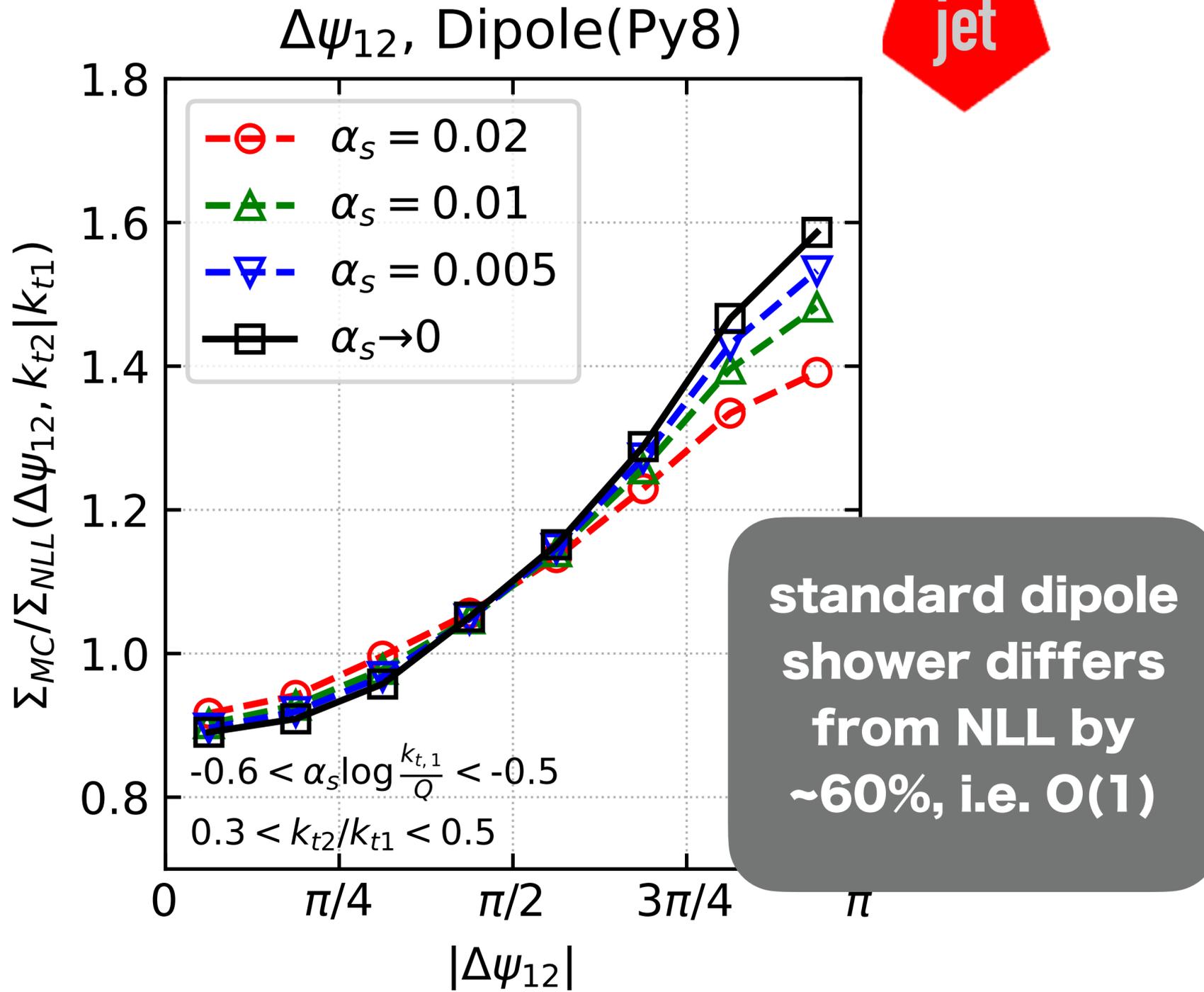
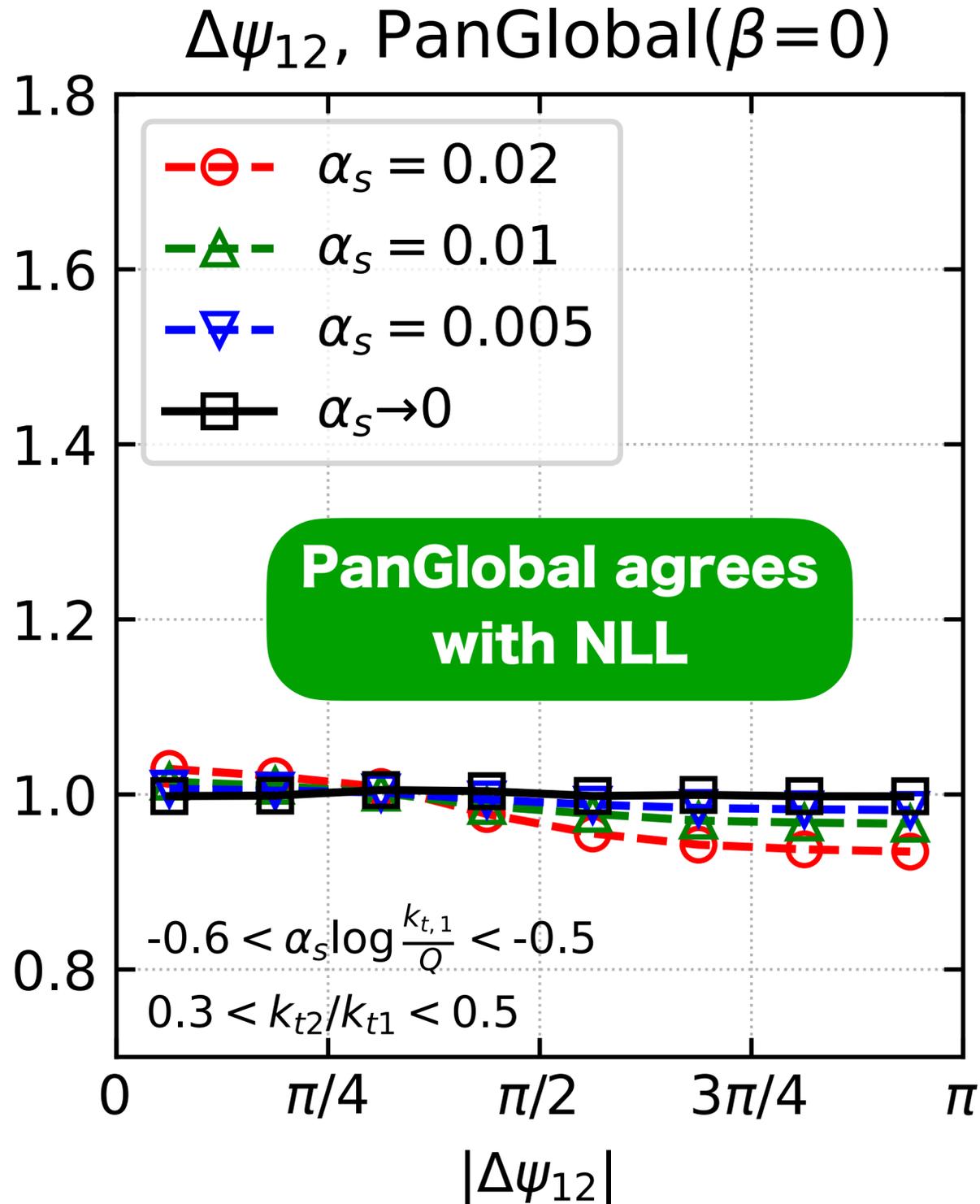
ratio to NLL



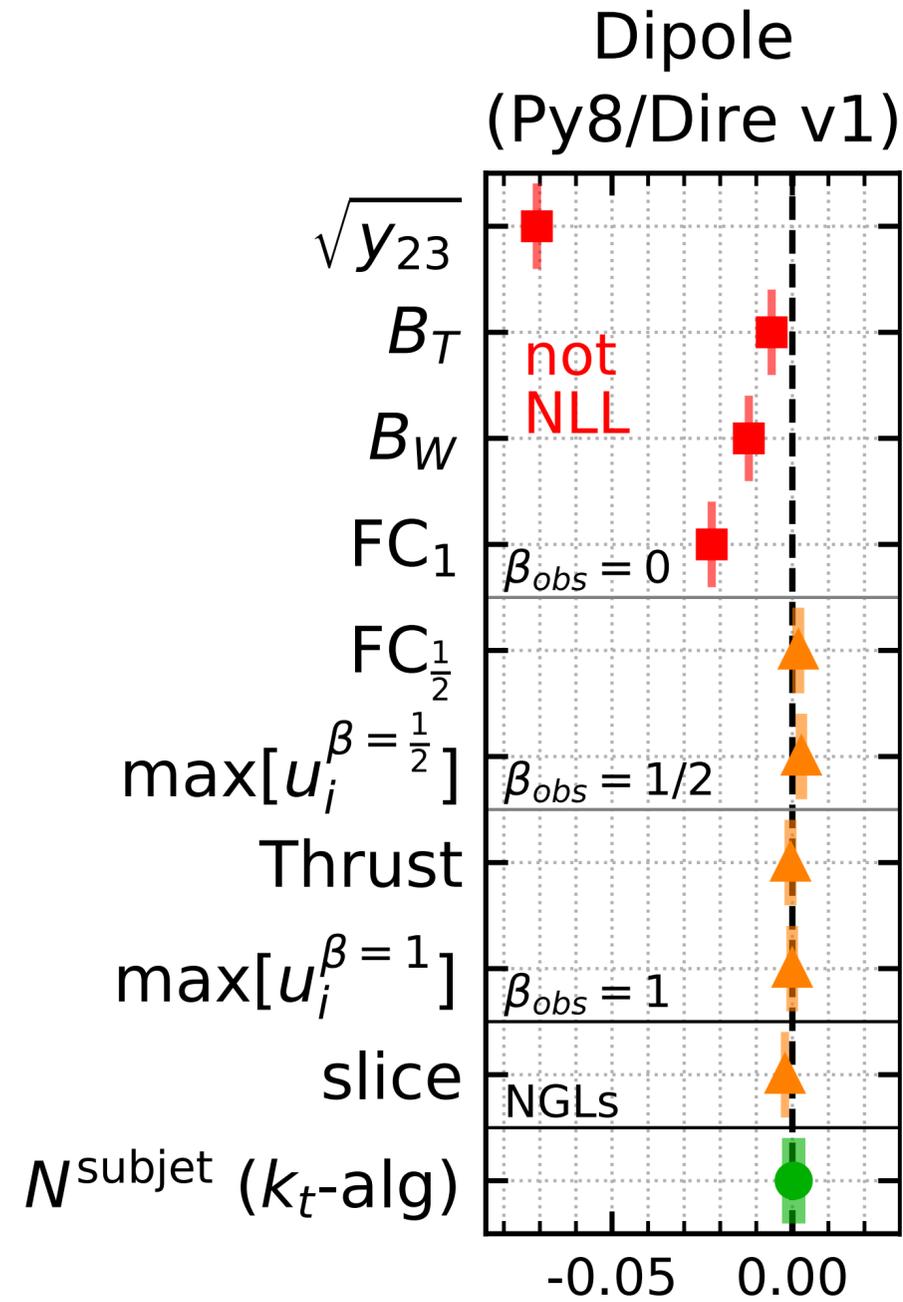
Test class 2: full shower v. all-order NLL



ratio to NLL

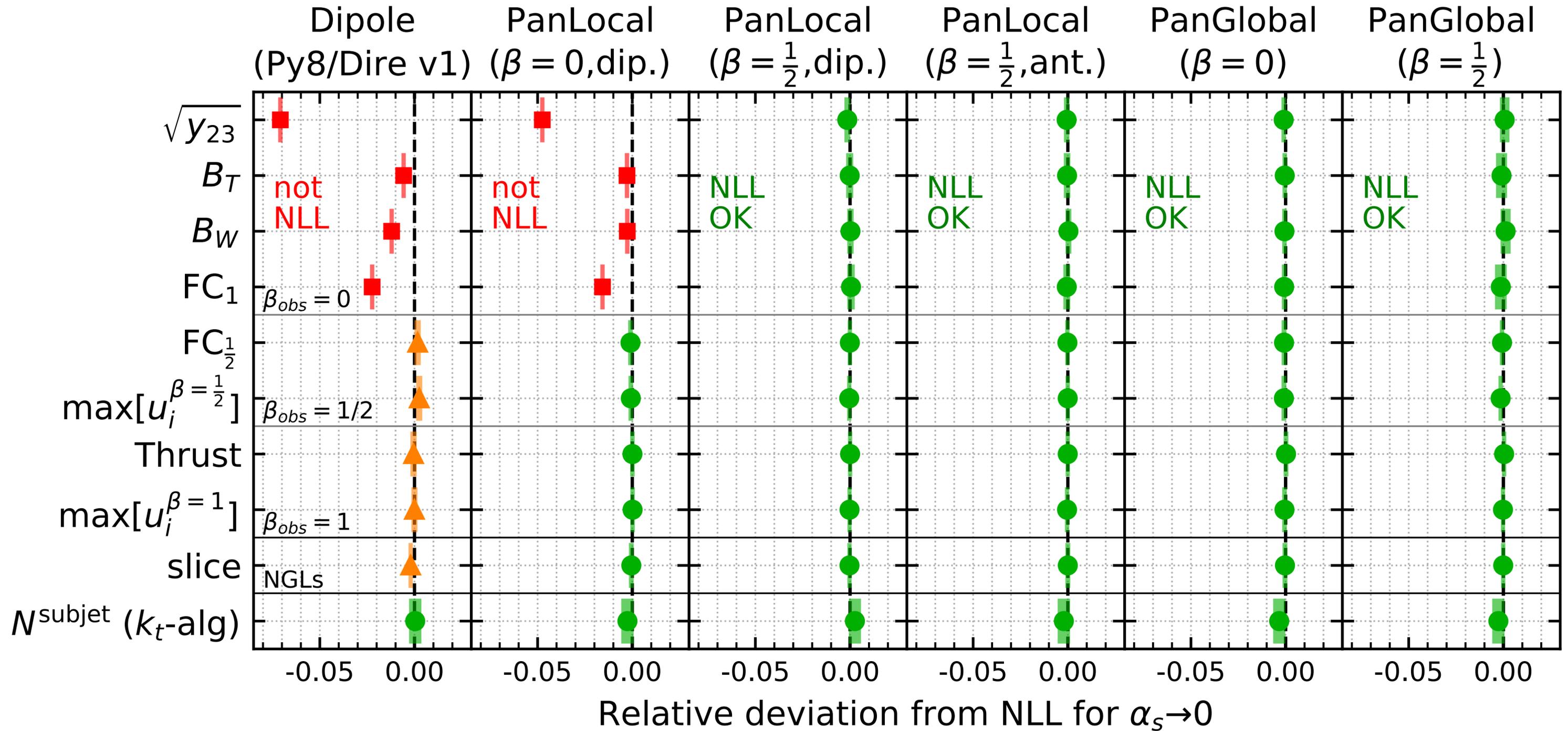


Test class 2: full shower v. all-order NLL — many observables



Relative deviation from NLL for $\alpha_s \rightarrow 0$

Test class 2: full shower v. all-order NLL — many observables



NLL accuracy is becoming the new standard

Logarithmic accuracy of parton showers: a fixed-order study

Dasgupta, Dreyer, Hamilton, Monni, Salam [1805.09327]

Parton showers beyond leading logarithmic accuracy

Dasgupta, Dreyer, Hamilton, Monni, Salam, Soyez [2002.11114]

A new approach to color-coherent parton evolution

Herren, Höche, Krauss, Reichelt, Schönherr [2208.06057]

New approach to QCD final-state evolution in Alaric processes with massive partons

Assi, Höche [2307.00728]

The Alaric parton shower for hadron colliders

Höche, Krauss, Reichelt [2404.14360]

A partitioned dipole-antenna shower with improved transverse recoil Apollo

Preuss [2403.19452]

Summation of large logarithms by parton showers Deductor

Nagy, Soper [2011.04773]

Summation by parton showers of large logarithms in electron-positron annihilation

Nagy, Soper [2011.04777]

Herwig

Initial state radiation in the Herwig 7 angular-ordered parton shower

Bewick, Ferrario Ravasio, Richardson, Seymour [2107.04051]

Spin correlations in final-state parton showers and jet observables

Karlberg, Salam, Scyboz, Verheyen [2103.16526]

PanScales parton showers for hadron collisions: formulation and fixed-order studies

van Beekveld, Ferrario Ravasio, Salam, Soto Ontoso, Soyez, Verheyen [2205.02237]

Next-to-leading-logarithmic PanScales showers for deep inelastic scattering and vector boson fusion

van Beekveld, Ferrario Ravasio [2305.08645]

PanScales

PanScales parton showers for hadron collisions: all-order validation

van Beekveld, Ferrario Ravasio, Hamilton, Salam, Soto Ontoso, Soyez, Verheyen [2207.09467]

Introduction to the PanScales framework, version 0.1

van Beekveld, Dasgupta, El-Menoufi, Ferrario Ravasio, Hamilton, Helliwell, Karlberg, Medves, Monnim Salam, Scyboz, Soto Ontoso, Soyez, Verheyen [2312.13275]

Building a consistent parton shower

Forshaw, Holquin, Plätzer [2003.06400]

Improvements on dipole shower colour

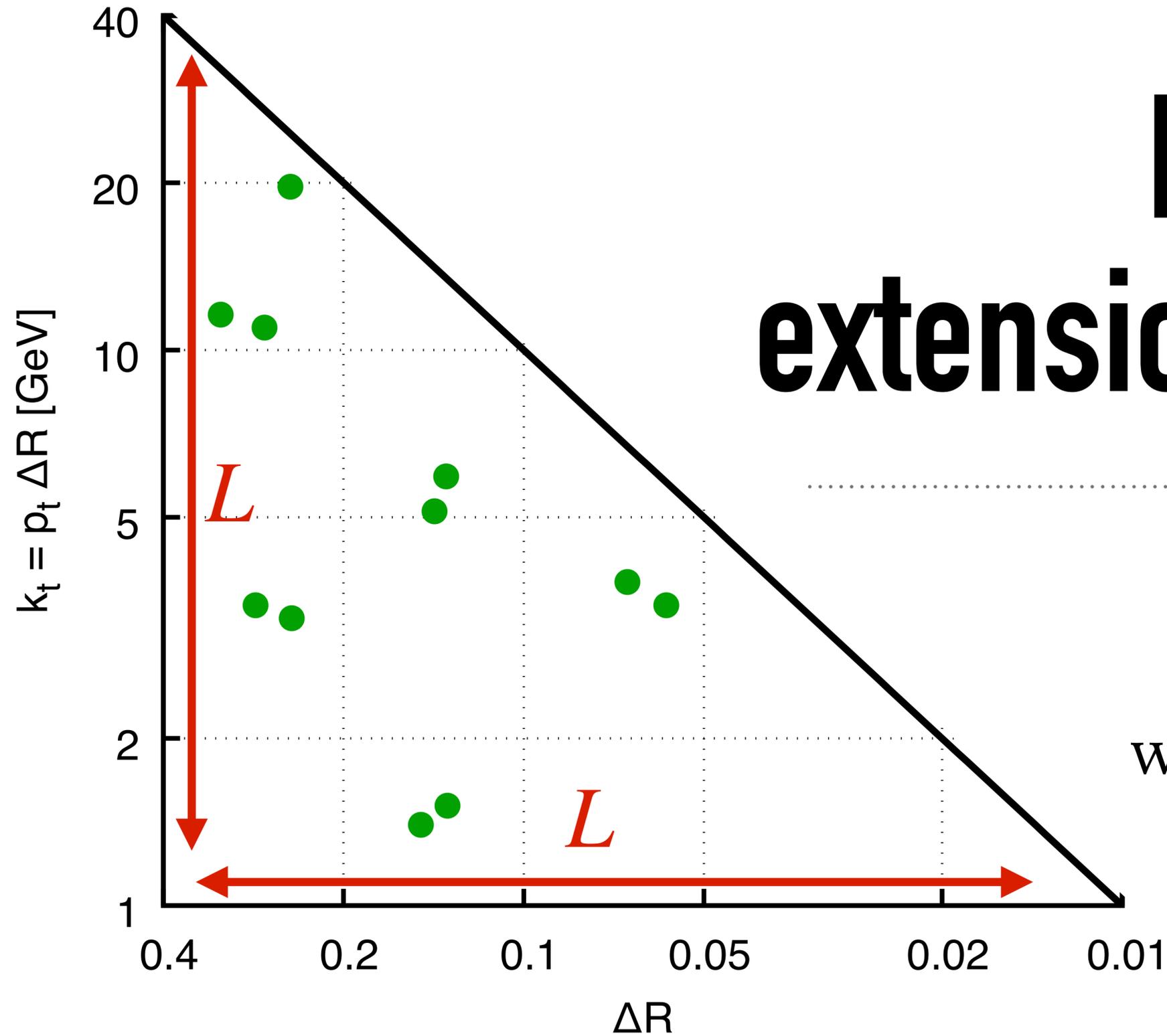
Forshaw, Holquin, Plätzer [2011.15087]

Logarithmic accuracy of angular-ordered parton showers

Bewick, Ferrario Ravasio, Richardson, Seymour [1904.11866]

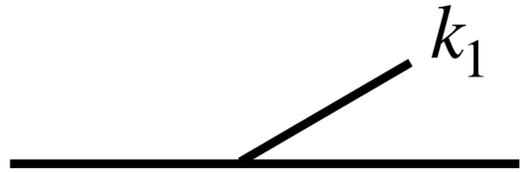
slide from M. van Beekveld

Element #3: extension to higher orders



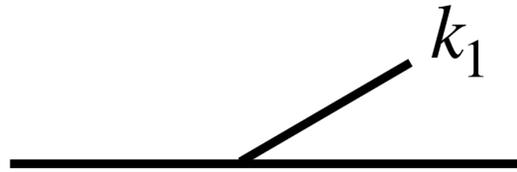
E.g. at NNLL, effective matrix element should be correct even where there are pairs of emissions close by in the Lund plane

Make each new emission's distribution conditional on **one** previous emission

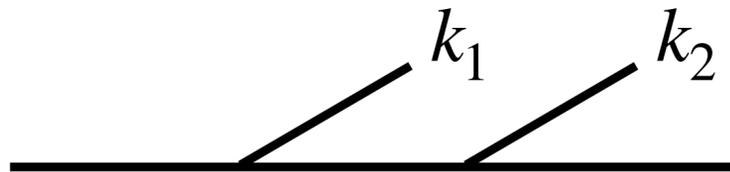


Distribute k_1 according to $M^2(k_1)$

Make each new emission's distribution conditional on **one** previous emission

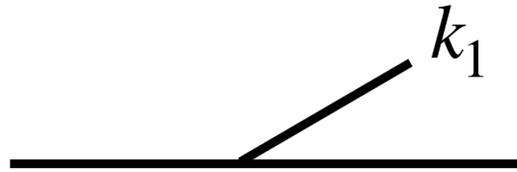


Distribute k_1 according to $M^2(k_1)$

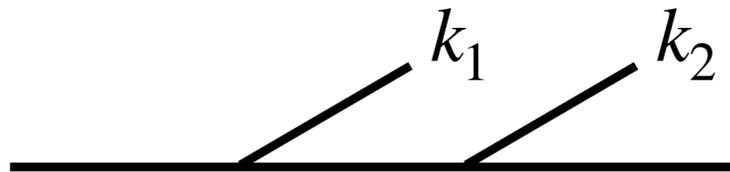


Distribute k_2 according to $M^2(k_1, k_2)/M^2(k_1)$

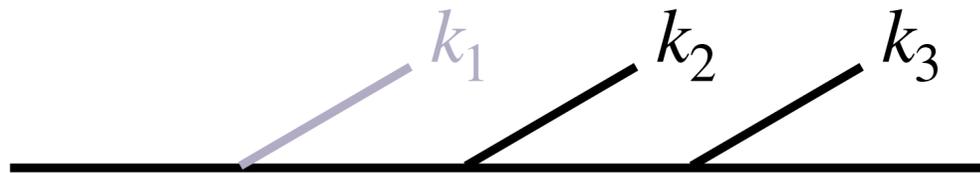
Make each new emission's distribution conditional on **one** previous emission



Distribute k_1 according to $M^2(k_1)$

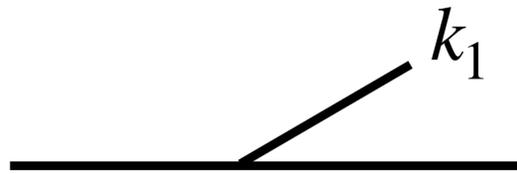


Distribute k_2 according to $M^2(k_1, k_2)/M^2(k_1)$

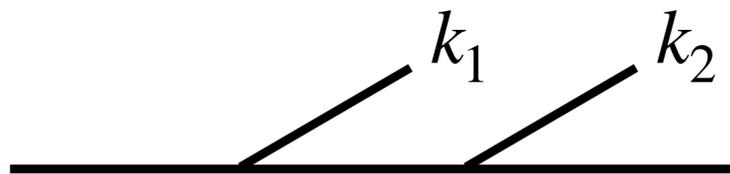


Distribute k_3 according to $M^2(k_2, k_3)/M^2(k_2)$

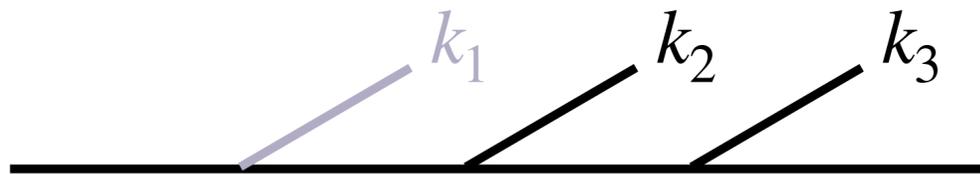
Make each new emission's distribution conditional on **one** previous emission



Distribute k_1 according to $M^2(k_1)$



Distribute k_2 according to $M^2(k_1, k_2)/M^2(k_1)$

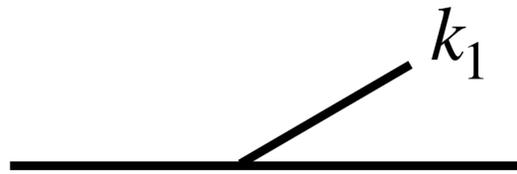


Distribute k_3 according to $M^2(k_2, k_3)/M^2(k_2)$

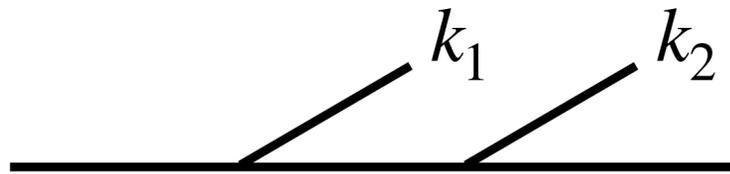


Distribute k_4 according to $M^2(k_3, k_4)/M^2(k_3)$

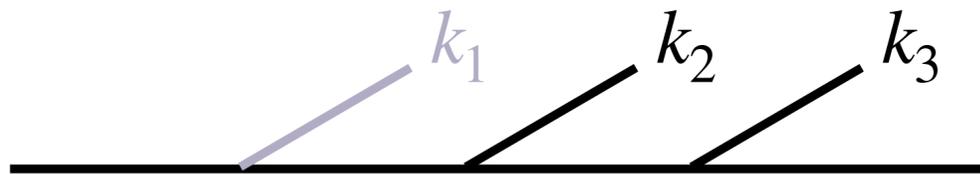
Make each new emission's distribution conditional on **one** previous emission



Distribute k_1 according to $M^2(k_1)$



Distribute k_2 according to $M^2(k_1, k_2)/M^2(k_1)$



Distribute k_3 according to $M^2(k_2, k_3)/M^2(k_2)$

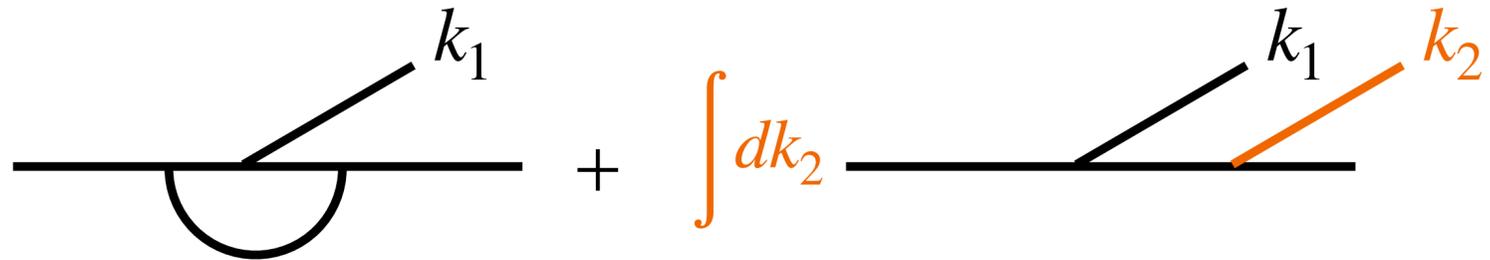


Distribute k_4 according to $M^2(k_3, k_4)/M^2(k_3)$

Relies on factorisation: e.g. $M^2(k_1, k_2, k_3, k_4)/M^2(k_1, k_2, k_3) \rightarrow M^2(k_3, k_4)/M^2(k_3)$
if 3 and 4 well separated in Lund plane from 1 and 2

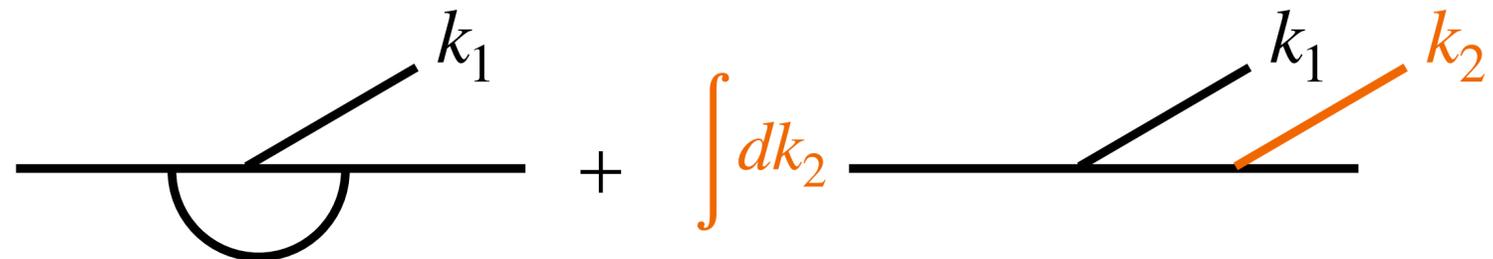
[factorised matrix elements given in Dokshitzer, Marchesini & Oriani '92, Campbell & Glover, [hep-ph/9710255](https://arxiv.org/abs/hep-ph/9710255),
Catani & Grazzini [hep-ph/9810389](https://arxiv.org/abs/hep-ph/9810389), etc.]

Account for virtual corrections associated with each emission

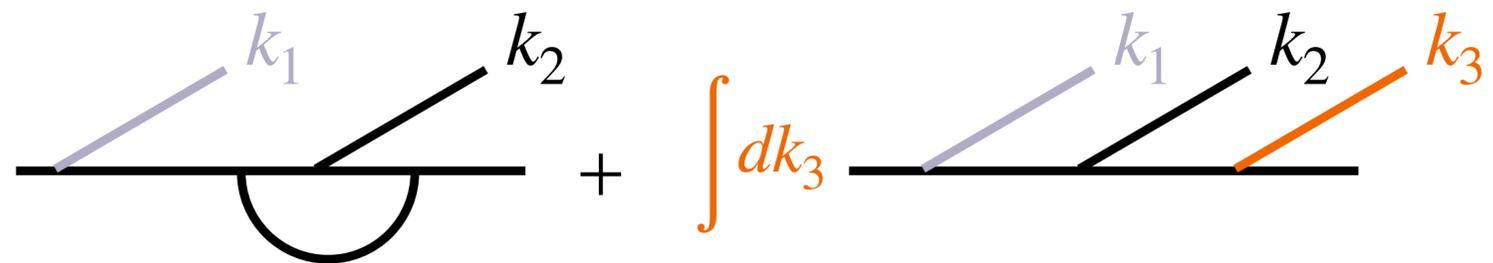


NLO correction to k_1 emission
intensity sums loop correction and all
possible scenarios for the next
emission

Account for virtual corrections associated with each emission



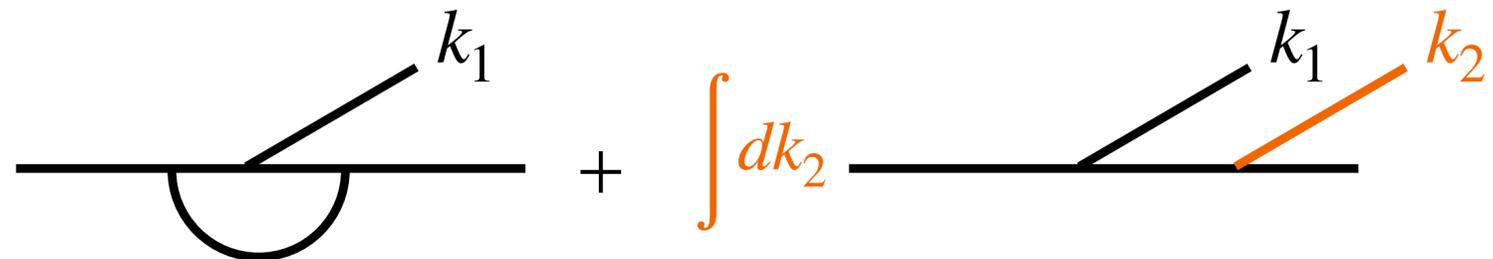
NLO correction to k_1 emission
intensity sums loop correction and all possible scenarios for the next emission



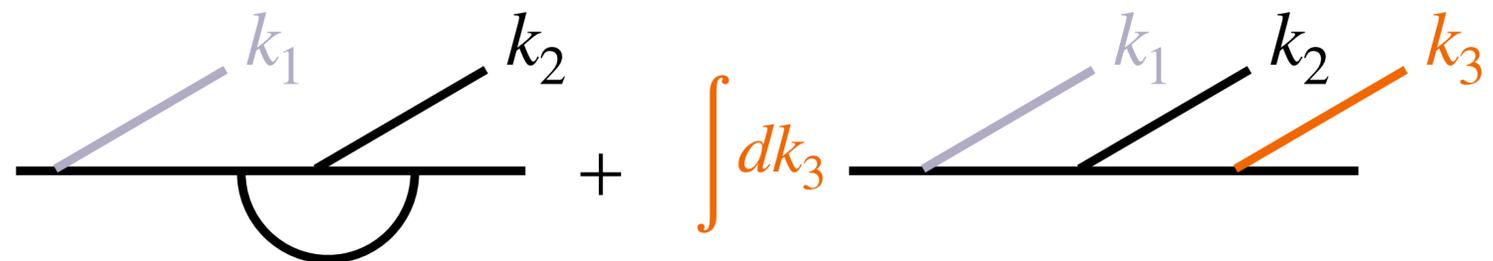
etc.

NLO correction to k_2 emission
intensity sums loop correction and all possible scenarios for the following emission

Account for virtual corrections associated with each emission



NLO correction to k_1 emission
intensity sums loop correction and all
possible scenarios for the next
emission



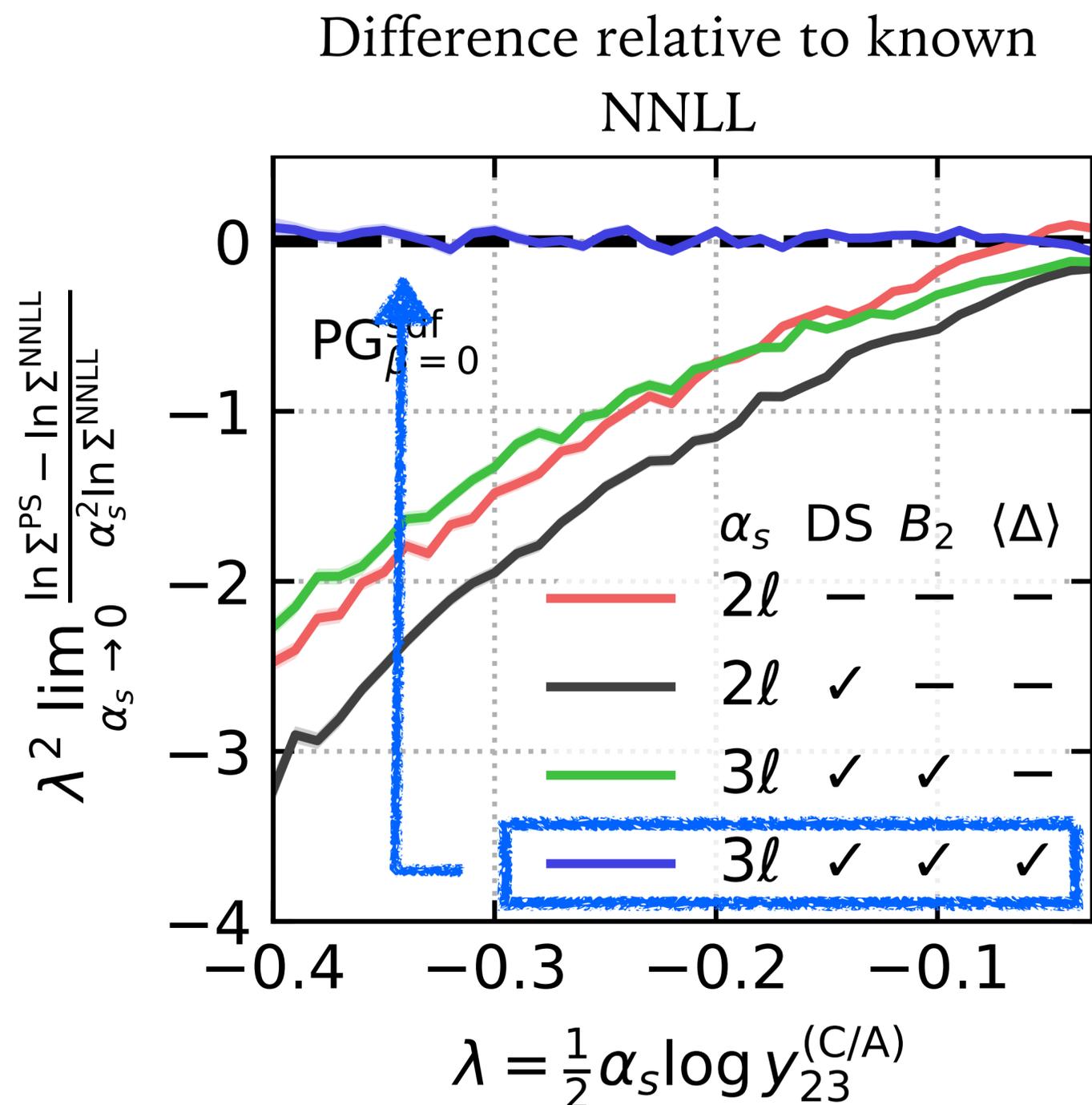
NLO correction to k_2 emission
intensity sums loop correction and all
possible scenarios for the following
emission

etc.

Again relies on factorisation, e.g. when 1 and 2 are well separated in the Lund plane
+ careful nesting, cf. Ferrario Ravasio et al, [2307.11142](#); van Beekveld, Dasgupta, El-Menoufi,
Helliwell, Monni, [GPS 2409.08316](#)

(see also Hartgring, Laenen & Skands, 1303.4974, Campbell et al [2108.07133](#) at fixed order)

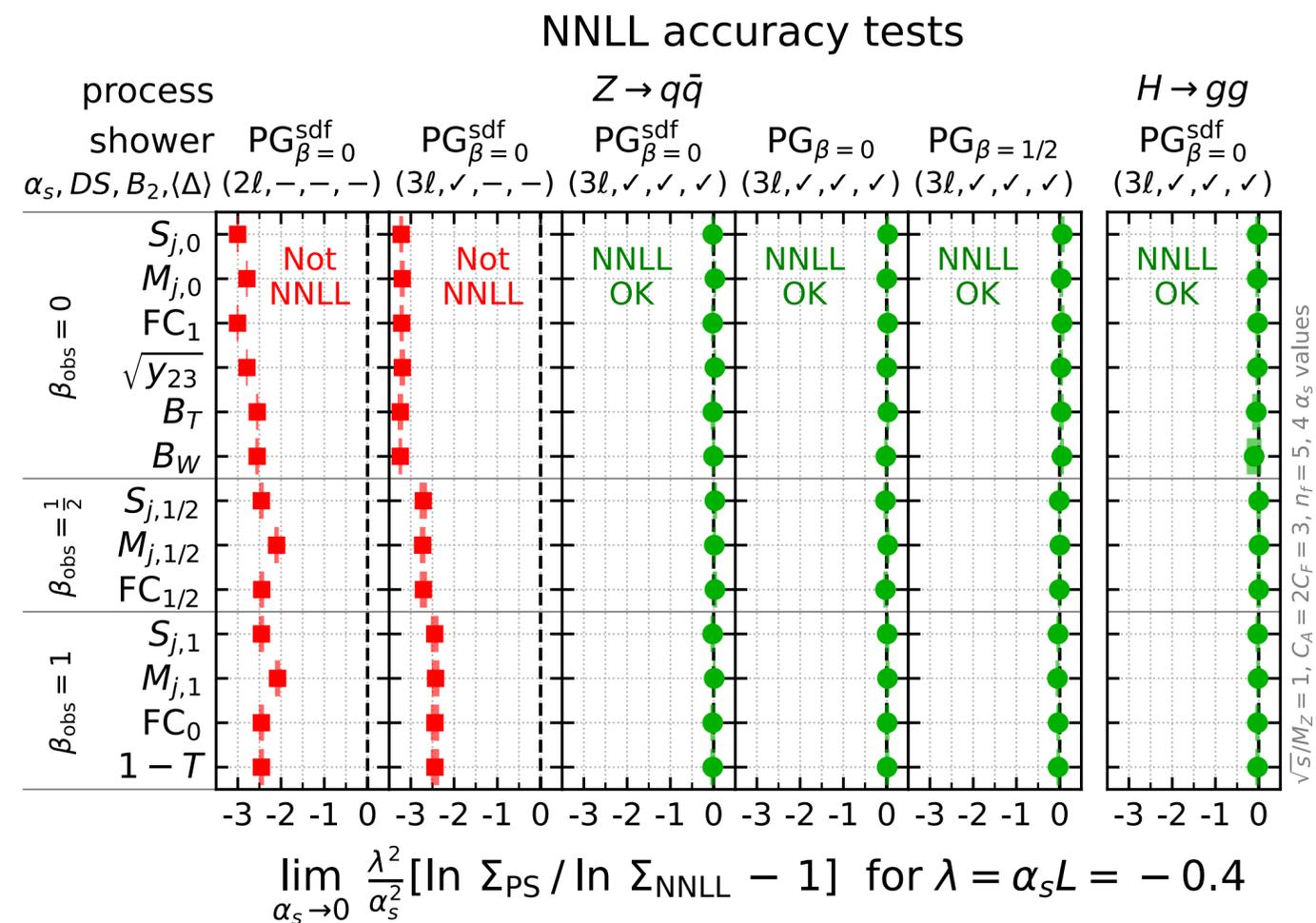
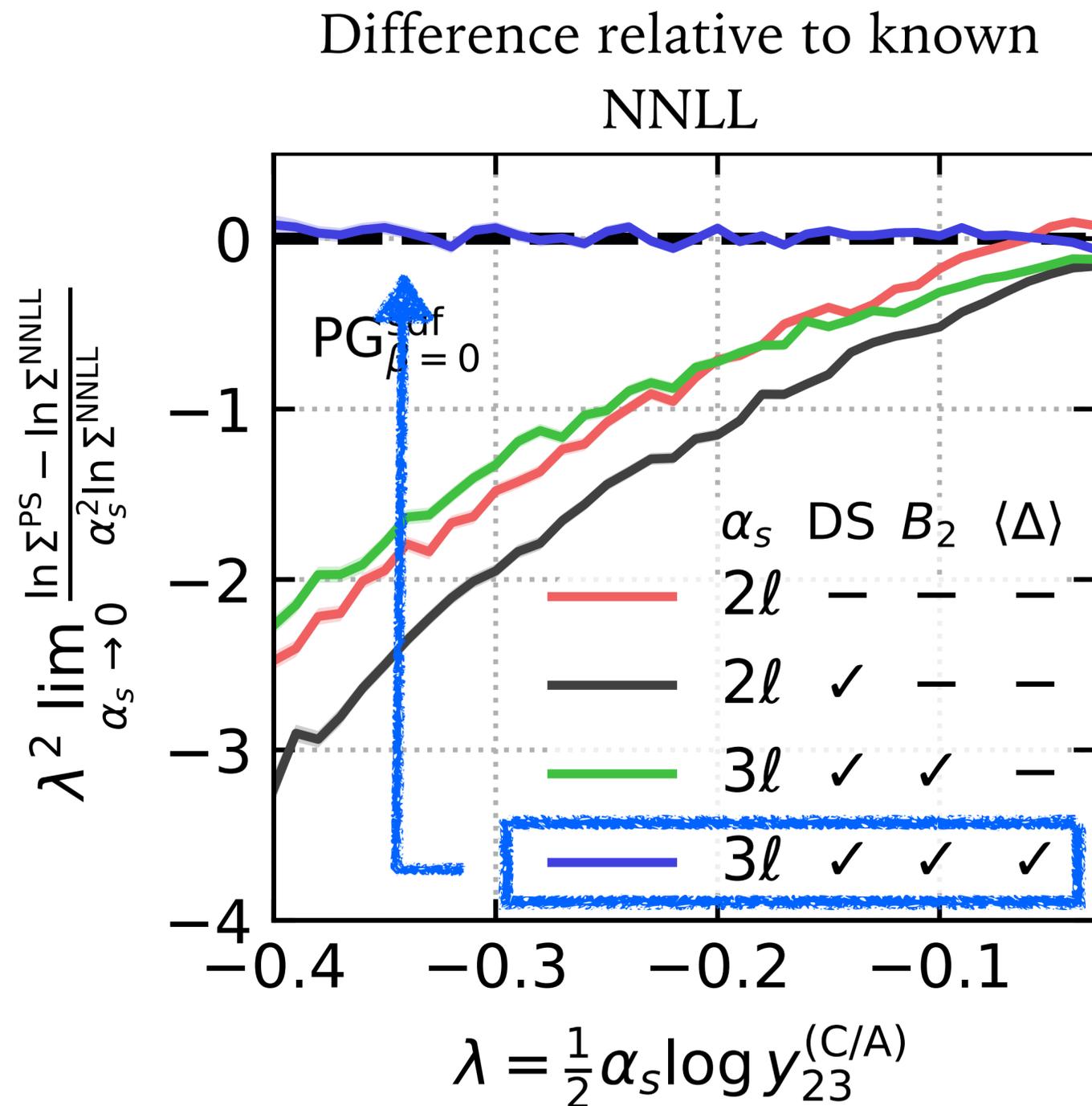
Testing NNLL for event shapes (so far only for e^+e^- collisions)



need to analyse and account for all possible sources of NNLL contribution

(some, which don't affect event shapes, are still work in progress)

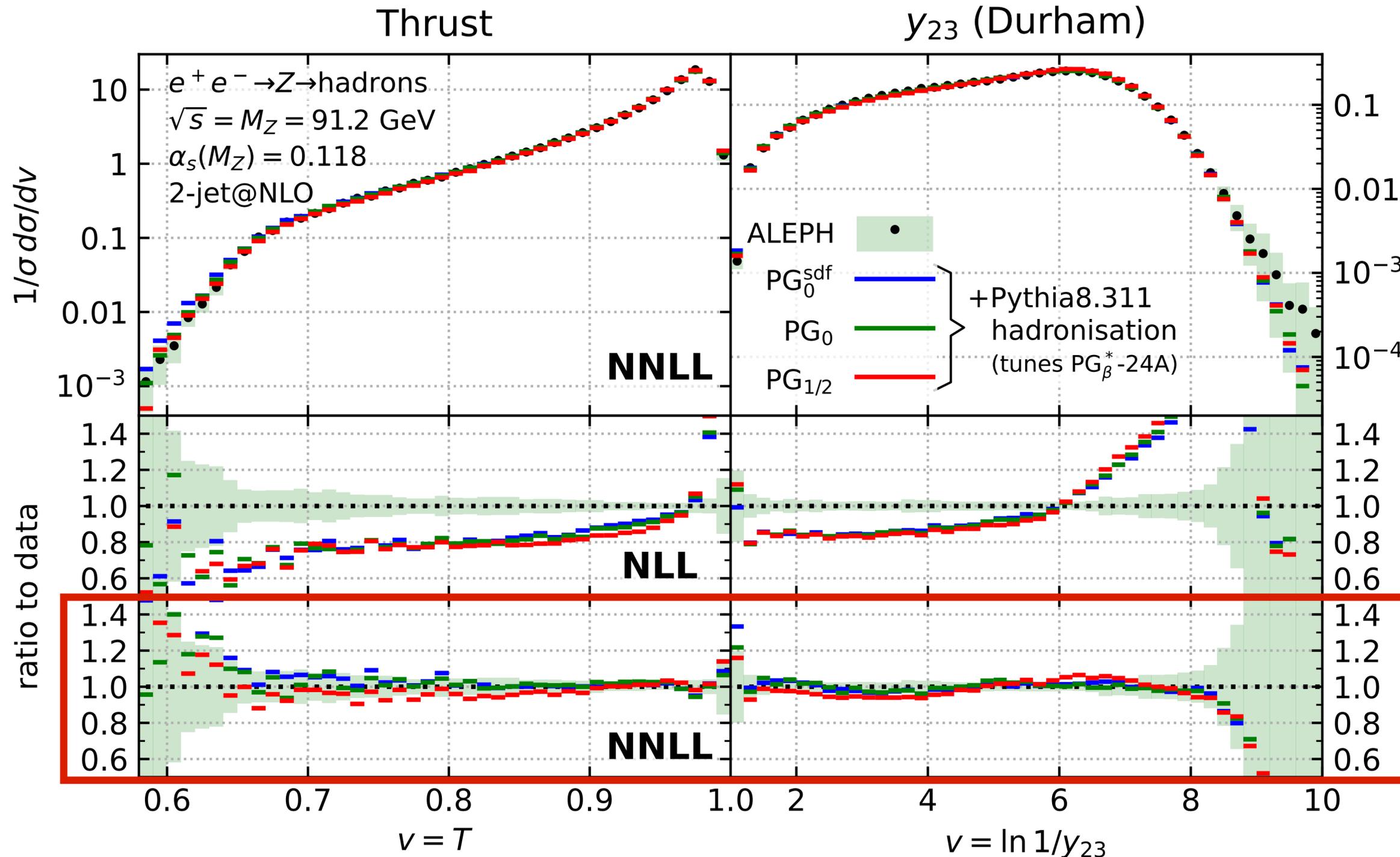
Testing NNLL for event shapes (so far only for e^+e^- collisions)



need to analyse and account for all possible sources of NNLL contribution

(some, which don't affect event shapes, are still work in progress)

Comparing to LEP event-shape data



NNLL brings 20% effects ($\sim \alpha_s$)

Dramatically improves agreement with data, using a “normal” $\alpha_s = 0.118$

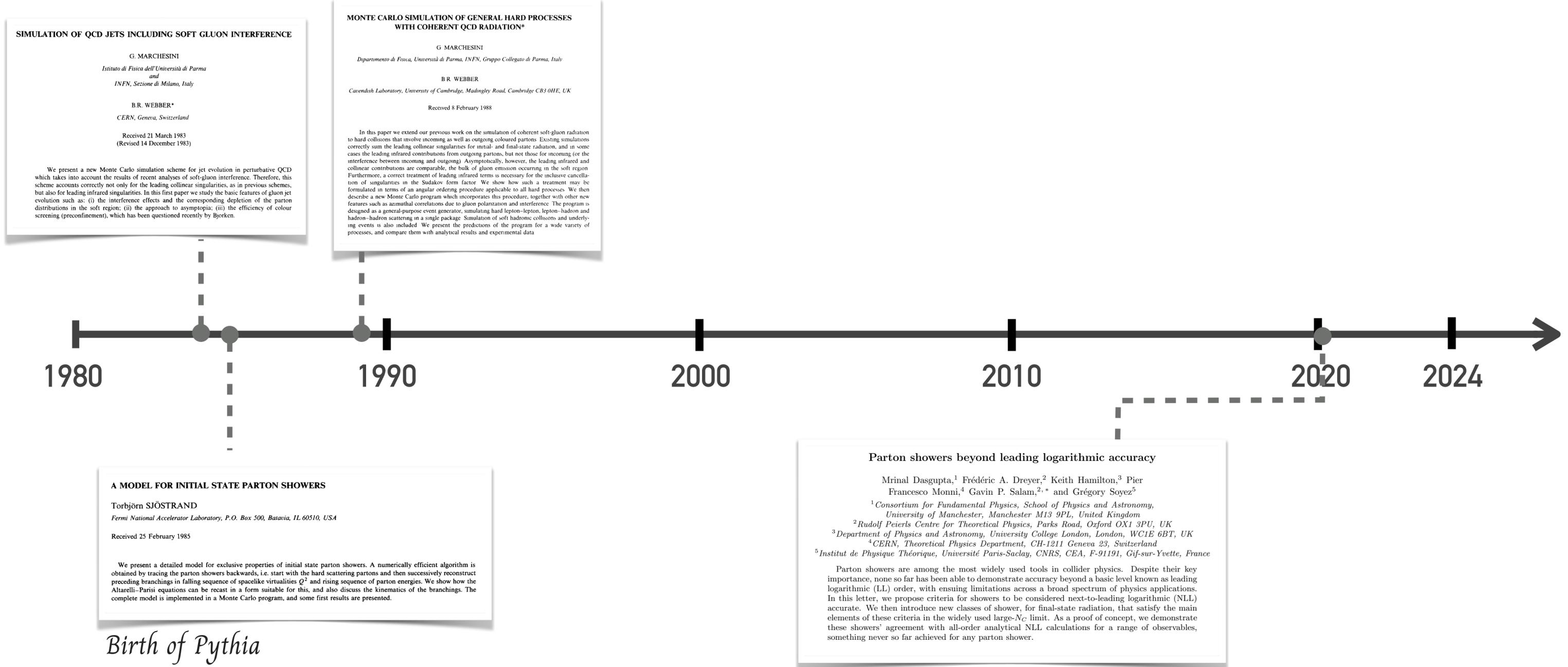
NB: 3-jet @ NLO still missing for robust pheno conclusions

← NNLL

Took about 35 years to reach full NLL since the birth of parton showers . . .

Birth of Herwig (with elements of NLL for global observables)

slide from Pier Monni



[ca. 800 papers on the subject of event generators

... key steps towards NNLL were just 0(5) years away

slide from Pier Monni

Birth of Herwig (with elements of NLL for global observables)

General principles for NNLL parton showers

SIMULATION OF QCD JETS INCLUDING SOFT GLUON INTERFERENCE

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Received 21 March 1983
(Revised 14 December 1983)

We present a new Monte Carlo simulation scheme for jet evolution in perturbative QCD which takes into account the results of recent analyses of soft-gluon interference. Therefore, this scheme accounts correctly not only for the leading collinear singularities, as in previous schemes, but also for leading infrared singularities. In this first paper we study the basic features of gluon jet evolution such as: (i) the interference effects and the corresponding depletion of the parton distributions in the soft region; (ii) the approach to asymptopia; (iii) the efficiency of colour screening (preconfinement), which has been questioned recently by Bjorken.

MONTE CARLO SIMULATION OF GENERAL HARD PROCESSES WITH COHERENT QCD RADIATION*

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Received 8 February 1988

In this paper we extend our previous work on the simulation of coherent soft-gluon radiation to hard collisions that involve incoming as well as outgoing coloured partons. Existing simulations correctly sum the leading collinear singularities for initial- and final-state radiation, and in some cases the leading infrared contributions from outgoing partons, but not those for incoming (or the interference between incoming and outgoing). Asymptotically, however, the leading infrared and collinear contributions are comparable, the bulk of gluon emission occurring in the soft region. Furthermore, a correct treatment of leading infrared terms is necessary for the inclusive cancellation of singularities in the Sudakov form factor. We show how such a treatment may be formulated in terms of an angular ordering procedure applicable to all hard processes. We then describe a new Monte Carlo program which incorporates this procedure, together with other new features such as azimuthal correlations due to gluon polarization and interference. The program is designed as a general-purpose event generator, simulating hard lepton-lepton, lepton-hadron and hadron-hadron scattering in a single package. Simulation of soft hadronic collisions and underlying events is also included. We present the predictions of the program for a wide variety of processes, and compare them with analytical results and experimental data.

A new standard for the logarithmic accuracy of parton showers

Melissa van Beekveld,¹ Mrinal Dasgupta,² Basem Kamal El-Menoufi,³ Silvia Ferrario Ravasio,⁴ Keith Hamilton,⁵ Jack Helliwell,⁶ Alexander Karlberg,⁴ Pier Francesco Monni,⁴ Gavin P. Salam,^{6,7} Ludovic Scyboz,³ Alba Soto-Ontoso,⁴ and Gregory Soyez⁸

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We report on a major milestone in the construction of logarithmically accurate final-state parton showers, achieving next-to-next-to-leading-logarithmic (NNLL) accuracy for the wide class of observables known as event shapes. The key to this advance lies in the identification of the relation between critical NNLL analytic resummation ingredients and their parton-shower counterparts. Our analytic discussion is supplemented with numerical tests of the logarithmic accuracy of three shower variants for more than a dozen distinct event-shape observables in $Z \rightarrow q\bar{q}$ and Higgs $\rightarrow gg$ decays. The NNLL terms are phenomenologically sizeable, as illustrated in comparisons to data.



A MODEL FOR INITIAL STATE PARTON SHOWERS

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Received 25 February 1985

We present a detailed model for exclusive properties of initial state parton showers. A numerically efficient algorithm is obtained by tracing the parton showers backwards, i.e. start with the hard scattering partons and then successively reconstruct preceding branchings in falling sequence of spacelike virtualities Q^2 and rising sequence of parton energies. We show how the Altarelli-Parisi equations can be recast in a form suitable for this, and also discuss the kinematics of the branchings. The complete model is implemented in a Monte Carlo program, and some first results are presented.

Birth of Pythia

Parton showers beyond leading logarithmic accuracy

Mrinal Dasgupta,¹ Frédéric A. Dreyer,² Keith Hamilton,³ Pier Francesco Monni,⁴ Gavin P. Salam,^{2,*} and Grégory Soyez⁵

¹Consortium for Fundamental Physics, School of Physics and Astronomy, University of Manchester, Manchester M13 9PL, United Kingdom
²Rudolf Peierls Centre for Theoretical Physics, Parks Road, Oxford OX1 3PU, UK
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⁴CERN, Theoretical Physics Department, CH-1211 Geneva 23, Switzerland
⁵Institut de Physique Théorique, Université Paris-Saclay, CNRS, CEA, F-91191, Gif-sur-Yvette, France

Parton showers are among the most widely used tools in collider physics. Despite their key importance, none so far has been able to demonstrate accuracy beyond a basic level known as leading logarithmic (LL) order, with ensuing limitations across a broad spectrum of physics applications. In this letter, we propose criteria for showers to be considered next-to-leading logarithmic (NLL) accurate. We then introduce new classes of shower, for final-state radiation, that satisfy the main elements of these criteria in the widely used large- N_C limit. As a proof of concept, we demonstrate these showers' agreement with all-order analytical NLL calculations for a range of observables, something never so far achieved for any parton shower.

Parton showering with higher-logarithmic accuracy for soft emissions

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²Department of Physics and Astronomy, University College London, London, WC1E 6BT, UK
³Rudolf Peierls Centre for Theoretical Physics, Clarendon Laboratory, Parks Road, Oxford OX1 3PU, UK
⁴All Souls College, Oxford OX1 4AL, UK
⁵IPhT, Université Paris-Saclay, CNRS UMR 3681, CEA Saclay, F-91191 Gif-sur-Yvette, France

The accuracy of parton-shower simulations is often a limiting factor in the interpretation of data from high-energy colliders. We present the first formulation of parton showers with accuracy one order beyond state-of-the-art next-to-leading logarithms, for classes of observable that are dominantly sensitive to low-energy (soft) emissions, specifically non-global observables and subject multiplicities. This represents a major step towards general next-to-next-to-leading logarithmic accuracy for parton showers.

General principles for a NLL parton shower (formulated for e^+e^- , many extensions will follow)

[ca. 800 papers on the subject of event generators]

Conclusions

Collider particle physics is a rich and diverse subject

Core exploration of the Higgs sector has only just started

- Many aspects are (hypothesized to be) crucial for the world around us
- Major targets for future colliders: e.g. triple-Higgs interaction \Leftrightarrow Higgs potential

Central to quantitative collider physics is the strong interaction

- quest for accuracy brings huge challenges & QCD is delivering on multiple fronts
- one of those fronts is the question of how to span disparate momentum scales in simulations: major conceptual steps over past years & soon to be available for practical use.